

# Heavy Photon Search Experiment at Jefferson Laboratory: proposal for 2014-2015 run

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## ABSTRACT

The Heavy Photon Search (HPS) is an experiment proposed for Jefferson Laboratory to search for new heavy vector boson(s), aka “heavy photons” or “dark photons” or “hidden sector photons,” in the mass range of 20 MeV/c<sup>2</sup> to 1000 MeV/c<sup>2</sup>. Such particles will arise if there are additional U(1) gauge bosons in nature, and they will couple, albeit weakly, to electric charge through kinetic mixing. HPS searches for electro-produced heavy photons with both invariant mass and separated decay vertex signatures using a compact, large acceptance forward spectrometer. The first stage of HPS, the HPS Test Run, was approved by the Jefferson Lab PAC37 on January 14, 2011, after which it was proposed to DOE HEP for funding and approved and funded by Summer 2011. The Test Run was built in 2011-2012, and installed and run at JLAB in Spring, 2012. PAC39 reviewed HPS in June, 2012, and on the basis of the successful run, granted it an “A” rating, a commissioning run with electron beams, and “C1” approval to proceed to the full experiment contingent on final approval from JLAB management. This proposal describes the second stage of our program, the full HPS experiment, which is capable of searching for heavy photons over a wide and uncharted region in parameter space, and of discovering “true muonium,” the QED  $\mu^+\mu^-$  atom. It reviews the scientific motivations for HPS; describes the proposed apparatus, data acquisition system, and trigger; summarizes key results of the HPS Test Run which demonstrate that the experiment is technically feasible and that electron beam backgrounds are well-understood; discusses expected performance and experimental reach; and concludes with budget and schedule information and a proposed run plan. Jefferson Laboratory management has encouraged HPS to be ready for a commissioning run in the Fall of 2014 and a data taking run beginning in Spring 2015. Keep alive funding is being used to do the advanced design and R&D that will make this schedule possible. With timely approval and funding from DOE HEP, the HPS Collaboration will complete design, build, test, and commission HPS in time to take advantage of this scheduling opportunity, and begin in earnest its search for spectacular new physics at the Intensity Frontier.

# Contents

<b>1. Introduction</b>	7
<b>2. Motivations for Searching for Heavy Photons</b>	10
2.1. Theory Update	11
2.1.1. Heavy Photons and Dark Matter	12
2.1.2. Heavy Photons and Muon $g - 2$	17
2.2. Update on Experimental Status	17
2.3. HPS physics with True Muonium	18
2.4. HPS Searches for Hidden Sectors	19
<b>3. Proposed Measurements</b>	20
3.1. Search for the heavy photon	21
3.2. Search for true muonium	25
3.3. Other searches for hidden sector particles	26
<b>4. Description of the HPS setup</b>	28
4.1. Overview	28
4.2. HPS beamline	29
4.2.1. Layout of the HPS setup	30
4.2.2. Running Conditions	34
4.2.3. Beam Diagnostics	36
4.2.4. Vacuum chambers	39
4.2.5. Beam dumps and shieldings	40
4.2.6. Targets	41
4.3. Slow Controls	43
4.3.1. Framework	43
4.3.2. Voltage controls	44
4.3.3. Motor controls	45
4.3.4. Magnet controls	46
4.3.5. Temperature control and monitoring	47

	5
4.4. Silicon vertex tracker	48
4.4.1. Layout	49
4.4.2. Module Design	50
4.4.3. Mechanical Support, Cooling and Services	54
4.5. Electromagnetic Calorimeter	56
4.5.1. Improvements to the existing calorimeter	58
4.6. Trigger and DAQ	60
4.6.1. SVT DAQ	61
4.6.2. ECal and Muon System FADC Readout	64
4.6.3. Trigger System	67
4.6.4. Event Size and Data Rates	72
4.7. Software	73
4.8. Offline Computing Model	74
<b>5. May-2012 Test Run</b>	<b>77</b>
5.1. HPS Test Run Apparatus	78
5.1.1. Test Run SVT	79
5.1.2. Test Run ECal	81
5.1.3. Test Run Data Acquisition	83
5.1.4. Test Run Trigger System	85
5.2. Multiple Coulomb Scattering Measurement	87
5.3. Test Run Apparatus Performance	91
5.3.1. SVT Performance	91
5.3.2. ECal performance	97
5.3.3. Trigger performance	98
<b>6. HPS Performance Studies</b>	<b>100</b>
6.1. Simulation of Backgrounds and Detector Occupancies	101
6.1.1. Simulation of Backgrounds	101
6.1.2. Simulated Tracker Occupancies	104
6.1.3. Simulated ECal Occupancies	107
6.2. ECal Trigger Rates	108
6.3. Track Reconstruction	112

6.4. Tracking Efficiency, Pattern Recognition and Fake Rates	113
6.5. Track Momentum and Spatial Resolution	113
<b>7. Experimental Reach</b>	120
7.1. Resonance Search	122
7.2. Displaced Vertex and Resonance Search	124
7.3. Reach in Mass-Coupling Parameter Space	125
<b>8. Run Plan and Beam Time Request</b>	126
<b>9. Schedule and Cost Baseline</b>	128
9.1. Cost	131
9.2. Schedule	135
9.3. Manpower	140
9.4. Project Management	141
<b>A. WBS Tables</b>	144
<b>B. Additions and improvements to the HPS setup using non-DOE sources of funding</b>	153
B.1. Improvements to ECal	153
B.2. Muon system	156
B.2.1. Conceptual Design	158
B.2.2. $\mu^+\mu^-$ Trigger	163
B.2.3. Muon system trigger rates	163
<b>C. Simulation Tools</b>	167
<b>D. Test Run ECal Calibration</b>	170
<b>References</b>	172

# 1 Introduction

Access to higher and higher luminosities and ever faster detection and recording techniques enables searches for new physics at otherwise well-explored energies. This fundamental premise of Intensity Frontier physics has already seen dramatic demonstration at the  $e^+e^-$  B factories, where high luminosities and impressive data handling capacities have allowed extensive exploration of CP violation in the quark sector. The same principle is being exploited in new proposals to explore neutrino masses, mixings, and CP violation by directing ever more intense neutrino beams at massive detectors to push sensitivity well beyond present limits. At the Intensity Frontier, searches for new physics often rely on the study of rare processes and the search for subtle effects which would indirectly indicate physics beyond the Standard Model. But this is not the rule. New studies of otherwise commonplace phenomena at electron machines, like trident production off heavy nuclear targets, can, with sufficient sensitivity, explore whole new worlds and directly search for hidden sector particles and forces, those without direct couplings to our Standard Model world. The Heavy Photon Search at Jefferson Laboratory does exactly this, utilizing the high duty factor CEBAF accelerator, intense beams, fast detectors, electronics and triggering, and state of the art data acquisition to explore a very common landscape in search of a most uncommon quarry.

Heavy photons, or “dark” or “hidden sector” photons, may well be part of our universe and related to the Dark Matter. Particles of dark matter, which interact very weakly with normal matter and account for a quarter of the universal mass-energy, are of course not yet detected. The Dark Matter may inhabit a “hidden sector” and interact very weakly with normal, baryonic matter. This sector could include a complex of new forces and other new particles with which we barely interact. Stimulated by the observation of very high energy electrons and positrons in the cosmic rays and the difficulty of understanding their production in terms of tried and true SUSY dark matter annihilation, several authors [1][2] realized that models in which massive dark matter particles annihilate to hidden sector photons, which in turn decay to high energy electron-positron pairs, could naturally account for the observations. These theories presume heavy photons couple to the dark matter, mediate its interactions, are produced in its annihilation, and weakly couple to electric

charge. Heavy photons in the mass range of 20 to 1000 MeV can reasonably account for the observed cosmic ray fluxes [3].

Many Beyond Standard Model theories generate extra U(1) gauge groups, and the associated gauge bosons could have masses over a very wide range. As Holdom [4] realized in the mid 80's, it is natural that such "heavy photons" kinetically mix with our own photon, leading to their induced coupling to electric charge. This mixing can be mediated by GUT level particles which carry both Standard Model hypercharge and its hidden sector analogue. Interestingly, the natural scale for this mixing results in heavy photons coupling to Standard model charged particles with couplings of order  $10^{-3}e$ . So heavy photons naturally couple to electrons, albeit with couplings much suppressed compared to those in standard QED. It follows that electrons will radiate heavy photons, and heavy photons will decay to electron-positron pairs or pairs of other kinematically accessible charged particles, but at rates significantly below QED trident production, and with lifetimes far longer than those expected from purely electromagnetic interactions.

HPS distinguishes heavy photons from the copious background of QED tridents by using both invariant mass and decay length signatures. With good mass resolution, heavy photons will appear as sharp resonances above the QED continuum. For suitable values of mass and coupling, heavy photons will have long lifetimes, resulting in discernible secondary decay vertices. The Heavy Photon Search employs a large acceptance forward magnetic spectrometer with precise momentum measurement and vertexing capability, followed by a highly segmented crystal Electromagnetic Calorimeter for fast triggering and electron identification. HPS depends on the 40 MHz readout capability of the silicon microstrip vertex tracker, 250 MHz FADC readout of the electromagnetic calorimeter, and very high rate triggering and data acquisition systems, to fully exploit CEBAF's essentially DC beams and high intensities. A planned upgrade for HPS, a muon identification system just downstream of the ECal, would significantly boost the experimental reach for heavy photon masses above the dimuon threshold and provide an independent trigger. The electron beam is transported in vacuum through the entire apparatus to eliminate beam gas backgrounds; and the apparatus is split top-bottom, to avoid electrons which have multiple Coulomb scattered or radiated in the target, and been subsequently dispersed by the magnetic field.

HPS probes a unique region of the mass-coupling parameter space where the heavy photon signal would be lost in the trident background without the vertex signature, and it



simultaneously accesses a region at higher coupling strength by relying on bump hunting alone. HPS is sensitive to a region of parameter space favored by accounting for the discrepancy between measured and calculated values for the muon's  $g-2$  with the existence of a heavy photon, and probes an extensive region suggested by parameters which could account for dark matter annihilations into heavy photons. In broader terms, HPS searches for heavy photons in a region suggested on very general theoretical grounds. As seen above, coupling strengths of order  $10^{-3}e$  are theoretically natural; masses of order  $\alpha m_W$  are expected in models where a Higgs mechanism is operative in the hidden sector. Interestingly, HPS is also sensitive to the production of "true muonium," the QED atom comprised of  $\mu^+\mu^-$ , which is produced with a well-defined (and detectable) cross-section, and decays with a well-defined (and observable) lifetime to  $e^+e^-$ . HPS should discover true muonium, measure some of its properties, and find it a useful calibration signal.

This proposal seeks funding for the Heavy Photon Search (HPS) Experiment at Thomas Jefferson National Accelerator Facility. This experiment is the second stage of a program that was initiated with the Heavy Photon Search Test Run Proposal [5, 6], which was approved by the Jefferson Laboratory Program Advisory Committee PAC37 in January, 2011, and approved and funded by DOE HEP in the late Spring of 2011. PAC37 also conditionally approved the full experiment, contingent upon the Test Run results. During the remainder of FY2011 and the first half of FY2012, the Test Run apparatus, data acquisition system, and system software were designed, constructed, and tested. On April 19, 2012, the newly constituted HPS Collaboration installed the experiment in Jefferson Lab's Hall B experimental area, and began commissioning the experiment parasitically, using the HDIce photon beam. Although the Jefferson Lab schedule did not accommodate the electron beam running which had been requested, the apparatus was fully commissioned by running parasitically in the photon beam. The trigger and data acquisition and storage systems worked well, and all systems performed as expected. Efficient track reconstruction in the Silicon Vertex Tracker has been demonstrated, measurements of shower energies and positions have been made in the Electromagnetic Calorimeter, and critical assumptions about background rates have been confirmed. The critical test run goals have been accomplished. A status report summarizing HPS's progress and results was submitted to PAC39 [7] along with a request for unconditional approval for the full experiment. At its June, 2012 meeting, PAC39 graded HPS physics with an "A" rating, approved a commissioning run with electrons, and granted

so-called “C1” approval, which gave Jefferson Laboratory management the final say in granting HPS the running time needed to search comprehensively for Heavy Photons. Since that approval, Laboratory management has urged HPS to be ready for both commissioning and data taking runs in 2014 and 2015, when the upgraded CEBAF 12 accelerator will have been completed, commissioned, and operational. CLAS12 is the large general purpose apparatus being constructed to exploit CEBAF 12 in Hall B. Scheduling and funding delays in the construction of the CLAS12 superconducting magnets are expected to delay CLAS12 installation, and provide HPS the opportunity for commissioning and first data taking. To take full advantage of this scheduling opportunity, the HPS Collaboration has re-visited the original HPS design, and simplified and improved it. The resulting simplifications make it possible to construct and test HPS in time for installation in the Fall of 2014. The resulting improvements extend the reach far beyond that of the Test Run experiment, maximize the Hall B physics output during this time period, and let HPS begin its search for heavy photons in a large and hitherto unexplored region of parameter space.

In the following, this proposal motivates and describes the new HPS Experiment, documents the experience and performance obtained with the Test Run Apparatus, demonstrates that the backgrounds expected in electron running are understood and manageable, reviews the performance and physics reach of the new experiment, and outlines the budget, schedule, and milestones for constructing and deploying it. It concludes with a request for beam time.

## 2 Motivations for Searching for Heavy Photons

HPS will search for heavy photons, called  $A'$ s, which are new hypothesized massive vector bosons that have a small coupling to electrically charged matter, including electrons. The existence of an  $A'$  is theoretically natural and could explain the discrepancy between the measured and observed anomalous magnetic moment of the muon and several intriguing dark matter-related anomalies. As discussed in the following section, HPS should also have the capability to make the first detection of *True Muonium*, a bound state of a  $\mu^+ - \mu^-$  pair predicted by Quantum Electrodynamics (QED). The search for  $A'$ s has generated enormous interest in the international physics community. This is evidenced, for example, by its inclusion in the recent Intensity Frontier Workshop [8, 9], many novel searches in colliding

beam and fixed-target data (see [10] for a recent summary of results), and by numerous new experiments (in addition to HPS) proposed to search for them, including APEX [11, 12], MAMI [13], and DarkLight [14]. We briefly review the theory and motivation for heavy photons and existing constraints on  $A'$ .

## 2.1 Theory Update

The  $A'$  is a new abelian  $U(1)$  gauge boson with a weak coupling to electrically charged particles induced by “kinetic mixing” with the photon [4, 15]. Kinetic mixing produces an effective parity-conserving interaction  $\epsilon e A'_\mu J_{EM}^\mu$  of the  $A'$  to the electromagnetic current  $J_{EM}^\mu$ , suppressed relative to the electron charge  $e$  by the parameter  $\epsilon$ , which can naturally be in the range  $10^{-12} - 10^{-2}$  [16–19].

More broadly, “kinetic mixing” of the photon with new forces offers one of the few portals with which ordinary matter can be used to search for light new forces beyond the Standard Model consistent with known symmetries. An  $A'$  would also allow ordinary matter to have a small coupling to new particles in a “hidden sector” that do not interact with the Standard Model’s strong, weak, or electromagnetic forces. There has been intense speculation over the past three decades about the existence of hidden sectors. Theoretical models with dark matter, supersymmetry, and string theory constructions often employ hidden sectors with new particle content to resolve various phenomenological questions [20–24] (see [8] for a recent review). The photon mixing with an  $A'$  could provide the only non-gravitational window into their existence.

While loop level effects can naturally generate  $\epsilon$  in an observable range, simple theory arguments offer less guidance for what range of  $A'$  mass to search for. Many mass generating mechanisms have been proposed –  $A'$  masses can arise, for example, via the Higgs mechanism as in the models of [25–28], or via a Stueckelberg mechanism, as often occurs in large volume string compactification models [8]. In models using a Higgs mechanism, a natural mass range for an  $A'$  is near (but beneath) the weak scale, in the MeV to GeV range. This mass range has received considerable attention in part because it may also allow  $A'$ s to resolve several anomalies (see below). Existing constraints are shown in Fig. 1. HPS will be sensitive to  $A'$  masses in between 20–1000 MeV.

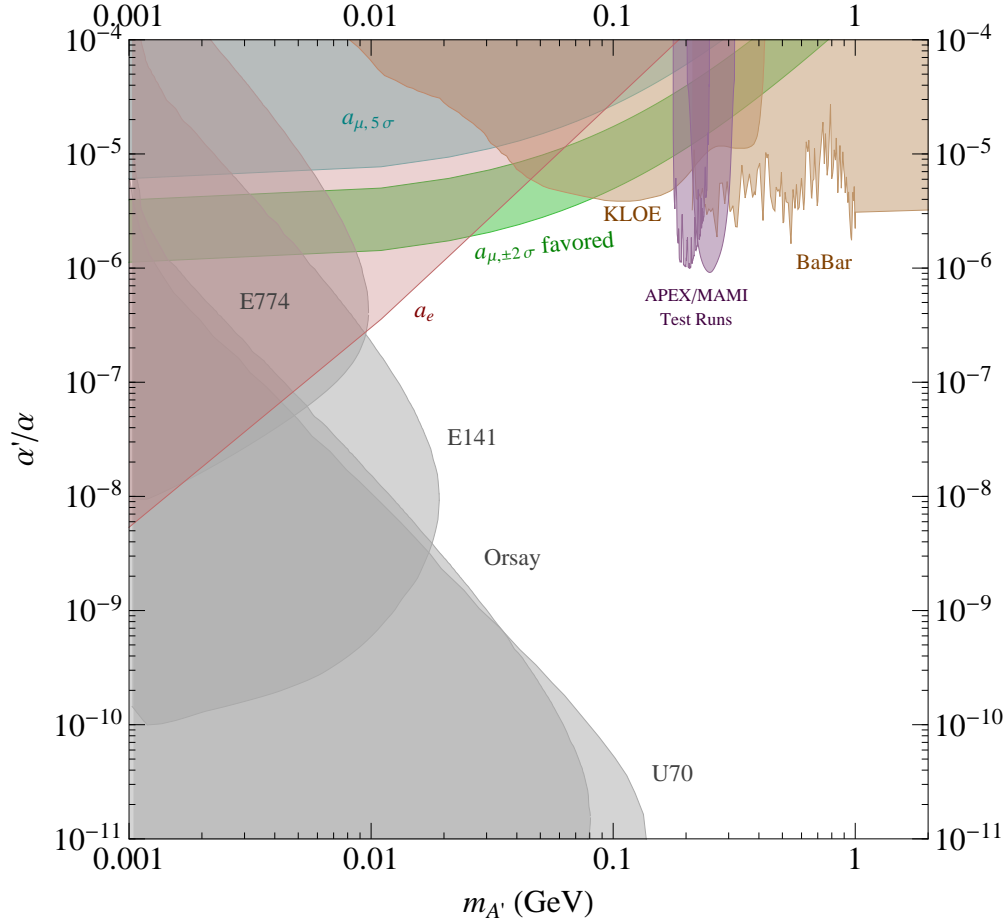


FIG. 1: Existing constraints on heavy photons ( $A'$ ). Shown are existing 90% confidence level limits from the beam dump experiments E141, E774, Orsay, and U70 [29–32, 35, 37, 38], the muon anomalous magnetic moment  $a_\mu$  [39], KLOE [40], the test run results reported by APEX [12] and MAMI [13], an updated estimate using a BaBar result [35, 41, 42], and an updated constraint from the electron anomalous magnetic moment [33, 34]. In the green band, the  $A'$  can explain the observed discrepancy between the calculated and measured muon anomalous magnetic moment [39] at 90% confidence level.

### 2.1.1 Heavy Photons and Dark Matter

The possible role of heavy photons in the physics of dark matter [1, 2] has provided an urgent impetus to search directly for heavy photons. Results from two classes of dark matter searches — “indirect” searches for galactic dark matter annihilation and “direct” searches for dark matter scattering off nuclei — have both been interpreted as potential signals of dark matter interacting through a heavy photon. Both areas have developed considerably in recent years, but not decisively. Here we briefly summarize the status of dark matter, the case for its interactions with heavy photons, and pertinent recent developments in both

observation and theory. The motivation to test these theories of dark matter in a controlled laboratory experiment remains strong.

The concordance model of big bang cosmology — the Lambda Cold Dark Matter ( $\Lambda$ CDM) model — explains all observations of the cosmic microwave background, large-scale structure formation, and supernovae, see e.g. [43]. This model suggests that Standard Model particles make up only about 5% of the energy density in the Universe, while “dark energy” and “dark matter” make up 68% and 27%, respectively, of the Universe’s energy density. The concordance model does not require dark matter to have any new interactions beyond gravity with Standard Model particles. However, an intriguing theoretical observation, dubbed the “WIMP miracle”, suggests that dark matter does have new interactions. In particular, if dark matter consists of 10 GeV to 10 TeV particles interacting via an electroweak-strength force (weakly interacting massive particles or WIMPs), they would automatically have the right relic abundance consistent with the  $\Lambda$ CDM model.

If dark matter does interact with ordinary matter, such interactions could produce at least two observable consequences: dark matter particles in the Milky Way Galaxy (and other bound astrophysical systems) can annihilate or decay into visible matter, which could be detectable as energetic cosmic rays and/or gamma rays at Earth (indirect detection). Dark matter passing through Earth can also scatter off nuclear targets, causing the target to recoil. This recoil is observable in radio-pure detectors with sufficiently low background rates of nuclear recoil (direct detection).

#### *a. Indirect Detection*

The satellites PAMELA [44] and Fermi [45], the balloon-borne detector ATIC [46], the ground-based Cherenkov telescope HESS [47, 48], and other experiments have all reported an excess in the cosmic-ray flux of electrons and/or positrons above backgrounds expected from normal astrophysical processes. The evidence for this excess has only grown, with new measurements of the cosmic-ray electron flux by PAMELA [49] and confirmation by Fermi and AMS-2 of the positron excess [50, 51]. While further data from AMS-2 may shed more light on the spectrum of these excess cosmic-rays, the origin of these excess positrons and electrons remains unknown. It may plausibly arise from any of three possibilities: pair creation in nearby pulsars, acceleration in supernova shocks, or dark matter annihilation or decay.

If the excess arises from dark matter annihilation, two features are incompatible with

annihilation of “conventional” thermal WIMP dark matter charged under the Standard Model weak interactions, but compatible with an alternative explanation, namely that dark matter is charged under a new  $U(1)'$  and annihilates into  $A'$  pairs, which decay directly into electrons and positrons, and/or into muons that decay into electrons and positrons (see e.g. [1, 2, 52–54]):

- The annihilation cross-section required to explain the electron signal is 50 – 1000 times larger than the cross-section favored for the “WIMP miracle”. This can be explained if dark matter interacts with an  $\mathcal{O}(\text{GeV})$ -mass  $A'$ , which mediates a new moderate range force and enhances the annihilation rate at low velocities (the relative velocity of dark matter in the Galactic Halo,  $v \sim 10^{-3}c$ , is much lower than in the early universe, and the relative velocity in self-bound dark matter subhalos is lower still). We refer the reader to [3, 55] for a recent discussion.
- The PAMELA satellite did not see any anti-proton excess [56], which implies that, if dark matter annihilation is responsible for the positron/electron signals, it does not produce baryons. This contradicts expectations for dark matter annihilating through Standard Model interactions, but is expected if dark matter decays into light  $A'$ , which (for  $m_{A'} \lesssim \text{GeV}$ ) are kinematically unable to decay into protons and anti-protons.

We emphasize that these cosmic-ray excesses do not point to a unique region in the  $\epsilon - m_{A'}$  parameter space. Firstly, the value of  $m_{A'}$  determines the branching ratios of the  $A'$  (and hence the dark matter, which here is assumed to annihilate to the  $A'$ ) to different Standard Model states, including  $e^+e^-$ ,  $\mu^+\mu^-$ , pions etc. Since one is trying to match the  $e^-$  and  $e^+$  flux on Earth from dark matter annihilation in the Milky-Way halo to the measured cosmic-ray spectra, the required dark matter mass and annihilation cross section is sensitive to the different branching ratios and, hence,  $m_{A'}$ . For example, for  $m_{A'} < 2m_\mu$ , the dark matter would almost exclusively annihilate to  $e^+e^-$ . However, for  $m_{A'} \sim 700 \text{ MeV}$  near the  $\rho$  or  $\omega'$  resonance, the dark matter would annihilate dominantly to pions, decreasing the energy and yield of  $e^+$  and  $e^-$  per annihilation event; this would require a larger dark matter annihilation cross section and larger dark matter mass to fit the cosmic-ray spectra. A large degeneracy thus exists between  $m_{A'}$  and the dark matter mass and cross section. The degeneracy can be lessened somewhat, but not removed completely, since various other constraints will prefer some regions over others (e.g.  $m_{A'} \sim 100 \text{ MeV}$  over  $700 \text{ MeV}$ ). A set

of benchmarks can be found in e.g. [3], but it will be of interest to cover the whole  $m_{A'}$  region as proposed by HPS (even higher masses near 900 MeV would be of interest).

The second parameter of interest is  $\epsilon$ . Unfortunately,  $\epsilon$  is almost completely unconstrained by the cosmic-ray data. The reason is that  $\epsilon$  determines the lifetime of the  $A'$ , but does not affect the dark matter annihilation cross section nor the decay branching ratios of the  $A'$  (and hence the dark matter) to Standard Model final states; and to explain the cosmic-ray anomalies, it is irrelevant if the  $A'$  decays promptly or only after traveling for thousands of kilometers.

If dark matter annihilation produces the high-energy  $e^+e^-$  excess, correlated gamma-ray fluxes are expected from more distant astrophysical systems where dark matter can annihilate; such fluxes are not expected for the other possible explanations of the cosmic-ray excesses. Such gamma-ray fluxes have not been seen by satellite or ground-based gamma-ray telescopes, like the the Fermi Gamma-ray Telescope, MAGIC, HESS, or VERITAS. Bounds on the gamma ray flux from dwarf spheroidals [57], the outer Milky Way [58], the Galactic Center (e.g. [59, 60] and references therein), and distant galaxies [60, 61] and clusters [62] can thus be used to constrain dark matter interpretations of the Pamela/FERMI excess. In a similar spirit, dark matter annihilation in the epoch of atomic recombination would leave an imprint in the cosmic microwave background radiation, which is similarly constrained [63], and the self-interaction of dark matter via  $A'$  exchange could affect the shape of galactic halos [64, 65]. Each of these systems can be used to constrain models of the PAMELA/Fermi excess, albeit with large theoretical uncertainties.

The present situation can perhaps be summed up as follows: corroborating evidence for an explanation of the cosmic-ray excesses in terms of annihilating dark matter *could* have shown up, but have not. However, the size of the expected corroborating signals is very uncertain, so that the present situation is still inconclusive. Perhaps the best hope for a more definitive statement on a dark matter origin of the cosmic-ray excesses will arise from new CMB polarization data expected from Planck, which will improve sensitivity to dark matter annihilation at the time of recombination by a factor of 10 over WMAP [63]. Planck should either find evidence for dark matter annihilation with a high cross-section (providing further support for dark matter interpretations of the  $e^+e^-$  excess), or more robustly constrain the minimal theories. We note that the recent release of the Planck data only included the temperature data, and the resulting constraints will be only minimally improved in comparison

to the previously available data from WMAP9, the Atacama Cosmology Telescope (ACT), and the South Pole Telescope (SPT). A significant improvement is expected next year, when the CMB polarization power spectrum from Planck, ACT, and SPT will become available.

A very important caveat to the above discussion is that we assumed that dark matter *annihilations* to  $A'$ 's are the origin of the excesses. Instead, dark matter *decays* to an  $A'$  and other light hidden sector particles are also a viable possibility [66, 67]. In this case, an  $A'$  mass below  $\sim 1$  GeV is again motivated by the absence of an antiproton signal, but the size of the  $e^+/e^-$  signal is set by the dark matter decay lifetime, and independent of the  $A'$  mass (recall that in the case of dark matter annihilations, the  $A'$  mass was an important ingredient in determining the size of the Sommerfeld enhancement and, thus, the annihilation cross section). Dark matter decays are less constrained than annihilations as a possible origin to the cosmic-ray excesses, as they produce a smaller corroborating gamma-ray signal (this signal is now proportional to the dark matter density  $\rho$  and not  $\rho^2$ ). Also, no evidence is expected to show up in the CMB data, since the required dark matter lifetime to explain the cosmic-ray excesses is  $\sim 10^{26}$  seconds, much larger than the time of the CMB formation ( $\sim 10^{13}$  seconds).

#### *b. Direct Detection*

The search for dark-matter-nuclear scattering has also seen considerable developments recently, but remains equally ambiguous. Four experiments have reported excesses that *may* be attributable to dark matter, although more mundane explanations are certainly possible: DAMA/Libra [68], CoGeNT [69], which also reported an annual modulation signal [70], CRESST [71], and CDMS-Silicon [72]. If all or a subset of these signals have a dark matter origin, they are most readily attributed to light dark matter ( $\sim 10$  GeV). However, results from CDMS-Germanium [73], XENON10 [74], and XENON 100 [75] appear to disfavor the same parameter regions. Experimental and detector uncertainties remain large enough that perhaps some model parameter space remains moderately consistent with all of these results [76, 80]. In fact, a recent re-analysis of the XENON10 constraint in [80] found a mistake in the original XENON10 publication [74], which weakens the published limit by a factor of a few. The situation is very fluid; more data is forthcoming and will shed light on the current situation. Though the evidence for light dark matter is controversial, it does raise a puzzle: dark matter with such low masses and high couplings cannot easily interact through Standard Model forces (such as  $Z$ -boson exchange), without being excluded by



measurements of the total  $Z$  width at LEP. If indeed dark matter is light, then it seems most likely to interact through a new mediator, a possibility that HPS will probe in the case of an  $A'$ .

We note that heavy inelastic dark matter ( $\sim 100 - 1000$  GeV) interacting with nuclei through  $A'$ -exchange was a possible explanation for these direct detection anomalies a few years ago, and its annihilation to  $A'$ 's could also have explained the cosmic-ray excesses. However, this possibility is now highly constrained by results from XENON100 [77] and CRESST [71]. Light dark matter, as mentioned above, is still viable. In order to have a unified dark matter explanation of the cosmic-ray excesses and direct detection anomalies, one would now likely need two components of dark matter, one light and one heavy component. Theoretical examples of such a possibility have been discussed in the literature, see e.g. [67].

### 2.1.2 Heavy Photons and Muon $g - 2$

Besides being theoretically natural and having a possible connection to dark matter, an  $A'$  could explain the discrepancy between the measured and calculated value of the anomalous magnetic moment of the muon ( $a_\mu = g - 2$ ) [39]. This long-standing puzzle has several possible resolutions, but among the simplest new physics explanations is the existence of a new force mediator that couples to muons, like the  $A'$ . The contribution to  $a_\mu$  of the  $A'$  is like that of the photon, but suppressed by the mixing parameter  $\epsilon^2$  and dependent on the  $A'$  mass. The green region in Fig. 1 is the  $2\sigma$  band in which the  $A'$  can explain the discrepancy. This is an intriguing region, which the HPS experiment will probe.

## 2.2 Update on Experimental Status

The most recent (as of October 2012) comprehensive update summarizing the experimental status of  $A'$  searches can be found in the presentations and summary talk of the Frascati “Dark 2012” workshop [10]. All relevant measurements and constraints, as of this workshop, are included in Fig. 1. One important change relative to a year ago is that an improved measurement of the Rydberg energy scale has allowed previous measurements of  $g - 2$  of the electron to constrain the allowed parameter space somewhat (in the low  $A'$  mass range) [33, 34]. Additionally, searches for  $A'$ 's in rare  $\phi$  decays at KLOE and rare

$\pi^0$  decays at WASA have slightly reduced the allowed parameter space on the high mass range [78]. Finally, improved theoretical calculations and modeling of the experimental acceptance have led to slightly revised constraints on the  $A'$  parameter space from past beam dump experiments sensitive to  $A'$  production and decay to  $e^+e^-$  pairs [79].

### 2.3 HPS physics with True Muonium

Positronium and muonium, bound states of  $(e^+e^-)$  and  $(\mu^+e^-)$  pairs, respectively, have been produced and studied [81–83], but True Muonium has not yet been detected (see e.g. [84–94]). Together with tauonium  $(\tau^+\tau^-)$  and tau-muonium  $(\tau^\pm\mu^\mp)$ , True Muonium is among the most compact pure QED systems. While  $(\tau^+\tau^-)$  and  $(\tau^\pm\mu^\mp)$  are difficult to detect since the  $\tau$  has a weak decay that competes with the QED decay, the  $\mu$  is very long lived so that the decay of True Muonium is purely a QED process.

The detection of True Muonium would be a significant discovery and would constitute a further important test of QED. A number of applications of True Muonium measurements have been highlighted in [86], designed to exploit True Muonium as a perturbative laboratory for QCD bound state physics. These include measuring dissociation cross-sections as a function of energy and lifetimes of the various states. More speculatively, the discrepancy between theory and experiment for  $g - 2$  of the muon [95] and the discrepant measurement of the charge radius of the proton using muon bound states [96] suggest that further measurements of muon properties would be useful to resolve these puzzles.

Studies of the production and dissociation of True Muonium suggest that the yields in HPS should be sufficient for observation [97], and are discussed further in section 3. That HPS is uniquely suited for detecting True Muonium is straightforward to understand. The triplet True Muonium states  $1^3S_1$ ,  $2^3S_1$ , and  $2^3P_2$  all eventually decay to  $e^+e^-$  final states, with lifetimes long enough to leave a detectable displaced vertex. In that important respect, triplet True Muonium states behave just like  $A'$ s. True Muonium production kinematics is a bit different. In HPS, True Muonium will be produced by electron scattering off the high- $Z$  nuclear target. The so-called “single-photon” production mechanisms gives rise to True Muonium states with kinematics extremely similar to  $A'$ s – HPS will be most sensitive to these. The “three-photon” production mechanism, which is typically larger, gives rise to True Muonium with characteristically lower energy.

In addition to primary production mechanisms, there are a variety of secondary mechanisms that are important in targets thicker than  $\sim 0.01\%$  radiation lengths. This was studied in some detail in [97], where it was shown that  $1^3S_1$  excitations (in the target) are the main source of  $2^3S_1$  and  $2^3P_2$  production. The  $2S$  and  $2P$  state are especially long-lived, so this finding suggests that HPS may first discover these states as they will comprise a sizable fraction of the  $e^+e^-$  decays with displaced vertex in the range of  $\sim 1$  cm to several cm. The  $1S$  state will be the main component of the decays in the region of  $\sim 1$  cm and below.

## 2.4 HPS Searches for Hidden Sectors

As highlighted in the Intensity Frontier Workshop report [8], a well-motivated class of beyond the Standard Model scenarios include new particles that interact indirectly or very weakly with Standard Model matter (hidden sectors), possibly associated with dark matter. Low-energy and high-intensity experiments offer an excellent tool for exploring these possibilities, complementary to the ongoing efforts at high energy colliders.

HPS is primarily designed to look for new sub-GeV  $A'$ s that decay into lepton pairs. But if an  $A'$  is part of a larger hidden sector, as is often assumed in the literature, some fraction of the decays could be more intricate. For example, an  $A'$  might decay into hidden sector particles, which in turn may decay back into Standard Model lepton or photons. These decays would typically have displaced vertices and multiple leptons or photons. The phenomenology of a variety of such scenarios have been considered in [19, 98] (and references therein). Search strategies to look for more general decays of  $A'$ s into hidden sector particles are actively being developed within HPS, and we comment in more detail on this physics opportunity in Section 3.3.

### 3 Proposed Measurements

The primary goal of the proposed experiment is to search for a heavy photon (dark photon) in the mass range from 20 MeV to 1000 MeV in at least two settings of beam energy 2.2 GeV and 6.6 GeV. HPS ultimately relies upon the precision measurement of two quantities: the invariant mass of the  $A'$  decay products and the position of the decay vertex. By placing a tracking and vertexing detector immediately downstream of the target inside an analyzing magnet, the complete kinematic information required for  $A'$  reconstruction can be obtained from a single system, whose proximity to the target naturally maximizes the acceptance of a relatively compact detector and provides excellent momentum and vertexing resolution. A finely segmented, fast electromagnetic calorimeter, just downstream of the tracker, provides a powerful high rate trigger, identifies electrons, and augments the electron energy measurement. A muon system consisting of scintillator hodoscopes sandwiched between iron absorbers is also currently being considered as a potential upgrade. The muon system would provide a trigger for  $(\mu^+\mu^-)$  detection and be used for muon identification. It would extend the search for high mass  $A'$  in di-muon decay mode. Very high rate data acquisition systems, for the tracker, Ecal and muon system, make it possible to trigger and transfer data at 10s of kHz, and run with negligible dead time.

The HPS experiment also has the potential to discover “true muonium”, a bound state of a  $\mu^+\mu^-$  pair and to search for non-minimal hidden sector final states.

HPS plans to execute the full experiment in two phases. The first phase will start with a commissioning run in 2014 which will include data taking for roughly 2 weeks (on the floor) each at 1.1 and 2.2 GeV beam energies. More extensive data taking will continue in 2015, with runs at 2.2 and 6.6 GeV. This first phase will use about one sixth of the total beam time that HPS has requested, roughly 4 weeks on the floor at each of 2.2 and 6.6 GeV. Assuming 50% combined uptime for the accelerator and detector, this corresponds six weeks total of (perfect) run time: 1 week at 1.1 GeV, 3 weeks at 2.2 GeV, and 2 weeks at 6.6 GeV. The second phase of HPS running, which will occur in 2016 and beyond, will use the additional run time to extend the search for heavy photons to the largest possible region of parameter space, and study the properties of true muonium in detail.

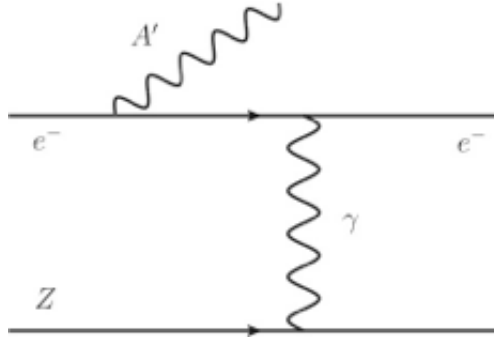


FIG. 2: Diagram of  $A'$  production by bremsstrahlung off of an incoming electron scattering with an atomic nucleus.

### 3.1 Search for the heavy photon

$A'$  particles are generated in electron collisions on a fixed target by a process analogous to ordinary photon bremsstrahlung, see Figure 2. The rate and kinematics of  $A'$  radiation differ from massless bremsstrahlung in several important ways:

- **Rate:** The total  $A'$  production rate is controlled by  $\alpha^3 \epsilon^2 / m_{A'}^2$ . Therefore, it is suppressed relative to photon bremsstrahlung by  $\sim \epsilon^2 m_e^2 / m_{A'}^2$ .
- **Angle:**  $A'$  is dominantly emitted at small angles ( $\theta_{A'}$ ). Near its median value, the cutoff emission angle is

$$\theta_{A',max} \sim \max \left( \frac{\sqrt{m_{A'} m_e}}{E_0}, (m_{A'}/E_0)^{3/2} \right), \quad (1)$$

which is parametrically smaller than the opening angle of the  $A'$  decay product,  $\sim m_{A'}/E_0$ . Although this opening angle is small, the backgrounds mimicking the signal dominate at even smaller angles.

- **Energy:**  $A'$  bremsstrahlung is sharply peaked at  $x = E_{A'}/E_{beam} \approx 1$ . When an  $A'$  is produced, it carries nearly the entire beam energy. In fact, the median value of  $(1-x)$  is  $\sim \max \left( \frac{m_e}{m_{A'}}, \frac{m_{A'}}{E_0} \right)$ .
- **Lifetime** For the ranges of  $\epsilon$  and  $m_{A'}$  probed by this experiment, the mean decay length  $l_0$  of the  $A'$  can be prompt or as large as tens of centimeters. The coupling and

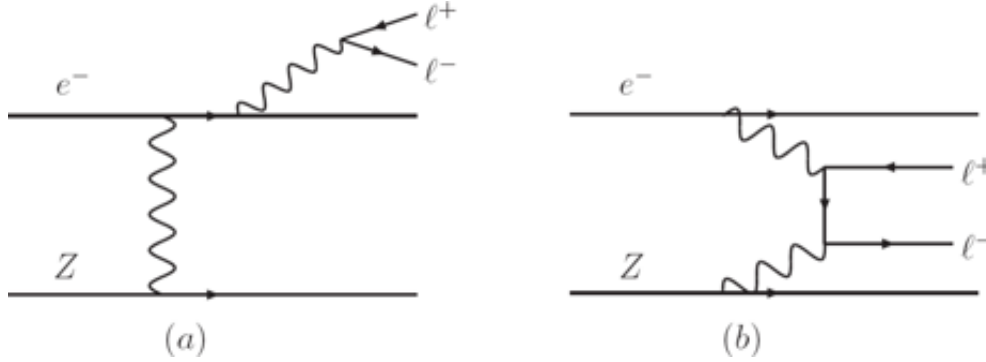


FIG. 3: Sample diagrams of (left) radiative trident ( $\gamma^*$ ) and (right) Bethe-Heitler trident reactions that comprise the primary background to the  $A' \rightarrow l^+l^-$  search.

mass dependence is:

$$l_0 \equiv \gamma c\tau \approx \frac{0.8cm}{N_{eff}} \left( \frac{E_0}{10GeV} \right) \left( \frac{10^{-4}}{\epsilon} \right)^2 \left( \frac{100MeV}{m_{A'}} \right)^2. \quad (2)$$

All of the background decays promptly at the target.

The latter three properties are quite important in resolving signal events from the main backgrounds, as discussed below.

The irreducible background rates are given by the diagrams shown in Figure 3. These trident events can be usefully separated into “radiative” diagrams (Figure 3 (a)), and “Bethe-Heitler” diagrams (Figure 3 (b)), that are separately gauge-invariant. These QED tridents dominate the final event sample, so we consider their properties in some detail here.

The contribution from the radiative diagrams (Figure 3 (a)) alone is also useful as a guide to the behavior of  $A'$  signals at various masses. Indeed, the kinematics of the  $A'$  signal events is identical to the distribution of radiative trident events restricted in an invariant mass window near the  $A'$  mass. Moreover, the rate of the  $A'$  signal is simply related to the radiative trident cross-section within the spectrometer acceptance and a mass window of width  $\delta m$  by

$$\frac{d\sigma(e^-Z \rightarrow e^-Z(A' \rightarrow l^+l^-))}{d\sigma(e^-Z \rightarrow e^-Z(\gamma^* \rightarrow l^+l^-))} = \frac{3\pi\epsilon^2}{2N_{eff}\alpha} \frac{m_{A'}}{\delta m}, \quad (3)$$

Where  $N_{eff}$  is the number of final states that are open to the  $A'$  to decay to. This exact analytic formula was also checked with a MC simulation of both the  $A'$  signal and the radiative trident background restricted to a small mass window  $\delta m$ , and we find nearly perfect agreement. Thus, the radiative subsample can be used to analyze the signal, which

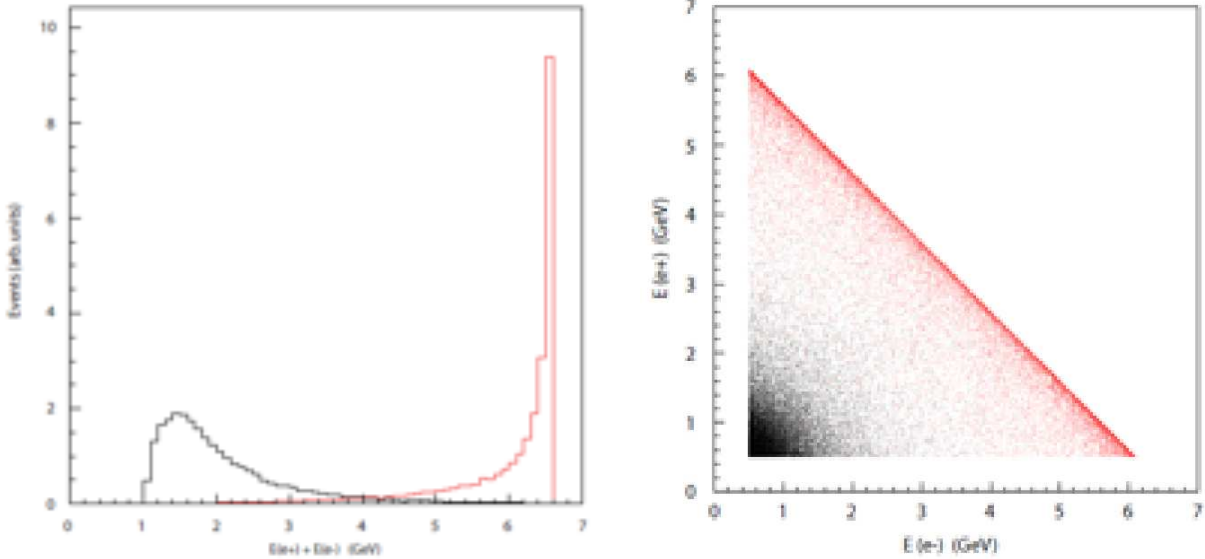


FIG. 4: Left: The distribution of Bethe-Heitler background events (black) and  $A'$  signal events (red) as a function of the sum of the electron and positron energy. Note that the  $A'$  events are peaked at high energies, while the Bethe-Heitler background is peaked at much lower energies. Right: The distribution of the positron versus electron energy for Bethe-Heitler events (black dots) and  $A'$  events (red dots). Note that in both plots the  $A'$  and Bethe-Heitler events are normalized to the same number. In reality, the number of Bethe-Heitler events is orders of magnitude larger (prior to kinematic cuts) than the number of  $A'$  events for couplings that HPS is sensitive to. Even after making a reasonable kinematic cut, the rate of Bethe-Heitler events is much larger than the  $A'$  rate. Also, note that the electron energy here refers to the energy of the electron produced in the reaction, not the recoiling beam electron.

simplifies the analysis considerably.

Although the Bethe-Heitler process has a much larger total cross-section than either the signal or the radiative trident background, it can be significantly reduced by exploiting its very different kinematics. In particular, the  $A'$  carries most of the beam energy while the recoiling electron is very soft and scatters to a wide angle. In contrast, the Bethe-Heitler process is not enhanced at high pair energies. Moreover, Bethe-Heitler processes have a forward singularity that strongly favors asymmetric configurations with one energetic, forward electron or positron and the other constituent of the pair much softer. These properties are discussed further in the Appendix of [4], and illustrated in Figure 4.

The radiative and Bethe-Heitler backgrounds show a smooth, continuous distribution in  $m_{e^+e^-}$  and occur promptly at the production target. The  $A'$ , however, will produce a peak at  $m_{e^+e^-} = m_{A'}$  and, at lower values of  $\epsilon$  have a vertex that is displaced from the target. These two characteristics have led us to design a detector with both good momentum and

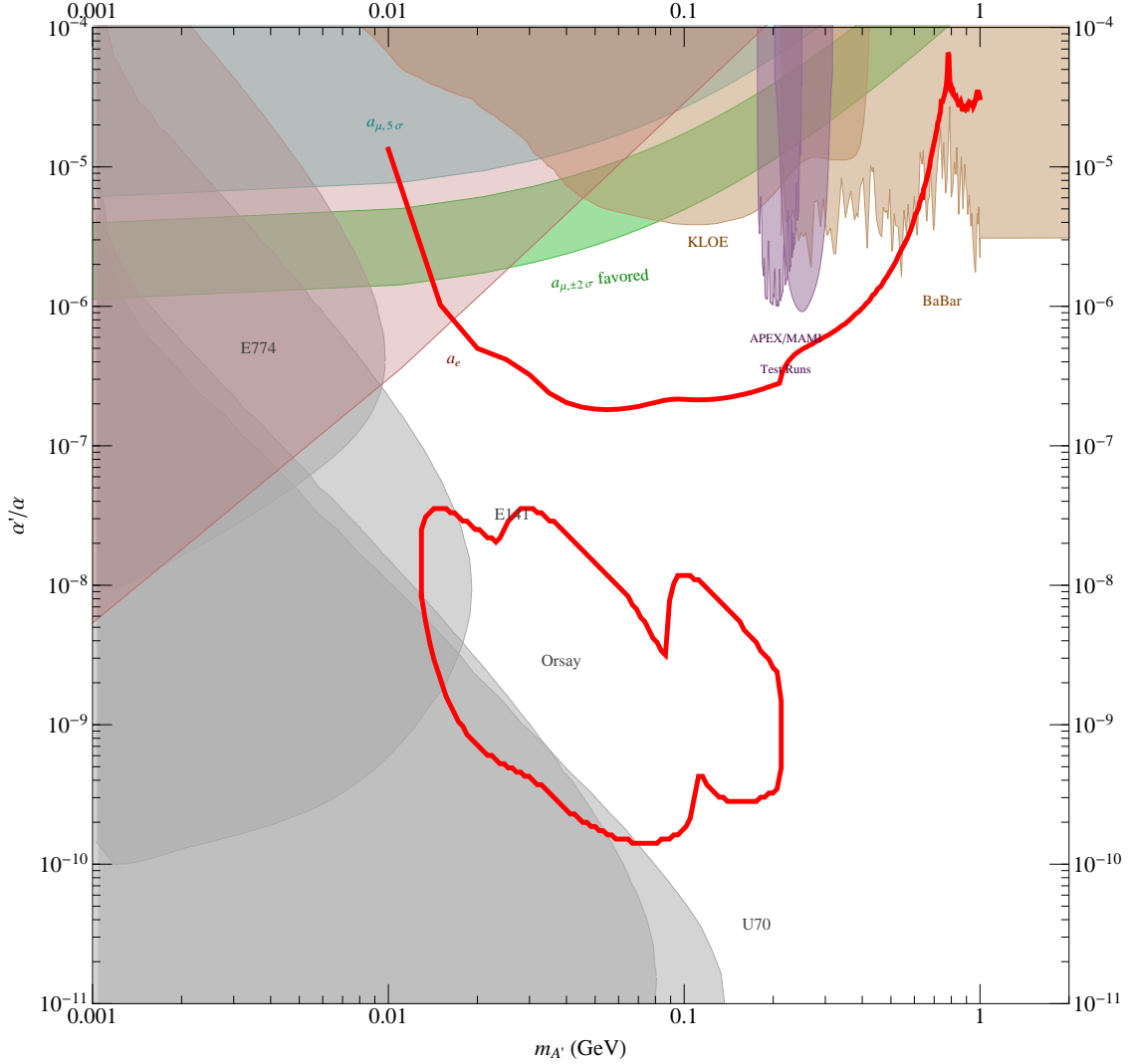


FIG. 5: Expected mass vs coupling parameter space reach full 2014-2015 running (solid red). Red line contour corresponds to 1 week of beam time at 1.1 GeV, and 3 weeks of beam time at 2.2 GeV and 6.6 GeV.

spatial resolution. The expected parameter reach in the first phase of the HPS is shown in Figure 5. The reach in mass-coupling parameter space is calculated using the simulated detector response as shown in Section 6. The plot shows two distinct regions, one at larger coupling corresponding to a purely bump-hunt region and another at lower coupling where the vertex of the  $A'$  decay is displaced.



### 3.2 Search for true muonium

The proposed HPS experiment has the potential to discover “true muonium”, a bound state of a  $\mu^+\mu^-$  pair, denoted here by  $(\mu^+\mu^-)$ . We expect that HPS will discover the 1S, 2S, and 2P true muonium bound states with its proposed run plan. The detection of these states should demonstrate the capability of the HPS experiment to identify rare separated vertex decays, and will provide a natural calibration tool for improving searches for heavy photons. The  $(\mu^+\mu^-)$  atom is hydrogen-like, and so has a set of excited states characterized by a principal quantum number  $n$ . The binding energy of these states is  $E = -1407 \text{ eV}/n^2$ . The  $(\mu^+\mu^-)$  “atom” can be produced by an electron beam incident on a target such as tungsten [84, 85].

With the existing proposal, HPS will search for true muonium just as it does for heavy photons with separated vertices, requiring a vertex cut at about 1.5 cm to reject almost all QED background events, then searching for a resonance at  $2 m_\mu$ . An additional cut on the total energy of the  $e^+e^-$  pair of  $E_{e^-} + E_{e^+} > 0.8 E_{beam}$  will also be required for triggering.

Based on [97], the total production yield for 1S, 2S, and 2P (including secondary production) leaving a target of thickness  $t_b$  (or larger) and satisfying the above requirements is,

$$N_{(\mu^+\mu^-)} = 200 \left( \frac{I}{450 \text{ nA}} \right) \left( \frac{t}{1 \text{ month}} \right) \sim 100 \text{ events} \quad (4)$$

where a beam energy  $E_{beam} = 6.6 \text{ GeV}$ , and the nominal conditions of 450 nA beam current for 2 weeks of beam-time on a single foil has been assumed. The vertices near the cut of 1.5 cm will be dominated by the 1S state, while a tail of vertices extending out beyond a few cm is dominated by 2S and 2P.

Accounting for all the efficiencies associated with a separated vertex search, we would expect to see about 10–20 true muonium events in 2 weeks of 6.6 GeV run (we caution that the acceptance parameterization here is uncertain at the 50% level). The HPS experiment should be able to identify enough events to claim a discovery, and in addition, should be able to measure the mass of true muonium. There are certainly other properties of true muonium that would be interesting to measure. A measurement of the lifetimes would be interesting, as the lifetimes are sensitive to physics that couples to leptonic currents. With enough statistics, it should be possible to perform a measurement of the lifetimes of the 1S,

2S, and 2P states; work is ongoing to investigate this possibility.

### 3.3 Other searches for hidden sector particles

While the primary motivation for the HPS experiment is a search for  $A'$  decaying to lepton pairs, following Arkani-Hamed *et al.*[1], it is useful to explore the sensitivity of HPS to other hidden sector particles, in particular those associated with Hidden Valley (HV) scenarios. As Strassler and Zurek[98] pointed out, HV scenarios provide many natural explanations of Dark Matter. The recent discovery of a boson that seems to have the properties of the Higgs boson of the Standard model also brings an old problem to the fore: why is the Higgs mass so light compare to the Planck scale? Searches at the LHC for supersymmetric partners and other particles with Standard Model couplings have so far been unsuccessful, pushing the range of allowed masses for such particles higher and higher. HPS may be the first in a series of experiments that complement the energy frontier searches by looking for new, light particles that couple to the SM particles through new, very weak forces. Just as one should be ready for the unexpected when exploring the energy frontier at the LHC, so should we be in exploring the coupling frontier at with HPS.

In a general HV scenario, the new fermions may be lighter than the  $A'$  and have their own QCD-like forces, which results in them forming hidden mesons and baryons. Some of these hidden particles may be part of the Dark Matter. These hidden mesons would decay into SM fermion pairs either democratically or with mass-dependent branchings. One interesting

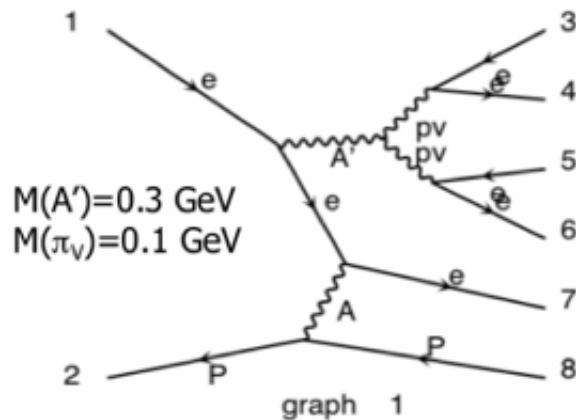


FIG. 6: Sample diagram of a non-Abelian hidden sector interaction.

case is where the  $A'$  decays into a pair of dark mesons ( $\pi_v$ ), which in turn promptly decay into electron pairs (see Fig 6). The high multiplicity of the typical final states makes an exclusive search for events such as these difficult to trigger on, but it also reduces the background significantly by providing a number of new constraints and possible invariant mass bumps to search for and find. Simulation studies are ongoing towards estimating the reach HPS has for such hidden mesons.

## 4 Description of the HPS setup

### 4.1 Overview

HPS will utilize a setup located at the upstream end of experimental Hall-B at Jefferson Lab. The setup will be based on a three-magnet chicane, the second dipole magnet serving as the analyzing magnet for our forward spectrometer. The detector package will include a silicon tracker, an electromagnetic calorimeter, and the necessary readout electronics. The overall design provides for a muon detector as well, which is not costed in the present proposal, but will be added with supplemental funding and/or an upgrade proposal in the future. High luminosities are needed to search for heavy photons with small couplings and masses in the 20 to 1000 MeV range. Utilizing CEBAF's essentially continuous duty cycle, the experiment can simultaneously maximize luminosity and minimize backgrounds by employing detectors with short live times and rapid readout. The HPS setup is designed to run with  $> 200$  nA electron beams at energies from 1.1 GeV to 6.6 GeV impinging on a tungsten target of up to  $0.0025 X_0$  located 10 cm upstream of the first layer of the tracker.

The HPS tracker consists of six double layer planes, 36 microstrip sensors in total. Placing the planes of the tracker in close proximity to the target means that the primary beam must pass directly through the middle of the tracking detector. This has necessitated that the sensors don't encroach on a "dead zone," where multiple Coulomb scattered beam particles and radiative secondaries are bent into the horizontal plane, the so-called "wall of flame." However, since the energy released in the decay of a low mass  $A'$  is small relative to its boost, the opening angle between decay daughters can be quite small. To maximize the acceptance for low mass  $A'$ 's, the vertical extent of the dead zone must be minimized and sensors placed as close as possible to the beam, so our design incorporates precision movers that can bring the silicon detectors to the required positions. Since interactions of the primary beam with air or even helium at atmospheric pressure gives rise to low-momentum secondaries that generate unacceptable occupancies, we have chosen to place the entire tracking and vertexing system in vacuum, in the Hall B pair spectrometer's magnet vacuum chamber. Silicon microstrip sensors are used in the tracker/vertexer because they collect ionization in 10's of nanoseconds and produce pulses as short as 50 – 100 ns. The sensors are read out

continuously at 40 MHz using the APV25 chip, developed for the CMS experiment at the LHC. Running the high speed silicon module readout in vacuum further requires a vacuum compatible cooling system, and data and power vacuum feedthroughs. All these features are incorporated in the design of the apparatus, as described below and have been tested in the May 2012 test run.

The electromagnetic calorimeter (Ecal) consists of 442  $\text{PbWO}_4$  crystals (reconfigured from the CLAS Inner Calorimeter) that are read out with APDs and amplifiers. The short pulse widths allow the ECal to run at very high rates. The Ecal data is digitized in the JLAB FADC250, a 250 MHz flash ADC developed for the JLAB 12 GeV Upgrade program. The full analogue information from the FADCs coupled with the fine spatial information of the calorimeter is available to the trigger, which uses energy deposition, position, timing, and energy-position correlations to reduce the trigger rate to a manageable  $\sim 30$  kHz. Like the tracker system, the electromagnetic calorimeter is split to avoid impinging on the “dead zone.” The beam and radiative secondaries pass between the upper and lower ECal modules, which are housed in temperature-controlled enclosures, needed to stabilize the energy calibration.

In the following, the various elements of the experiment are discussed in more detail, beginning with the beamline and experimental controls, continuing with the tracker/vertexer and electromagnetic calorimeter, then data acquisition and trigger systems, and offline computing at the end. The muon upgrade is described in Appendix B.2.

## 4.2 HPS beamline

The Hall-B beam line, magnetic elements, beam profile monitors, and beam position and current monitors upstream of the HPS setup will be used as is (after slight modifications for 12 GeV). The only modification needed for the upstream part of the beam line is the addition of a collimator upstream of the Hall-B tagger magnet. The role of the collimator is to prevent the beam from directly hitting the Si-tracker in an event in which the high intensity beam can move up or down from its nominal position. The collimator, which can consist of a 1 cm thick tungsten plate with different size oval holes (to match the beam profile), can be incorporated into the Hall-B photon tagger radiator ladder to provide vertical alignment. Horizontal alignment of the whole system will also be needed for fine tuning of the collimator

position relative to the beam. Downstream of the HPS setup, there will be two beamlines, the electron beam line that will transport electron beam to the Hall-B electron beam dump, and a photon beam line that will transport the photon beam generated in the target to a photon beam dump mounted on the space frame. The electron beam will be transported in vacuum all the way through to the beam dump. Following the vacuum chamber in the last chicane dipole, the photon beam will go to the dump in Helium bag. There will be an H-corrector installed on the electron line after the HPS chicane to compensate any possible mis-steering of the beam in the chicane and to make sure that the electron beam stays on the original beamline to the dump. A YAG viewer will be used to monitor beam position just before the dump. The beam position on the dump monitor must stay unchanged before and after energizing the chicane.

#### 4.2.1 Layout of the HPS setup

The HPS experiment will use the same three magnet chicane that was used for the CLAS Two Photon Exchange experiment (TPE). The layout of the beam line and the chicane is shown in Figure 7. The Hall B pair spectrometer magnet, an 18D36 (pole length 91.44 cm, gap  $45.72 \times 15.24 \text{ cm}^2$ , max-field 1.5 T), will serve as the analyzing magnet. The dipole field direction (Y) is perpendicular to the horizontal (XZ) plane. The Hall B “Frascati” H magnets (pole length 50 cm, gap  $21 \times 8.25 \text{ cm}^2$ , max-field 1.2 T) will be used as the first and the last dipoles of the chicane. The analyzing magnet will be operated at a 0.25 T-m field for 1.1 GeV, a 0.5 T-m field for 2.2 GeV, and a 1.5 T-m field for 6.6 GeV running. The two bending magnets will be set to 0.125 T-m, 0.25 T-m, and 0.75 T-m fields, respectively. The distance between the centers of the magnets will be about 50 cm bigger than it was for the TPE run, to accommodate the muon system. The location of the analyzing magnet along the beam will be exactly the same as it was for the TPE run.

The analyzing magnet will be displaced to beam left by  $\sim 11 \text{ cm}$  in order to optimize the detector acceptance for  $e^+$  and  $e^-$ , see Figure 8.

The HPS target is positioned at the upstream edge of the analyzing magnet’s pole. The distance from the target to the first layer of the silicon tracker is 10 cm, and to the face of the electromagnetic calorimeter  $\sim 137 \text{ cm}$ . There is continuous vacuum for the electron beam throughout the entire setup ending in the Hall B electron beam dump. The Si-tracker and

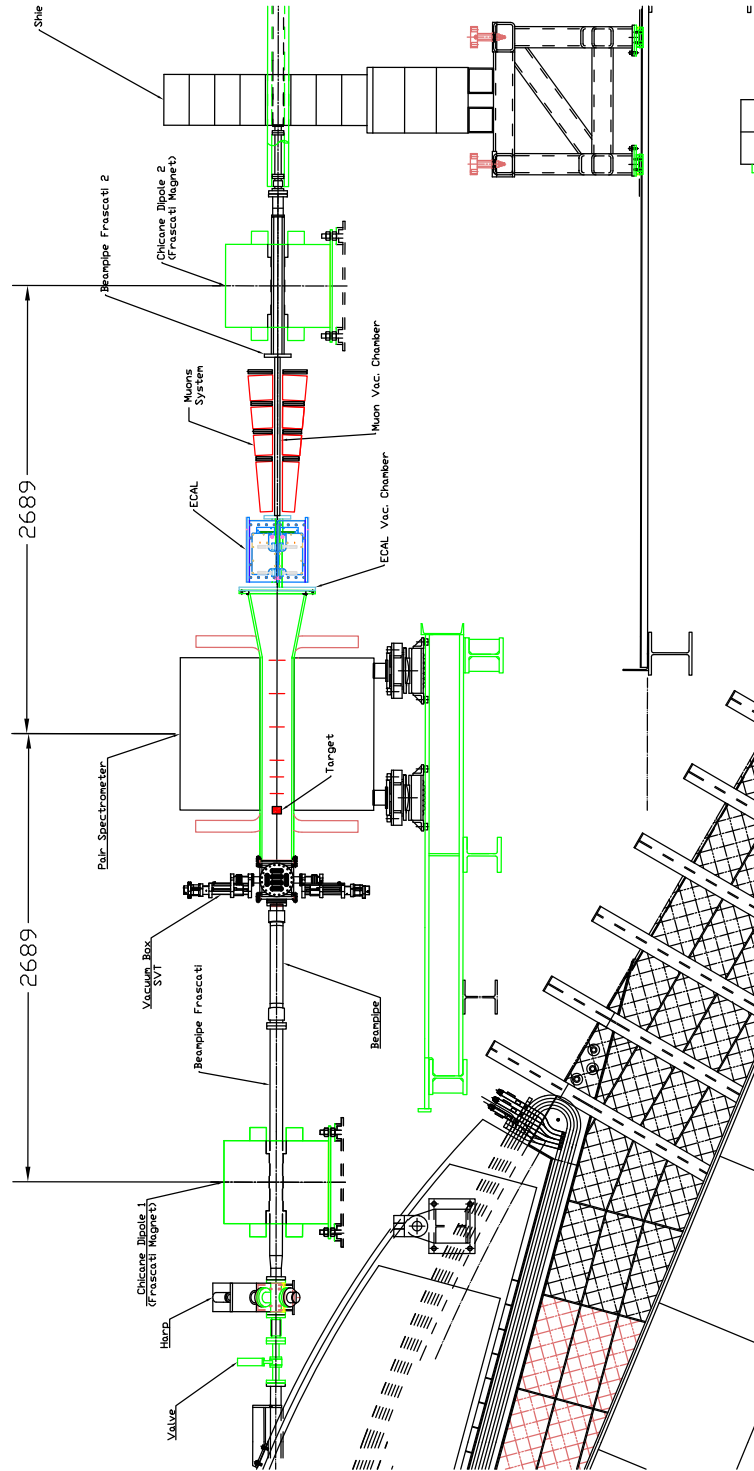


FIG. 7: Beam line configuration for the HPS test run with electron beams. The chicane configuration is similar to a previously run CLAS experiment.

the target will be located inside the Hall-B pair spectrometer vacuum chamber. The SVT vacuum box is mounted on the upstream end of the analyzing magnet vacuum chamber to

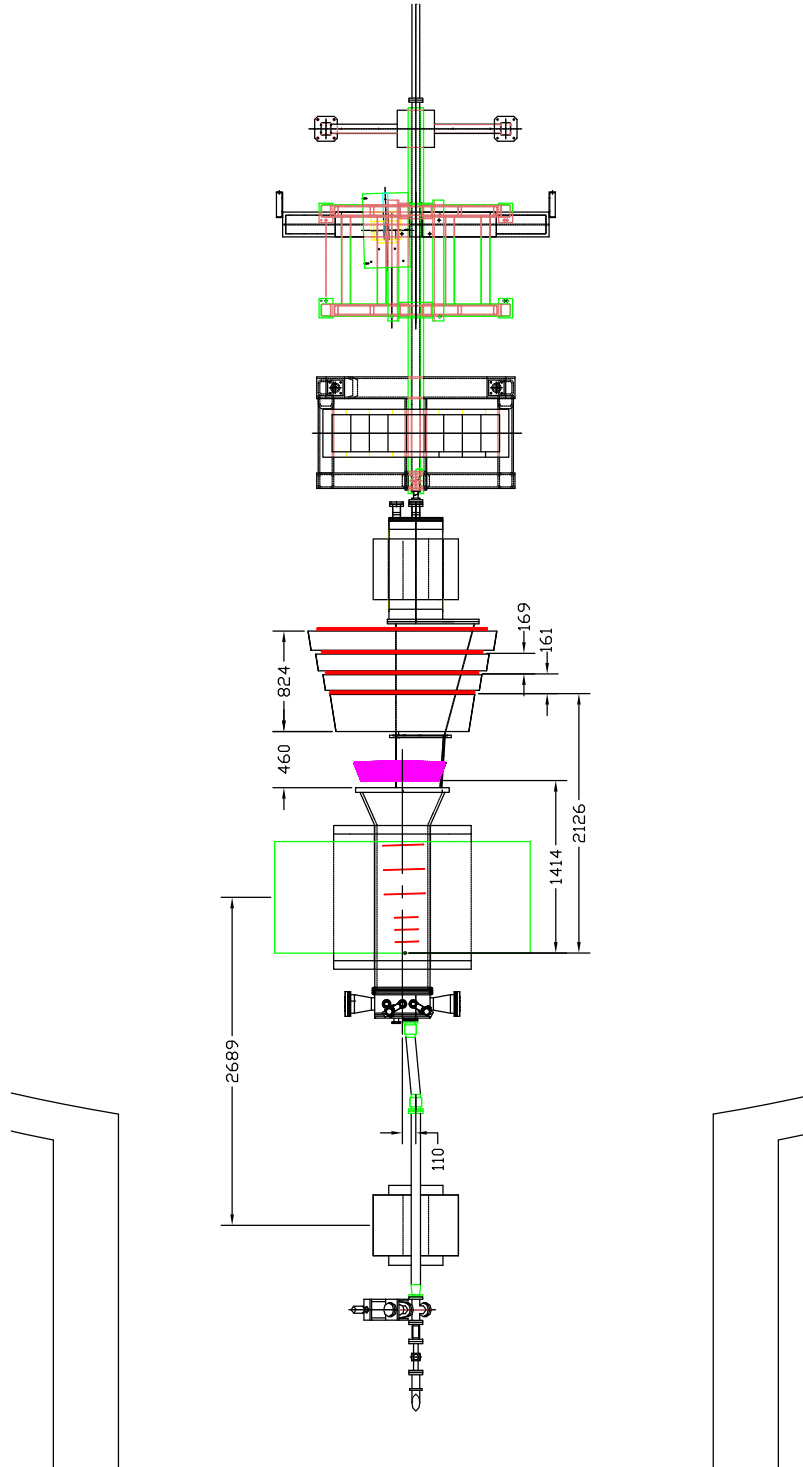


FIG. 8: Top view of the beam line configuration for the HPS test run with electron beams. The analyzing magnet is shifted by 4 inches (110 mm) to beam's left to get optimal acceptance for both  $e^+$ s and  $e^-$ s.



provide connections for the SVT motion system, the cooling system, power and signal cables, and the target motion system. The Ecal vacuum chamber is attached to the downstream end of the analyzing magnet vacuum chamber, above and below which are placed the Ecal modules. Downstream of the Ecal vacuum chamber, another vacuum chamber is attached, leading to the downstream chicane magnet.

The analyzing magnet, the Hall B pair spectrometer dipole, has its own power supply. The “Frascati” H magnets will use one common power supply and will be powered by the Hall B “mini-torus” power supply. There will be a shunt installed between the two “Frascati” magnets to allow independent small changes in currents on those two magnets if necessary (as it was done during the TPE experiment, although never used). Both power supplies are bipolar, so the magnets can be degaussed when needed. From available field map data at 900 A, the  $\int Bdl$  of Frascati H magnets along the path of the electron beam is 0.663 T-m. The specified max current for these magnets is 950 A. In order to get 0.75 T-m an additional 10% increase in field value will be needed. From initial evaluation of the magnet design and characteristics, it should not be a problem to run them at 10% higher currents over their specified max current. If this is not possible, reducing the gap by 1/3 inch can yield the desired  $\int Bdl$ .

### 4.2.2 Running Conditions

The HPS will use  $\sim 1.1$  GeV,  $\sim 2.2$  GeV, and  $\sim 6.6$  GeV electron beams of up to 500 nA incident on a thin tungsten (W) target. Operational experience (with 6 GeV machine) showed that the CEBAF beam is very clean, and is contained within  $\pm 0.5$  mm with halo at the level of less than  $10^{-5}$ . It is expected that the beams from the 12 GeV machine will be of comparable quality, at least for up to 3-pass beams (up  $\sim 6.6$  GeV), so the primary electron beam should cleanly pass through the “dead zone” gap of the HPS setup.

For optimizing the vertexing performance and acquiring physics data, an asymmetric beam profile is desirable. Since the vertex resolution in the non-bend plane will be high, beam sizes of  $< 50 \mu\text{m}$  in the Y direction are preferable. The momentum measurement will not benefit from small beam sizes in the X direction, and if the beam sizes in both dimensions are too small, the target foil will overheat. For these reasons the required beam sizes for HPS will be  $\sigma_X \sim 250\mu\text{m}$  and  $\sigma_Y < 50\mu\text{m}$ . The HPS beam parameter requirements are presented in Table I.

Parameter	Requirement			Unit
E	1100	2200	6600	MeV
$\delta E/E$	$< 10^{-4}$			
Current	$< 200$	$< 400$	$< 500$	nA
Current Instability	$< 5$			%
$\sigma_x$	$< 300$			$\mu\text{m}$
$\sigma_y$	$< 50$			$\mu\text{m}$
Position Stability	$< 30$			$\mu\text{m}$
Divergence	$< 100$			$\mu\text{rad}$
Beam Halo ( $> 5\sigma_Y$ )	$< 10^{-5}$			

TABLE I: Required beam parameters.

The B-line optics in the 6 GeV era was checked using simulation and a beam test of the system. The optics program ELEGANT [99] was used to determine the optimized B-line parameters needed to achieve an asymmetric beam size,  $\sigma_X \approx 250\mu\text{m}$  and  $\sigma_Y \approx 20\mu\text{m}$ , at the HPS test run target location. Beam tests were conducted in Hall B to validate these optics simulations during the Two Photon Exchange experiment when 2.2 GeV beam was available (February of 2011). Parameters were set for a beam profile of  $\sigma_X \approx 300 \mu\text{m}$  and  $\sigma_Y \approx 10\mu\text{m}$  at the Hall B “tagger” beam profiler ( $\sim 8$  meters upstream of the proposed HPS target location). Several beam profile scans with different scanner and data readout speeds

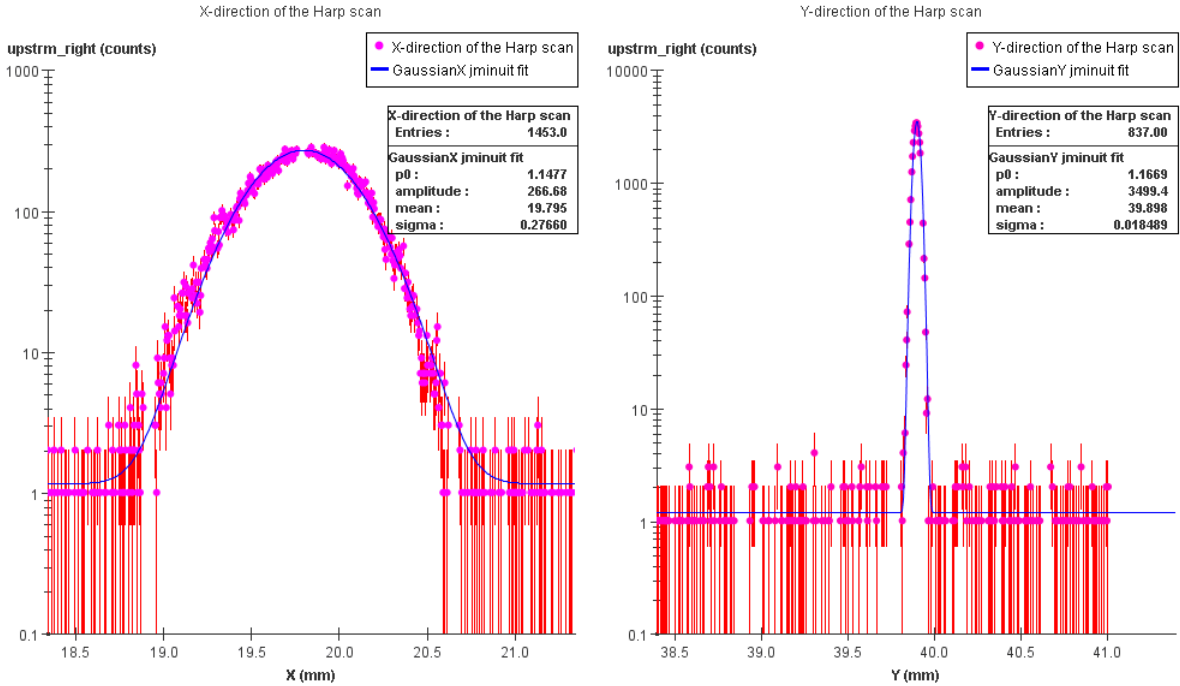


FIG. 9: Wire harp scan after loading optics parameters from the ELEGANT program. The wire scan speed was 0.1mm/s, readout speed is 15Hz. Based on the width of the Y-profile, the beam position stability is  $< 20\mu\text{m}$ . Note: any beam motion with more than  $10\mu\text{m}$  amplitude and faster than 1Hz is included in the scan.

were performed to check the beam stability and the systematics in the measurements. One of the scans is shown in Figure 9. As can be seen from the figure, the required profile can be reliably achieved. Several scans performed over two hours resulted in a consistent and stable beam profile. Based on the width of the Y-profile, beam position stability is  $< 20\mu\text{m}$ . Note that any beam motion with more than  $10\mu\text{m}$  amplitude and faster than 1Hz is included in the scan.

The beamline optimizations have been performed for the 12 GeV CEBAF machine including the proposed changes for Hall-B/CLAS12 operations. Using the program ELEGANT and inputting the new locations of magnetic elements and their field maps, the beam profile was optimized at the HPS target location. In Figure 10 the beam sizes and the beam angles are shown for 6.7 GeV setup. The required beam size is achievable within the operating specifications of all the quadrupoles. Since HPS will run at beam energies  $< 6.7$  GeV it is straight forward to scale (linearly) the magnets down to the other energies. The beam size/angle (beam transport) remains the same for 1.1, 2.2, and 4.4 GeV energies with the exception of the small emittance increase at 6.7 GeV.

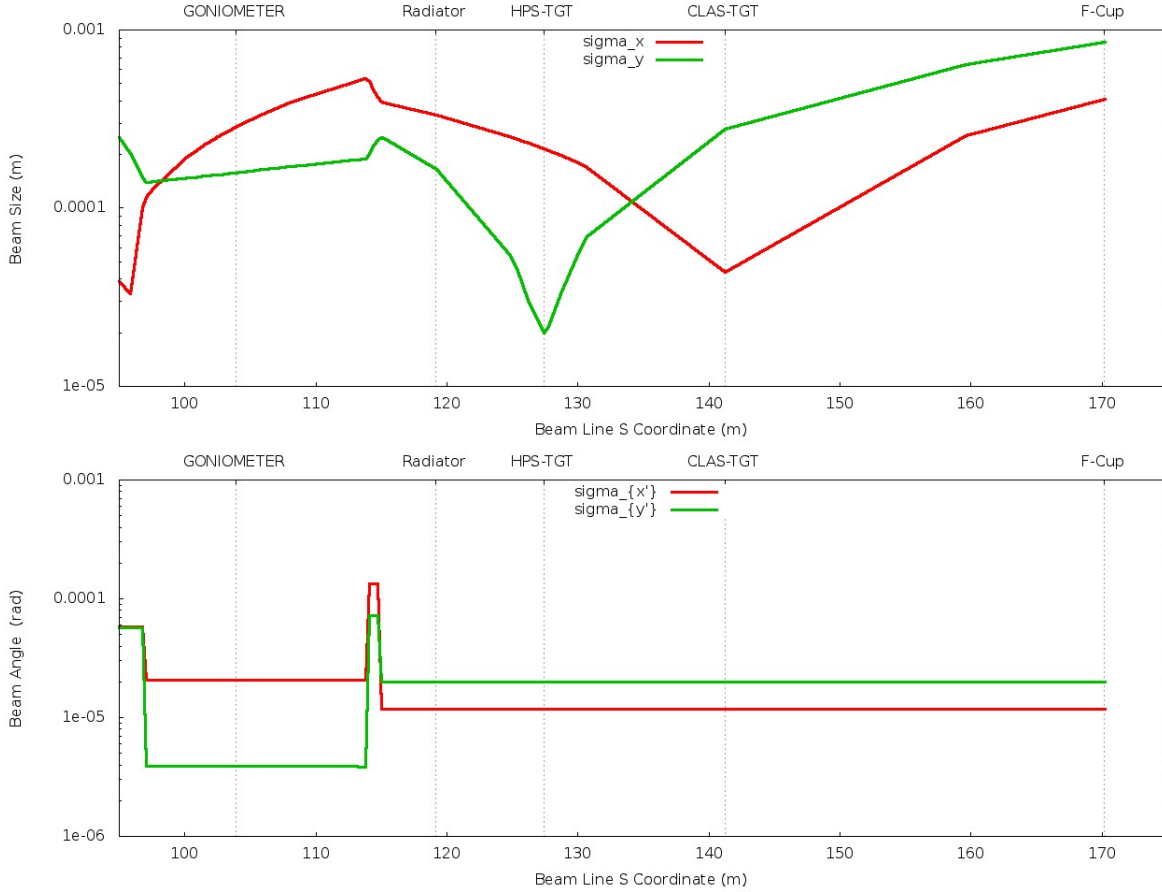


FIG. 10: Beam sizes in X and Y along the B-line in the upstream tunnel and in the region of the HPS test run setup. At the HPS target an asymmetric beam profile  $\sigma_X = 300\mu\text{m}$  and  $\sigma_Y = 20\mu\text{m}$  can be achieved with existing B-line optics.

### 4.2.3 Beam Diagnostics

Beam position and current will be controlled using inputs from two sets of cavity beam position monitors (BPMs), that are located in the upstream tunnel (see Figure 11). Sets of corrector dipoles and quadrupoles are routinely used to tune the beam for Hall B (2C21 to 2C24). A pair of BPMs, 2C21 and 2C24, will define the incoming trajectory of the beam and are included in the fast feedback loop. Readings from these BPMs will be used to maintain stable beam positions and currents. The stability of beam positions at two different locations also ensures the stability of the beam inclination.

The beam profile will be measured using three wire scanners. Two are installed in the tunnel, the first one at 2C23, and the second one before the Hall-B tagger magnet, (2c24 harp, called “tagger harp”), about 8 meters upstream of the HPS target. The third wire

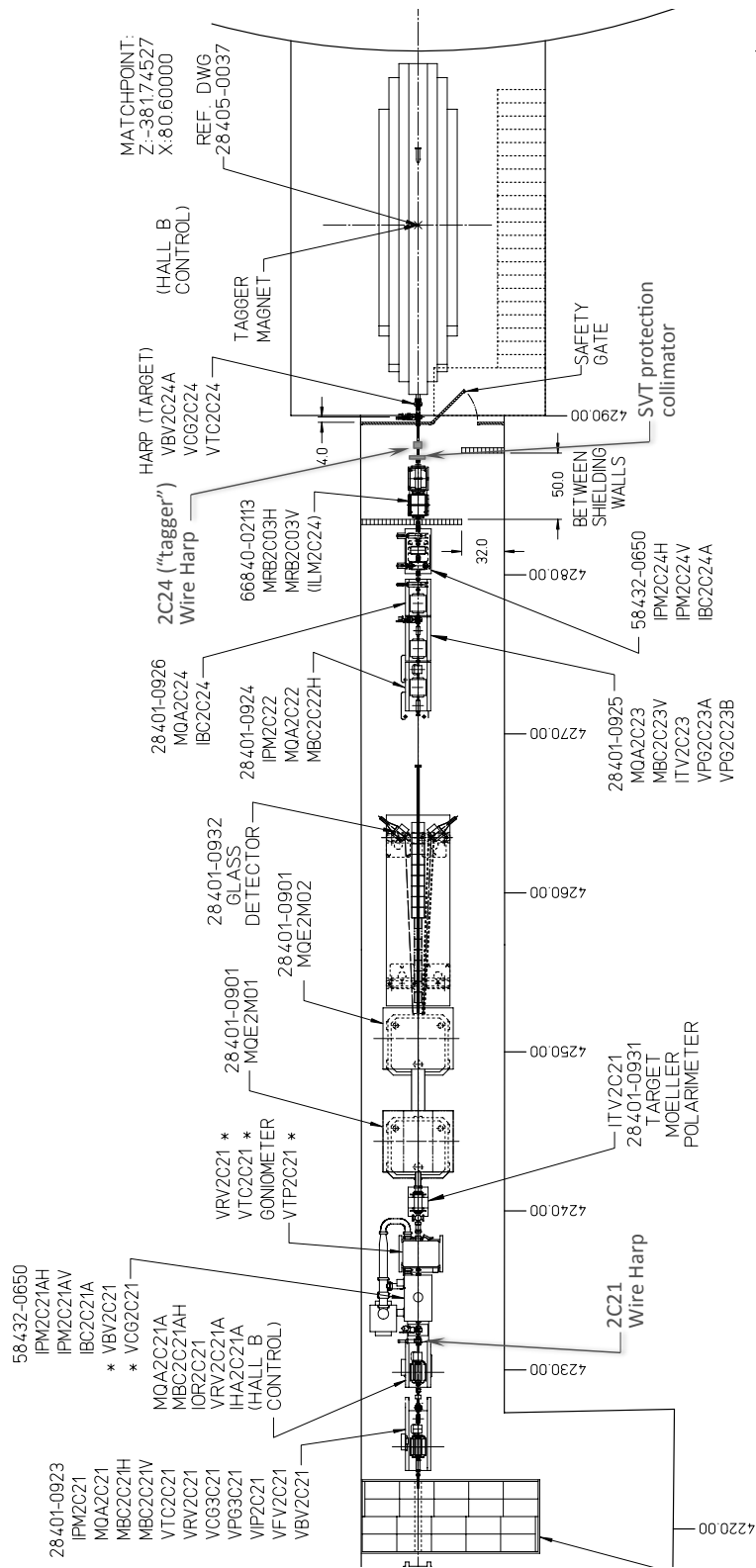


FIG. 11: Upstream beam line configuration for HPS.

harp, 2H00 harp, will be located just before the first chicane dipole. The first two profilers

will be used to establish the required beam parameters during the initial setup. The Hall-B tagger magnet will be energized when beam tune is in progress, diverting the beam away from the HPS apparatus. After an acceptable beam profile is achieved, the tagger magnet will be degaussed and turned off, and the beam will be put through the HPS system and the beam profile will be checked using the 2H00 wire harp. The backgrounds in the HPS silicon tracker from beam profiling using the 2H00 harp have been simulated. At 5 nA beam current, the radiation damage is equivalent to about 10 sec. of production beam current on the HPS target, so is not a concern.

A set of tungsten beam-fiducial wires will be installed immediately in front of the silicon detectors in the experiment's analyzing magnet. One horizontal wire, 20 micron diameter, and one 30 micron wire at 9 degrees to the horizontal, will be mounted on a frame attached to the upper movable silicon support plate, and similarly for the bottom plate. The frames for the wires are wide enough that they do not occlude the silicon active area. The wires can be used to locate the position of the beam relative to the silicon. To accomplish this safely, the vertical separation between the front silicon sensor and its nearest wire is 8 mm. This separation, and the small wire diameters, also means that, when the sensors are positioned for data taking, the wires will have a negligible effect on acceptance. The wires are also available for use as a fairly conventional wire scanner. In particular they can provide information about the minor and major axes, and the tip angle (roll), of a strongly elliptical beam.

An insertable YAG screen beam viewer will be installed in the downstream alcove of Hall-B, before the Faraday cup,  $\sim 40$  meters downstream of the HPS target. Both the position and profile of the beam will be used to setup the chicane magnets and to monitor beam quality during the run. A set of beam halo counters mounted along the beam line provides continuous and fast monitoring of the beam conditions. These counters are like those used for beam profile measurements. Excess noise in the beam halo counters triggers the machine fast shutdown system (FSD) in order to terminate beam in the event of beam excursions which could damage the HPS detectors. The FSD will occur in less than 40  $\mu$ s. In addition to halo counters, a beam offset monitor (BOM) will be installed upstream of the 2H00 wire harp. It is similar to the BOM used in CLAS. A short quartz cylinder, about 6 mm OD and 4 mm ID, with optical fibers attached around the edge will be centered on the beam. Even a few electrons in the beam tail will generate light in the cylinder that

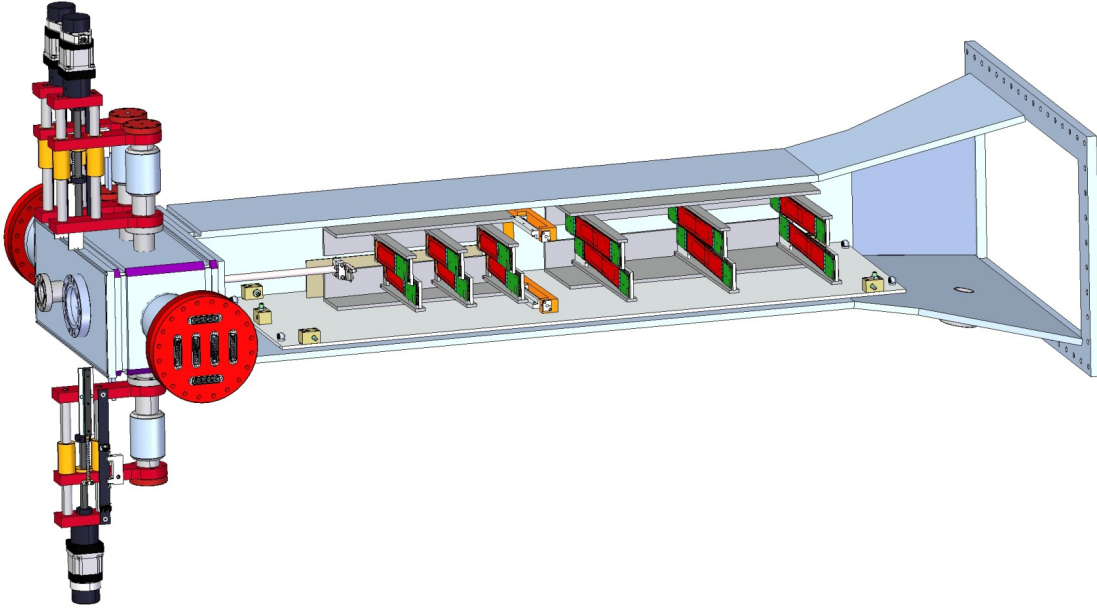


FIG. 12: Rendering of the SVT inside the Hall-B pair spectrometer vacuum chamber and the upstream vacuum box with SVT and target connections.

will be detected in a multi-anode PMT attached to the readout fibers. Errant beam motion towards the collimator located upstream of the tagger magnet will generate more light and increase the counts in the quartz cylinder, signalling a potential problem. The BOM will be wired to FSD as part of the equipment protection system.

#### 4.2.4 Vacuum chambers

The SVT vacuum box will be attached to the existing magnet vacuum chamber as shown in Figure 12. Power, high voltage, and data signals to and from the hybrids are connected through two 8" flanges on the sides of the vacuum box. Two vertical linear motion mechanisms driven by stepper motors are used to position the SVT upper and lower modules with a precision of  $1.25 \mu\text{m}/\text{step}$ . A third linear motion mechanism is used to position the target on or off the beam. All the stepper motors are placed at a large enough distance from the magnet to avoid any ill effect from the magnetic fringe field. An existing stepper motor driver and EPICS-based control software will be used.

The scattering chamber between the top and bottom parts of the ECal is a critical beamline element. In order to keep the calorimeter as close as possible to the beam plane,

include sufficient thermal insulation for the ECal, and maintain as wide a vacuum gap as possible, the top and bottom plates of the scattering chamber must be quite thin. At the location where the primary beams ( $e^-$  and  $\gamma$ ) exit, the openings in the chamber have been enlarged. In Figure 13 a rendering of the scattering chamber in between the two halves of the ECal is shown. The front flange of the chamber connects directly to the magnet vacuum chamber. Vacuum is maintained only on the electron side (beam right) since the backgrounds on the positron side are negligible. This design is based on detailed GEANT4 simulations of the background rates and acceptance of the ECal. It places crystals within 20 mm from the beam plane to maximize low-mass acceptance. In order to avoid excessive deformation of the thin walls of the vacuum chamber, an aluminum honeycomb support is inserted between the upper and lower walls, to beam's right.

The ECal vacuum chamber will be connected to the muon system vacuum chamber, located between the two halves of the muon system. Special openings for the photon and electron beams are not needed in the muon system vacuum chamber. The gaps for the radiated secondary electrons are essentially projections of those in the ECal vacuum chamber. At its upstream end, the muon vacuum chamber will have a gap of  $\sim 5$  cm. At the downstream end that gap will be  $\sim 7.5$  cm.

The last vacuum chamber, which passes through the third dipole, does not need to have a narrow opening. It will have size of the Frascati H magnet gap. At the downstream end of this chamber, there will be flange with two exit windows, a Kapton window for the photon beam to exit the chamber and go to the photon beam dump through a Helium bag, and a vacuum continuation to the standard beam line for the electron beam to go to the Hall-B electron beam dump.

#### 4.2.5 Beam dumps and shieldings

The Hall-B electron beam dump will be used to terminate the electron beam. Due to its high intensity, the beam will not be dumped in the Faraday cup (FC); instead, the existing beam blocker before the FC will be used to terminate the beam. The photon beam will be dumped in a photon beam dump, which will be a hut made of lead bricks located on the space frame. There will be a shielding wall after the last chicane magnet to prevent radiation from reaching the detector systems on the downstream side of the Hall.



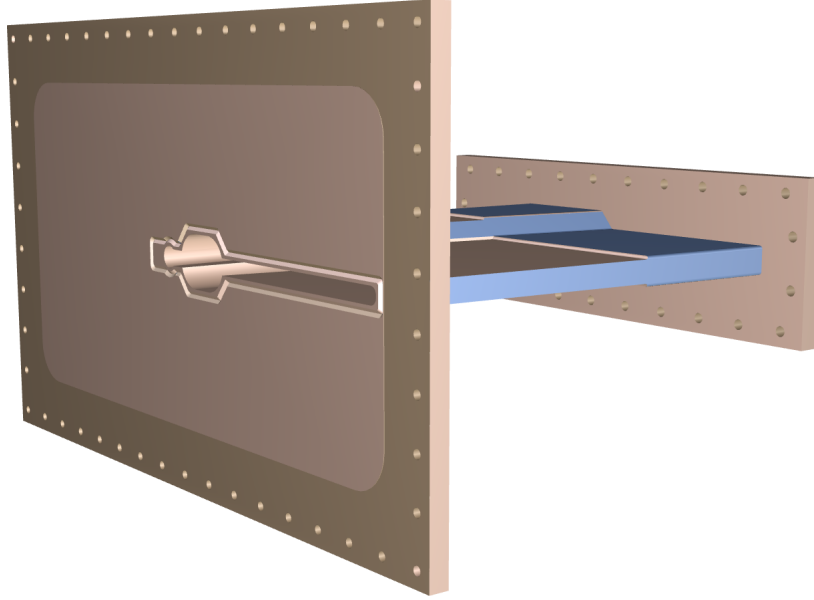


FIG. 13: Rendering of the ECal and the ECal vacuum chamber.

#### 4.2.6 Targets

A thin tungsten foil is used as the target. High  $Z$  material is chosen for its short radiation length, to minimize the hadronic production relative to the electromagnetic trident and  $A'$  production. The target is located 10 cm in front of the first plane of silicon strip detectors.

The primary target, 10 mm square, is 0.00125 radiation lengths (approximately  $4 \mu\text{m}$  tungsten). Mounted immediately above it is a similar area of 0.0025 radiation lengths, available for some of the data taking, adjusting the beam current as appropriate. The foil can be fully retracted from the beam, and is inserted on to the beam line from above, using a stepping motor linear actuator. The bottom edge of the foil is free-standing so there is no thick support frame to trip the beam when the target is inserted. Its position is adjustable vertically allowing either thickness to be selected, and different sections of the tungsten can be used in the event of beam damage. The support frame on the beam-right side of the target is made thin enough to prevent radiation damage to the silicon in the event of an errant beam caused, for example, by an upstream chicane magnet trip.

The target is intended to operate with beam currents up to 500 nA, which produce strong local heating. The strength of tungsten drops by an order of magnitude with temperature increases in the range of 1000 C. In addition, the material re-crystallizes above this range, which increases the tendency for cracking where thermal expansion has caused temporary

dimpling. For these reasons, it was decided to keep the temperature rise less than about 1000 degrees, which is accomplished by selecting an adequately large beam spot area. For example at 200 nA the rms beam radii will be held above 20 by 250  $\mu\text{m}$ , or 40 by 250  $\mu\text{m}$  for 400 nA. Simulations have shown that these beam spot sizes do not diminish the pair reconstruction resolution of the experiment.

## 4.3 Slow Controls

### 4.3.1 Framework

As any complex high energy and nuclear physics experiment HPS needs a sophisticated experimental controls system which will incorporate the existing slow controls components in Hall B, as well as newly developed subsystems. The slow controls for HPS will be based on the Experimental Physics and Industrial Control System (EPICS) framework. This choice is primarily driven by the necessity to be compatible with the existing Hall B and CEBAF slow controls systems. It will allow us to control hardware used in the experiment from the counting house over the Ethernet, and we will benefit from numerous existing software components developed by the EPICS community, such as alarm systems, archivers, and setpoint save-and-restore software. EPICS Input/Output Controllers (IOC's) will run on the rack-mountable Linux servers in the rack-room of the Hall B counting house and on the VME controllers in the Hall running either VxWorks or Linux operating systems. These IOC's will serve EPICS process variables to multiple clients over Ethernet using ChannelAccess (CA) protocol. The operators in the counting house will be able to monitor and change the parameters of the controlled hardware whenever necessary using CA-clients that have direct access to the slow controls network of Hall B. For control and alarm screens HPS will use Controls System Studio (CSS) framework which will also be used by the CLAS12 slow controls group. CSS provides a modern Java-based display management tool called BOY that allows the developer to quickly crate very dynamic control screens. We also currently plan to use the alarm framework developed by the controls group of Spallation Neutron Source at ORNL and is already integrated with CSS. The HPS experiment control system includes multiple components such as control and monitoring of SVT, ECal, and the muon system's power supplies, beamline components and temperature monitoring. Each of these subsystems needs to have an EPICS interface to be controlled remotely from the counting house during the running. The subsections below briefly describe the main HPS slow control components.

### 4.3.2 Voltage controls

Operation of the HPS silicon detectors requires bias voltages and three different types of low voltages: DVDD, AVDD and V125 (see Sec. 4.4 and Ref. [5]). Bias voltages will be adjustable up to 500 V, while the other three will be fixed at preset values. All four types of power will be provided by two different models of Wiener MPOD based boards: the MPV8008 8-channel low voltage boards and the EHS F205-F 16 channel high voltage boards. Because HPS silicon detectors use 36 hybrids and each hybrid requires three low voltage channels and one high voltage channel, we will need total of 14 MPV8008 boards and four EHS F205-F boards. The MPV8008 boards allow for voltage regulation at the desired sensor locations to account for the voltage drop due to the current draw in the low voltage circuits. Bias voltages can only be regulated only at the terminals, but their current draws are negligible. Wiener MPOD chassis can be remotely controlled and monitored over Ethernet using standard SNMP protocol. EPICS support for these boards and chassis already exists and has been successfully used at Jefferson Lab. It allows for automatic generation of the list of EPICS variables by reading out the content of the MPOD chassis during IOC initialization stage. An EPICS application running on a rack-mountable Linux server will continuously monitor and modify the parameters of the boards using SNMP protocol. The set-points and readback values of the parameters will be archived using EPICS archiver developed and maintained by the the controls group of Jefferson Lab Accelerator division. HPS-specific control and alarm screens will be developed similar to those used during 2012 HPS test run in Hall B, see Figure 14.

The power for APD's of the ECal and for the PMTs of the muon system will be provided by the CAEN SY1527 high voltage chassis. CAEN A1520 boards will power the ECal APDs while CAEN A1535 boards will provide high voltage to the PMTs of the muon system. CAEN SY1527 chassis can be remotely controlled over Ethernet using a proprietary protocol. EPICS interface for the SY1527 chassis already exists and it has been successfully used at Jefferson Lab and other laboratories world-wide. The EPICS IOC for the ECal and muon voltage control will run on a rack-mountable computer in the counting house and serve the voltage-related process variables to other ChannelAccess clients. High voltage control screens will be designed similar to those used for CLAS IC and EC voltage control GUIs. The alarm system will be configured to alert the shift personnel and annunciate deviation



FIG. 14: Voltage and temperature control screens during 2012 test run in Hall B. Similar control screens will be developed for 2014 run using CSS BOY display management framework.

of the HV channels readback from its demand voltage values as well as of the trip state of the channels set by the chassis itself.

### 4.3.3 Motor controls

The HPS experiment will use eight applications requiring motion control. Three of them are the standard Hall B PMT-based beamline harps at 2C21, 2C24 and 2H00 locations, and the controls and analysis software for them will not change. HPS will use the “long radiator” ladder of CLAS to align the SVT collimator with the electron beamline as described in Sec. 4.2, and only minor modifications to the existing software will be required to control the position of the collimator. Another Hall B piece of equipment that was used by a number of experiments is the beam blocker that is inserted into the beam line to protect the Faraday cup in the electron beam dump from excessive beam currents. The controls hardware and software for the beam blocker will be identical to that used by CLAS.

HPS will require development of three additional stepper motor-based applications. First, as it was discussed in Sec. 4.2.3, the two mounting frames of the silicon detectors will hold four wires for SVT alignment with respect to the electron beam. Second, the two thin tungsten targets used in HPS will be changeable depending on the beam and run conditions. The target positions will be controlled by a single stepper motor, while the positions of the upper and lower modules will be independently controlled by two separate stepper motors.

All motors used by HPS will be powered by Oregon Micro Systems (OMS) PMD4 drivers housed in custom-designed chassis that contain four PMD4 boards. The chassis provide connections for the power, limit and home switches and, the encoder. The computer interface is provided by OMS VS4 VME boards, each capable of controlling four motion axes. This is the standard configuration that was used for all of the motor-based applications in Hall B. Therefore, HPS does not need to develop software for basic motor control with EPICS, and will simply take advantage of the existing framework used by CLAS.

Higher tier controls and online analysis software required for the SVT wire scans will be developed on the basis of the existing software for the Hall B harp scans. The relative position of the silicon detectors with respect to the beamline will be determined by slowly moving a wire in the transverse direction across the beam and by measuring at different wire positions the scaler rates from the beam halo counters mounted around the beam-pipe downstream of the silicon detectors. The speed of the wire, the motion range, and the sampling rate will be tuned during the commissioning of the experiment. The scan software will perform an online analysis of the acquired data and the resulting plots from each scan will look similar to the plots presented in Fig. 9. Using the locations of the peaks in the counting rates determined by a fit to a Gaussian distribution, one should be able to optimize the alignment of silicon detectors with respect to the electron beam.

#### 4.3.4 Magnet controls

During the HPS run in 2014 we plan to use existing magnets that were used in previous Hall B experiments: the tagger magnet to bend the electrons into the tagger dump during beam tuning; Hall B pair spectrometer dipole magnet as the analyzing magnet, and two H-type chicane magnets. As described in Section 4.2.1, all four magnets have fully operational power supplies with controls hardware and software already installed. Therefore, no

additional effort will be required to be able to control the HPS magnets.

#### 4.3.5 Temperature control and monitoring

Both ECal and SVT readout electronics will need temperature stabilization, therefore both components will have cooling systems requiring two separate chillers. ECal will utilize the existing chiller that was used for CLAS Inner Calorimeter and that does not need an external control. For the SVT cooling system HPS will purchase a new chiller which will allow for a remote control and monitoring through a serial connection. HPS slow controls group will develop EPICS support for the new SVT chiller to be able to monitor its status and change the temperature setpoint remotely during the experiment.

The temperature monitoring of ECal will use thermocouples mounted on the ECal electronics. HPS slow controls group will develop EPICS readout software for these thermocouples based on the prototype system used during the 2012 HPS beam test. The SVT temperature monitoring will be done using the internal temperature sensors read out by the SVT data acquisition system. An EPICS IOC will serve records to hold the SVT temperature values while a client on the SVT DAQ will regularly update the values of these records. In addition, the temperature at the entrance and the exit of the vacuum chamber will also be measured using thermocouples similar to those used for ECal which will allow us to continuously monitor the SVT ambient temperature independent of the SVT DAQ status.

In order to safely operate silicon detectors a number of interlocks will need to be set similar to those used during 2012 HPS beam test. Because the silicon detectors will operate in vacuum it is crucial to maintain coolant flow through the system when the silicon detectors are powered. The flow of the coolant has to be turned off in the event of deteriorating vacuum, as a precaution to prevent any potential fluid leakage into the vacuum chamber. If the readings for the temperature or pressure in the vacuum chamber exceeds its respective threshold value, the power to the SVT and the flow in the SVT cooling system will be shut off using a hardware interlock system. The mass flow of the coolant, the vacuum pressure, and the status of the interlock components will be continuously readout and monitored via EPICS.

## 4.4 Silicon vertex tracker

Achieving the best possible acceptance at low  $A'$  mass requires positioning the silicon as close as possible above and below the primary beam, where radiation and occupancy are both limiting factors. As a result, the silicon must be actively cooled to retard radiation damage, hits must be read out quickly and tagged with the best possible time resolution to reduce effective occupancies, and the tracker must operate in a vacuum to eliminate beam-gas secondaries. At the limit of feasibility from these considerations, the silicon in the first layer is only 0.5 mm from the center of the beam, so prudence dictates that the tracker be retractable from the beam. Meanwhile, achieving the best possible acceptance at high mass means that the active volume of the tracker fill as much of the magnet bore as possible. Finally, both the mass and vertex resolution that determine the experimental sensitivity are dominated by multiple scattering effects, so minimizing material in the tracker is the principal design goal. A more complete discussion of the key requirements and design principles of the HPS Silicon Vertex Tracker (SVT) may be found in the initial proposal to the JLab PAC [5].

The HPS Test Run SVT, described and discussed in Chapter 5, achieved these goals with the minimal possible apparatus capable of delivering  $A'$  physics during a short run. Unlike the initial proposal for the full experiment, this design used a single type of silicon microstrip module with small stereo angle arranged into five layers and compromised redundancy, precision, and longevity in order to compress the project timeline and reduce the budget. In the process of developing this design, it was found that this simple system was capable of delivering a surprising fraction of the physics potential anticipated for the full experiment. With this in mind, we now propose a design for the HPS SVT that builds upon the HPS Test SVT, principally by addressing the compromises made for HPS Test to ensure the best possible performance for  $A'$  physics within the envelope of the existing beam line layout and analyzing magnet. This design uses the exact same sensors and readout chips, retaining the most successful elements of the HPS Test design and addressing the weaknesses identified during assembly and operation to ensure the success of the experiment.



#### 4.4.1 Layout

The layout of the HPS SVT is summarized in Table II and shown in Figure 15. There are six measurement stations, or “layers,” placed immediately downstream of the target. Each layer comprises a pair of closely-spaced planes and each plane is responsible for measuring a single coordinate, or “view.” Introduction of a stereo angle between the two planes of each layer enables three-dimensional tracking and vertexing.

Layer number	1	2	3	4	5	6
nominal $z$ , from target (cm)	10	20	30	50	70	90
Stereo Angle (mrad)	100	100	100	50	50	50
Bend-plane resolution ( $\mu m$ )	$\approx 60$	$\approx 60$	$\approx 60$	$\approx 120$	$\approx 120$	$\approx 120$
Non-bend resolution ( $\mu m$ )	$\approx 6$	$\approx 6$	$\approx 6$	$\approx 6$	$\approx 6$	$\approx 6$
Number of sensors	4	4	4	8	8	8
Nominal dead zone in $y$ (mm)	$\pm 1.5$	$\pm 3.0$	$\pm 4.5$	$\pm 7.5$	$\pm 10.5$	$\pm 13.5$
Module power consumption (W)	6.9	6.9	6.9	13.8	13.8	13.8

TABLE II: Layout of the HPS SVT. The angle of stereo sensors is relative to the axial sensors that have their readout strips parallel to the edge of the dead zone.

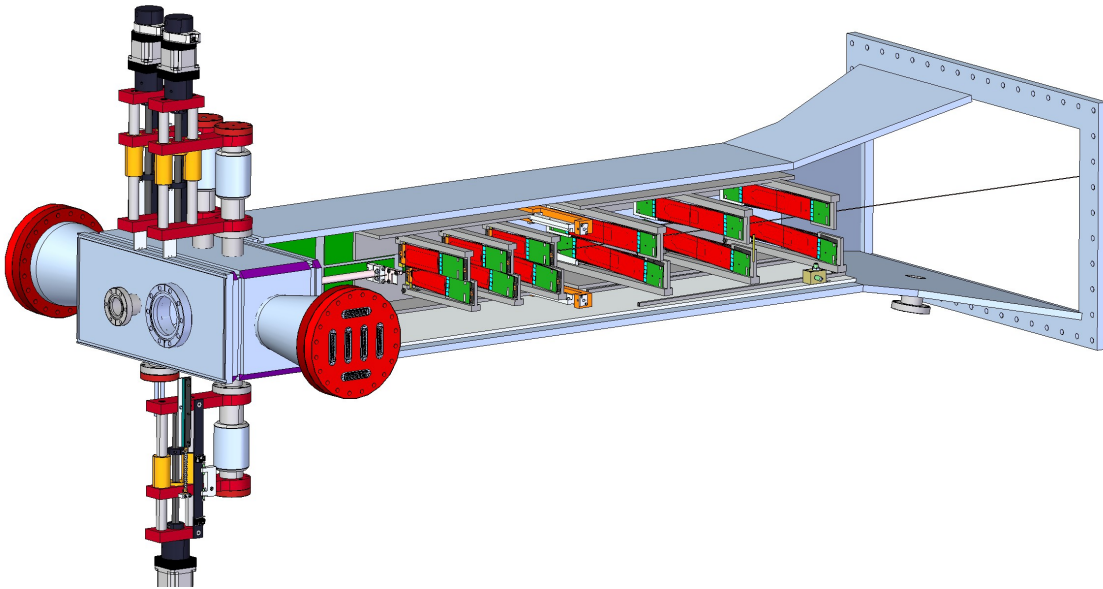


FIG. 15: A rendering of the concept for the HPS SVT. The beam enters from the left through the vacuum box providing services to the SVT. The silicon is shown in red and the hybrid readout boards in green.

The layout of the first three layers is the same as in the HPS Test SVT, with a single sensor of coverage both above and below the beam and 100 milliradian stereo angle to balance acceptance against the resolution required for vertexing. The last three layers are two sensors wide in the bend direction to better match the ECal acceptance and use 50 milliradian stereo, as in HPS Test, to break the tracking degeneracy that creates fake tracks from ghost hits in layers with the same stereo angle. The first five layers cover the full acceptance of the ECal with one redundant layer. The sixth layer has only slightly less acceptance than the ECal and results in an extra safety factor in tracking purity and improved momentum resolution for the vast majority of tracks. Altogether, this layout comprises 36 sensors and 180 readout chips for a total of 23040 readout channels.

Acceptance for larger  $A'$  masses is limited by the size of the magnet but low-mass sensitivity depends on reconstructing tracks as close to the primary beam as possible; minimizing the so-called “dead zone” surrounding the degraded primary beam in the mid-plane of the detector. For the tracker, there are a number of considerations including proximity to beam halo and radiation damage in the first layer, ability to resolve time-overlapping hits, and the wall of pattern recognition errors at very high occupancies. With sensors capable of operation to  $1.5 \times 10^{15}$  1 MeV neq., readout with single-hit resolution of  $\sim 2$  ns and two-hit resolution of  $\sim 50$  ns, and pattern recognition robust to 2% occupancy; the closest tolerable position of Layer 1 results in tracking acceptance outside of a 15 mrad dead zone around the beam plane. In this configuration, the edge of the silicon in Layer 1 is 0.5 mm away from the center of the beam where occupancy from beam-gas curlers would be unacceptable. Therefore, along with the drive to minimize multiple scattering errors, the desire to maximize low-mass acceptance creates the principal design challenges for the SVT: a lightweight, movable, liquid-cooled tracker with superior time resolution and operated in vacuum.

#### 4.4.2 Module Design

One strength of the HPS Test SVT is exceptionally low material budget, an average of 0.7%  $X_0$  per double-sided layer in the tracking volume with only 10% of this from support material. The HPS Test sensor modules achieve this figure by compromising the straightness, mechanical stability and cooling of the sensors. The module design for HPS aims to maintain this material budget but eliminate these compromises by retaining the design of the carbon

fiber half-modules but mounting them in a more robust way. Furthermore, this design allows the existing half-modules built for HPS Test to be reused for the first three layers of HPS, enabling the development of this module concept and the assembly of HPS to commence immediately and with a very small initial investment.

A half module for the HPS Test SVT consists of a single microstrip sensor and a hybrid electronic readout board glued to a polyimide-laminated carbon fiber composite backing, as shown in Figure 16. A window is machined in the carbon fiber leaving only a frame around

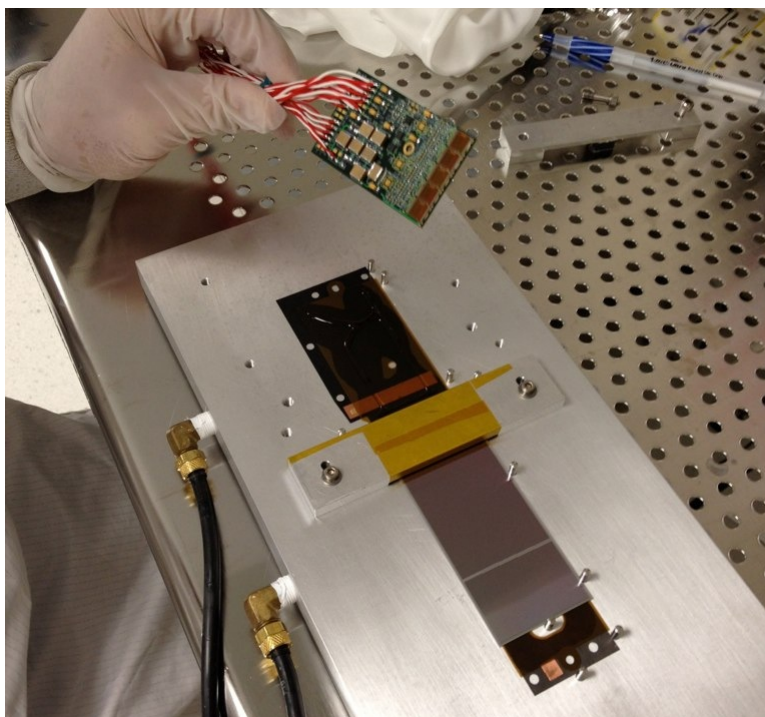


FIG. 16: A half module being assembled for Layers 1-3 of the HPS SVT. A silicon sensor and an APV25 readout hybrid are being glued to the Kapton laminated carbon fiber support frame.

the periphery of the silicon to minimize material. A 50  $\mu\text{m}$  sheet of polyimide is laminated to the surface of the carbon fiber with 1 mm overhang at all openings to ensure good isolation between the back side of the sensor, carrying high-voltage bias, and the carbon fiber which is held near ground. The sensors are single-sided, radiation tolerant,  $p^+$  in  $n$  bulk, AC coupled, poly-biased sensors fabricated on  $\langle 100 \rangle$  silicon manufactured by the Hamamatsu Photonics Corporation for the RunIIb upgrade of the  $D\phi$  silicon detector. The sensors have 60(30)  $\mu\text{m}$  readout(sense) pitch and are read out by APV25 chips operating in multi-peak mode, allowing single hit position resolution of  $\sim 6 \mu\text{m}$  and tagging of individual hit times with a precision of approximately 2 ns. Although specified to achieve 350 V bias, approximately

90% of sensors break down above 1000 V, sufficient to operate layer one of the SVT with full efficiency for six months at full beam intensity.

For HPS Test, the half-modules were sandwiched back-to-back around an aluminum cooling block at the hybrid end and a similar PEEK spacer block at the other. To allow for module rework, the modules were assembled with hardware and thermal compound instead of adhesive and have no stiffening material between the two sensors. For simplicity, only the cooling block at the hybrid end is supported, resulting in deviations in the planarity of the sensors as large as a few hundred microns at the cantilevered end. Furthermore, the lack of cooling at the unsupported end where there is no heat source from readout electronics limits the temperature achievable at the most highly irradiated portion of the sensor, a very small spot along the edge of the sensor adjacent to the dead zone. Improved cooling is necessary to achieve the longevity required for the longer running time envisioned for the HPS detector.

For layers 1-3 of HPS, these same half-module structures will be tensioned, like bicycle spokes, between a pair of cooling blocks held by a grooved aluminum base, as shown in Figure 17. Rather than building manifolds to provide cooling to these blocks and attempt

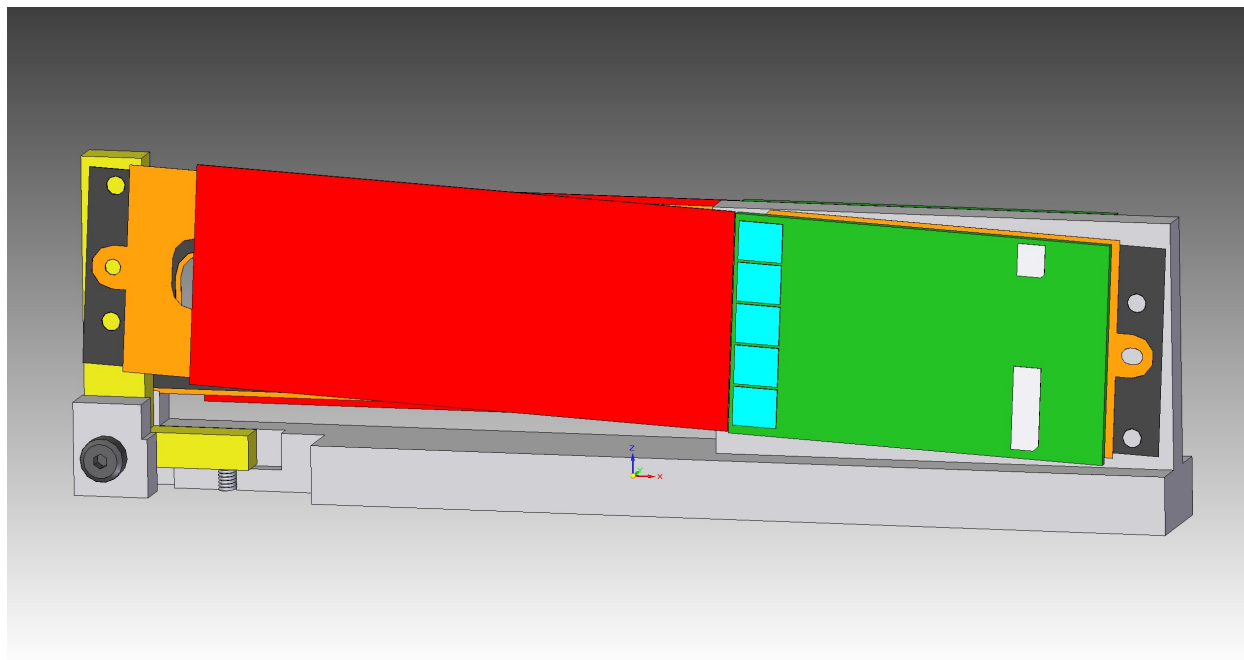


FIG. 17: A rendering of the concept for the new modules for Layers 1-3 of the HPS SVT. A cutaway at the left shows the spring and lever mechanism that maintains tension on the carbon fiber of the half modules.

to isolate them thermally from the underlying support structure as in HPS Test, the entire

aluminum support structure will be cooled. The block at the hybrid end of the module is fixed, while the other pivots on a shoulder screw with a small stainless compression spring providing tension of approximately 40 N and taking up the 60 micron differential contraction between the half module and the support structure during a 30 °C cool down. The same low-viscosity thermal compound used in HPS Test will provide the thermal contact in the pivot joint between the grooved base and the pivoting block and generates a negligible temperature drop across the gap. This arrangement results in much flatter silicon, much better mechanical stability and much better cooling that provided by the previous design for layers 1-3. With temperature stability during running better than 1 °C, dimensional stability of the tracker will be better than intrinsic measurement resolutions and more than an order of magnitude better than the resolution limitation from multiple scattering effects.

More importantly, this scheme allows the ultra-thin design to be extended to the longer, double-sensor half-modules of layers 4-6 that have a pair of silicon sensors in the middle and a readout hybrid at each end, as shown in Figure 18. With a larger heat load to transmit,

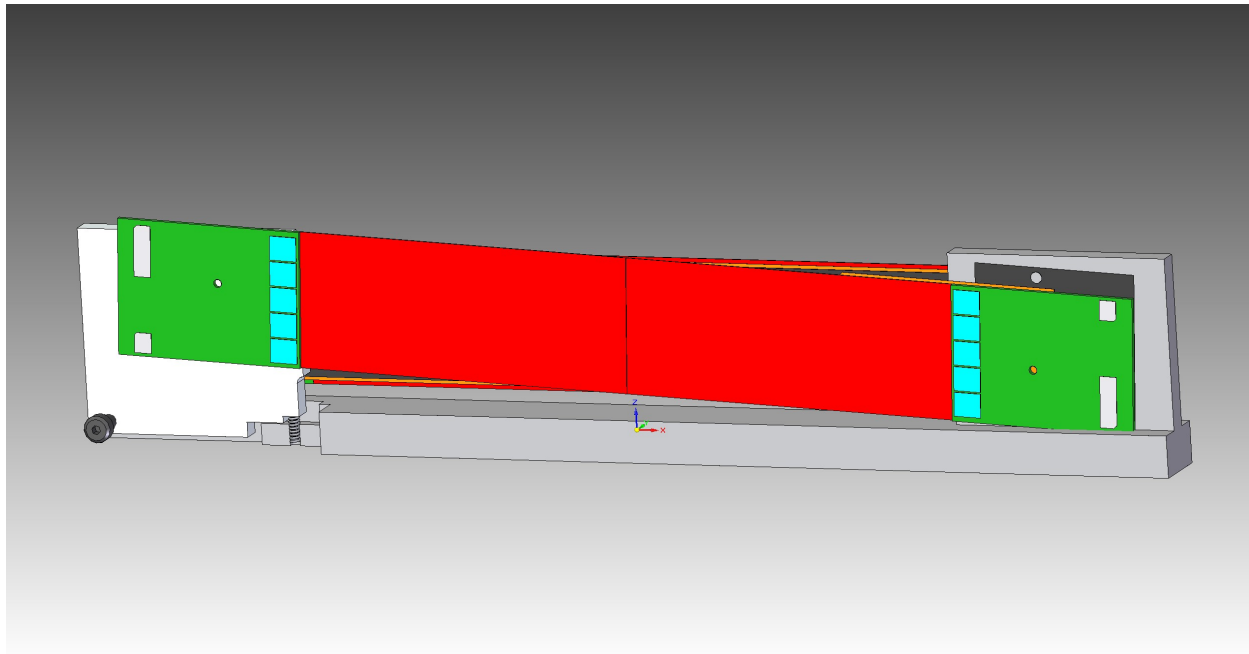


FIG. 18: A rendering of the concept for the new modules for Layers 4-6 of the HPS SVT. A cutaway at the left shows the mechanism responsible for keeping the half modules under tension.

the temperature drop through the pivot joint of the moving block will be approximately 0.5 °C. To accommodate the length of these double modules across the width of the vacuum chamber, the hybrids will be shortened by approximately 20%. Through the use of polyimide

flex cables instead of twisted pair and the elimination of superfluous circuitry, this footprint can be achieved with little effort. Aside from this, only minor changes to the conceptual design of the half module concept of HPS Test are envisioned to reduce assembly effort.

#### 4.4.3 Mechanical Support, Cooling and Services

Sag of the aluminum support plates, when subjected to the bending load of the long motion levers, was the largest source of mechanical imprecision in the HPS Test SVT. For HPS, this motion system will be reused but with changes to eliminate this weakness. First, only layers 1-3 will retract, reducing the length of the support plate by a factor of two. Replacing the twisted pair readout with flex cables eliminates the largest external load on the plate. Finally, the 0.5 inch plate will be replaced by a 0.25 inch plate with 0.25 inch sides, forming a u-channel for increased stiffness, as shown in Figure 19. The walls of this

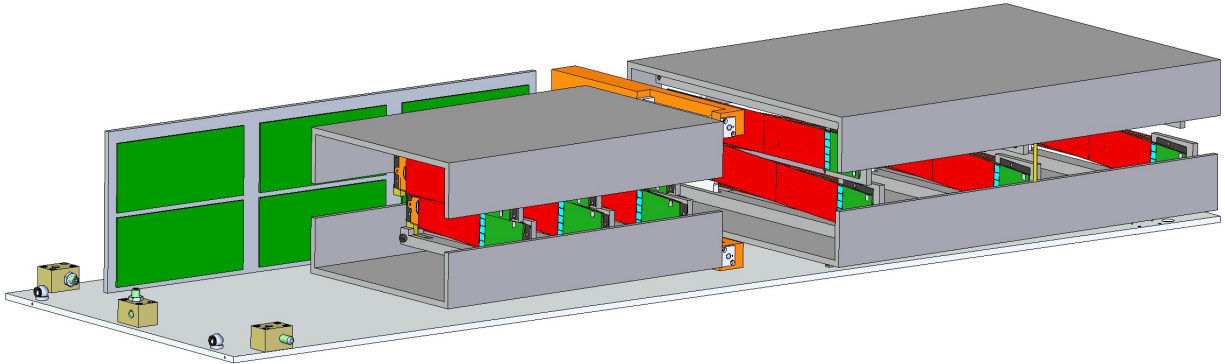


FIG. 19: A rendering of the support concept the HPS SVT. The modules are mounted in cooled channels. The channels for Layers 1-3 pivot on a downstream “C-support” and are moved by lever extending upstream to linear shifts on the vacuum box, as in HPS Test. Layers 4-6 are fixed in place. The DAQ boards, shown in green, are mounted to a separate cooling plate located in a low-neutron region upstream on the positron side of the detector.

support channel will extend almost to the dead zone and the entire structure will be cooled by large, embedded copper tubes. Surrounding the modules over most of the solid angle, these support channels will nearly eliminate thermal radiation from the walls of the vacuum chamber, the primary heat load on the sensors. Layers 4-6 will be mounted inside similar cooled channels but will be fixed since they are already far enough from the dead zone for safety during beam tuning. These two support structures will be mounted to a single

baseplate as before, with complete adjustability relative to the vacuum chamber.

Readout and power for HPS Test required 30 conductors per hybrid, or a total of 600 lines, with even that count requiring elimination of sense for DVDD on the hybrids. It does not appear feasible to scale this solution by nearly a factor of two for HPS. Instead, we plan to provide digitization and optical readout of the APV25 data as well as shared power for the hybrids on boards placed inside the vacuum chamber, as described in Section 4.6.1. In consideration of the DAQ requirements and the environment inside of the vacuum chamber, it appears that there is a natural location for support and cooling of the necessary boards adjacent to layers 1-3 on the positron side. This structure consists of a single vertically-oriented plate with an embedded cooling loop, shown in Figure 19. By separating the readout boards on the outside of this plate by the same 20 cm separation of layers 4-6, a single cable solution can be employed to connect the hybrids of layers 4-6 to these boards. This leaves layers 1-3 equidistant from the pair of readout boards on the inner side, where a second cable type can connect to the existing pigtail connectors for those hybrids. The feedthroughs required for power and data in this design fit easily on one of the two flanges in the existing vacuum box, leaving the other flange for additional cooling, eliminating the need for a cooling manifold inside of the vacuum chamber.

## 4.5 Electromagnetic Calorimeter

The Ecal, depicted in Figure 20, consists of 442 lead-tungstate  $\text{PbWO}_4$  crystals with avalanche photodiode (APD) readout and amplifiers. The short output pulse widths permit operation at very high rates. Indeed,  $(\text{PbWO}_4)$  crystals with APD readout are ideally suited to deal with the expected high radiation and high rate environment and they can operate in the fringe field of the HPS magnetic field as well. The lead-tungstate modules, see Figure 21, are taken from the Inner Calorimeter (IC) of the JLab CLAS detector, which was built by IPN Orsay (France) and other Hall B collaborators and was used in a series of high energy electroproduction experiments. Orsay played a key role in the design and fabrication of the support frames, thermal enclosure, amplifiers, and motherboards of the IC, and is playing a similar role with the HPS ECal. The  $\text{PbWO}_4$  crystals are 16 cm long and tapered. The cross section of the front face is  $1.3 \times 1.3 \text{ cm}^2$ , and  $1.6 \times 1.6 \text{ cm}^2$  at the back. Modules in the ECal are arranged in two formations, as shown in Figure 20. There are 5 layers in each formation; four layers have 46 crystals and one has 37. The ECal is mounted downstream of the analyzing dipole magnet at the distance of about 137 cm from the upstream edge of the magnet. The two ECal modules are positioned just above and below the ECal vacuum chamber, through which the beam, radiated photons, and the wall of flame will pass unimpeded. The innermost edge of the crystals is just 2 cm from the beam. In order to stabilize the calorimeter's performance, the crystals, APDs, and amplifiers are enclosed within a temperature controlled environment, held constant at the level of  $1^\circ \text{ F}$ . The energy resolution of the system, expected from operational experience with the IC, is  $\sigma_E/E \sim 4.5\%/\sqrt{E}$  (GeV). As in the IC,  $\text{PbWO}_4$  modules are connected to a motherboard that provides power and transmits signals from individual APDs and amplifier boards to the data acquisition system. The ECal data is digitized using the JLAB FADC250, a 250 MHz flash ADC developed for the 12 GeV Upgrade. Pulse height information and spatial and timing information from each crystal are available for the trigger decision, which uses this information to reduce the trigger rate to a manageable  $\sim 30 \text{ kHz}$  (see Section 4.6.3 for details).

The HPS calorimeter described above was built and used during the HPS test run in April-May of 2012. This was the first time that a readout and trigger system utilizing the JLAB FADC250s had been used in a real experiment. The trigger algorithm was designed to



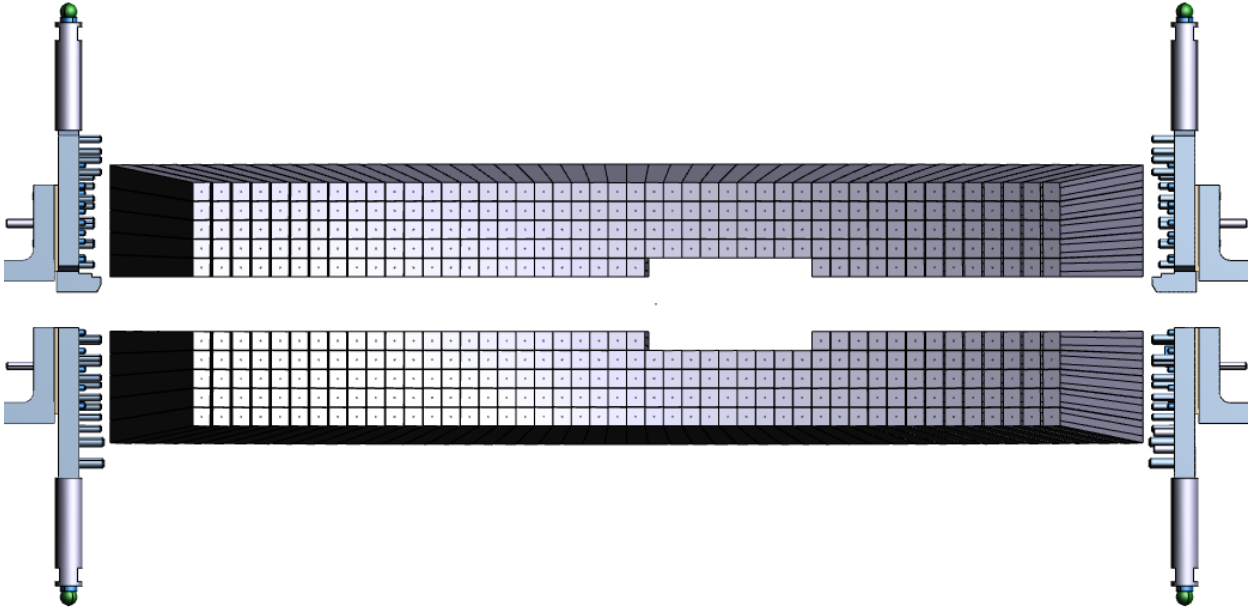


FIG. 20: Arrangement of ECal crystals. The two modules are positioned above and below the beam plane. Each module has 5 layers. There are 46 crystals in each layer, with the exception of the layers closest to the beam plane in which 9 crystals are removed to allow a larger opening for the outgoing electron and photon beams.

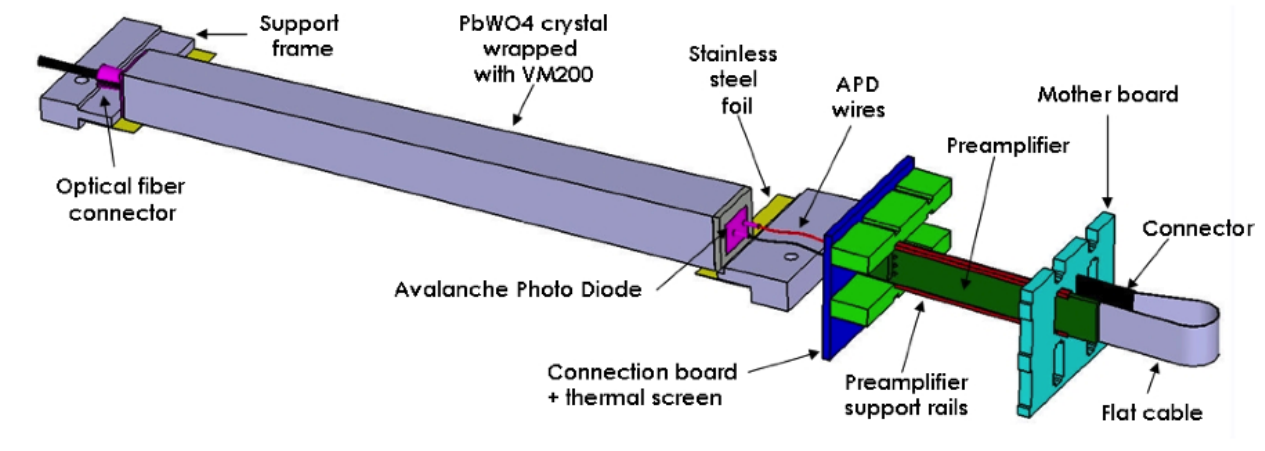


FIG. 21: The ECal module is composed of a 16 cm long lead-tungstate crystal, Avalanche Photo Diode, and a amplifier board.

satisfy the HPS event selection criteria and was implemented with newly developed FPGA-based trigger processors. Not all aspects of the trigger system were tested in the HPS Test Run because of the low interaction and background rates associated with photon running. But the ECal and its readout performed well and the critical goals for the Test Run run were achieved (see Section 5 for details). While the ECal performance during the test run

was satisfactory, several aspects are in need of improvement, as described below.

#### 4.5.1 Improvements to the existing calorimeter

There are no plans for major mechanical changes to the ECal proper. The crystals, support frames, and thermal enclosure operated as designed and will stay unchanged. Most of the needed changes are related to the signal readout chain and problems encountered in the Test Run with the ECal and FADC250. Plans for modifications and/or improvements are as follows:

1. **Replacing the ECal mounting system** - During the test run the ECal was supported by the Hall-B pair spectrometer mount rails which also support the pair spectrometer hodoscopes. Photon running prevented the installation of the ECal vacuum chamber and relaxed the requirements for precise ECal alignment. Consequently, the ECal was simply hung from the mount rails using long threaded rods. This system must be replaced with a more robust and finely adjustable (both horizontally and vertically) support mechanism to align the ECal correctly with the ECal vacuum chamber.
2. **Modification of motherboards** - One of the issues faced during the test run was excess noise on some channels of the motherboards. Most missing channels, those which are absent on the performance figures in Section 5, were in fact switched off because they were very noisy and there was no time to debug them. New motherboards will be designed and built to resolve these noise issues. One of options under discussion is to replace long motherboards with shorter ones with power and signal connectors located on the top (for the top module) and the bottom (for the bottom module) of the thermal enclosure.
3. **Signal splitting** - From the experience gained with the JLAB FADC250 by the HPS group and others, it is evident that the FADCs can be used both for precision time measurements and as real-time scalers simply by developing the appropriate firmware. Precise pulse timing will allow tighter coincidence windows and lower backgrounds, and real time scalers will provide good online monitoring of detector performance. The present ECal readout configuration uses signal splitters to divide the preamplifier

signal in a 2:1 ratio, sending  $2/3$  of the signal to a discriminator that has built-in scalers and feeds the TDC channel. The other  $1/3$  of the signal goes to the FADC for the energy measurement. Removing the splitter will increase the signal on the FADC input by  $\times 3$ . This will allow use of a new lower gain amplifiers with much improved noise level. This effort will make use of newly developed readout system for the CLAS12 Forward Tagger calorimeter. This development is a joint venture of the IPN Orsay and INFN Genova groups, which are members of the HPS Collaboration.

## 4.6 Trigger and DAQ

The HPS experiment data acquisition (DAQ) handles the acquisition of data for the three sub-detectors: the SVT, ECal and the Muon System. HPS employs two DAQ architectures: the SVT is readout with Advanced Telecom Communications Architecture (ATCA) hardware while the ECal and Muon System use VXS based hardware. The trigger system receives input from the ECal and Muon System, and distributes a trigger signal to all detector sub-systems to read out a selected event. Figure 22 gives a schematic block diagram of the DAQ system.

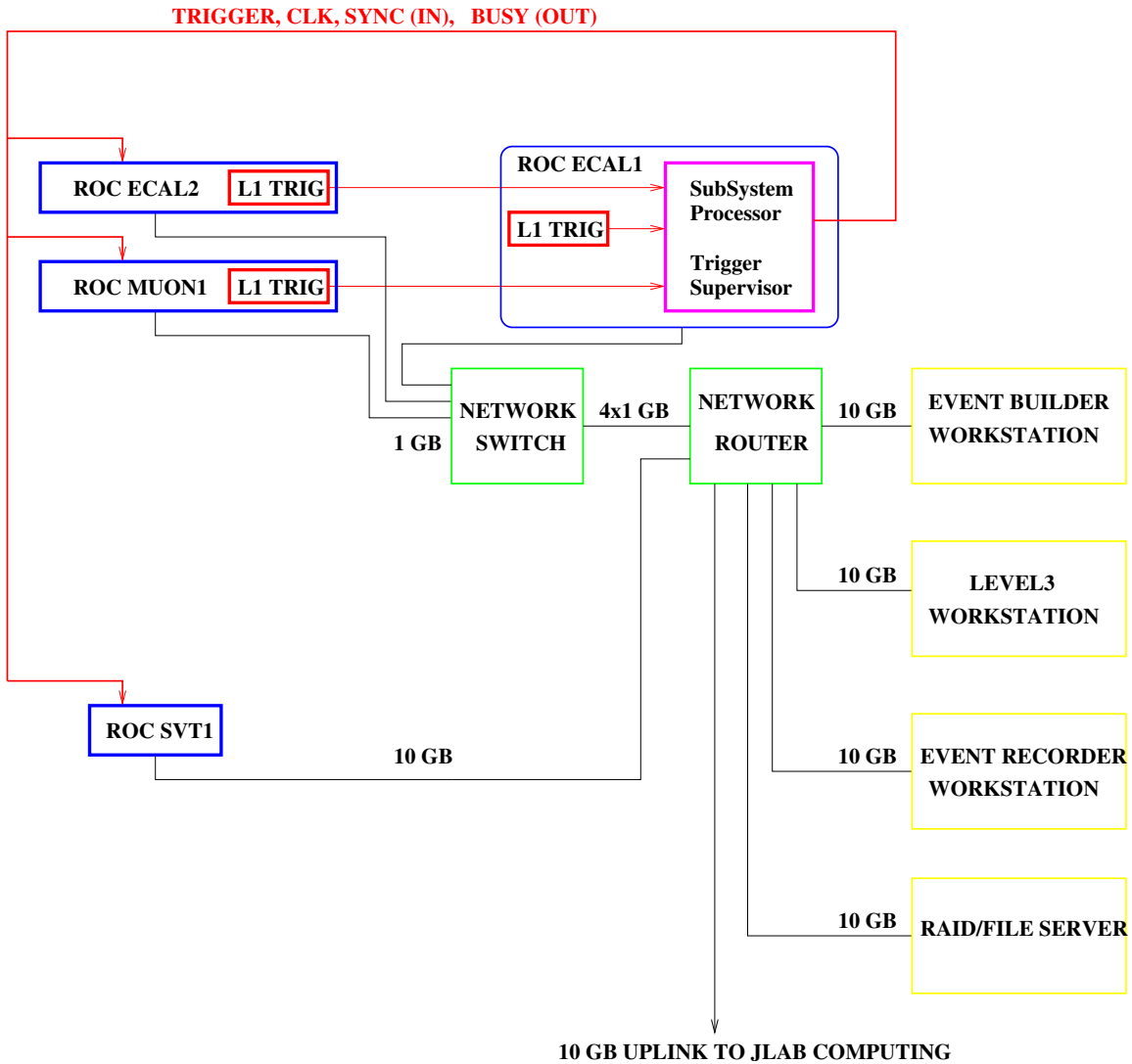


FIG. 22: Schematic block diagram of the data acquisition system.

For the ECal and Muon System, every VXS crate contains a Readout Controller (ROC)

that collects digitized information, processes it, and sends it on to the Event Builder (EB). The ROC is a single blade Intel-based CPU module running DAQ software under CentOS Linux OS. For the SVT ATCA system, the ROC application runs on an embedded processor situated on the ATCA main board. The EB assembles information from the SVT, ECal and Muon System ROCs into a single event which is passed to the Event Recorder (ER) that writes it to a RAID5-based data storage system capable of handling up to 100 MB/s. The EB and other critical components run on multicore Intel-based multi-CPU servers. The DAQ network system is a Foundry router providing high-speed connections between the DAQ components and to the JLab computing facility. The SVT ROC, which must handle large data volumes, has a 10 Gbit/s link to the Foundry router, while a 1 Gbit/s link is adequate for the ECal and Muon System. A 10 Gbit/s uplink to the JLab computing facility is used for long-term storage.

Section 4.6.1 describes the SVT DAQ in more detail while the VXS-based readout for the ECal and Muon System is described in Sec. 4.6.2. The trigger system is explained in more detail in Sec. 4.6.3.

#### 4.6.1 SVT DAQ

The goal of the SVT DAQ is to support the continuous 40 MHz readout and processing of signals from the 36 silicon strip sensors of the SVT. It also selects and transfers those events that were identified by the trigger system to the JLab DAQ for further event processing at rates approaching 50 kHz. High occupancy in the detector, pile-up from multiple bunches, and sampling pulse heights in six consecutive time buckets for each hit in order to facilitate reconstruction of the hit time to high accuracy result in large data volumes.

The system adopted is an evolution of the SVT DAQ used for the HPS Test run described in Sec. 5.1.3. Several features changed in response to the new SVT design and the evolution of SLAC's ATCA system. The new SVT has nearly twice the number of sensors as the Test Run detector, necessitating a more compact way to transfer data and power to the individual sensor modules through the vacuum flanges. Accordingly, the new system incorporates a front end board within the vacuum volume for power distribution and signal digitization, allowing many fewer vacuum connections per sensor and less interference within the vacuum volume. Problems encountered with reflections on long twisted pair data lines, although

ultimately overcome, have been avoided altogether by incorporating a flange board just outside of the vacuum and adopting an optical link. The ATCA system has evolved to using optical input, so this change lets HPS optimally piggyback on SLAC's ATCA system development.

Each of the 36 silicon strip sensors is connected to a hybrid board incorporating five 128-channel APV25 front-end ASICs [100, 101]. Figure 37 in Sec. 5.1.3 shows a picture of a hybrid board from the 2012 Test Run. The APV25 ASIC, initially developed for the Compact Muon Solenoid silicon tracker at the Large Hadron Collider at CERN, was chosen because it provides excellent signal to noise, analog output for optimal spatial resolution, and signal pulse shape sampling capability for good hit time resolution. Each hybrid board has five analog output lines (one for each of the APV25 ASICs) which are sent to the front-end readout board using low power LVDS differential current signals over about 1 m of flex cable. At the front-end readout board, a preamplifier scales the APV25 differential current output to match the range of a 14-bit Analog to Digital Converter (ADC). Each front-end board services four hybrids. The ADC operates at the system clock frequency of 41.667 MHz. The digitized output from the front-end board is sent through compact 8-pair mini-SAS cables to the vacuum flanges to connect to the external DAQ which resides outside the vacuum chamber. The front-end readout board houses a FPGA and buffers to allow for control of the distribution of clock, trigger and I<sup>2</sup>C communication with the APV25 ASICs. To further simplify the services and minimize cabling that enter through the vacuum flanges, it contains linear regulators to distribute and regulate three low voltage power lines to each of the APV25 ASICs and the high voltage bias. Figure 23 shows a schematic layout of this part of the readout chain.

The digitized signals are converted to optical signals just outside the vacuum flange on custom built flange boards. Each flange board houses optical drivers to handle the electrical-optical conversion and to transmit the optical signals over  $\sim 10$  m fibers to the ATCA crate. The flange board also interfaces the low- and high voltage power transmission from the Wiener MPOD power supplies to the front-end boards located inside the vacuum chamber.

The main SVT DAQ uses the ATCA system for high speed data transfer. The optical signals from four hybrids, one half flange board, are received at one of four sections of the Rear Transition Module (RTM) board in the ATCA crate as shown schematically in Fig. 24. Each section of the RTM connects to one of four Data Processing Processing Modules (DPM)

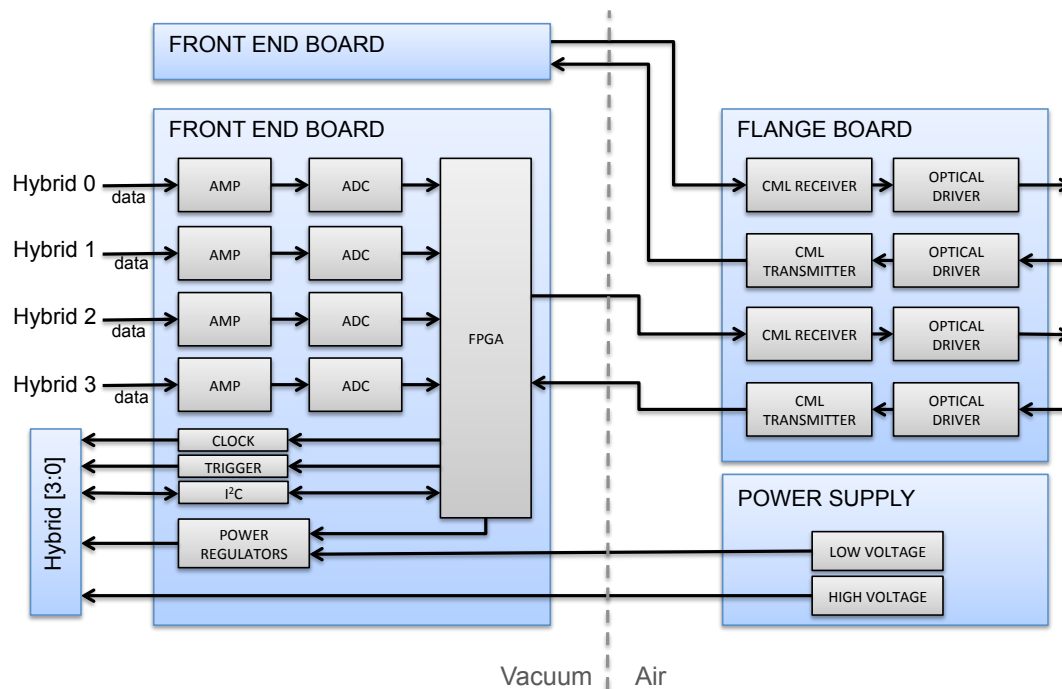


FIG. 23: Schematic overview of the front end and flange boards of the downstream part of SVT DAQ.

on the main ATCA board, the COB (Cluster On Board). The modular ATCA design permits HPS to re-use architecture and functionality from other DAQ systems such as the ATLAS muon system whose components are similar to those used by HPS. Figure 38 shows the boards designed and used for the HPS Test run. In order to minimize the complexity of the system inside the vacuum chamber, all signal processing is done at the DPM. Each DPM consists of two RCEs (Reconfigurable Cluster Elements) which are generic computational building block based on Xilinx Zynq 7 System-On-Chip technology running a dual core 1 GHz ARM processor with 1 GB of DDR3 memory tightly integrated with on-chip programmable logic (FPGA).

Each data DPM receives the digitized signals from the RTM, applies thresholds for data reduction and organizes the sample data into ethernet frames. One of the DPMs functions as the trigger interface which receives trigger signals from the optical fiber module on the RTM, distributes clock and trigger signals, and handles communication with the JLab trigger supervisor and the RCEs. Four COBs housed in two ATCA crates are sufficient to handle the 36 hybrids of the SVT.

One of the RCEs receives and buffers ethernet frames from the data and trigger DPMs

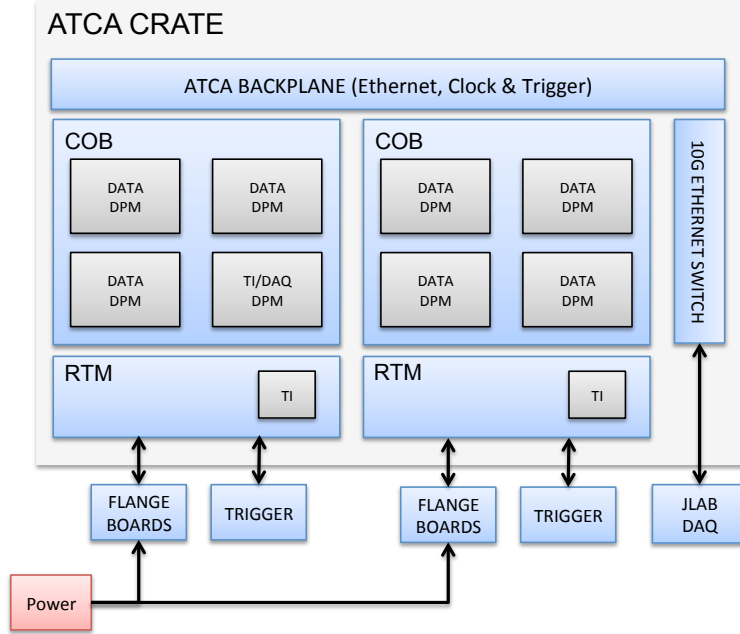


FIG. 24: Schematic block diagrams of the SVT data acquisition system.

and assembles them into full event frames. This RCE runs an implementation of the JLab ROC application that integrates the SVT event frames into the JLab DAQ system described above. The RCE node transfers data to the JLab DAQ through a 10 Gbit/s Ethernet backend interface. The maximum readout rate of the SVT is approximately 50 kHz, limited by the APV25 readout rate.

#### 4.6.2 ECal and Muon System FADC Readout

The analog signals from the individual APD's of the ECal (shaped and amplified as described in Sec. 4.5) and phototubes of the Muon System are input to a single channel on the 16-channel JLab FADC250 VXS module (FADC), shown in Fig. 25.

Three 20-slot VXS crates are needed to accommodate the system: one for each half of the ECal with 221 channels and one for the Muon System with a total of 232 channels.

The FADCs store 12-bit digitized samples at 250 MHz in 8  $\mu$ s deep pipelines. When a trigger is received, the appropriate part of the pipeline is accessed. If a FADC signal exceeds a predefined threshold within that time window, the integrated amplitude of a pre-defined number of samples before (NSB) and after (NSA) the signal passed threshold, in addition to the time, are recorded as explained in Fig. 26. This scheme significantly compresses the





FIG. 25: A Jefferson Lab FADC250 VXS module.

data input to the FADC. During data analysis, a pedestal value is subtracted to obtain the actual summed energy.

The main characteristics of the FADC are:

- 12-bit digitizer with sampling rate of 250 Msps,
- $50\Omega$  termination input,
- front-end input range:  $-0.5\text{V}$ ,  $-1\text{V}$  or  $-2\text{V}$  (sufficient to avoid signal clipping for large pulse heights),
- nominal charge resolution between 10-39 fC per ADC (see Tab. III).

Input range (V)	Nominal charge resolution (fC per ADC count)
-0.5	9.76
-1.0	19.53
-2.0	39.06

TABLE III: Nominal FADC charge resolution for different front-end input ranges.

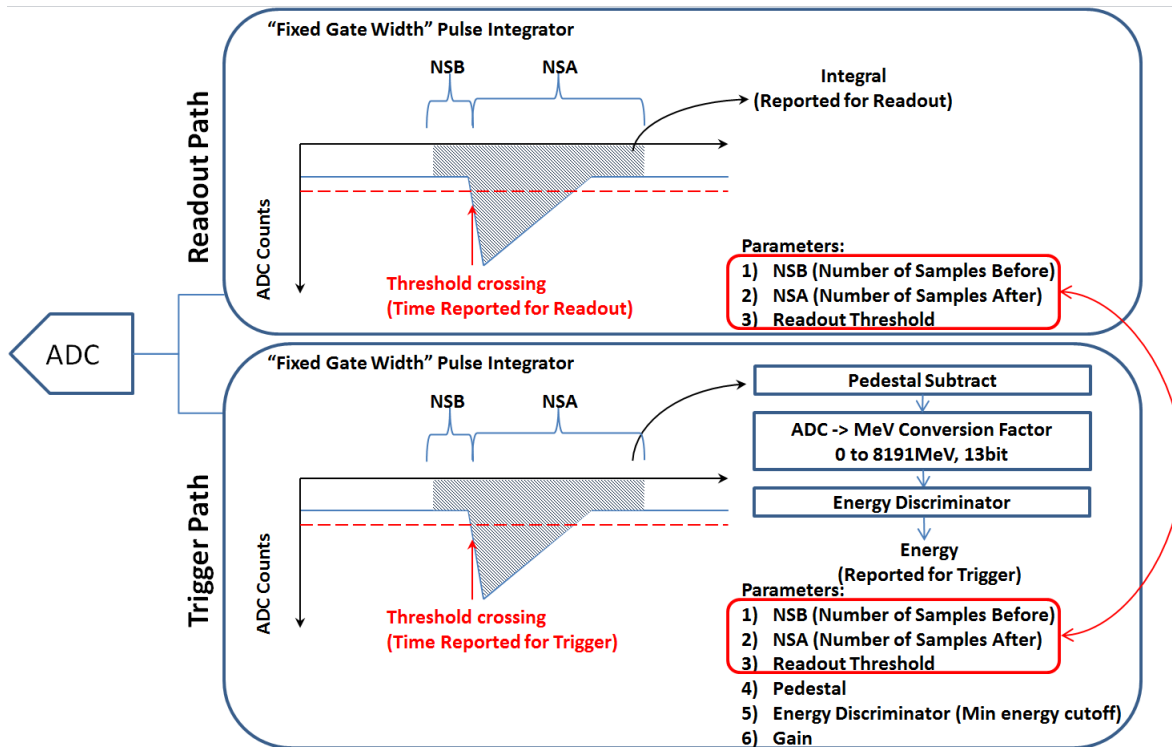


FIG. 26: FADC data paths

As shown in Fig. 26, the FADC has two parallel data paths: the readout and trigger paths. The trigger path runs continuously to report hits to the trigger system. The readout path only reports hits to the DAQ when the FADC receives a trigger.

For the readout path, every FADC has the following parameters:

- the number of samples integrated before the signal crossed threshold (NSB),
- the number of samples integrated after the signal crossed threshold (NSA),
- the readout threshold, measured in ADC counts.

The number of samples for a given channel integration is the sum of NSB+NSA samples. It is a fixed gate width pulse integration with no pedestal subtraction where the sum is stored in a 17-bit register for readout (pedestal subtraction happens offline).

For the trigger path, every channel has, in addition to NSB and NSA:

- trigger threshold, measured in ADC counts,
- a pedestal,

- a conversion factor (gain) that converts the ADC counts to energy in MeV (with 13 bits: from 0 to 8191 MeV),
- an energy discriminator threshold (minimum energy cutoff).

Note that the threshold for the trigger path can be set independently from the readout threshold. The pedestal value is subtracted from the integrated sum over NSB+NSA samples and converted to MeV units using a supplied gain conversion factor. The energy discriminator can be used to cut off low energy pulses before reporting to the Crate Trigger Processor (CTP). The values reported to the CTP are the 13-bit pulse energy and the time at which the pulse crossed the threshold. Data for every channel is sent to the CTP every 32 ns (if there is no hit a 0 energy pulse is sent) which sets a worst case double pulse resolution of 32 ns for individual channels, but less if pulses occur in adjacent 32 ns windows.

### 4.6.3 Trigger System

The trigger system is designed to efficiently select  $e^+e^-$  and  $\mu^+\mu^-$  events by using information from the ECal and Muon System. For  $e^+e^-$  events, the trigger looks for time coincidences of clusters in the top and bottom half of the ECal. The clusters also have to satisfy loose kinematic selections optimized on  $A'$  events to further reduce the rate. For  $\mu^+\mu^-$  events<sup>1</sup>, signals from at least the first two layers of the muon hodoscopes are combined with an ECal signal consistent with a minimum ionizing particle (MIP).

As described above in Sec. 4.6.2, the first stage of the trigger logic is incorporated into the FPGA's on the FADC boards which sends crystal energy and time information to the CTP. With the available 3-bit time information, the CTP can in principle look for time coincidence of crystal signals with 4 ns resolution (HPS will use a 8 ns time coincidence interval). The final trigger decision is made in the CTP and Sub-System Processor (SSP). The Trigger Supervisor generates all necessary signals and controls the entire DAQ system readout through the Trigger Interface (TI) units. The TI units are installed in every crate that participate in the readout process.

The trigger system is free-running and driven by the 250 MHz global clock and has essentially zero dead time at occupancies expected by HPS. The Trigger Supervisor can

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<sup>1</sup>see Appendix B.2 for details.

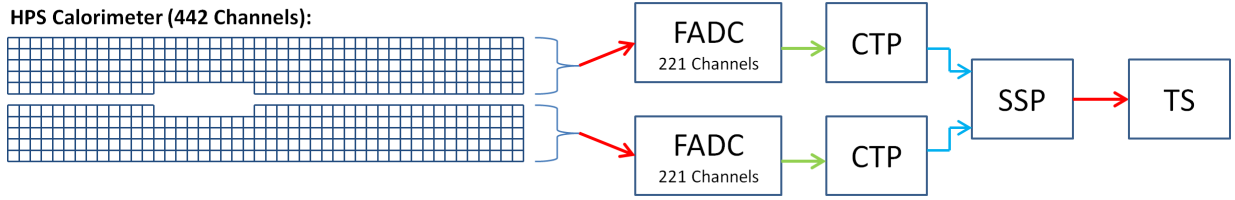


FIG. 27: Block diagram of the ECAL trigger system consisting of the FADC that samples and digitizes the detector channel signals and sends them for cluster finding in the CTP. The CTP clusters are sent to the SSP where the final trigger decision is taken based on pairs of clusters in both halves of the ECal. The decision is sent to the Trigger Supervisor (TS) that generates the necessary signals to readout the sub-detectors.

apply dead time if necessary, for example on a ‘busy’ or ‘full’ condition from front-end electronics. The system is designed to handle trigger rates above 50 kHz and a latency set to  $\approx 3 \mu\text{s}$  to match that required by the SVT APV25 chip.

### $e^+e^-$ Trigger

The trigger system for  $e^+e^-$  events can be broken down into three sections (see Fig. 27):

- FADC (pulse finding): samples and digitizes the signal pulses from each detector channel. Sends the measured pulse energy and arrival time to the CTP.
- CTP (cluster finding): groups FADC pulses from each half of the ECal into clusters. The cluster energy, arrival time, and hit pattern are sent to the SSP.
- SSP (cluster pair finding): searches for time coincidence of pairs of clusters from the top and bottom half of the ECal and applies topological selections.

The time coincidence window of pairs of clusters in the top and bottom half of the ECal are programmable with 4 ns resolution. The cluster finding algorithm is very fast and makes use of the parallel processing nature of FPGA’s by simultaneously searching for 125 clusters, up to 3x3 in size, across the calorimeter crystal array, see Fig. 28. It performs the following tasks:

- Adds energy from hits together for every 3x3 square of channels in ECal.
- Hits are added together if they occur (leading edge) within a programmable number of 4 ns clock cycles of each other (HPS will use 8 ns time coincidence time interval).

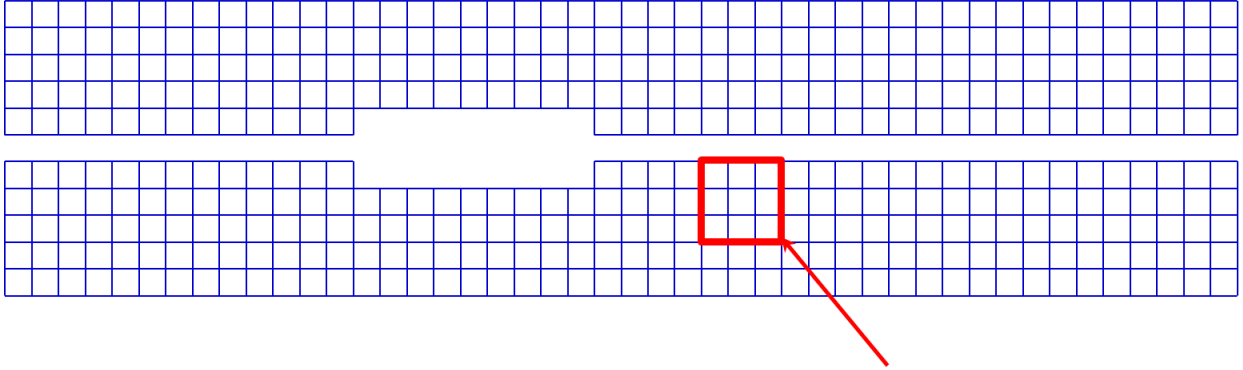


FIG. 28: Cluster finding algorithm.

- If the 3x3 energy sum is larger than the programmable cluster energy threshold and the sum is greater than any neighboring 3x3 windows, the CTP reports the cluster parameters to the SSP.

The CTP evaluates all hits in its half of the calorimeter every 4 ns. A programmable time window is used to allow hits that are out of time with each other to be considered as part of a cluster sum. This is done by reporting hits when they occur and then reporting them again for the next  $N$  number of 4 ns clock cycles, where  $N \in [0, 7]$ . This is useful to deal with skew and jitter that develop from the detector, cabling, and electronics. As described above, the CTP only selects the 3x3 window with the highest energy sum of its neighbors. This filtering is applied to deal with overlapping clusters and cases where the cluster is larger than a 3x3 window.

The CTP sends the following information about the clusters to the SSP:

- 13-bit cluster energy (in MeV)
- Cluster position (crystal index: x,y)
- Cluster time (with 4 ns resolution)
- Cluster 3x3 hit pattern (the detector channels reporting a hit in the cluster)

The cluster position is the coordinate of the peak crystal energy as seen from a 3x3 view. The 3x3 cluster hit pattern can be used by the SSP to help filter strange cluster patterns and/or make a low resolution cluster centroid computation. The SSP collects the cluster information from the two halves of the calorimeter and applies selections optimized to further reduce background rates with small impact on the A' signal:

- Energy sum,  $E_{min} \leq E_{top} + E_{bottom} \leq E_{max}$
- Pair time coincidence,  $|t_{top} - t_{bottom}| \leq \Delta t_{max}$
- Energy difference,  $|E_{top} - E_{bottom}| \leq \Delta E_{max}$
- Energy slope,  $E_{cluster\_with\_min\_energy} + R_{cluster\_with\_min\_energy} \times F_{energy} \leq Threshold_{slope}$
- Co-planarity,  $|\tan^{-1}(\frac{X_{top}}{Y_{top}}) - \tan^{-1}(\frac{X_{bottom}}{Y_{bottom}})| \leq Coplanarity_{angle}$
- Number of hits in 3x3 window,  $\#hits_{3 \times 3} \geq HitThreshold$

where  $E_{max}$ ,  $\Delta t_{max}$ ,  $\Delta E_{max}$ ,  $Threshold_{slope}$ ,  $F_{energy}$ ,  $Coplanarity_{angle}$  and  $HitThreshold$  are programmable parameters.

The SSP can also create a trigger decision based on the existence of a single cluster in the ECal exceeding the energy threshold which is useful for commissioning and calibration runs.

Online event analysis will be provided to be compared against trigger event data for immediate verification (on each trigger cut: cluster energies, positions, timing, energy slope, coplanarity and hit threshold). With identical trigger and data readout paths and high energy resolution, very precise agreement can be expected between trigger and readout.

## Diagnostic Tools

Previous experience with similar (but much simpler) trigger systems showed that diagnostic tools are necessary to make sure that the calorimeter and trigger electronics work as expected.

Scalers will be implemented for every ECal channel. An example of such a diagnostic tool is presented in Fig. 29 from the previous version of the ECal. Hot or dead channels are easily identified and disabled online. A diagnostic scope permits analyzing the trigger logic online. The goal is to have a Two-Dimensional Analyzer to provide a remote debug interface to identify bad channels, verify cluster finding algorithms and check timing. The details of this logic analyzer are as following:

- Runs in parallel, non-intrusively to the calorimeter trigger
- Can setup trigger on any ECal pixel arrangement and/or cluster count

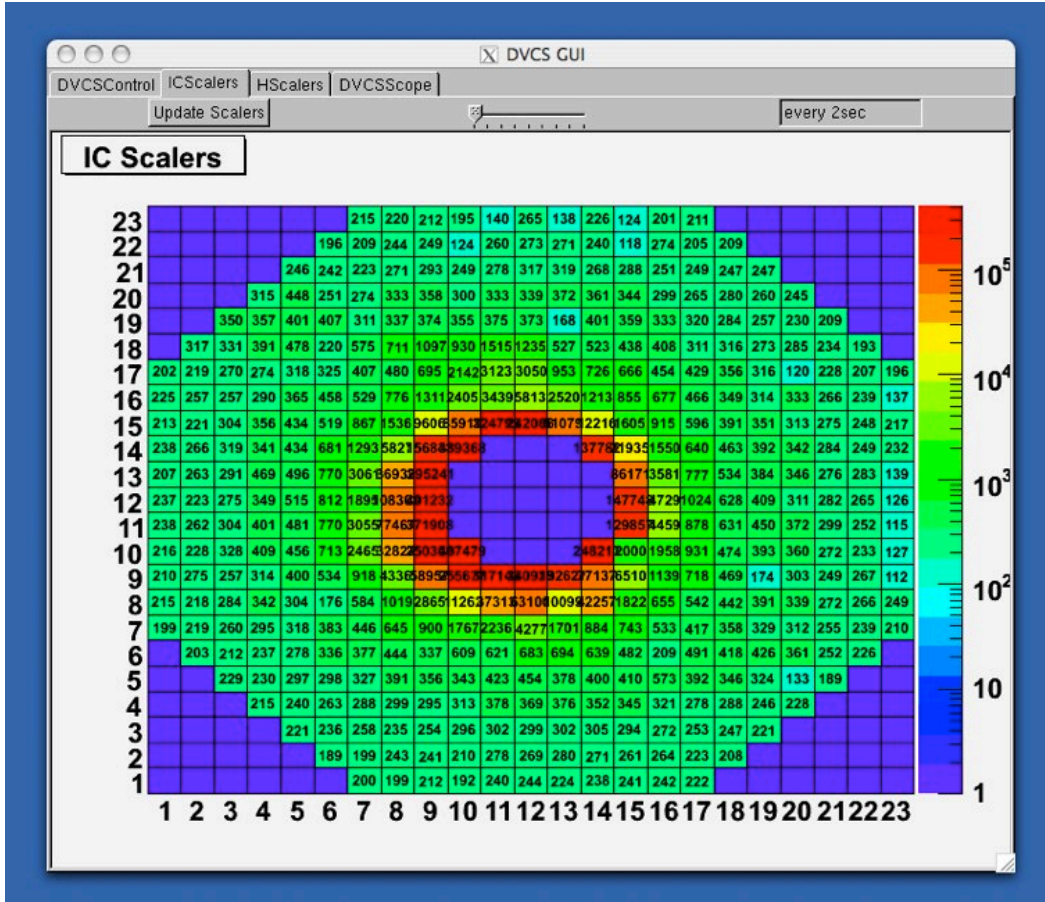


FIG. 29: Scalers (example from the previous version of the calorimeter).

- Can move forward/backward in time by 250 ns to see timing details
- Will be customized for HPS geometry and hardware

An example of the 2D analyzer is presented in Fig. 30. Two clusters are displayed in the picture. The red color displays the hits in the calorimeter and the center of clusters is displayed in yellow.

In addition to scalars, the distributions of individual ADC channel pulse energies will be monitored and cluster hit position and energy from the SSP processor will be histogrammed as well. Two histograms (accepted and rejected) will be provided for each trigger cut used in the trigger logic.

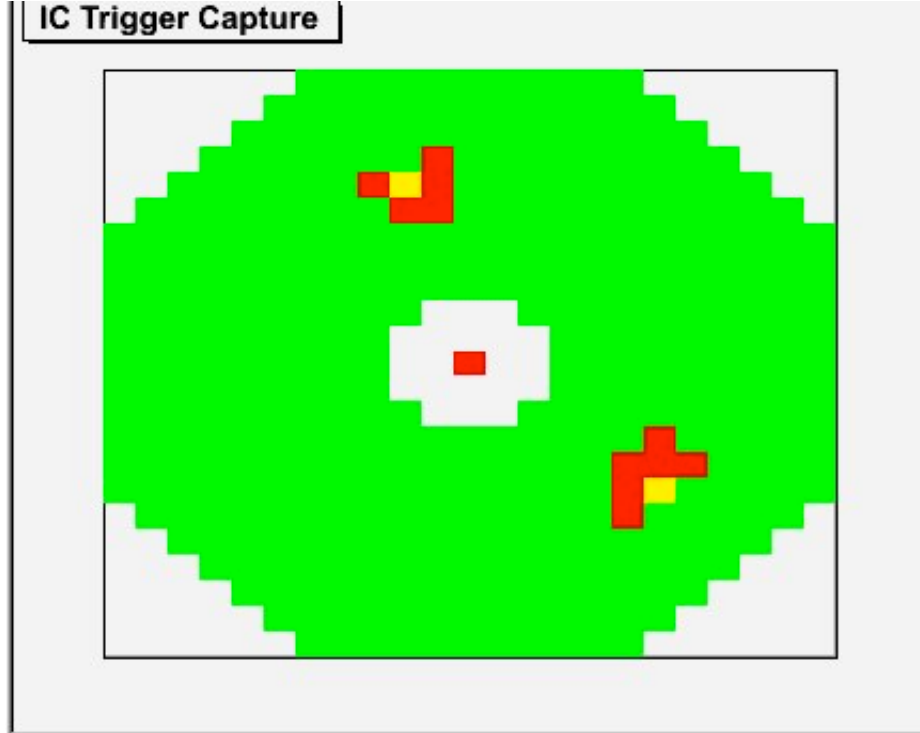


FIG. 30: Diagnostic scope (example with two clusters found from the previous version of the calorimeter). Green - no hits, red - tower with hit, yellow - cluster found.

#### 4.6.4 Event Size and Data Rates

The high occupancies in the detector require a high readout bandwidth to be able to transfer hits from the detectors to disk. The event sizes and rates are based on estimates from full Geant4-based simulations including all known backgrounds. As expected the SVT dominates the expected rates. The noise hit occupancy in the SVT is kept low by requiring that three of the six samples are above a threshold of twice the noise level, and that the signal rises either from the second to the third sample, or from the third to the fourth. The resulting occupancy due to noise hits is roughly 0.02%, or an average of 3 hits. The dominant contribution to the occupancy is then the estimated high rate of beam background hits. This is estimated from detailed full simulation resulting in an occupancy of around 0.4% or an average of 80 channels above threshold. Each SVT hit has, in addition to the six digitized samples, header information that identifies the the channel number and its chip address. The complete SVT event size also include the overhead from each FPGA and the JLab data stream bank header. The maximum average event size increased with decreasing beam energy since a larger fraction of backgrounds get larger opening angles and



	Occupancy(%)			Event size (kB)			Data rate (MB/s)		
Beam energy (GeV)	1.1	2.2	6.6	1.1	2.2	6.6	1.1	2.2	6.6
SVT	0.4	0.4	0.3	1.8	1.8	1.6	33.1	29.2	19.8
ECal	3.0	4.1	4.5	0.4	0.5	0.5	7.0	7.9	6.7
Total	-			3.0	2.2	2.0	40.1	37.1	26.4

TABLE IV: Summary of the occupancy, event size and data rate expected for the runs at the three beam energies in the run plan.

thus potentially higher than the 15 mrad vertical dead zone angle. For a beam energy of 1.1 GeV, the average SVT event size is 1.8 kB and the rate is 33 MB/s, well within the SVT DAQ capabilities.

The two VXS crates in the ECal system contribute a total of approximately 0.5 kB to each event, and maximum rate across the crates of about 7 MB/s, well below the limit of 100 MB/s per crate.

The Muon System is still under study, but data rates are expected to be less than those of the ECal since the Muon System uses only one VXS crate, uses the same data format as the ECal and is expected to have similar or lower occupancy. As this implies a 10% contribution to data rates (at most), we omit the Muon System from our estimates. Table IV summarizes the event size and data rates. The highest overall rate, for a 1.1 GeV run, that needs to be written to disk is 40 MB/s which is within the current DAQ system design limit of 100 MB/s.

## 4.7 Software

The main HPS data analysis software is built onto the `org.lcsim` framework, a set of software tools written in Java originally for detector studies for the International Linear Collider (ILC). The detector simulation is done with SLIC, a GEANT4 based Monte Carlo simulation that allows for a very flexible geometry setup, which is identical to the geometry used by the analysis software. The MC output or the actual raw data are analyzed for tracks and particle identification using a dedicated reconstruction code written using the `org.lcsim` framework. The output is in the LCIO format and can be further analyzed for physics signals directly or by transforming it to a set of ROOT based data summary tape (DST).

For successful data taking, besides the DAQ system described in section 4.6, a number

of monitoring and calibration programs will be required. The Event Transport ring (ET), part of the Jefferson Lab DAQ system, allows a number of client programs to attach to the raw data stream and receive a predetermined fraction of the events, or events on demand. Several client monitoring programs will look at the low level raw data, including a standalone raw data event display and data quality monitors. An interface from the ET ring to the full analysis system also exists, allowing for high level histograms which can also monitor the data quality and a Wired based event display. The underlying histogramming system automatically provides live updates of the histograms. Specific monitoring for the sub-detectors will be contributed by the detector sub-groups, as will the needed calibration code.

The HPS code base is maintained by the HPS Software group, chaired by M. Holtrop, and the Data Analysis Group, chaired by M. Graham. Around ten members of the HPS collaboration participate very actively in the development of the HPS software.

## 4.8 Offline Computing Model

The following is an outline of the offline computing model envisioned for satisfying the analysis needs of the HPS experiment. The raw data collected over the running periods must be processed through calibration passes, reconstructed into calibrated objects useful for analysis and separated into analysis streams. Corresponding Monte Carlo will also need to be produced and separated into the same analysis streams.

The raw data must be processed to produce physics data objects that can be analyzed. This reconstruction process will also include filters to select events of physics interest. We use the event size estimates in Table V, which are based on Table IV from the previous section and object sizes in EVIO (raw data) and LCIO (processed data) formats.

Beam energy	Raw (EVIO) event size (kB)	Reconstructed (LCIO) event size (kB)
1.1 GeV	2.2	4.8
2.2 GeV	2.3	5.0
6.6 GeV	2.1	4.0

TABLE V: Data event sizes.

Table VI shows the expected number of triggered events and the total amount of data expected over the different runs. We assume that the experiment collects data for all of

its available beam time and the time allocated for detector commissioning, even though the experiment reach only assumes 50% availability; this gives a conservative estimate of computing requirements. For trigger rate estimates, we use the ECal trigger rate from Section 6.2; based on Appendix B.2.3, the muon system trigger rate is expected to be negligible.

Run	$E_{beam}$ (GeV)	Time (days)	Events ( $\times 10^9$ )	Raw data (TB)	Processed data (TB)
2014	1.1	21	33	73	159
2014	2.2	21	29	67	145
Total	-	42	62	140	304
2015	2.2	35	48	112	241
2015	6.6	35	38	80	153
Total	-	70	86	192	394

TABLE VI: Summary of the raw and processed data expected from the HPS runs.

For modeling signals, estimating backgrounds and confirming the understanding of the detector performance, extensive Monte Carlo simulation is needed. Three types of events will be simulated: general beam background, trident background, and A' events.

General beam background events will be generated by fully simulating beam background as described in 6.1.1 and simulating the HPS trigger. Because this is a compute-intensive process, only 1 million triggered events will be simulated at each beam energy; this is adequate for trigger and DAQ studies.

Trident background and A' events will be generated by using MadGraph to make trident or A' events with enhanced trigger probability, overlaying beam background, and simulating the trigger. The amount of triggered trident events to be simulated is 10% of the amount expected in actual data; the number of triggered A' events to be simulated is 100 million at each of 10 mass points at each beam energy. These will be used to test the analysis.

In total 472 (618) TB of storage for data (raw, reconstructed and simulated) is needed for the 2014 (2015) run. Tape is currently the only economical storage solution for storing all of the raw, simulated and processed data.

The processing of the raw data is foreseen to occur at JLab. Given a typical bandwidth between sites of 3 to 4 TB/day, only data summaries of events satisfying pre-selection criteria for targeted analyses will be exported to remote sites. Likewise, the size of the simulated data samples suggests that the simulation should be processed and stored at JLAB and that only data summaries or small samples of the full data will be exported.

Analyses needing access to the hit level information will need to be run at JLab or run on small samples of exported data unless they can take advantage of the data summaries.

Data summaries will be written as ROOT trees. 65 (89) TB of DSTs will be generated in the 2014 (2015) run. These will be generated and stored on tape at JLab, and mirrored on tape at SLAC.

Disk space at JLAB will be needed for code releases and scratch areas. Disk space will also be needed at SLAC for staging, code releases and scratch areas. Both needs are covered by existing computing infrastructure.

The HPS storage requirements are summarized in Tab. VII.

Storage category	2014 (TB)	2015 (TB)
Raw data	140	192
Reconstructed data	304	394
Simulated data (raw and reconstructed)	27	31
Total data	472	618
DST (run data)	62	86
DST (simulated data)	3	3
Total DST	65	89

TABLE VII: Data storage summary; data storage is at JLab only, while DST storage is common to JLab and SLAC.

Simulation production and data reconstruction will be done on the batch farm at JLab. The CPU requirements are summarized in Tab. VIII. The requirements correspond to roughly a month of running each year on the JLab analysis farm using all of the existing cores, but this will be spread throughout the years.

Computing category	2014	2015
Raw data processing ( $\times 10^6$ CPUh)	0.26	0.36
Simulation production ( $\times 10^6$ CPUh)	0.84	0.99
Total ( $\times 10^6$ CPUh)	1.10	1.35

TABLE VIII: Computing needs summary in CPU hours using typical 2.4 GHz cores.

The Jefferson Lab Computing Center provides computing and storage for experiments at JLab. An updated request will be submitted for data storage (tape and disk), computing resources (CPU hours for simulation and production), and data transfers to/from JLab. The needs are close to those previously submitted.

## 5 May-2012 Test Run

The HPS Test run was proposed to DOE early in 2011 as the first stage of the HPS experiment. Its purposes included demonstrating that the apparatus and data acquisition systems are feasible and that the trigger rates and occupancies encountered in electron-beam running are as simulated. It also provided valuable experience to the HPS Collaboration in all aspects of designing, building, installing, and running the experiment at JLab. Furthermore, had electron running been possible, and had the detector performed adequately with intense electron beams, the HPS Test Run apparatus was capable of beginning the search for heavy photons. The HPS Test apparatus was installed on April 19, 2012, and ran parasitically with the HDice experiment, using a photon beam, until May 18. The JLab run schedule precluded any dedicated electron beam running, but the HPS Test Run was allowed a short and valuable dedicated run with the photon beam.

This section briefly reviews the HPS Test Run apparatus, a simplified version of that planned for the full HPS experiment, and demonstrates the feasibility of the detector technologies proposed for silicon tracker, ECal, and data acquisition systems. It documents the performance of the trigger, data acquisition, silicon tracker, and ECal and shows that the performance assumed in calculating the physics reach of the experiment is realistic. Of particular importance, data from dedicated photon beam running has been used to compare the observed trigger rates with that expected in simulation. The trigger rate is almost entirely due to photons which have converted to  $e^+e^-$  upstream of HPS and is sensitive to the multiple Coulomb scattering of electrons and positrons in the conversion target. Since scattered primary beam is the dominant source of occupancy in running HPS in an electron beam, good agreement between data and simulation confirm the background simulation used to benchmark the physics reach of the HPS experiment.

In addition to this important test of our background simulation, the test run accomplished the following goals which are explained below in Sec. 5.3:

1. More than 97% of SVT channels functioned properly.
2. SVT readout signal to noise is 25.5, sufficient to achieve the expected spatial and temporal resolutions.

3. SVT hit time resolution is 2.6 ns, proving that hit time reconstruction will work for HPS.
4. SVT hit efficiency is greater than 98%.
5. Survey-based SVT alignment performed as expected and will allow track-based alignment.
6. 87% of ECal crystals functioned properly, with defects to be corrected by planned ECal upgrades.
7. The ECal can be calibrated using SVT tracks.
8. The SVT and JLab data acquisition were successfully integrated.
9. The trigger functioned as designed; FADC trigger rate was tested to greater than 100 kHz.

The Test Run allowed us to test many aspects of the HPS experiment and its expected performance, and has given us full confidence that HPS will work as designed. In the following we briefly review the test run apparatus, data acquisition, and detector performance, and we discuss confirmation of occupancy and trigger rate simulations using Test Run data.

## 5.1 HPS Test Run Apparatus

In Figure 31, the layout for parasitic running is shown. The silicon vertex tracker was installed inside the Hall B pair spectrometer magnet vacuum chamber with the electromagnetic calorimeter mounted downstream. Both the tracker and the ECal were retracted off the beam plane to allow clean passage of the photon beam through the system.

For dedicated HPS running the photon beam was generated in the interaction of the 5.5 GeV electrons with a gold radiator of  $10^{-4}$  r.l., located  $\approx 8$  meters upstream of the pair spectrometer magnet. After collimation ( $D = 6.4$  mm), the photon beam is directed toward the pair spectrometer, where it passes through a thin converter foil located 77 cm upstream of the first layer of the silicon vertex tracker and then through the HPS system. Data were taken on different converters (empty,  $1.8 \times 10^{-3}$  r.l.,  $4.5 \times 10^{-3}$  r.l., and  $1.6 \times 10^{-2}$  r.l.). These measurements were repeated for the reverse field setting of the pair spectrometer dipole.

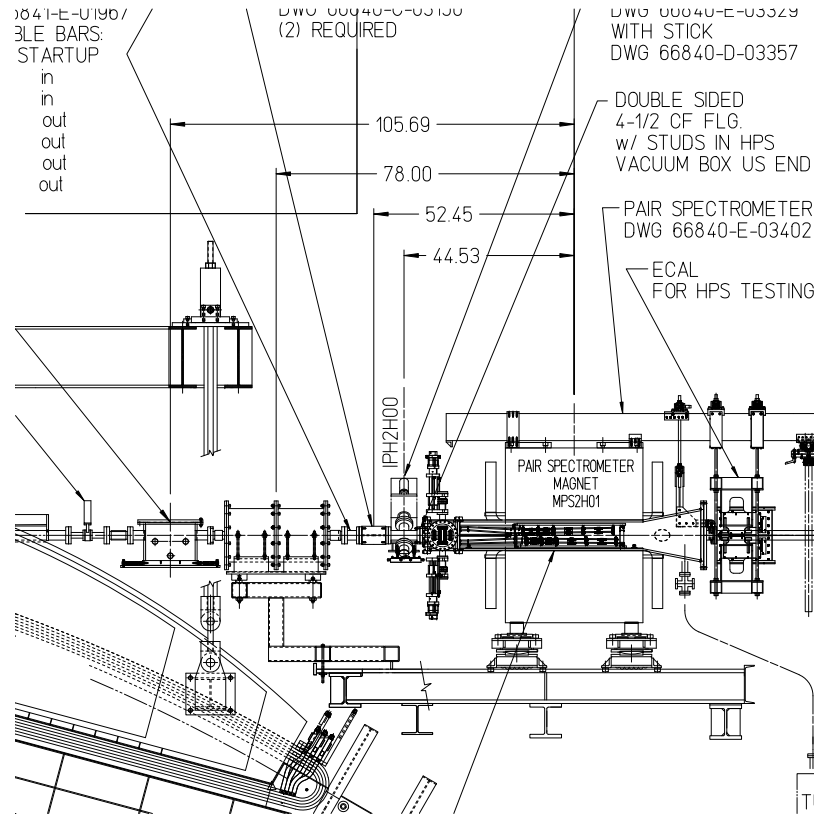


FIG. 31: Layout of the HPS parasitic run.

### 5.1.1 Test Run SVT

The silicon tracking and vertexing detector for HPS Test, or SVT, operates in an existing vacuum chamber inside the pair spectrometer analyzing magnet in Hall B at JLab. The design principles of the SVT are described in further detail in the HPS Test Run Proposal [6]. There are five measurement stations, or “layers,” placed immediately downstream of the target. Each layer comprises a pair of closely-spaced planes and each plane is responsible for measuring a single coordinate, or “view.” Introduction of a stereo angle between the two planes of each layer provides three-dimensional tracking and vertexing throughout the acceptance of the detector with one redundant layer.

In order to accommodate the dead zone, the SVT is built in two halves that are mirror reflections of one another about the plane of the nominal electron beam. Each half consists of five double-sided modules mounted on a support plate that provides services to the modules and allows them to be moved as a group relative to the dead zone. Each module places a pair of silicon microstrip sensors back-to-back at a specified stereo angle with independent

cooling and support.

Modules with 100 milliradian stereo are used in the first three layers to provide higher-resolution 3-d space points for vertexing. The 50 milliradian stereo of the last two layers breaks the tracking degeneracy of having five identical layers and minimizes fakes from ghost hits, improving pattern recognition while still providing sufficient pointing resolution into Layer 3 for robust hit association in the denser environment there. These stereo angles are the same as those proposed in Section 4.4 for the new SVT. The details of the five layers are shown in Table IX and a rendering of the detector layout is shown in Figure 32. Figure 33 shows a photograph of both completed detector halves prior to final assembly. Altogether, this layout comprises 20 sensors and hybrids and 100 APV25 chips for a total of 12780 readout channels.

Layer	1	2	3	4	5
Nominal $z$ , from target (cm)	10	20	30	50	70
Stereo angle (mrad)	100	100	100	50	50
Bend plane resolution ( $\mu m$ )	$\approx 60$	$\approx 60$	$\approx 60$	$\approx 120$	$\approx 120$
Non-bend plane resolution ( $\mu m$ )	$\approx 6$	$\approx 6$	$\approx 6$	$\approx 6$	$\approx 6$
# sensors	4	4	4	4	4
Nominal dead zone (mm)	$\pm 1.5$	$\pm 3.0$	$\pm 4.5$	$\pm 7.5$	$\pm 10.5$
Power consumption (W)	6.9	6.9	6.9	6.9	6.9

TABLE IX: Layout of the HPS Test SVT.

Power is provided to each hybrid using CAEN power supplies on loan from Fermilab. Three low voltages are supplied for the APV25 and one high voltage to reverse bias the sensor. The supplies that provide sensor bias are capable of 500V operation and can be used to test operation at high voltage in close proximity to an electron beam. The total power consumption of each hybrid during normal operation is approximately 1.7 W, which is removed by the cooling system. Care was exercised in selecting power and data cables to ensure vacuum compatibility and sufficient radiation hardness. A custom junction box interfaces the CAEN power supply output channels to the SVT hybrids. Control of the supplies is provided via an EPICS graphical user interface, which allows monitoring of the detector and interlock protection.

The linear shifts that define the opening of the SVT are controlled by a pair of stepper motors located in low field regions at the ends of the analyzing magnet. For photon running,



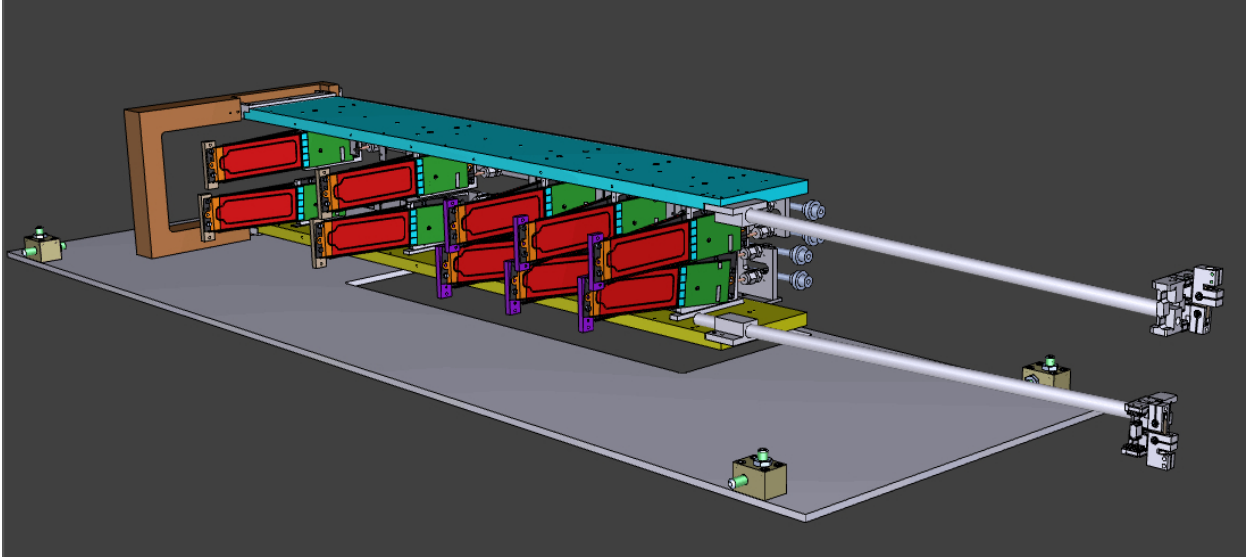


FIG. 32: A partial rendering of the HPS Test SVT solid model showing the modules of the upper and lower half-detectors on their support plates, the hinged C-support, the motion levers, the cooling manifolds on their strain relief plate and the baseplate with its adjusters. The sensors are shown in red and the hybrids in green. The beam enters from the right.

these are locked in the open position, but for electron running they will be connected and controlled through EPICS so that the distance between the beam and the sensors can be adjusted to balance detector occupancies and acceptance.

### 5.1.2 Test Run ECal

The electromagnetic calorimeter (ECal) for HPS, as described in Section 4.5, was built and tested in the test run. The only differences between the test run ECal and what is proposed here for HPS are in the position and the vacuum chamber. The vacuum chamber between the two ECal modules was not used for the photon test run; instead a 2" beam pipe was used to transport photon beam from the pair spectrometer vacuum chamber to the HDIce target. The ECal was mounted downstream of the analyzing dipole magnet at the distance of  $\sim 148$  cm from the upstream edge of the magnet. The two ECal modules were positioned just above and below the beam pipe such that the edge of the crystal closest to the beam was 3.7 cm from it.

For the test run, the ECal made use of the existing low and high voltage systems from the CLAS IC, as well as signal cables and splitters. Connectors on the existing signal cables were rearranged to accommodate the new layout of the channels.



FIG. 33: Both halves of the HPS Test SVT fully assembled at SLAC.



FIG. 34: Assembly of the ECal bottom module.

Assembly of the bottom half to the ECal is shown in Fig. 34. Figure 35 shows the ECal in its installed position for the parasitic run with photon beam.

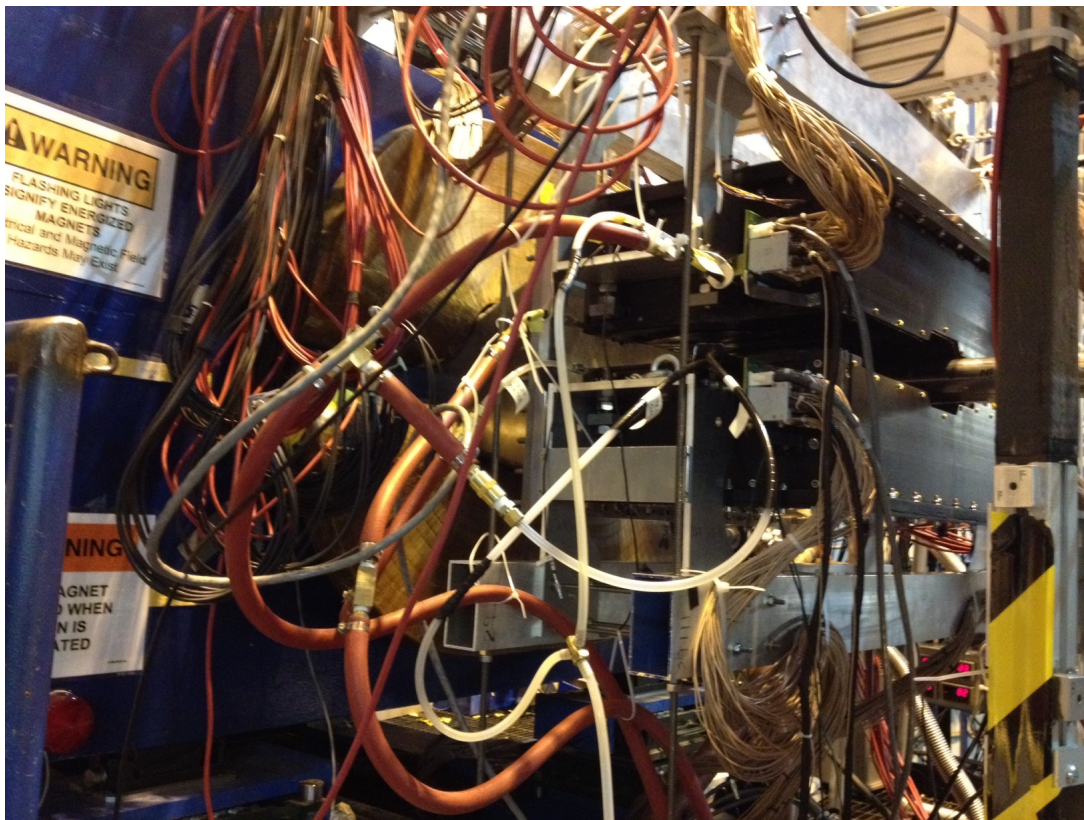


FIG. 35: ECal mounted downstream of the Hall-B pair spectrometer for the parasitic run with photon beams. Hoses for the cooling system, and the power and signal cables for beam-right side of both modules are visible.

### 5.1.3 Test Run Data Acquisition

The data acquisition system (DAQ) for the HPS Test run was a somewhat simplified version of the DAQ proposed for HPS in Sec. 4.6. In addition to simpler trigger logic, the primary difference for the ECal is a different signal processing for the trigger which results in slightly worse single-crystal energy resolution and no possibility to calibrate individual crystal gains at the trigger level. For the SVT, the smaller number of channels eliminated the need for optical conversion and aggregation of data and detector power inside the vacuum chamber. Finally, most data links had 1 Gbit bandwidth, sufficient for the purposes of the test run.

In other respects, the HPS Test run DAQ was essentially identical to that proposed for HPS and used to verify the overall technical approach of the system. The ECal provides data to the FADC-based Level 1 trigger system. Accepted events are read out from the ECal and SVT DAQ and processed by an event builder before output to disk. For the

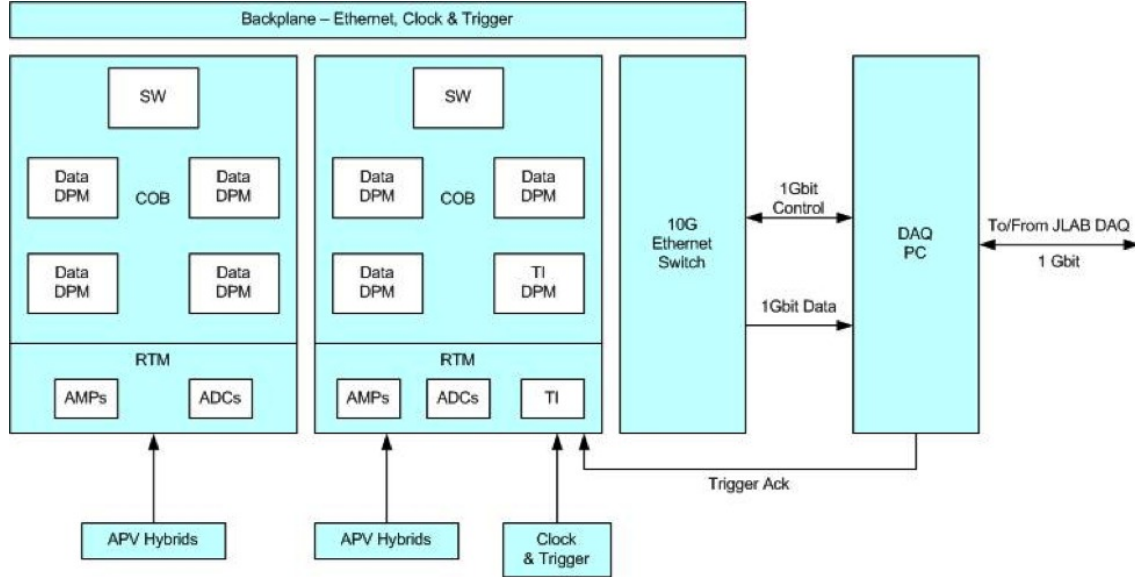


FIG. 36: Schematic of the SVT DAQ used for the test run. Note that the hybrids are connected directly to the RTM and that an external DAQ PC is used for control and transfer of data to the JLab DAQ.

ECal, the Readout Crate Controllers (ROCs) were the same as those proposed for HPS. For simplicity, a hybrid approach was used for the SVT DAQ in the test run where the ROC ran on an external PC connected to the ATCA crate. Similar to that proposed for HPS, a Foundry Router was used as the backbone of the DAQ system and to the JLAB Computer Center, which was sufficient for the test run. The RAID5 storage system had 100 MB/s capability, sufficient to handle the anticipated data rates for electron running.

The SVT DAQ system was based on the same overall architecture as that described in Sec. 4.6.1, see Fig. 36 for a schematic layout. Being only half the size of the HPS SVT, the largest difference is the provision for individual power and data for each sensor and readout chip from the power supplies and DAQ outside of the vacuum chamber. This simplification reduced development time and cost at the expense of a large number of connections and vacuum feedthroughs. With a total of 20 silicon microstrip sensors, each one connected to an onboard hybrid readout card hosting five 128-channel APV25 ASICs (shown in Figure 37), 600 lines for power and data are required. Without optical conversion inside the chamber as proposed for HPS, analog signals from the APV25 chips are carried directly via multi-twisted-pair cable to ADCs on the Rear Transition Module (RTM), see Fig. 38, in the ATCA crate located outside the vacuum chamber. The ATCA main board, the Cluster on Board or COB (see Fig. 38), is similar to the one described for the HPS DAQ with the important

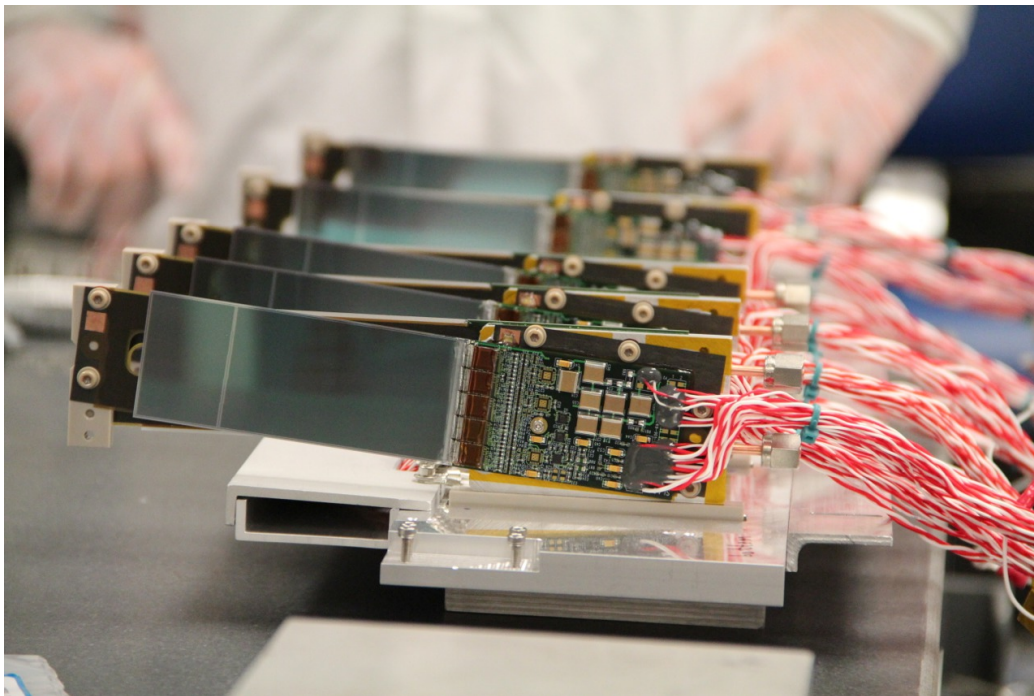


FIG. 37: View from upstream of one half of the SVT modules mounted on the support plate. Signal, power and control are soldered, and potted, to pads at one end while the five APV25 chips are wirebonded to the silicon sensor at the other.

exception that one of the Data Processing Modules (DPMs) functions as the trigger interface and there is no Reconfigurable Cluster Element (RCE) module. Instead, the DPMs package and send the data from the hybrids through a 1 Gbit ethernet connection to an external PC which serves the same purpose as the RCE module in the HPS DAQ. The external PC also supports slow control and monitoring, communication with JLab DAQ, and trigger acknowledge interface to the trigger DPM in the ATCA crate.

The ECal DAQ system used in the test run is very similar to that described for HPS in Sec. 4.6.2. The only significant difference is that in the test run, the signals from the ECal modules were sent to a signal splitter. One of the outputs of the splitter is fed to a discriminator that also has an internal scaler, and then to a TDC channel. The other output is sent to the JLab FADC250 VXS module, shown in Fig. 25.

#### 5.1.4 Test Run Trigger System

The trigger system used in the test run is described in the HPS status update to PAC39 [7] in full detail; only the chief differences will be described here. It is generally similar to



FIG. 38: Picture of a RTM (top) and COB board (bottom) used in the HPS test run 2012.

that described in Section 4.6.3; the same hardware (FADC, CTP, and SSP) was used.

The key differences were in the FADC integration algorithm, the energy resolution reported by the FADC to the CTP, and the SSP trigger decision algorithm.

In the test run, the FADC used a time-over-threshold algorithm to integrate hit energy for use in trigger decision, where only samples above threshold were summed. This algorithm is shown in Figure 39.

The FADC reported hits to the CTP in an 8-bit format consisting of a 5-bit channel sum and a 3-bit timestamp. This meant that the pulse integral had to be truncated to fit in 5 bits; consequently some energy resolution was lost.

For the test run, the simplified trigger logic in the SSP was a simple threshold: the trigger was configured to fire on a single cluster with energy exceeding a given threshold. However, the full trigger described in Ref. [6] was implemented and tested.

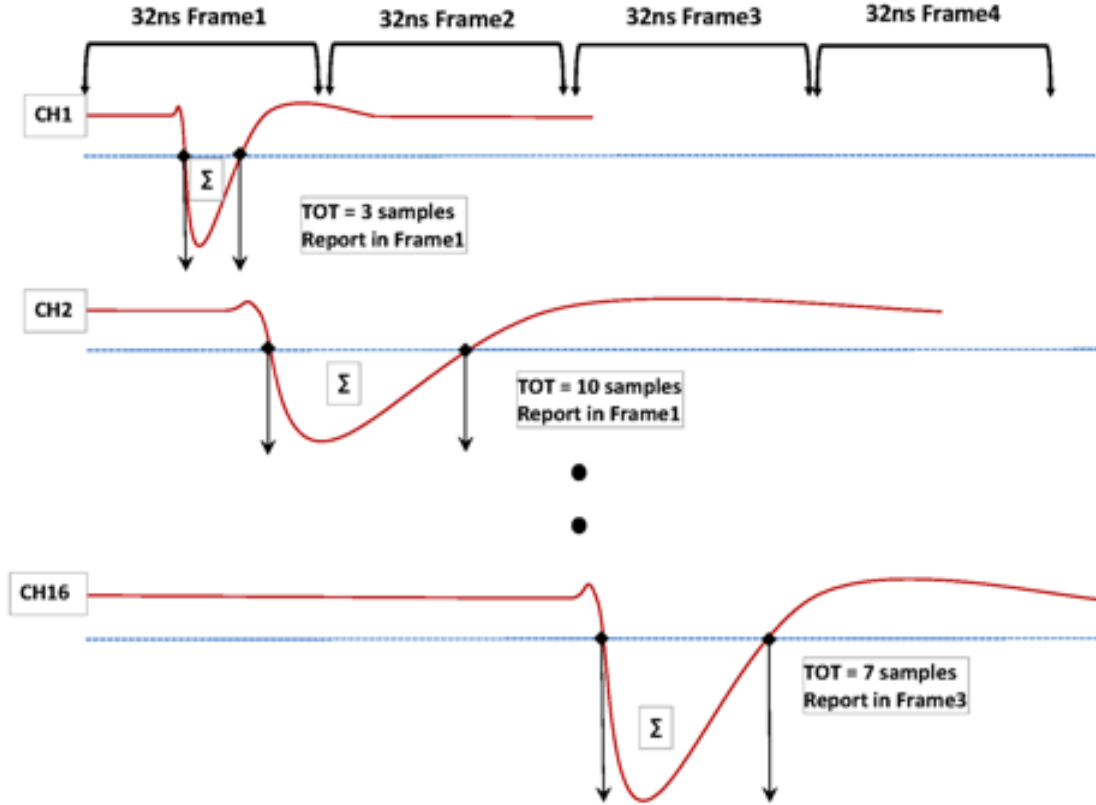


FIG. 39: Example of input signals, and how they are integrated and digitized for the test run trigger.

## 5.2 Multiple Coulomb Scattering Measurement

Occupancies close to the beam create many of the key challenges in the HPS experiment and determine the limits of sensitivity to low  $A'$  masses. These occupancies are dominated by electrons which have multiple Coulomb scattered to relatively large angles in the converter. Because HPS is sensitive to scattering angles far out on the tail of the multiple Coulomb scattering distribution, well beyond the angles important in other experiments, care must be taken to ensure our simulations are correct in this regime. In particular, Geant4 overestimates the multiple Coulomb scattering rate by a factor of two at large angles as explained in detail in the appendix (see Fig. 90). One of the main goals of the test run in 2012 was to evaluate the description of the tails of the multiple Coulomb scattering in order to gain further confidence in our expected detector occupancy. As will be shown below, data from the test run can be used to confirm our model of multiple Coulomb scattering despite the

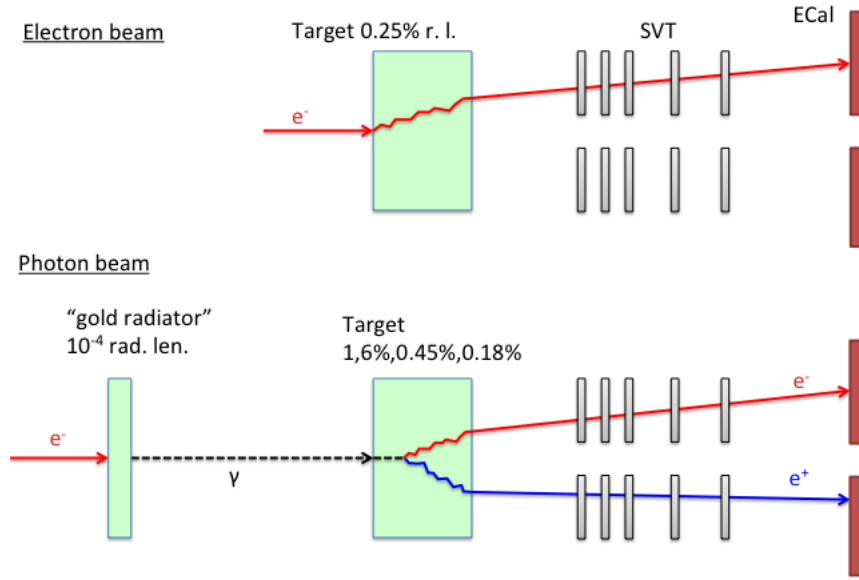


FIG. 40: Schematic comparison of the the setup in the test run photon beam compared to the HPS electron beam.

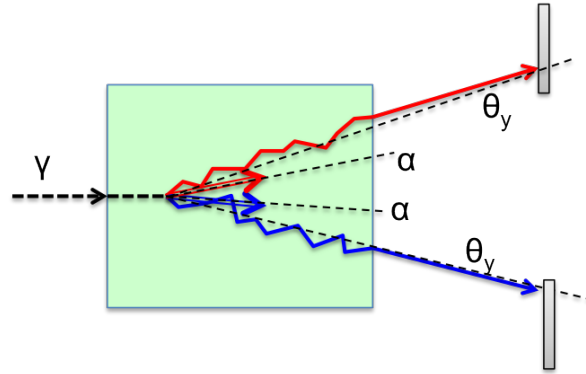


FIG. 41: Schematic description of the relevant angles for pair production in the test run.

fact that all data were taken with a photon beam.

Figure 40 gives a schematic view of the main differences between the photon and electron beam setup. In particular, the angular distribution of the pair produced electron and positions emerging from the converter has comparable contributions from *i*) the pair production angle and *ii*) the multiple Coulomb scattering of the electron and positron in the converter after production, see Fig. 41.

For this analysis, we measure the angular distribution of electrons and positrons using the



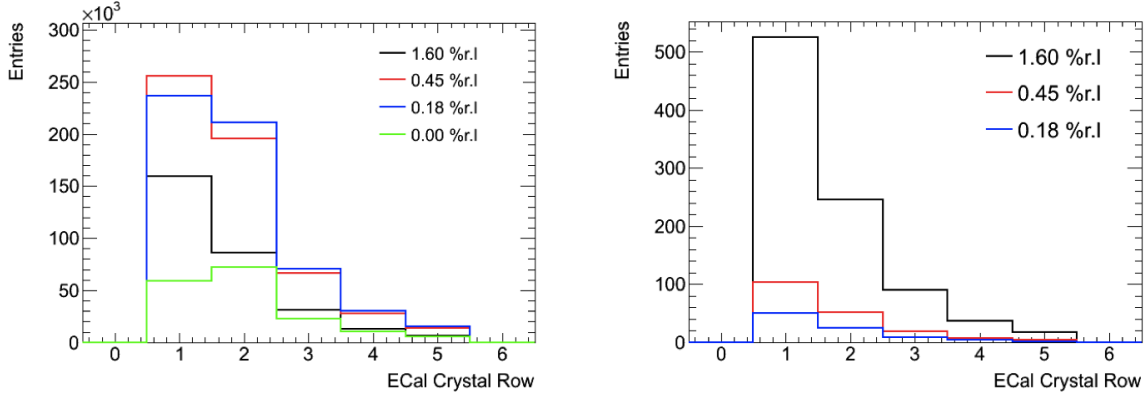


FIG. 42: Measured raw vertical angular distributions before (left) and after (right) normalization and background subtraction.

Converter thickness (%r.l.)	Duration (s)	$e^-$ on converter ( $\mu\text{C}$ )
1.6	911	24.4
0.18	2640	193.5
0.45	2149	140.7
0	1279	88.1

TABLE X: Measured integrated currents for the runs used to measure the angular distributions.

ECal. Cluster reconstruction was done using the algorithm described in [5]: build clusters around seed hits (hits above a “seed” energy threshold and with greater energy than any neighboring hits), and add all neighboring hits above an “add” energy threshold. Hit energy is calibrated by matching track momentum to cluster energy, as described in Appendix D.

The measured angular distribution in the ECal for the three converter thicknesses are shown in Fig. 42 (left). The photon beam line during the test run produced a relatively large fraction of pairs originating upstream of the converter. This contribution was measured during data taking with “empty” converter runs i.e. removing the converter but with all other conditions the same. The upstream background measured in the “empty” converter runs was subtracted from the other runs, properly normalized using the measured integrated currents detailed in Tab. X. The background fraction for the three converter thicknesses was 16%, 52% and 71% for converter thicknesses of 1.6%, 0.45% and 0.18%, respectively. The background fraction was also cross-checked by pointing back tracks reconstructed in the SVT to identify the fraction of tracks not emanating from the converter. This can be seen in Fig. 50 (bottom) where small satellite peaks at  $\pm 10$  mm can be identified as tracks from the upstream background. The angular distributions, after normalization and subtracting

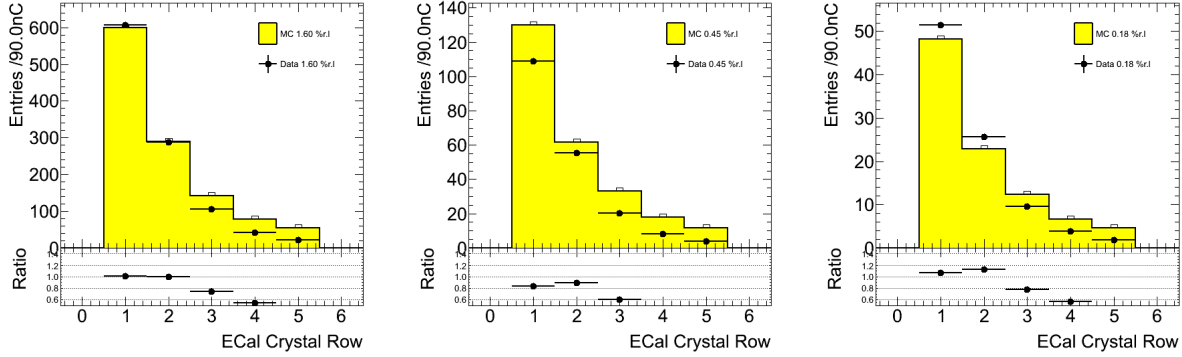


FIG. 43: Comparison between the observed and predicted angular distribution using EGS5 for converter thickness of 1.6% (left), 0.45% (middle) and 0.18% (right). Only statistical uncertainties are included.

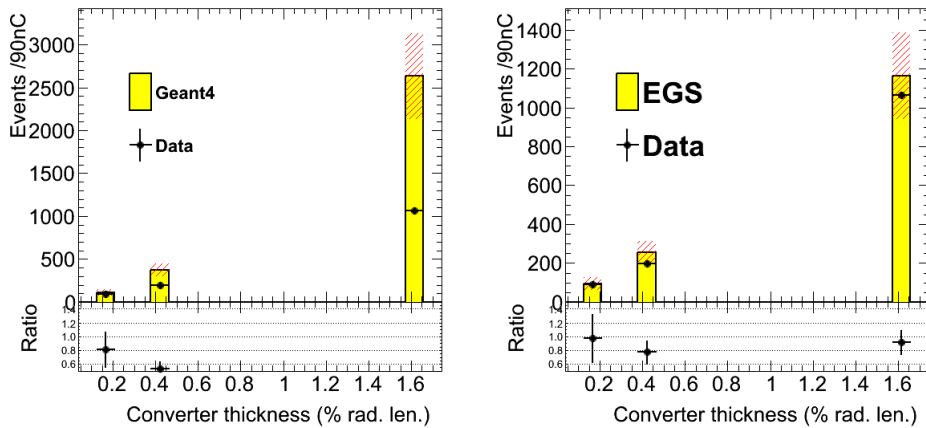


FIG. 44: The measured rate as a function of converter thickness comparing GEANT4 (left) and EGS5 (right).

the upstream background, are shown in Fig. 42 (right). We also checked that the contribution from photons to our triggered sample was less than 2% (without angular selections which would further reduce the contribution).

These measured angular distributions are compared to simulation to validate the modeling of the multiple Coulomb scattering. As described in more detail in Appendix C, EGS5 is used to generate the electromagnetic interactions in the converter while GEANT4 is used to simulate the particles after the converter. Figure 43 shows the angular distribution comparing data and EGS5 normalized to 1 s of beam at 90 nA beam current. The total rate measurements are in Fig. 44 and summarized in Tab. XI. The total systematic uncertainty was estimated to be between 10-18% depending on the run including: a 5% uncertainty

Converter (% r.l.)	1.60	0.45	0.18
EGS5	$1162 \pm 112$	$255 \pm 28$	$94 \pm 17$
GEANT4	$2633 \pm 250$	$371 \pm 38$	$114 \pm 18$
Observed	$1064 \pm 2$	$196 \pm 1$	$92 \pm 1$

TABLE XI: Observed and predicted number of events for 1 s of beam at 90 nA for three different converter thicknesses. The uncertainty on the prediction includes systematic uncertainties. The uncertainty on the observation is statistical.

on the integrated current normalization, alignment of the ECal, uncertainty from the background normalization, and limited Monte Carlo statistics.

In summary, the accurate modeling of the multiple Coulomb scattering is fundamental to estimate occupancies and trigger rates for HPS. EGS5 predicts the correct angular distribution across all converter thicknesses while GEANT4 overestimates the rates; with the disagreement increasing with larger converter thickness. This preliminary result verifies our modeling of the multiple Coulomb scattering using EGS5 for HPS.

### 5.3 Test Run Apparatus Performance

As previously noted, all running of the HPS Test apparatus was with photon beams, using the Hall-B pair spectrometer (PS) pair converter as a target. This target, located  $\sim 77$  cm from the first layer of the tracker, produced a sufficient flux of electrons and positrons to test the principles of running the HPS experiment. This section will report on a few selected preliminary results that demonstrate our understanding of the system.

#### 5.3.1 SVT Performance

During the duration of the dedicated photon run all SVT hybrids and APV25 read-out chips were configured to their nominal operating points [100] while all sensors were reverse-bias at 180 V. The sensors were operated within a temperature range of 20 to 24°C throughout the test run. Multiple calibration runs established a noise level of 60-68 ADC counts ( $\approx 750 - 850$  electrons) which was stable across all hybrids. With a linear gain up to  $\approx 3$  MIPs, the cluster charge for hits associated with a track follow the characteristic Landau shape with a mean of about 25,500  $e^-$  as expected, see Fig. 45.

One of the important tests of the SVT was the operation of the APV25 chips in multi-

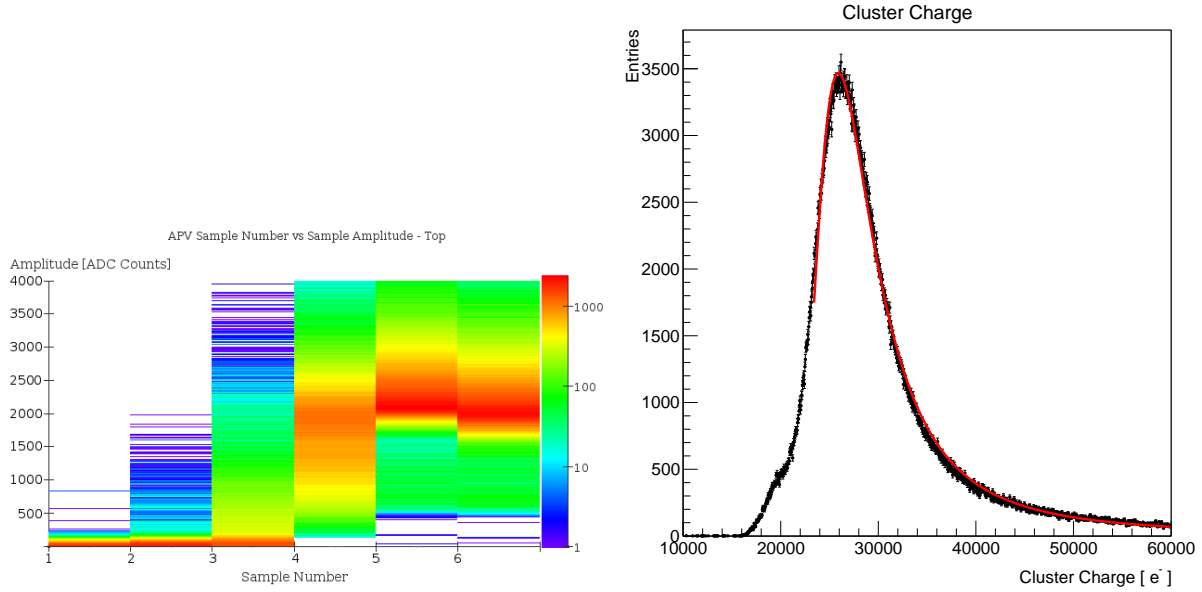


FIG. 45: The six pedestal subtracted samples associated with a hit on a track are shown on the left plot along with a distribution of the cluster charge exhibiting the characteristic Landau shape on the right.

peak readout mode discussed in Sec. 4.4. The six samples of the APV25 pulse shaper output are fitted to an ideal  $CR - RC$  function to extract the amplitude and the  $t_0$  of the hit. The typical pulse shape obtained is shown in Figure 45 also demonstrates that the SVT was well timed in to the trigger with the rise of the pulse at the 3rd sampling point. After clustering hits on a sensor, the hit time for each cluster is computed as the amplitude-weighted average of the fitted  $t_0$  channel times. The  $t_0$ -resolution is studied by comparing the cluster hit time with the average of all cluster hit times, the “track time”. Figure 46 shows the track time, with the expected jitter due to clock phase and trigger, and the residual to the individual cluster times. After correcting for offsets from each sensor (time-of-flight, clock phase) the extracted  $t_0$  resolution is 2.6 ns. This is somewhat worse than the  $\approx 2$  ns resolution expected in Section 6 which we attribute to the true pulse shape differing from our idealized fit function; work is in progress to use the true pulse shape in the fit. Reducing the APV25 pulse shaping time will also improve time resolution. In short, these results show that we can achieve time resolution adequate for pileup rejection during electron running.

Throughout the duration of the test run, approximately 97% of the 12,780 SVT channels were found to be operating normally. The fraction of dead or noisy channels varied from 2.4% to 4.7%; most of these were due to misconfigured readout chips (2–4 misconfigured

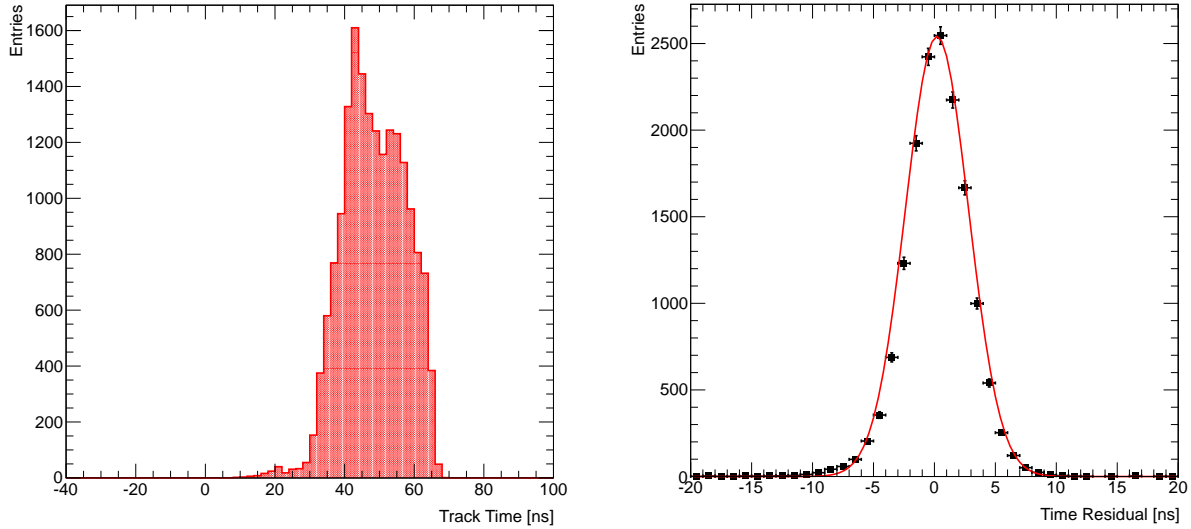


FIG. 46: Track time distribution (left) and cluster time residual (right). The track time is measured relative to the APV25 clock. The width of the distribution is due to trigger jitter (24 ns jitter in the tracker readout clock, plus 16 ns jitter in the trigger system). The cluster time residual is for a representative sensor relative to the track time.

chips out of 100 total) a known noisy half-module, and a couple of known noisy readout chips. These issues will be resolved for future running.

This resulted in occupancies and data rates that were higher than what were expected from simulation; the maximum data rate observed in the SVT was 4.1 MB/s. However, after masking out all known noisy channels found during the commissioning of the SVT, good agreement between simulation and test run occupancies was achieved as shown on Figure 47.

Similarly, the hit efficiency was measured to be above 98% for known good layers, see Figure 48.

The spatial resolution of similar microstrip sensors is well established by test beam data, against which the charge deposition model in the lcsim Monte Carlo is validated. This resolution can be parameterized as a function of the total signal to single-strip noise ( $S/N$ ) and the crossing angle of tracks through the sensor. The single-hit resolution for charged particles with  $S/N > 20$ , as demonstrated here, is relatively constant at approximately 6  $\mu\text{m}$  for tracks that are close to normal to the sensors as in HPS.

The SVT was aligned using a combination of optical, laser and touch probe surveys at SLAC and JLab. The optical survey of individual modules with precision of a few microns are combined with a touch-probe survey of the overall SVT support structure, with 25-

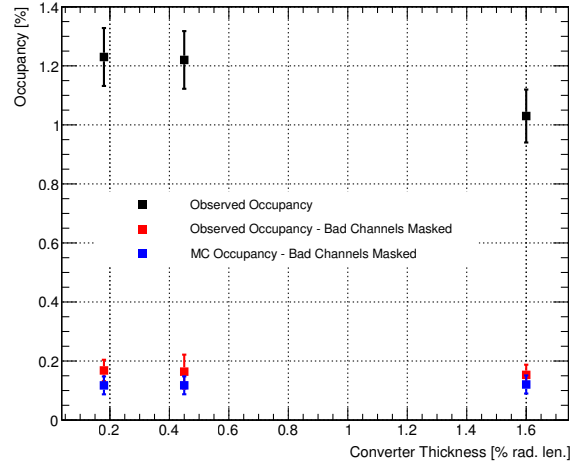


FIG. 47: Comparison of SVT occupancy between test run data (before and after masking noisy channels) and that expected from simulation.

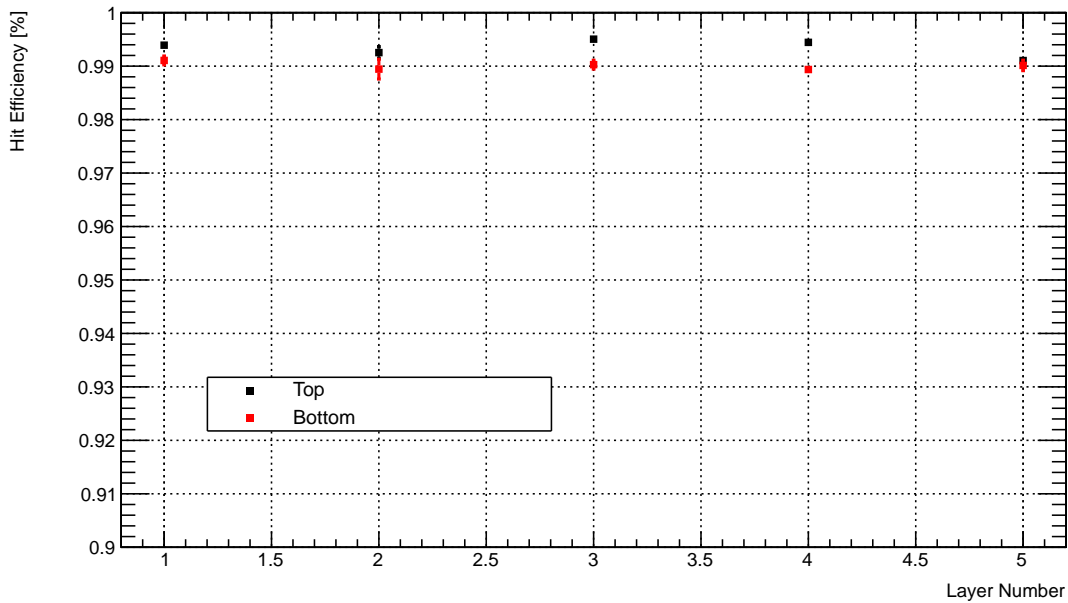


FIG. 48: The hit reconstruction efficiency as a function of detector layer. The variation across the layers can be explained by known DAQ issues.

100 microns precision, to locate the silicon sensor layers with respect to the support plates and the mechanical survey balls on the base plate. After full assembly and installation of the SVT at JLab, a mechanical survey of the SVT base plate position inside the pair spectrometer vacuum chamber is used to determine the global position of the SVT with respect to CEBAF beam line. The resulting survey-based alignment has the position of the

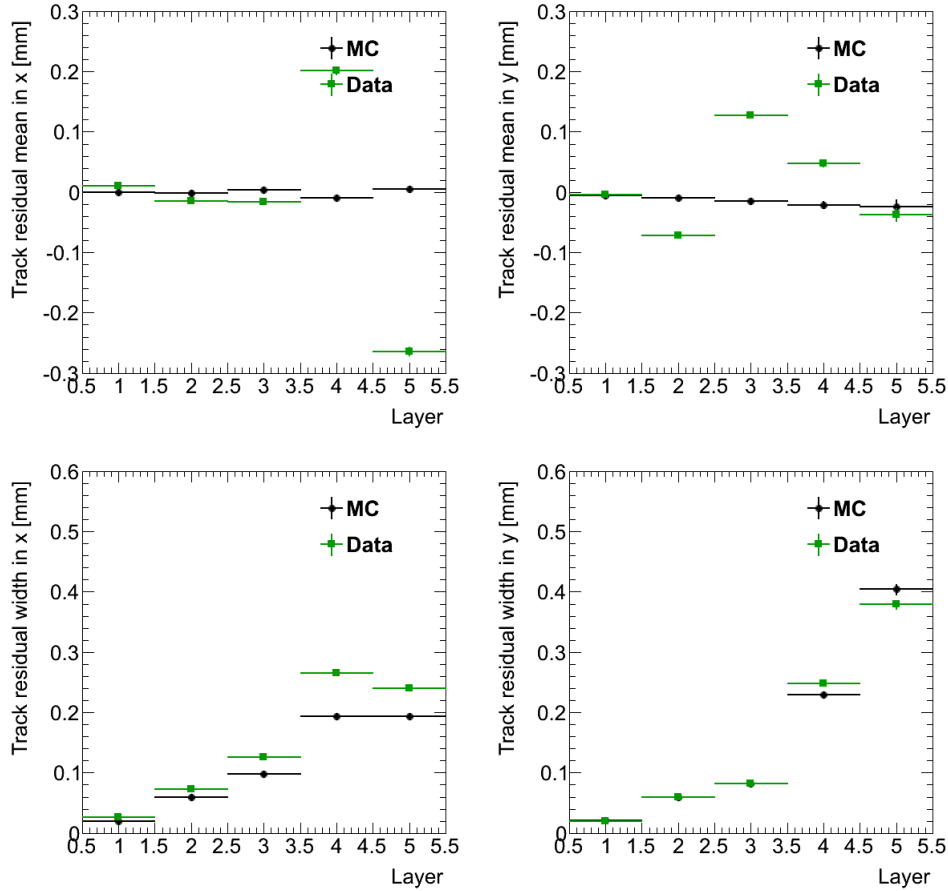


FIG. 49: Mean (top) and standard deviation (bottom) of biased residuals (i.e. hits are included in the track fit) between the actual hit position and the predicted position from the reconstructed tracks in the bend (left) and non-bend (right) plane in the top half of the SVT after mechanical survey. The smaller width for the 5th layer in the bend plane is an effect from mixing tracks with four or five hit tracks.

silicon sensors correct to within a few hundred microns as shown in the mean of the biased track residuals in Fig. 49. The large multiple scattering contribution can be seen by the large increase in the width of the residuals in the later layers. The agreement with simulation is reasonable; a slight track reconstruction algorithm bias can be seen in the mean for the simulation in later layers which will be fixed in the future.

We also extrapolate the reconstructed tracks back to the converter located  $\approx 77$  cm from our first silicon layer to understand the tracker alignment w.r.t. to the other components on the beam line. Figure 50 shows good agreement of the reconstructed track position at the converter with that predicted from simulation using the measured field map of the analyzing magnet to take into account the fringe field. The offset of the horizontal position simply

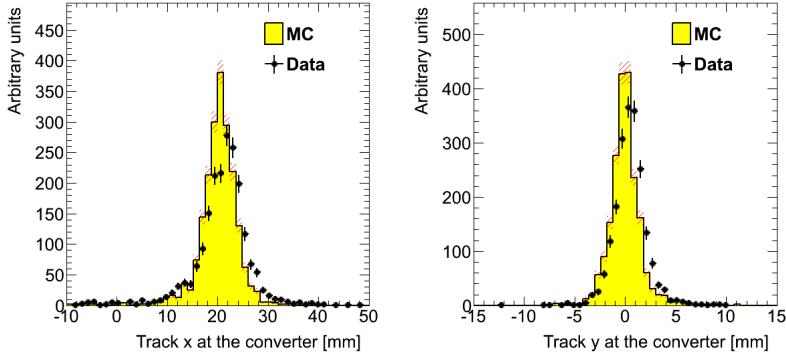


FIG. 50: Extrapolated track positions for reconstructed  $e^+e^-$  pairs in the SVT taking into account the measured fringe field of the analyzing magnet in data and a simulation with ideal geometry.

reflects the fact that the positions are reconstructed in an SVT-centered coordinate system, which is tilted with respect to the beam coordinate system.

With initial residuals less than  $\sim 500 \mu\text{m}$  across all layers of the tracker and a reconstructed beam profile similar to that expected from simulation, it appears these survey techniques are adequate to bootstrap the SVT alignment. For HPS, we are developing a more sophisticated global track-based alignment technique to reach the final alignment precision. This framework will also enable us to explore and understand important details such as weak modes and how dedicated alignment runs (e.g. with magnetic field off or with different targets) may shape operational procedures during HPS running.

By selecting  $e^+e^-$  pairs from the triggered events we are able to study basic distributions of pair production kinematics and in particular those related to our vertex performance. Pairs of opposite charge tracks, one in the top and one in the bottom half of the SVT, with larger than 400 MeV were selected. The pair production kinematics are relatively well reproduced given the alignment of the tracker; Fig. 51 shows the invariant mass and ratio of electron momentum over the sum of electron and positron.

Referring again to Figure 50, it is apparent that the extrapolated track resolution is as predicted by Monte Carlo. The track resolution is several mm at the conversion target location, but that target is  $\sim 67$  cm upstream of the nominal HPS target position, so the resolution is severely degraded by the long extrapolation.

The good agreement confirms the Monte Carlo description, which predicts that the resolution at the actual HPS target location, just 10 cm upstream of the first silicon layer, will be in the  $\sim 100$  micron range, depending on track momentum. This agreement at the



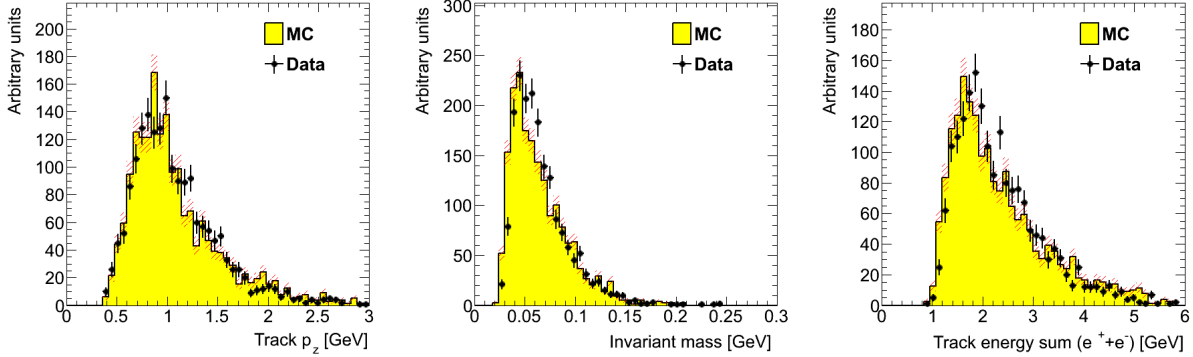


FIG. 51: Kinematic distributions for  $e^+e^-$  pairs selected by opposite charged tracks in the top and bottom half of the tracker: track momentum in the top half of the SVT (left), invariant mass (middle) and the sum of the track momentum for the pair.

individual track level, which was obtained without extensive track-based alignment which will be performed in the future, also makes us confident that the HPS vertexing resolution will be as designed.

### 5.3.2 ECal performance

Of 442 crystals/channels, 39 were disabled or disconnected and were not read out by the DAQ. 13 of these were not read out because of a shortage of FADC readout boards. The remainder either had no HV bias on the APD, or were disabled in the FADC software due to noise. In the data, we identified two types of abnormal channels. One FADC was not sending trigger signals correctly, resulting in low efficiency. This affected the 13 channels read out by that FADC. 5 channels were diagnosed as noisy because they had a high incidence of hits out of coincidence with the trigger. A large number of channels were originally misidentified as noisy because they had much higher hit occupancy than neighboring channels. Gain calibration shows that these channels have high gain (and thus lower energy threshold) but are otherwise normal. The abnormal channels were ignored in analysis in order to simplify comparison with Monte Carlo. This leaves 385 useful channels—87% of the ECal.

Each ECal crystal was calibrated for pedestal, noise and gain. This is described in Appendix D.

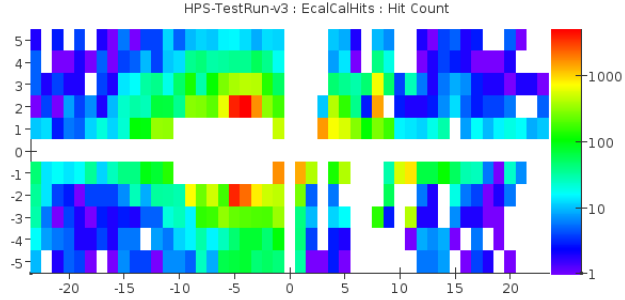


FIG. 52: Hit rates (above an energy threshold, to cut out effect of gain variations) vary smoothly with position.

### 5.3.3 Trigger performance

As described in Section 5.1.4, the trigger and DAQ integrate pulses differently to measure hit energy. The trigger integrates using a time-over-threshold window, and the DAQ readout integrates using a constant window (5 samples before and 30 samples after a threshold crossing). For every event, the trigger reports as a bitmask the trigger decision (top trigger, bottom trigger, or both) and the time the trigger fires.

We study trigger performance by simulating the trigger for each event and comparing to how events were actually triggered. First, we simulate the FADC to convert from readout hits (constant integration window) to trigger hits (time-over-threshold integration). We then simulate the CTP clustering algorithm and the trigger decision (described in Section 5.1.4), and compare the trigger decision and trigger time reported by the simulation to what was reported by the real trigger.

To eliminate trigger bias in checking the trigger decision, we use a tag and probe method: to check trigger performance in one half of the ECal, we tag events where there was a trigger in the other half, and exactly one probe cluster in the ECal half under test. We then measure trigger efficiency (proportion of tagged events where there was a trigger) as a function of ADC counts and energy of the probe cluster. These turn-on curves are shown for the top half of the ECal in Figure 53. The trigger threshold is seen to be 1280 ADC counts as expected. The threshold is not perfectly sharp in this analysis because of uncertainties in the conversion from readout to trigger hits but based on comparisons with Monte Carlo simulation we believe the trigger worked exactly as specified. The trigger threshold in terms of cluster energy is very uneven for two reasons; gain variations between different ECal

crystals lead to threshold variations and the nonlinearity of the time-over-threshold integral means that the effective threshold is higher for clusters that span multiple crystals. Overall

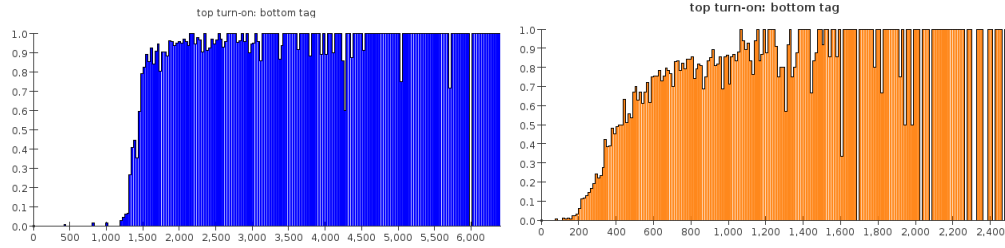


FIG. 53: Trigger turn-on as a function of probe cluster ADC counts (left) and probe cluster energy in MeV (right). Both plots are for the top half of the ECal; bottom is similar. Energy is not corrected for sampling fraction.

the trigger appears to have functioned exactly as intended. Changes planned for the next run (constant integration window and per-crystal gain calibration constants for the trigger) will solve both of the issues that led to threshold variations in the test run.

## 6 HPS Performance Studies

We use the HPS detector simulation system based on SLAC's `org.lcsim` infrastructure for full GEANT4 simulation of the passage and interaction of charged and neutral particles through the SVT and the ECal to the muon detector. In the SVT, the simulation creates realistic energy deposits in the silicon microstrip detectors, accounts for dead material, simulates APV25 signal sampling every 25 ns, creates clusters, and performs track finding and reconstruction. In the ECal, the geometry for the flange and vacuum chamber is based on a tessellated representation imported directly from the CAD drawings. It creates energy deposits in individual trapezoidal-shape  $PbWO_4$  crystals, simulates FADC signal time evolution and sampling every 4 ns, and generates triggers based on the FPGA trigger algorithm implementation. To maintain the chicane beamline configuration, the field strength of the chicane magnets must scale with the beam energy. The performance studies were made using the field strength of the analyzing magnet of 0.25 Tesla at 1.1 GeV, 0.5 Tesla at 2.2 GeV, and 1.5 Tesla at 6.6 GeV. Figure 54 shows a `lcsim` rendering of the HPS detector.

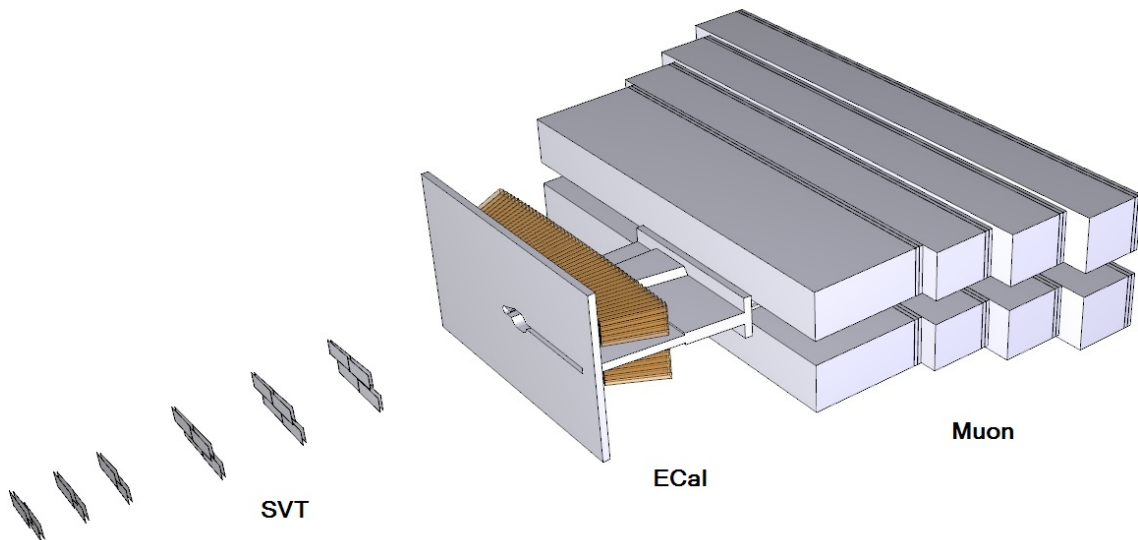


FIG. 54: Rendering of the HPS detector simulation.

## 6.1 Simulation of Backgrounds and Detector Occupancies

### 6.1.1 Simulation of Backgrounds

The multiple Coulomb scattering and bremsstrahlung processes in the target will generate high intensity fluxes of electrons and photons in the very forward direction, while the large Møller interaction cross section with atomic electrons will generate high intensity low energy electrons. We use the high energy interaction simulation tools GEANT4 and EGS5 to simulate these backgrounds. In the original HPS proposal to JLab PAC37 [5], we described a significant disagreement between these tools. GEANT4 predicted a broader angular distribution of multiple scattered electrons than EGS5, resulting in twice the occupancy in the tracker near the dead zone and much higher ECal trigger rates. The HPS Test run was motivated in part by the need to resolve this discrepancy, and the outcome of the test run is described in the previous section. The algorithms used in the codes to simulate the multiple scattering have been studied, and the findings are summarized in Appendix C. The test run result and the algorithm studies have confirmed that EGS5 can describe the multiple scattering tails more accurately than GEANT4. All the electromagnetic interactions in the target are simulated with EGS5.

When bound electrons in the target are ionized by incoming electrons or secondary photons, outer shell electrons will fill the vacancy and characteristic X-rays are emitted. These X-rays can contribute background hits in the SVT when a conversion takes place in the silicon sensors via the photoelectric effect or Compton scattering. Since X-ray production from electrons interactions is not fully simulated in EGS5, we estimate the X-ray intensity at SVT Layer 1 using the impact ionization cross section,  $\sigma_I$ , [102], the fluorescence yield,  $\omega$ , [103], the photoabsorption length in Tungsten,  $\lambda_W$ , to account for the self-absorption, and the solid angle of the SVT Layer 1. Table XII summarizes these parameters and the expected X-ray fluxes at SVT Layer 1 for 0.25%  $X_0$  Tungsten and 100 nA beam current in 8 ns time window. X-ray backgrounds are not a concern.

Hadrons are also produced in the target. Hadron production is at least three orders of magnitude smaller than the electromagnetic interaction. The polar angle for hadron production is predominantly larger than 100 mrad, whereas the HPS detector acceptance is limited to less than 100 mrad. Furthermore, the hadron energy spectrum is soft as they are produced

	Energy (keV)	$\sigma_I$ (barns)	$\omega$	$\lambda_W$ ( $\mu\text{m}$ )	$N_\gamma$ at Layer 1 in 8 ns
K-shell	60	40	0.95	100	0.5
L-shell	10	1000	0.30	5	2
M-shell	2	20000	0.02	0.2	0.1

TABLE XII: X-ray intensities.

from the  $1/k$  bremsstrahlung spectrum and more than 90% of the hadrons are swept away by the analysing magnet before reaching the ECal. The hadron production is simulated using GEANT4 and FLUKA. In a target thinner than about 5%  $X_0$ , the “virtual” photon interaction is dominant [104]. The inclusive hadron production  $\sigma(eA \rightarrow X)$  is simulated from the photonuclear process  $\sigma(\gamma A \rightarrow X)$  using the equivalent photon approximation,

$$\sigma(eA \rightarrow X) = \int \sigma_k(\gamma A \rightarrow X) dn(k),$$

where  $dn(k)$  is the number of equivalent photons with energy  $k$  [105] and there are approximately  $8 \times 10^{10}$  photons/sec in 6.6 GeV 100 nA beam. Table XIII summarizes the pion single rates from 1%  $X_0$  Tungsten target and 6.6 GeV 100 nA beam. While pion production is larger in GEANT4, the energy spectrum is softer and consequently the single rate of pions reaching the ECal is lower in GEANT4. While pions look like a minimum ionizing particle in the ECal most of the time, they can deposit significant energy when  $\pi^0$  are produced in the ECal crystals, and together with the beam background, they contribute accidental coincident triggers.

	Total production rate (kHz)	Single rate reaching the ECal (kHz)
GEANT4	410	8
FLUKA	240	15

TABLE XIII: Pion single rates from 1%  $X_0$  Tungsten target at 6.6 GeV 100 nA.

Other beam induced background we considered are:

- Beam halo

Beam halo was measured using a large dynamic range halo monitor during the 6 GeV era. The beam halo that extends to 2 mm was found at the level of  $10^{-7}$ . At this level, the halo contribution in the SVT occupancy is negligible. It is expected that the behavior of the 12 GeV machine will be understood at the same level.

- Synchrotron radiation

Synchrotron radiation is produced from the last dipole magnet in the beam line in the vertical plane, and from the chicane magnets in the horizontal plane. Since the characteristic energy is proportional to  $E_{beam}^2$ , synchrotron radiation is of concern only at 6.6 GeV. The characteristic energy ( $k_c$ ), the average energy ( $k_{ave}$ ), and the power of the synchrotron radiation is summarized in Table XIV. None of the radiation from the last dipole will enter the HPS detector as the radiation will be intersected by the beamline collimator. The radiation from the chicane magnets is in the dead zone, and none of the detector components are designed to intersect the beam plane.

Source	$k_c$ (keV)	$k_{ave}$ (keV)	$N_\gamma$ per e-	Power (mW) at 100 nA
Vertical bend	19	5.9	4.0	2.4
Frascati Magnet	52	16	4.6	7.4
PS magnet	44	14	9.3	13

TABLE XIV: Synchrotron radiations at 6.6 GeV.

- EM induced backgrounds

Electromagnetic fields induced by the high intensity beam could interfere with the SVT and its electronics as the detector is located as close as 0.5 mm from the beam. We have evaluated the direct beam field and its wake field, the diffraction radiations from the beamline apertures, and the transition radiations from the target. The intensities of these EM induced backgrounds are small and no interference with the SVT is expected. The beam charge per 2ns CEBAF bunch is only few thousand electrons, many orders of magnitude lower than that in other experiments which have chosen to shield against it.

### 6.1.2 Simulated Tracker Occupancies

Figure 55 shows the distribution of charged particle hits in the SVT Layer 1 which is located 10 cm from the target. The beam energy is 6.6 GeV, and the target thickness is 0.25%  $X_0$ . Multiple Coulomb scattered beam electrons are confined to within 0.5 cm of the beam axis ( $x=y=0$ ), while the low energy Møller electrons are distributed in a parabolic shape. There are very few positrons. From these distributions, the detector occupancy in the horizontal silicon strip in the 8 ns time window is calculated for a 400 nA beam current and five different target thicknesses, 1.0%  $X_0$ , 0.5%  $X_0$ , 0.25%  $X_0$ , 0.1%  $X_0$ , and 0.05%  $X_0$ , as shown in Fig. 56. For a 0.25%  $X_0$  target and 430 nA beam, the occupancy is 1% at a distance of 1.5mm from the beam in Layer 1, which corresponds to a dead zone of  $\pm 15$  mrad. As long as the product of target thickness ( $T$ ) and beam current ( $I$ ) is constant, the same  $A'$  production rate is maintained. Since multiple scattering and hence the effective beam size is reduced in a thinner target, it is advantageous to use a thinner target and a higher current. Using the constraint that the occupancy is 1% at 15 mrad, we find the beam current  $I$  which gives this occupancy for each of several potential target thicknesses  $T$ . The quantity  $(I \cdot T)^{1/2}$ , which is approximately proportional to the sensitivity  $S/\sqrt{B}$ , is given in Table XV, showing how the sensitivity improves as the target thickness decreases.

Target thickness (% $X_0$ )	Beam Current (nA)	$\propto S/\sqrt{B}$
1.0	60	7.7
0.5	170	9.1
0.25	430	10.4
0.10	1330	11.6
0.05	2860	11.9

TABLE XV: Beam current yielding 1% occupancy in SVT layer 1 for various target thicknesses at 6.6 GeV, and the relative experimental sensitivities which result.

The run conditions for other possible beam energies are studied using the same criterion that the maximum occupancy in SVT Layer 1 does not exceed 1%. Table XVI summarizes the target thickness and proposed beam current.



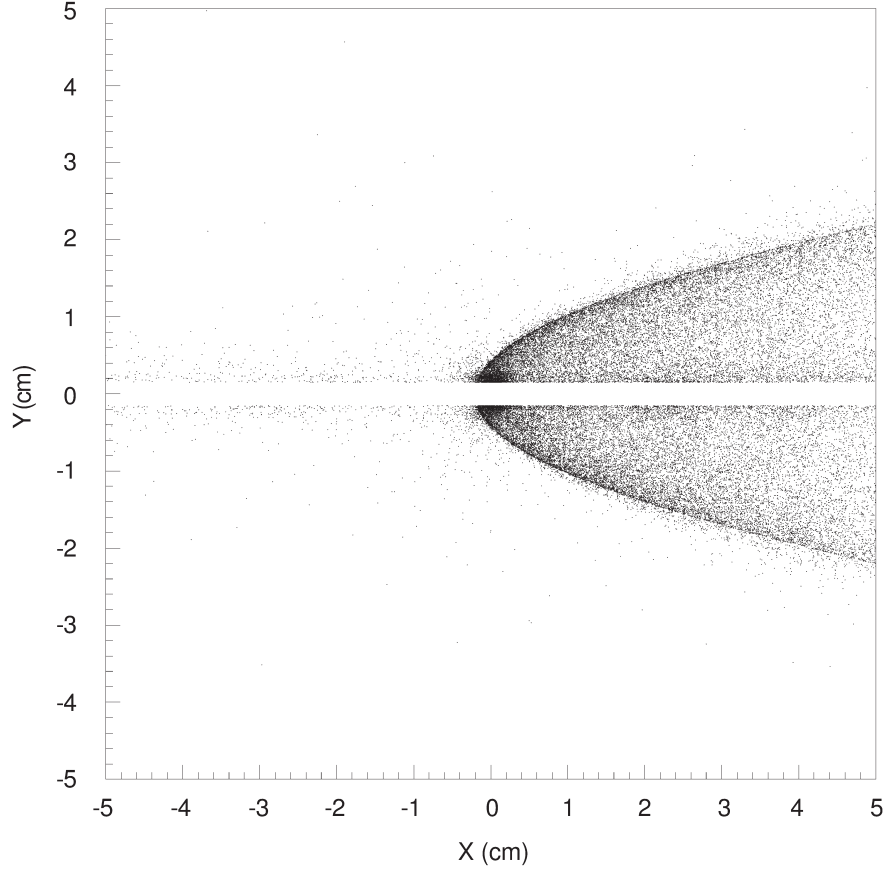


FIG. 55: Charged particle distribution in SVT Layer 1.

Beam Energy	Target thickness ( $\% X_0$ )	Beam Current (nA)
1.1	0.125	50
2.2	0.125	200
4.4	0.25	300
6.6	0.25	450

TABLE XVI: Run conditions.

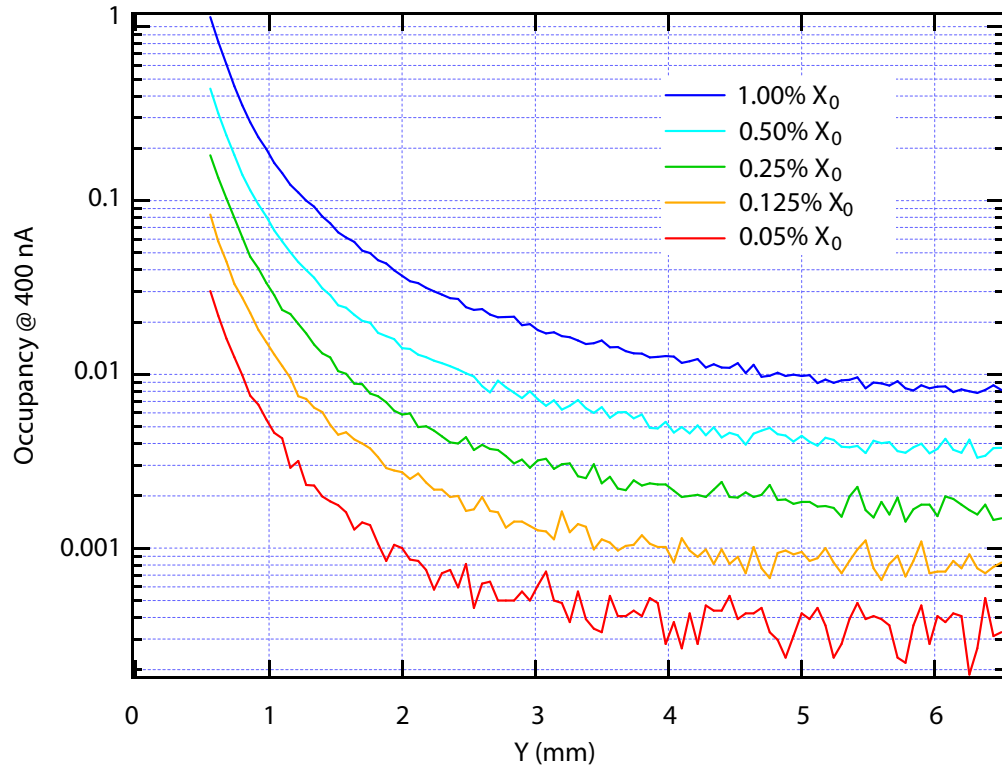


FIG. 56: Layer 1 silicon strip occupancy at 400 nA vs. distance from the beam in mm.

### 6.1.3 Simulated ECal Occupancies

There are two factors limiting the allowable ECal occupancy. First, the ECal readout algorithm uses a window of fixed size to integrate hit energy. This window was set to 140 ns ( $35 \times 4$  ns) for the test run, and so the number of hits above readout threshold in a 140-ns time window should be well below 1. Figure 57 shows that the maximum rate in any crystal is 500 kHz, which translates to 0.07 hits in 140 ns. Second, because the FADC only reads out on a rising threshold crossing, each hit above threshold causes dead time for that crystal until the pre-amplifier output falls back below threshold. Figure 58 shows the fraction of time each crystal spends above threshold. The maximum dead time is 0.03, meaning that even the hottest crystal is sensitive to new hits 97% of the time.

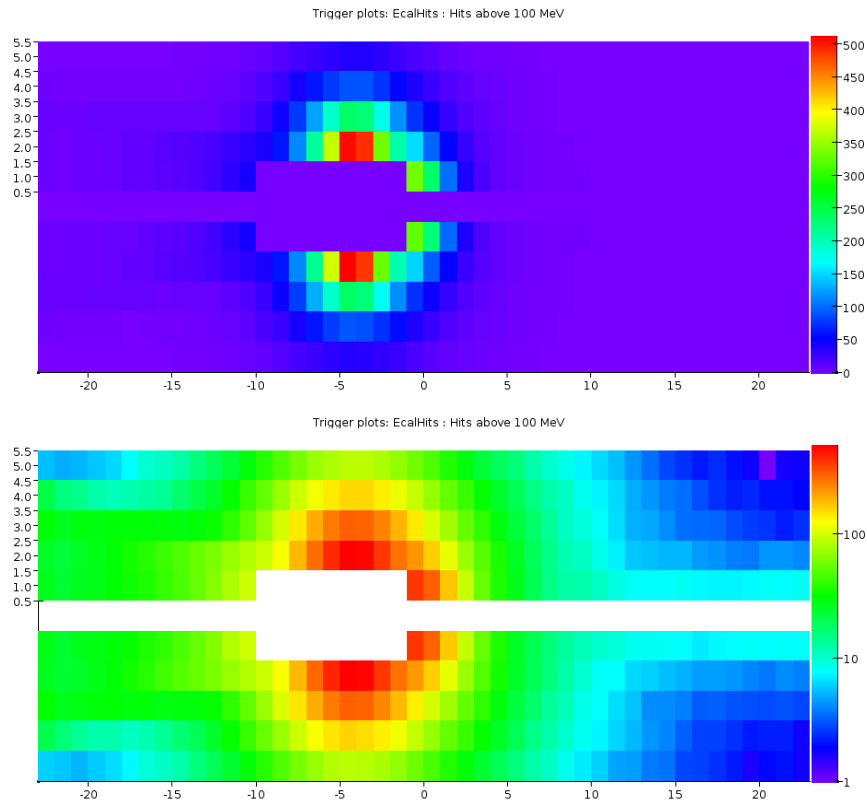


FIG. 57: Rate of hits over 100 MeV (units of kHz) per crystal (X and Y axes are the crystal index), for 2.2 GeV beam at 200 nA. Top plot uses linear scale for the Z-axis; bottom plot is log scale.

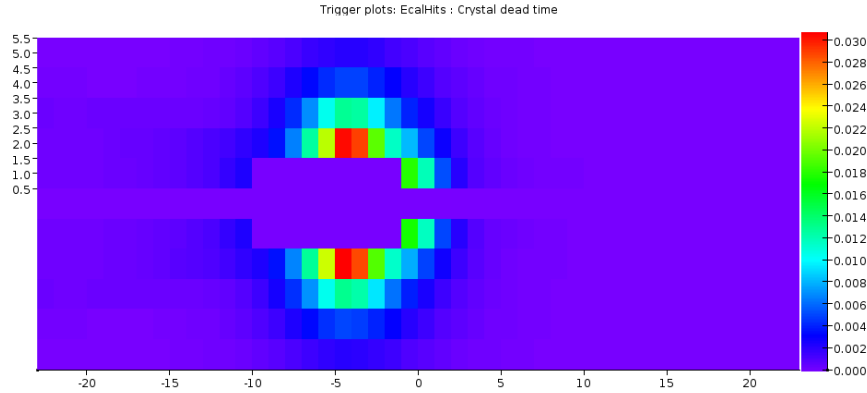


FIG. 58: ECal readout deadtime fraction for 2.2 GeV beam at 200 nA, with a threshold of 75 MeV for each crystal.

## 6.2 ECal Trigger Rates

The proposed ECal trigger was simulated to test trigger selections, verify that the trigger has acceptable efficiency for  $A'$  events, and verify that the trigger rate is compatible with the HPS DAQ in all running conditions.

The CEBAF beam bunch structure was simulated by sending one bunch equivalent of electrons, 625 (1.1 GeV), 2,500 (2.2 GeV) and 5,625 electrons (6.6 GeV), through the target. A total of 50 million bunches of beam background (equivalent to 100 ms of beam) were generated at each beam energy. The details of the target interactions are given in Sec. 6.1.1. Since the trident production process could not be done with EGS5, trident events were generated with MadGraph/MadEvent and overlaid on the beam background bunches with average rate expected from the trident cross section. For the trigger acceptance studies,  $A'$  events were generated with MadGraph/MadEvent at beam energies of 1.1, 2.2, and 6.6 GeV.

The complete chain of signal evolution in the ECal crystals and signal processing through the trigger system was simulated by following closely the ECal trigger description in Sec. 4.6.3. Starting from the energy deposits in the ECal crystals, signals were generated using the CR-RC shaper function with a time constant of 15 ns measured with the ECal crystals, amplitudes were sampled and pulse data evaluated every 4 ns (simulating FADC), and the cluster finding algorithm and trigger logic were applied (simulating CTP and SSP). The simulation has been tested against the actual performance of the test run detector and DAQ: see Sec. 5.3.3.

The cluster energy used in the trigger decision is not corrected for sampling fraction.

The trigger parameters described in Sec. 4.6.3 are chosen by running the simulation and plotting the relevant variables for beam background and A' events. This is done for each beam energy and a set of A' masses for each beam energy. Figure 59 shows the coplanarity angle vs. the azimuthal angle of the lower-energy cluster, indicating that A' events tend to have small coplanarity angles. Figure 60 shows the distance from beam axis vs. energy of the lower-energy cluster, indicating that the energy-distance cut can reduce the beam background effectively. Figure 61 shows the cluster energy difference vs. energy sum, indicating that the energy sum cut can retain A' events effectively.

These cuts are chosen to lie between the loosest reasonable values (accept as many A' events as possible) and the tightest (reject as many background events as possible). In some cases this leads to different cut values at different beam energies—for example, the coplanarity cut is looser at 1.1 GeV because the background events are clustered at large deviations from coplanarity and a relatively loose cut rejects most of them, but the cut is tighter at higher beam energies.

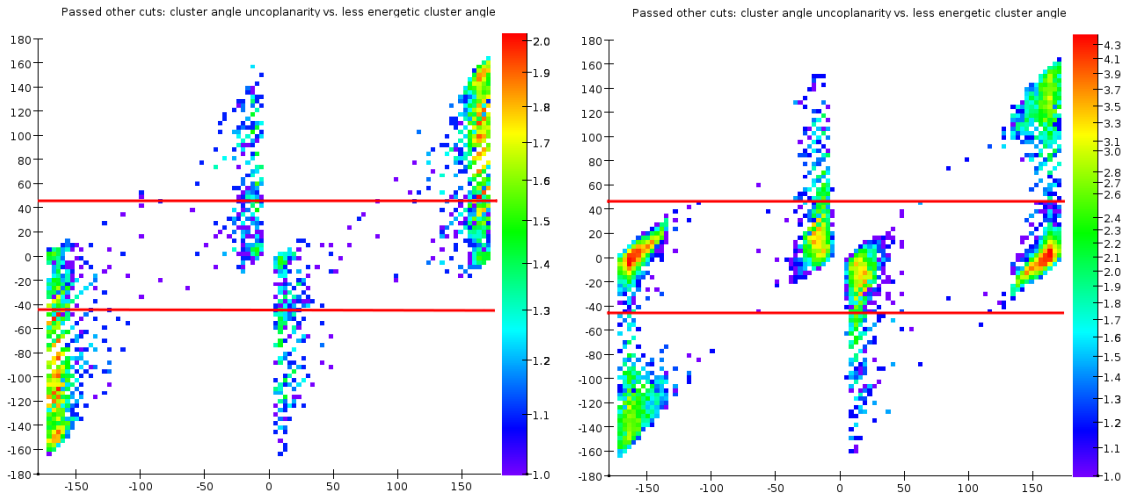


FIG. 59: Deviation of cluster pairs from coplanarity (units of degrees) for 2.2 GeV beam; background and 75 MeV A' tridents are shown. The X-axis is the azimuth around the beam axis ( $\phi_1$ ) of the lower-energy cluster, such that 0 degrees is the positron side of the detector and 180 degrees is the electron side; the Y-axis is the difference between the azimuth angles ( $\phi_1 - \phi_2 - 180$ ) of the two clusters. The coplanarity acceptance region is the space between the red lines.

The following trigger parameters were determined to be independent of beam energy:

- Minimum cluster energy ( $E_{min}$ ): 0.1 GeV

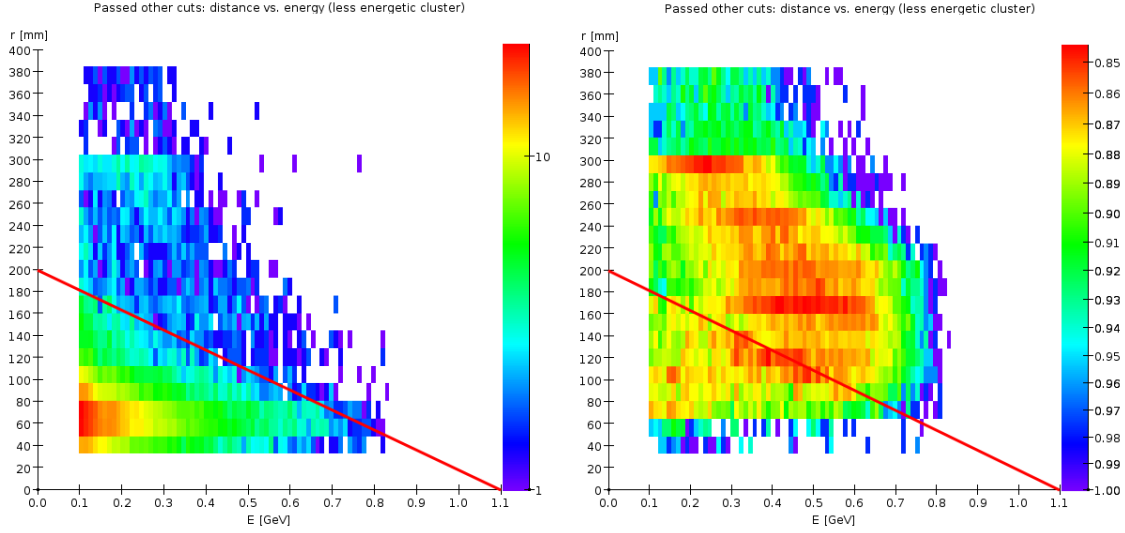


FIG. 60: Energy and distance from beam axis of the lower-energy cluster, for 2.2 GeV beam; background and 75 MeV A' tridents are shown. Cluster energy is not corrected for sampling fraction. The energy-distance acceptance region is above the red line.

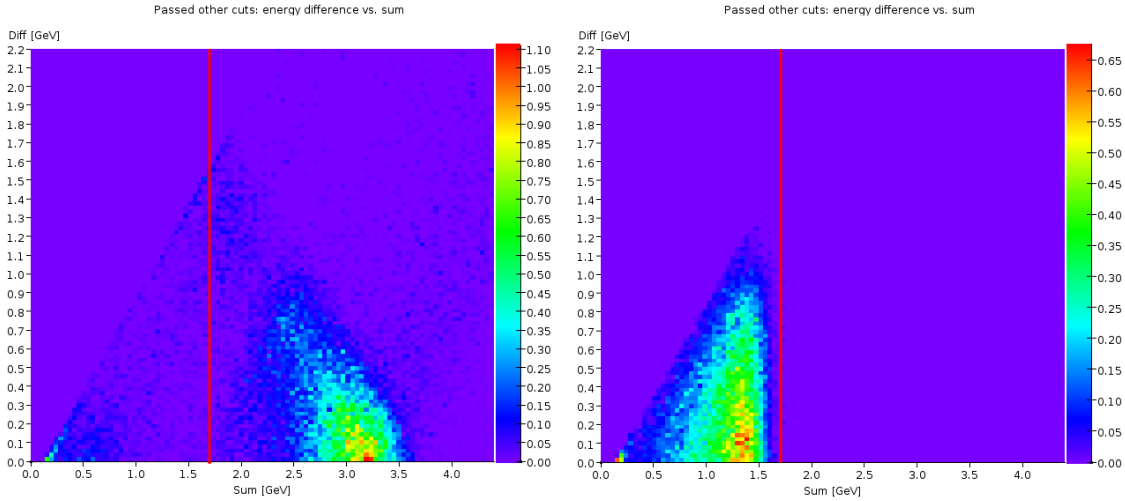


FIG. 61: Energy sum and difference of cluster pairs, for 2.2 GeV beam; background and 75 MeV A' tridents are shown. Cluster energy is not corrected for sampling fraction. The energy sum acceptance region is left of the red line.

- Distance ( $r_{edist}$ ) in the energy-distance cut: 200 mm
- Energy ( $E_{edist}$ ) in the energy-distance cut:  $0.5 \times E_{beam}$

Table XVII summarizes the trigger parameters that were found dependent on the beam energy. The remaining trigger parameters given in Sec. 4.6.3 did not have a significant effect on specificity of the trigger—neither the background trigger rate nor the trigger efficiency

Beam energy [GeV]	$E_{max}$ [GeV]	$E_{sum_{max}}$ [GeV]	$\Delta\phi_{max}$ [°]
1.1	0.7	0.8	90
2.2	1.6	1.7	45
6.6	5.0	5.5	60

TABLE XVII: Trigger parameters optimized for different beam energies.

Sample	Rate (kHz)
1.1 GeV beam background	$15.7 \pm 0.4$
1.1 GeV beam background+tridents	$18.3 \pm 0.4$
2.2 GeV beam background	$11.2 \pm 0.3$
2.2 GeV beam background+tridents	$15.8 \pm 0.4$
6.6 GeV beam background	$10.2 \pm 0.3$
6.6 GeV beam background+tridents	$12.6 \pm 0.4$
6.6 GeV beam background+tridents+pions (FLUKA)	$13.4 \pm 0.4$
6.6 GeV beam background+tridents+pions (G4)	$13.5 \pm 0.4$

TABLE XVIII: Trigger rates using various background samples, with statistical uncertainties.

is sensitive to those parameters..

Trigger rates are shown in Table XVIII. These rates are safely under the maximum readout rate of 43 kHz set by the SVT DAQ. Furthermore, tightening the coplanarity and energy-distance cuts lowers trigger rates to  $\approx 10$  kHz at 1.1 and 2.2 GeV and  $\approx 5$  kHz at 6.6 GeV, while reducing the  $A'$  efficiency by no more than 2 percentage points; this provides further safety margin in case trigger or data rates are higher than expected. The addition of pions to the 6.6 GeV background sample has only a small effect on the trigger rate.

Trigger efficiency for  $A'$  events is defined as the fraction of  $A'$  tridents (generated without fiducial cuts) that produce a trigger.

To evaluate the performance of the experiment, we are interested in the combined efficiency of the trigger and tracker: the fraction of  $A'$  tridents that produce a trigger and leave enough hits in the tracker for a pair of tracks to be reconstructed. We simulate charge deposition and readout of the tracker (turning off the generation of noise hits), and check each sensor for hits. If the DAQ reads out hits in four stereo pairs in each half of the tracker, the event is considered to be in the combined acceptance.

Figure 62 shows the ECal trigger efficiency and the ECal/SVT-combined efficiencies for  $A'$  events at 1.1, 2.2, and 6.6 GeV. Both trigger and tracker acceptances are dominated by the geometric acceptances of the ECal and tracker.  $A'$  prompt decays are assumed. The efficiency for decays displaced from the target decreases slowly (a few percent at 5 cm) with

decay distance for lower mass  $A'$ 's, but increases for higher masses. Some of the efficiency lost at long decay distances can be recovered by dropping the requirement that the first stereo pair has a hit, provided the subsequent track would miss the first layer.

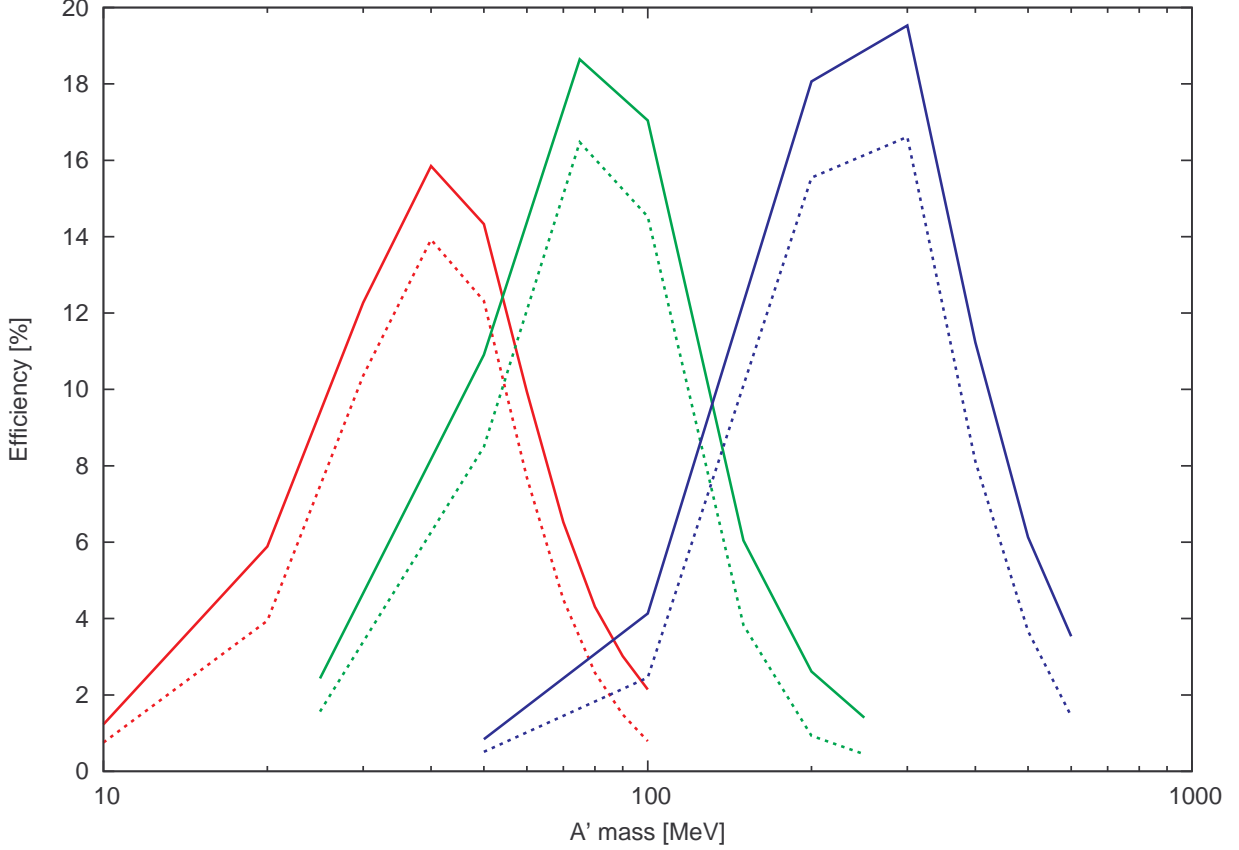


FIG. 62: Trigger efficiency (solid lines) and combined efficiency (dashed lines) as a function of  $A'$  mass, at beam energies of 1.1, 2.2 and 6.6 GeV (red, green and blue respectively).

### 6.3 Track Reconstruction

In order to study the tracking performance of the detector, we use samples of  $A'$  events at a variety of energies and decay lengths. On top of each event, we overlay backgrounds produced by the passage of beam electrons equivalent to our optimized run conditions at different beam energies and with a W target and a beamspot with a Gaussian sigma of  $40\mu\text{m}$  in the vertical direction and  $200\mu\text{m}$  in the horizontal. The beam energies, currents, target thickness and analyzing magnetic field used for these simulations are:

- 50nA at 1.1 GeV with  $X_0 = 0.125\%$  and  $B=0.25$  T



- 200nA at 2.2 GeV with  $X_0 = 0.125\%$  and  $B=0.5$  T
- 300nA at 4.4 GeV with  $X_0 = 0.25\%$  and  $B=1.0$  T
- 450nA at 6.6 GeV with  $X_0 = 0.25\%$  and  $B=1.5$  T

At each energy, we evaluate momentum, invariant mass, and vertex resolution. The plots shown in the following section typically use the 2.2 GeV beam as an example.

## 6.4 Tracking Efficiency, Pattern Recognition and Fake Rates

Due to the requirements imposed on the tracks, the efficiency for finding tracks in the geometric acceptance is not 1. The average track reconstruction efficiency is 98% (Fig. 63) and the bulk of the inefficiency comes from the cut on the total  $\chi^2$  of the track. Of the reconstructed tracks, a small percentage include a hit that is not from the correct electron. These “bad” hits may be from one of the high energy beam electrons scattered from the target into the detector or from a lower energy secondary. The left plot of Fig. 64 shows the number of bad hits/track for both the electron and positron from the  $A'$  decay. The number of tracks with 0 bad hits is  $> 98\%$ . The right plot of Fig. 64 shows the layer number of the bad hit. The rate of mishits are slightly higher in the downstream 3 modules due to the larger stereo angle. We'll show how these bad hits affect the track parameters in the next section.

## 6.5 Track Momentum and Spatial Resolution

The momentum resolution is shown in Fig. 65 as a function of momentum for tracks with 0 bad hits and for tracks with one or more. The momentum resolution for well-reconstructed tracks is  $\delta p/p = 4.5\%$  for  $B=0.5$ T (appropriate for a beam energy of 2.2 GeV) and is roughly inversely proportional to  $B$ .

One quantity we use to determine track quality is the distance of closest approach (DOCA) to the beam axis. We use this instead of the DOCA to the target beam spot since we are interested in long-lived decays and tracks from those will not point back to the target. We separate the distance into the bend plane (XOCA) and non-bend plane (YOCA) distances. Below, in Fig. 66, is the resolution of these quantities as a function of momentum.

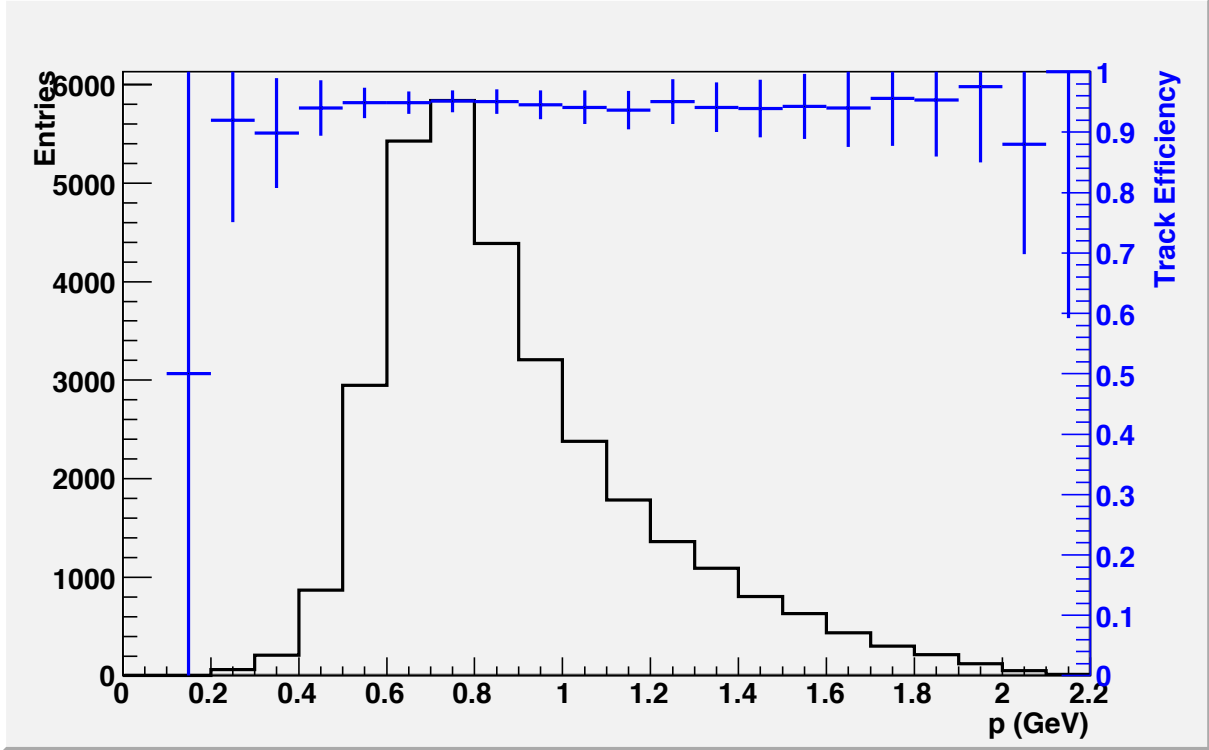


FIG. 63: Track reconstruction efficiency versus track momentum (right axis). The black histogram (left axis) show the track momentum distribution.

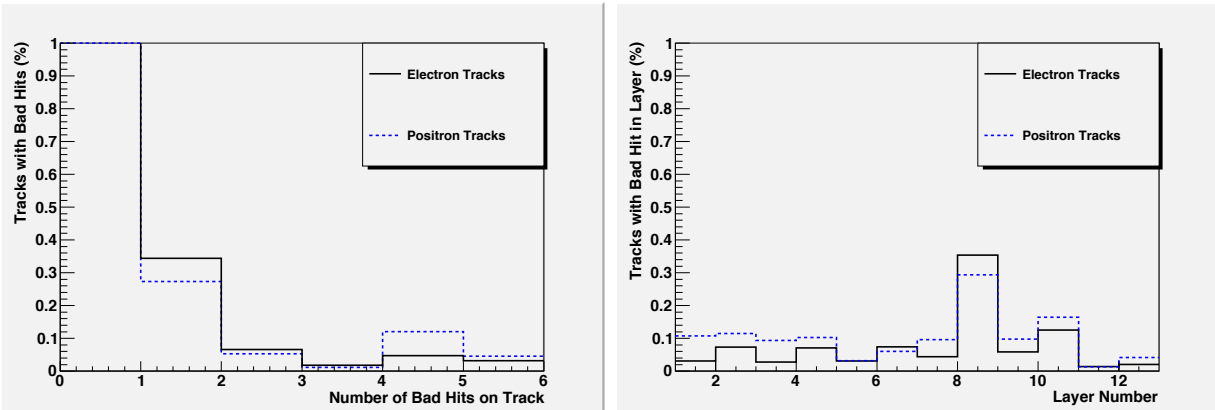


FIG. 64: The number of bad hits (left) and the layer number of the bad hit (right) for electron (black) and positron (blue) tracks.

The resolution is, on average, about  $200\mu\text{m}$  ( $400\mu\text{m}$ ) in the non-bend (bend) direction but increases significantly at low momentum. The position resolution for tracks with one or more bad hits is somewhat worse, depending on which layer the bad hit is. Tracks with bad hits in layers 1 or 2 are a major contribution to the tail of the vertex position distribution.

For long lived  $A'$  decays, the position of the decay vertex is an important discriminating

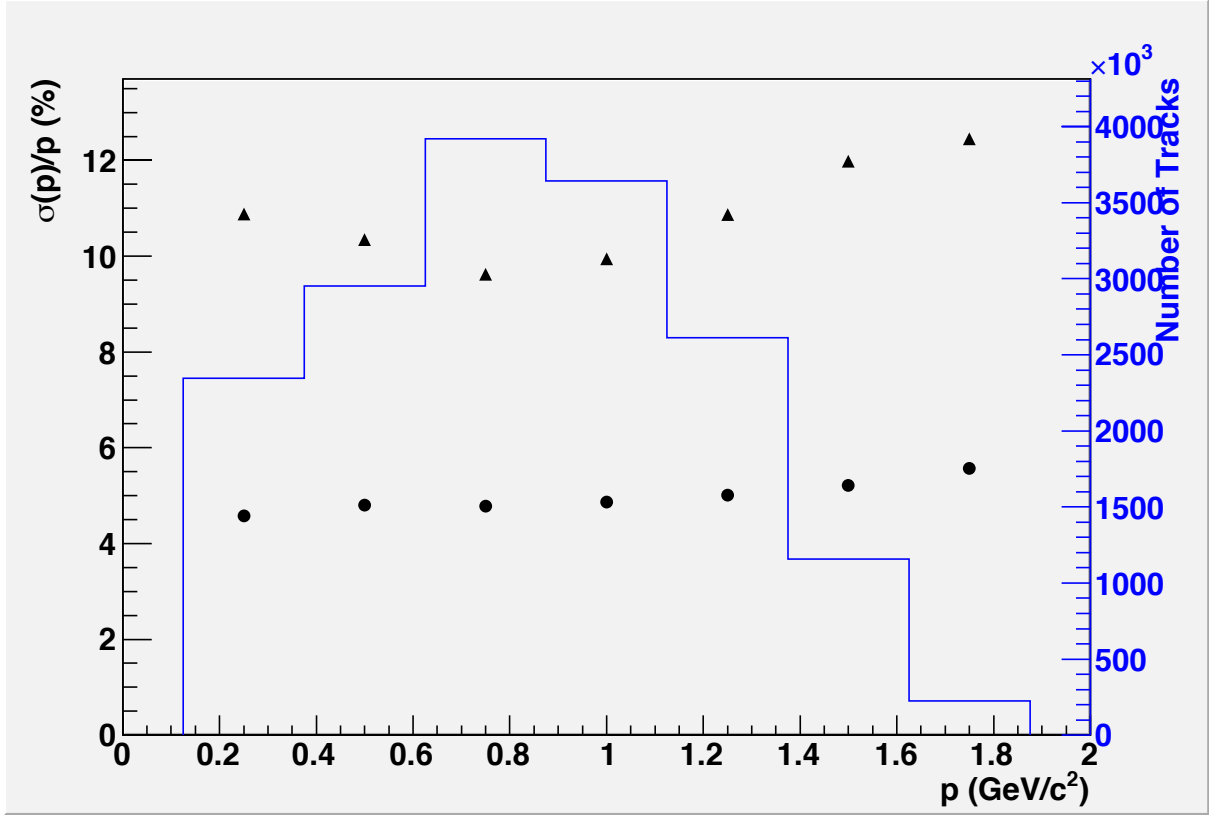


FIG. 65: Fractional momentum resolution versus momentum for a beam energy of 2.2 GeV and a magnetic field of 0.5 T. The dots represent tracks with 0 bad hits and the triangles with one or more.

variable. The dominant background to  $A'$  production is radiative events which originate in the target. Distinguishing  $A'$  decays from the background therefore depends on the vertex resolution and in particular on the tails of the vertex distribution. In order to study the tails, we use large samples of  $A'$  events decaying promptly overlaid on top of the simulated beam background events.

Each pair of oppositely charged tracks is fit to a common vertex using a Kalman filtering method first suggested by Billoir [106], [107] and used in many experiments. The method uses the measured helix parameters and their correlations to determine the most likely decay position of the  $A'$  and also returns fitted momenta for each particle. We actually fit each pair twice with different hypotheses of their origin. We constrain either the vertex to be consistent with an  $A'$ :

- which originates in the  $200\mu\text{m} \times 40\mu\text{m}$  beamspot at the target, and moves off in the direction given by the measured  $A'$  momentum. This fit will be used for the vertexing

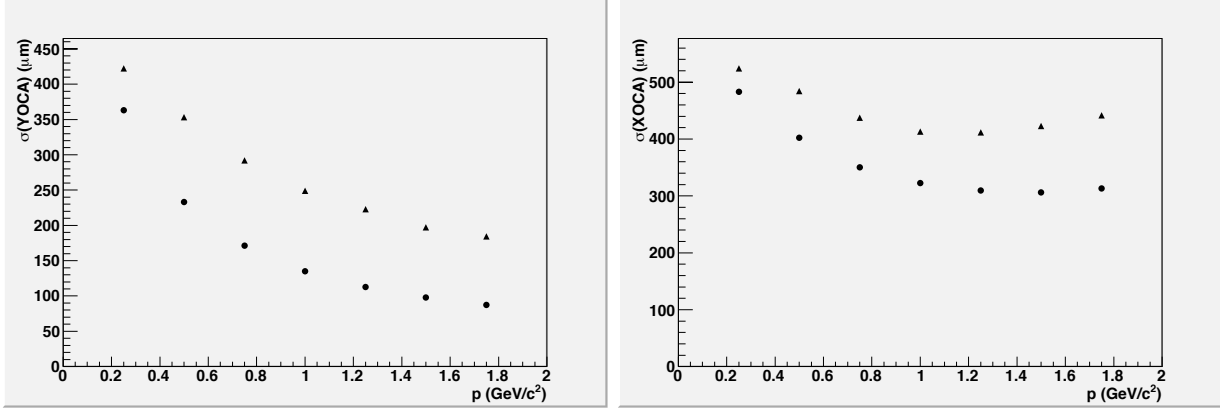


FIG. 66: The resolution of the position of closest approach to the beam axis versus track momentum in the (left) non-bend direction and (right) bend direction. The dots represent tracks with 0 bad hits and triangles with one or more.

search.

- which originates and decays at the target within the  $200\mu\text{m} \times 40\mu\text{m}$  beamspot. This fit will be used for the bump-hunt only search.

For each electron/positron pair reconstructed in the tracker, we compute the invariant mass based on the fitted momenta of the tracks. The mass resolution depends on the invariant mass of the pair and is shown in Fig. 67. The closed circles in Fig. 67 shows the improvement in the resolution for the second fit, where the decay is assumed to occur in the target.

Even for prompt decays, the  $z$  vertex position ( $V_z$ ) distribution of all reconstructed  $e^+e^-$  pairs (solid black histogram, Fig. 69) shows a long tail, still significant beyond 5cm. This tail is primarily comprised of events where one or both of the tracks use one or more bad hits. Fortunately there are a number of quantities we can use to minimize the tails. Namely, for purposes of this proposal, we make the following cuts:

- The  $\chi^2$  of each track is less than 20
- The total momentum of the  $A'$  candidate is less than the beam energy
- A very loose cut on the reconstructed vertex position  $|V_x| < 400\mu\text{m}$  and  $|V_y| < 400\mu\text{m}$
- The clusters in layer 1 of each track must be isolated from the next closest cluster by at least  $500\mu\text{m}$

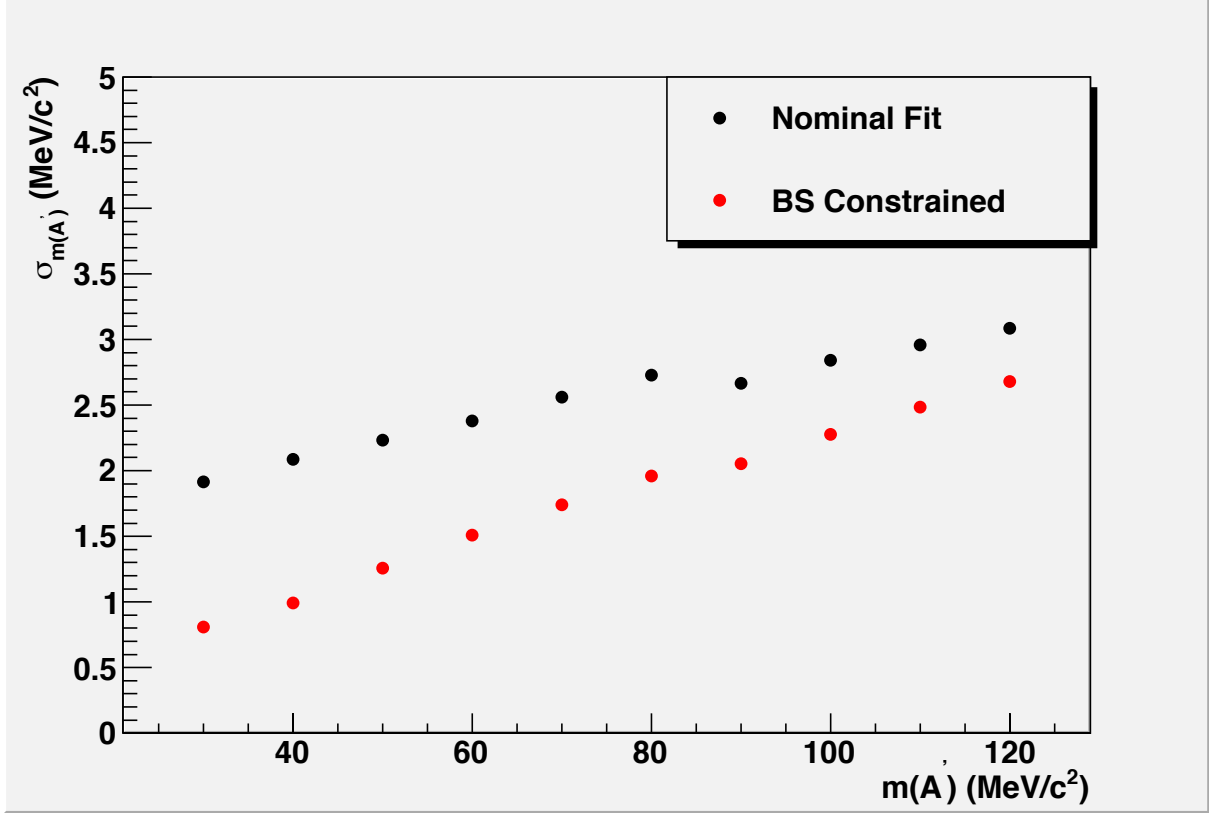


FIG. 67: The gaussian width of the mass distributions (MeV/\$c^2\$) vs generated \$A'\$ mass (MeV/\$c^2\$). The open circles are the resolutions when the decay is constrained to the target beamspot and the closed circles are without this constraint.

- A  $\chi^2$  cut on the vertex fit of less than 5.

The vertex resolution depends on the invariant mass of the particles being vertexed. Lower masses have worse Gaussian resolutions as shown in Fig. 68. This is expected since the error on the opening angle ( $\theta$ ), due to multiple scattering, scales like:  $\sigma(\theta)/\theta \sim (1/E)/(m/E) \sim 1/m$ .

Figure 69 shows the vertex resolution for samples of 80 MeV and 160 MeV  $A'$  events. The cuts above remove almost all of the tail past 1.5cm (points with errors in Fig. 69) while retaining 50% of the  $e^+e^-$  pairs from the  $A'$  candidate. The events on the tail are enhanced with vertices where there are one or more bad hits on the track (represented by the blue histogram in Fig. 69), although there is still a contribution from well-reconstructed tracks. The rejection of tracks with bad hits depends strongly on the precision of the virtual  $A'$  trajectory, which in turn depends on the size of the beamspot. Having a beamspot significantly smaller than the intrinsic tracker resolution,  $100\mu\text{m}$  in the non-bend and  $300\mu\text{m}$

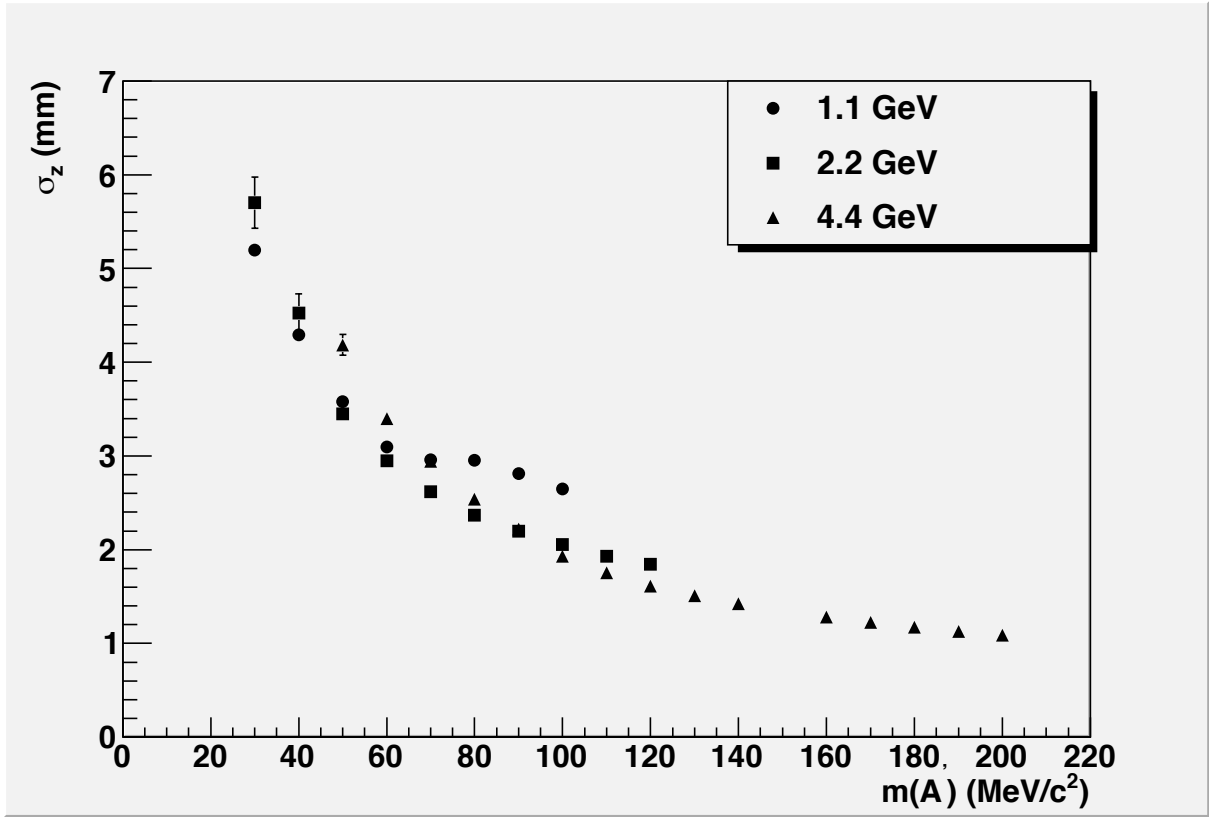


FIG. 68: The Gaussian resolution dependence versus  $A'$  mass for signal-only events.

in the bend directions, is important.

In practice, there is much more we can do to clean up the vertex and mass resolution both at the track level (e.g. remove hits that are clearly from scattered beam electrons) and at the vertex level. These will be pursued in the near future.

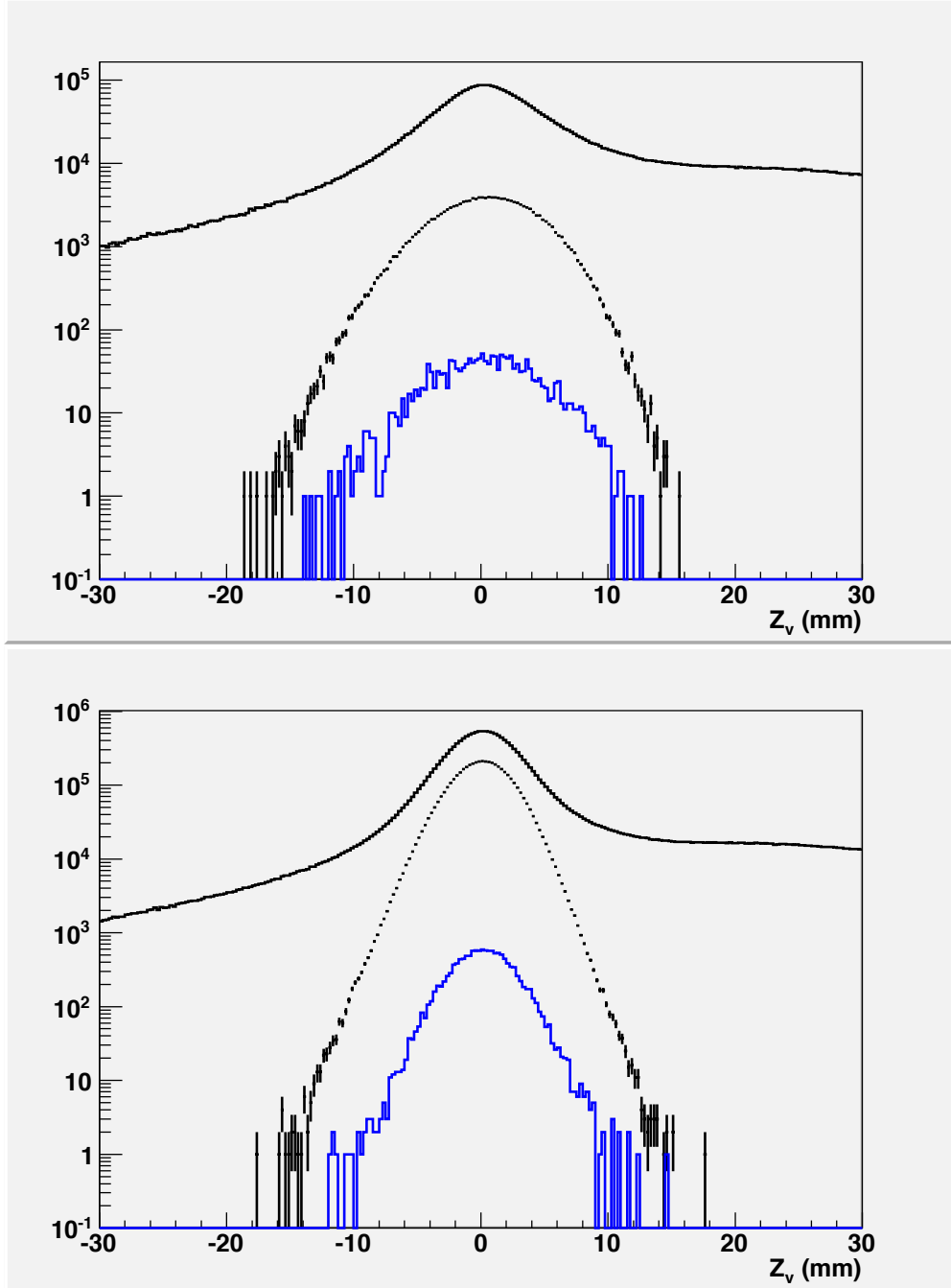


FIG. 69: Distribution of the reconstructed vertex position along the beam axis for 2.2GeV 40MeV (top) and 80MeV (bottom)  $A'$  events before (solid black) and after (points with errors) selection. The blue histogram shows the distribution for pairs that have at least one bad hit after selection.

## 7 Experimental Reach

The primary search channel for this experiment is  $A' \rightarrow e^+e^-$ , with or without a displaced vertex, depending on the magnitude of the coupling  $\alpha'$ . As such, the primary irreducible background is QED trident production, with rate given by the diagrams shown in Figure 2. Trident events can be usefully separated into “radiative” diagrams (Figure 3(a)), and “Bethe-Heitler” diagrams (Figure 3(b)), that are separately gauge-invariant. The expected parameter reach of the experiment is shown in Figure 5. Below, we discuss how this was calculated.

The contribution from the radiative diagrams (Figure 3(a)) alone is a useful guide to the behavior of  $A'$  signals at various masses. In particular, the kinematics of  $A'$  signal events is identical to that of radiative trident events restricted to an invariant mass window near the  $A'$  mass. Moreover, the rate of the  $A'$  signal is simply related to the radiative trident cross-section within a small mass window of width  $\delta m_{A'}$  by [1],

$$\frac{d\sigma(e^-Z \rightarrow e^-Z(A' \rightarrow e^+e^-))}{d\sigma(e^-Z \rightarrow e^-Z(\gamma^* \rightarrow e^+e^-))} = \left( \frac{3\pi\epsilon^2}{2N_{eff}\alpha} \right) \left( \frac{m_{A'}}{\delta m_{A'}} \right) \quad (5)$$

where  $N_{eff}$  counts the number of available decay states. A fraction  $\epsilon_{bin}$  of signal events will have reconstructed masses within the mass window, because of the finite mass resolution (for a  $2.5 \times \sigma$  mass resolution window,  $\epsilon_{bin} = 0.8$ ). Equation 5 corrected for  $\epsilon_{bin}$  allows us to conveniently express the sensitivity to  $A'$  signals in terms of the radiative portion of the total QED trident statistics, which we will do shortly.

The Bethe-Heitler process has a much larger total cross-section than either the signal or the radiative trident backgrounds, but exploiting its different kinematics can significantly reduce it. In particular, the  $A'$  carries most of the beam energy (see the discussion in Section 3.1) while the recoiling electron is very soft and scatters to a wide angle. In contrast, the Bethe-Heitler process is not enhanced at high pair energy. Moreover, Bethe-Heitler processes have a forward singularity that strongly favors asymmetric configurations with one energetic, forward electron or positron and the other constituent of the pair much softer. The geometric acceptance and trigger requirements select the region of phase space where signal is dominant, and the Bethe-Heitler background is smallest, as illustrated by Figure 4 (it should be emphasized, however, that even in this region the Bethe-Heitler background



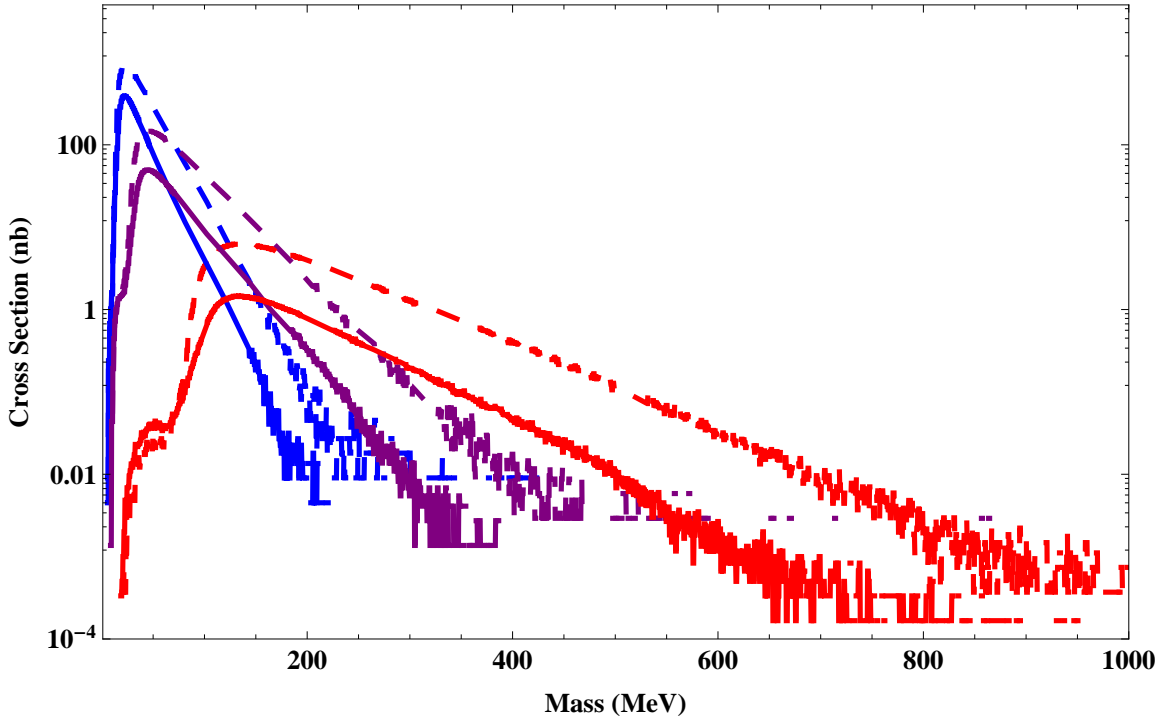


FIG. 70: The Bethe-Heitler (dashed) and radiative (solid) cross-section/MeV after acceptance for 1.1 (blue), 2.2 (purple) and 6.6 GeV (red).

rate exceeds that of radiative tridents by roughly a factor of 5). The radiative and Bethe-Heitler cross-sections, after accounting for acceptance in the HPS detector, are shown in Figure 70.

To compute the reach of the HPS experiment, we simulate the production of irreducible trident reactions in the detector. We additionally apply a mock-up of the geometric acceptance for the tracking and of the trigger requirements. In addition, high-statistics Monte Carlo samples at particular invariant masses have been used to estimate the background rejection efficiency for a vertex-based search. We produce generator-level events using MadGraph and MadEvent [108] to compute the full matrix elements for  $e^-Z \rightarrow e^-(e^+e^-)Z$  in leading order QED, but neglecting the effect of nuclear excitations on the kinematics in inelastic processes. We use the QED nuclear elastic and inelastic electric form-factors in [109]. The MadEvent code was modified to properly account for the masses of the incoming nucleus and electron in event kinematics.

We use a “reduced-interference” approximation that simplifies our analysis and is much less computationally intensive. In this approximation, we treat the recoiling  $e^-$  and the  $e^-$  from the produced pair as distinguishable. Furthermore, we separate trident processes into

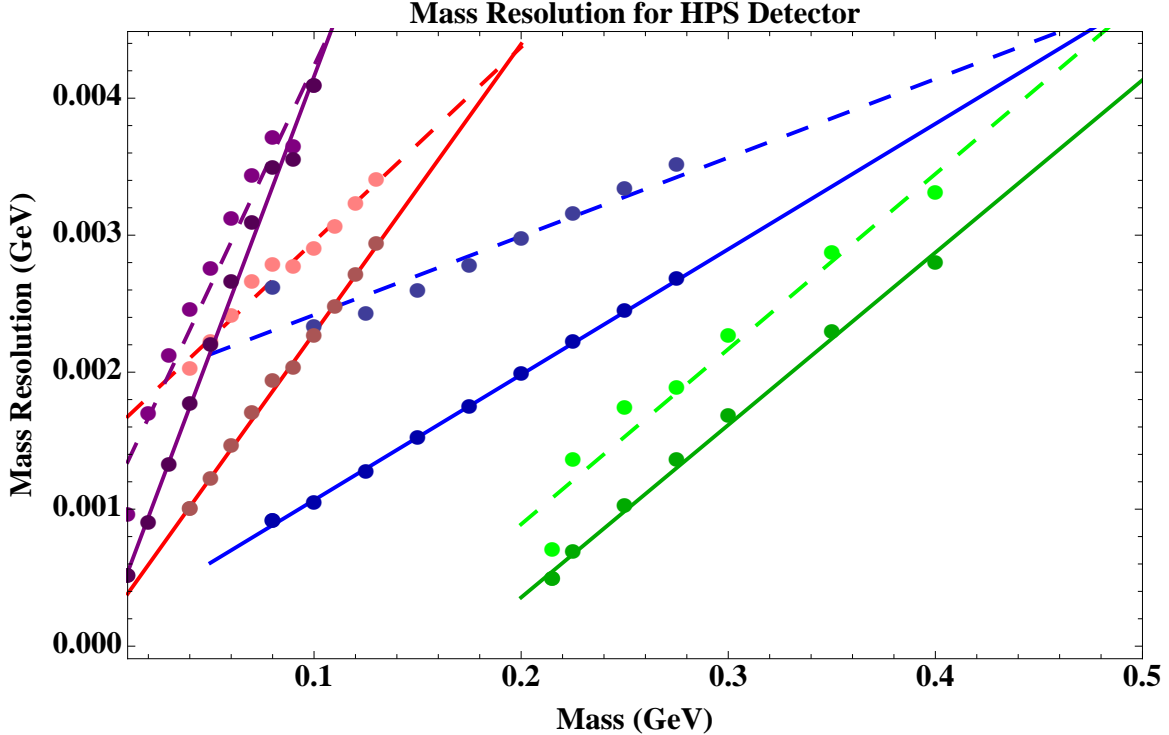


FIG. 71: The invariant mass resolution versus mass for 1.1 (purple), 2.2 (red), 6.6 GeV electron decays (blue) and 6.6 GeV muon decays (green). The points are from simulated data at various masses while the lines are linear fits to the points. For each energy, the points with worse resolution (dashed line) are from the vertex fit without constraining the decay to the target while the better resolution (solid line) require the decay to be prompt. The fitted curves are used in the reach calculation; the resolution from the dashed lines are used for the displaced vertex search while the bump-hunt search uses the resolution from the solid lines.

the radiative diagrams and the Bethe-Heitler diagrams, and we calculate the cross-section for both of these diagrams separately. Within the acceptance and signal region for the HPS experiment, the Bethe-Heitler reactions dominate the trident rate by 4:1. We have checked that the “reduced-interference” approximation does not correct the trident cross-section by more than 10% in a representative kinematic region [4].

## 7.1 Resonance Search

Equation 5 is used to compute the reach for a resonance search in the  $e^+e^-$  or  $\mu^+\mu^-$  final state. We start by simulating radiative and Bethe-Heitler trident events and require that  $e^+e^-$  or  $\mu^+\mu^-$  pairs pass the detector acceptance cuts. We additionally require that the total energy exceed 80% of the beam energy and that each track have at least 0.5 GeV

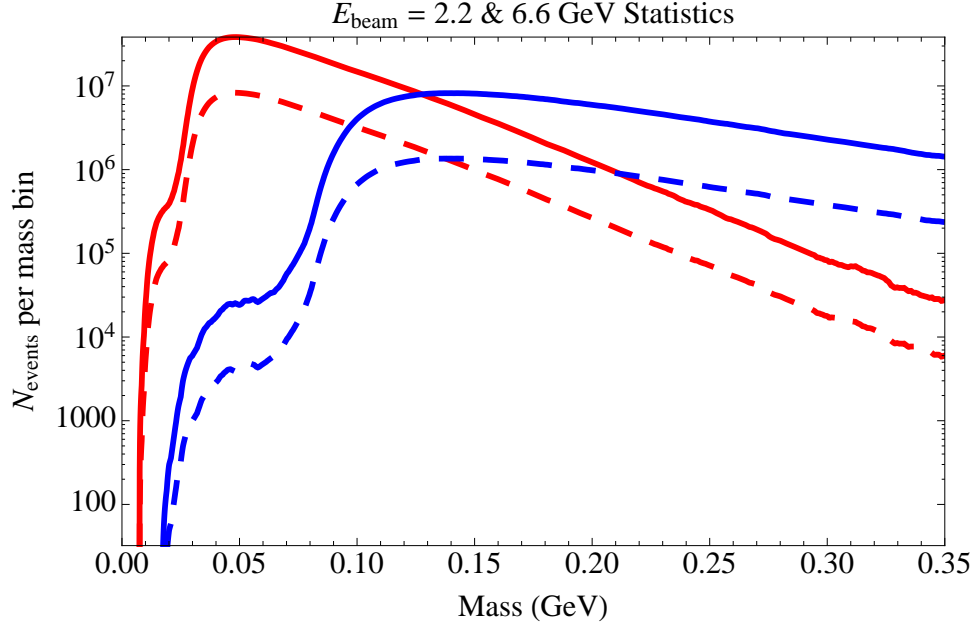


FIG. 72: Right: Distribution of statistics in full resonance search. The solid (dashed) curves indicate the number of background QED trident (radiative) events expected in a resolution-limited mass window of width  $\delta m_{(A')} = 2.5\sigma(m_{A'})$ . The blue curves correspond to the distributions of statistics for a one month run at 6.6 GeV beam energy, with a current of 450 nA on a 0.25%  $X_0$  target, while the red curves correspond to six weeks at 2.2 GeV beam energy and a 0.125%  $X_0$  target.

of energy. We will refer to these cuts collectively as the “detector/trigger mock-up”. We compute the total differential cross section, as a function of invariant mass, for radiative and Bethe-Heitler trident events to pass the detector/trigger mock-up cuts, and from this the final statistics are computed assuming a run duration according to the run plan laid out in Section 3, and conditions in Section 6.3. The mass resolution and the background statistics expected in each resolution-limited mass window are shown in Figures 71 and 72.

To quantify statistical sensitivity, we assume that the trident background in the resonance search can be modeled by a smoothly varying function and subtracted off. The significance is then determined by the ratio of the signal within an invariant mass window to  $\sqrt{N_{bin}}$ , where  $N_{bin}$  is the total background statistics in the same window. Using equation 5, the sensitivity for a resonance search is determined by

$$\left(\frac{S}{\sqrt{B}}\right)_{bin} = \left(\frac{N_{radiative}}{N_{total}}\right) \sqrt{N_{bin}} \left(\frac{3\pi\epsilon^2}{2N_{eff}\alpha}\right) \left(\frac{m_{A'}}{\delta m_{A'}}\right) \epsilon_{bin} \quad (6)$$

Here,  $\left(\frac{N_{radiative}}{N_{total}}\right)$  is the fraction of radiative reactions among all QED trident events in the

search region. This quantity is determined by simulation as described below.  $N_{bin}$  is the total number of QED trident events residing in a given invariant mass search bin, and is determined by

$$N_{bin} \equiv \epsilon_{reco}^2 \times \epsilon_{stat}(m_{A'}) \times \sigma_{trigger} \times L. \quad (7)$$

Here  $L$  is the integrated luminosity,  $\sigma_{trigger}$  is the trigger cross section,  $\epsilon_{stat}(m_{A'})$  is the fraction of the total statistics in an invariant mass window centered on  $m_{A'}$  of size  $\delta m_{A'} = 2.5\sigma(m_{A'})$ , and  $\epsilon_{reco} \cong 0.85$  is the efficiency for reconstructing each track that is within the geometric acceptance of the detector.

## 7.2 Displaced Vertex and Resonance Search

A search for resonances that decay with cm-scale displaced vertices opens up sensitivity to much smaller couplings than can be observed through a resonance search alone. The vertex reconstruction and quality selection is discussed in Section 6.3. For the purpose of computing reach, the vertex quality requirements reduce the signal efficiency by a factor  $\sigma_{sig} \sim 0.5$ . We use the high-statistics Monte Carlo studies described in Section 6.3 to model the tails of the vertex distribution for decays at the target. These vertex distributions have been generated at a few different masses for each beam energy. Away from these masses, we parameterize the background rejection factor  $\epsilon_{rejection}(zcut)$ , the fraction of events with a fake vertex beyond a beam line distance of  $zcut$ , by a smooth interpolation.

Because the fake vertex distribution falls quite rapidly, the greatest sensitivity is achieved far on the vertex tail, where less than one background event is expected. For the purpose of computing reach, we have determined a mass-dependent choice of  $zcut(m_{A'})$  such that the expected background in each resolution-limited mass window  $\delta m_{A'}$ , with reconstructed vertices beyond this cut, does not exceed 0.5 events in the run period. This requires rejection  $\epsilon_{rejection}(zcut)$  of background events from the target at the level of  $10^{-6}$  to  $10^{-7}$ , achieved for  $zcut \sim 5\text{-}30$  mm (see Figure 73).

The geometric acceptance falls off at decay lengths greater than 10 cm. For simplicity we compute reach using the geometric acceptance for  $z=0$ , but only considering decays with

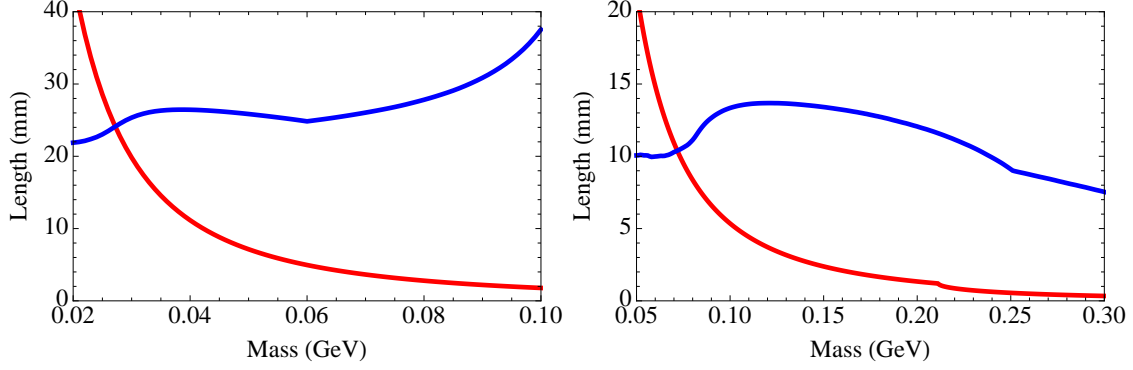


FIG. 73: In blue: the value of the minimum vertex displacement  $z_{cut}$  (in mm) along the beamline, required for the vertex-based resonance search at 2.2 GeV (left) and 6.6 GeV (right). These are chosen to bring the expected background to 0.5 events in each resolution-limited mass window. In red: the  $\gamma c\tau$  assuming  $\epsilon = 10^{-4}$ .

$z < z_{max} = 10$  cm, so that the fraction of signal events included in the vertex search is

$$\epsilon_{sig}(z_{cut}) \cong \epsilon_{vtx} \times \left( e^{-\frac{z_{cut}}{\gamma c\tau}} - e^{-\frac{z_{max}}{\gamma c\tau}} \right). \quad (8)$$

Accounting for the reduced acceptance of both signal and background events, the statistical significance expected can be computed from that of the pure resonance search as:

$$\left( \frac{S}{\sqrt{B}} \right)_{bin, z_{cut}} = \left( \frac{S}{\sqrt{B}} \right)_{bin} \frac{\epsilon_{sig}(z_{cut})}{\sqrt{\epsilon_{rejection}(z_{cut})}} \quad (9)$$

where  $\left( \frac{S}{\sqrt{B}} \right)_{bin}$  is given by (6). For the small expected background rate (0.5 events/bin), however, this formula becomes irrelevant, as the exclusion sensitivity of the experiment is limited by the probability of a downward fluctuation in the signal. Thus, for the vertex reach contours in Figure 5, we additionally require an expected signal

$$S_{bin, z_{cut}} = \left( \frac{N_{radiative}}{N_{total}} \right) N_{bin} \left( \frac{3\pi\epsilon^2}{2N_{eff}\alpha} \right) \left( \frac{m_{A'}}{\delta m_{A'}} \right) \epsilon_{sig}(z_{cut}) > 2.4 \text{ events} \quad (10)$$

### 7.3 Reach in Mass-Coupling Parameter Space

Using the  $S/\sqrt{(B)}$  for the bump-hunt and displaced vertex searches as described above, we estimate to cover the regions of coupling vs mass parameter space shown in Figure 74. The contours in the plot show the the two-sigma exclusion regions for:

- purple, dashed: 1 week of 50nA, 1.1 GeV beam on a 0.125% target
- blue, dashed: 1 week of 200nA, 2.2 GeV beam on a 0.125% target
- blue, solid: 3 weeks of 200nA, 2.2 GeV beam on a 0.125% target
- dark green: 2 weeks of 450nA, 6.6 GeV beam on a 0.25% target, detecting  $A' \rightarrow e^+e^-$
- light green: 2 weeks of 450nA, 6.6 GeV beam on a 0.25% target, detecting  $A' \rightarrow \mu^+\mu^-$
- red: the statistical combination of all of the above
- green shaded: 3 months each of 2.2 GeV and 6.6 GeV (same currents and thicknesses as above)

## 8 Run Plan and Beam Time Request

The Jefferson Laboratory PAC39 graded HPS physics with an “A,” approved a commissioning run with electrons, and granted so-called “C1” approval for the full HPS experiment. The total beam time requested in our original proposal to PAC37 is 180 days. Anticipating early running in Hall-B, we propose to conduct HPS in two phases. The first phase, expected to run in 2014-2015, will complete the commissioning run and begin the production running. For the upcoming commissioning run in 2014 we have requested 3 weeks; for the data taking runs in 2015, we have requested 5 weeks. The second phase of HPS, which will use the remaining beam time, can be scheduled later in 2015 or in 2016 and beyond, and will continue runs at 2.2 and 6.6 GeV and possibly other energies.

We plan to execute the first phase of the experiment in two run periods using the apparatus described above. First, we will perform a commissioning run which should produce our first physics output, and then, after a month or two of down time, continue with a longer run at multiple beam energies to cover as much parameter space as possible. The experimental apparatus, if funded on time, will be ready to be commissioned and take physics data in the fall of 2014 when the first physics quality beams should be available in Hall-B. The proposed run plan for phase one is as follows:

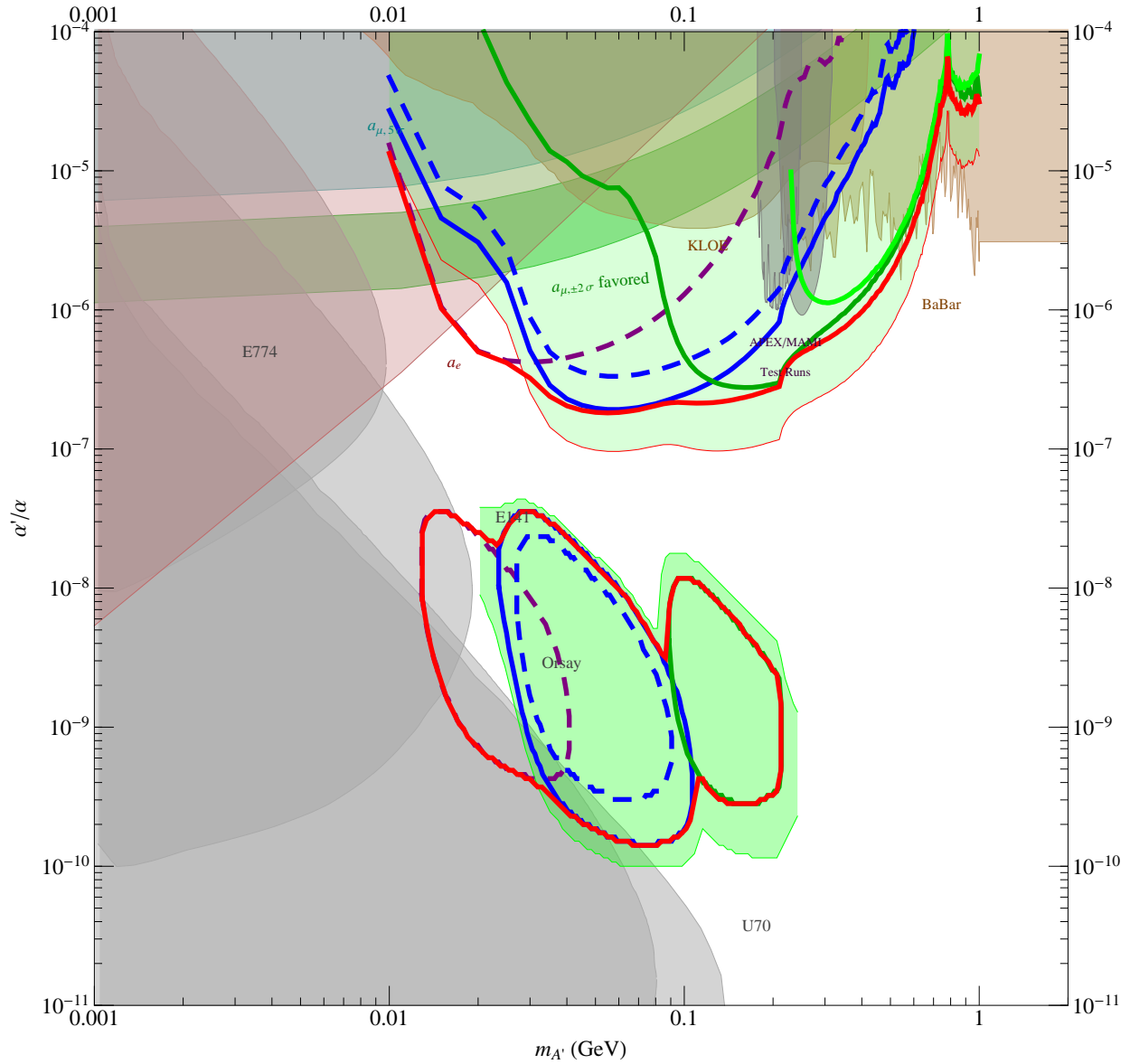


FIG. 74: The expected reach in mass-coupling parameter space. See text for details about different features.

- **Commissioning run in 2014, total of 3 weeks of beam time (6 weeks on the floor assuming 50% for combined efficiency of the accelerator and the detector):**
  - 1 week of detector commissioning
  - 1 week of physics run at 2.2 GeV
  - 1 week of physics run at 1.1 GeV
- **Physics run in 2015, total of 5 weeks of time beam (10 weeks on the floor**

**assuming 50% for combined efficiency of the accelerator and the detector):**

- 1 week of detector commissioning
- 2 weeks of physics run at 2.2 GeV
- 2 weeks of physics run at 6.6 GeV

If more beam time is available in 2015, HPS will continue data taking at these or other energies.

The proposed run plan will cover the remaining region of parameter space favored by the muon  $g-2$  anomaly, and will explore a significant region of parameter space, not only at moderate couplings ( $\alpha'/\alpha > 10^{-7}$ ), but also in the regions of small couplings, down to  $\alpha'/\alpha \sim 10^{-10}$ . This small coupling region is not accessible to any other proposed experiment. The excellent vertexing capability of the Si-tracker uniquely enables HPS to cover the small coupling region.

In the proposed run plan we have assumed running at a non-standard energy for 12 GeV CEBAF,  $E = 1.1$  GeV, in 2014. In case this will not be possible due to scheduling conflicts in 2014, we will continue to run at 2.2 GeV instead. We would then expect to complete the 1.1 GeV running at some time in 2015 by reducing the time at 2.2 GeV. The gap between the run periods in 2014 and 2015, on the order of two months, will be used to improve, correct or fix apparatus as necessary.

In summary, we request time for the first phase of HPS experiment, first in the fall of 2014 for a total of 3 weeks of beam time (6 weeks on the floor), with beam energies 1.1 GeV (two weeks) and 2.2 GeV (two weeks) in addition to commissioning time. Second, we request 5 weeks of beam (10 weeks on the floor) that will be equally shared between beam energies of 2.2 GeV and 6.6 GeV with some time for re-commissioning after the down. We expect that the second part of the run will be in 2015.

## 9 Schedule and Cost Baseline

Cost estimates for engineering, designing, fabricating, assembling, testing, and installing the Heavy Photon Search detector are given below. The costs assume considerable savings from the reuse of many parts of the Test Run, which was already assembled from silicon



microstrip sensors donated by Fermilab, made use of DAQ crates and equipment from SLAC, and utilized many contributions from JLab, including  $\text{PbWO}_4$  calorimeter crystals, the chicane and analyzing magnets, magnet power supplies, and beam diagnostic apparatus. Much of the calorimeter readout electronics utilizes designs which are already in place for the Hall B 12 GeV upgrade, eliminating engineering and design expense. Very significant cost savings come from utilizing the FADCs and data acquisition system being developed for the upgraded CLAS12 detector, which will be available free of charge to HPS. The SVT DAQ benefits from SLAC's development of an ATCA readout system, and incorporates many of its existing designs. The Orsay group has contributed engineering and design efforts for the ECal support structure, enclosure, and vacuum chamber, affording additional savings. It will contribute manpower for Ecal reassembly and test.

HPS costs are given in an accompanying WBS summary table, Table XIX, which itemizes the major items subsystem by subsystem, and indicates whether JLab (J) or SLAC (S) takes responsibility for construction. Table XX shows the cost breakdown between SLAC and JLab. The full WBS is given in Appendix A. Engineering, design, and technician labor rates are fully loaded, including benefits and lab overheads, which differ between the two laboratories. Our DAQ and beamline cost estimates have been made by engineering groups at SLAC and JLab which are experienced in cost estimation and actively involved in many related projects including the recent HPS Test run experiment. The SVT estimates came from HPS physicists and engineers with considerable experience in designing and fabricating silicon detector systems including those produced for the HPS Test Run. The Ecal estimates come from physicists and engineers at JLab and Orsay who have constructed a similar system, the CLAS IC, in the recent past, and assembled the Ecal for the Test Run. Scheduling and budgeting the Test Run provided the entire HPS team valuable experience and a good reality check.

The HPS Schedule is discussed below, including a brief description of the schedule for the each different subsystems. The overall schedule contingency is about 20%, and depends critically on the assumption that funding is available by beginning FY2014. Keep-alive funding is currently being used to advance the engineering design and R&D to the point that DOE support in FY2014 will be adequate to maintain project readiness for the fall of 2014. The HPS construction project has been organized into a Work Breakdown Structure (WBS) for purposes of planning, managing and reporting project activities. Work elements

TABLE XIX: Summary of HPS Budget.

WBS	Task Name	Labor	Material	Total	Labor w/Cont.	Material w/Cont.	Total w/Cont.	Spares	Proto.	Ops	Infra.	Capital Eq.
1	HPS	\$1367K	\$1039K	\$2408K	\$1761K	\$1211K	\$2972K	\$78K	\$21K	\$928K	\$248K	\$1796K
1.1	Beamline (J)	\$70K	\$101K	\$171K	\$91K	\$131K	\$223K	\$0	\$0	\$0	\$6K	\$217K
1.2	SVT (S)	\$346K	\$161K	\$507K	\$452K	\$204K	\$656K	\$8K	\$10K	\$75K	\$43K	\$539K
1.3	SVT DAQ (S)	\$342K	\$264K	\$608K	\$431K	\$352K	\$782K	\$70K	\$11K	\$80K	\$162K	\$540K
1.4	ECAL (J)	\$28K	\$9K	\$37K	\$36K	\$12K	\$48K	\$0	\$0	\$0	\$0	\$48K
1.5	TDAQ (J)	\$116K	\$10K	\$126K	\$151K	\$10K	\$161K	\$0	\$0	\$0	\$10K	\$151K
1.6	Slow Control (J)	\$76K	\$31K	\$107K	\$94K	\$39K	\$134K	\$0	\$0	\$0	\$28K	\$106K
1.7	Installation & Commissioning (S)	\$45K	\$0	\$45K	\$58K	\$0	\$58K	\$0	\$0	\$55K	\$0	\$3K
1.9	SLAC Travel Meetings (S)	\$104K	\$0	\$104K	\$125K	\$0	\$125K	\$0	\$0	\$125K	\$0	\$0
1.10	SLAC Travel for Commissioning and Running (S)	\$112K	\$0	\$112K	\$156K	\$0	\$156K	\$0	\$0	\$130K	\$0	\$26K
1.12	Project management (S)	\$128K	\$0	\$128K	\$167K	\$0	\$167K	\$0	\$0	\$0	\$0	\$167K
1.13	UCSC Funds	\$0	\$463K	\$463K	\$0	\$463K	\$463K	\$0	\$0	\$463K	\$0	\$0

TABLE XX: Cost breakdown between SLAC and JLab.

	Operations	Infrastructure	Capital Equipments	Total
SLAC	\$899K	\$242K	\$1332K	\$2473K
JLAB	\$29K	\$6K	\$464K	\$499K

are defined to be consistent with discrete increments of project work. Project Management efforts are distributed throughout the project, including conceptual design and R&D. The HPS has 12 WBS Level-2 elements, see Table XXI.

TABLE XXI: Project WBS structure.

WBS	NAME
1.1	Beamline
1.2	SVT
1.3	SVT DAQ
1.4	ECAL
1.5	TDAQ
1.6	Slow Controls
1.7	Installation & Commissioning
1.8	Electron Running
1.9	SLAC Travel Meetings
1.10	SLAC Travel for Commissioning and Running
1.11	Project Management
1.12	UCSC

## 9.1 Cost

The costs include Labor and M&S. The labor includes only engineering or technician manpower in professional centers at SLAC or JLab. It does not include labor provided by physicists, which is the dominant contribution to the project. Labor rates have been applied following the official shop rates at SLAC and JLab, which include already  $\sim 31\%$  or  $\sim 57\%$  fringe benefits, respectively. M&S have been determined from a best estimation of the commercially available parts, benefiting from our experience with the actual costs of the HPS Test Run. The overheads have been added to both labor and M&S, being respectively  $53\%$  and  $7.65\%$  at SLAC,  $49\%$  for both labor and M&S at JLab. SLAC travel includes  $53\%$  overheads. Contingencies have been set at  $10\%$  for catalogue items,  $20\text{-}25\%$  for items similar

to previous design, and 30-50% if the design is new. Since the project is staged over three years, an annual inflation rate of 2.5% is included in the FY15 and FY16 costs.

Most aspects of HPS are costed in the WBS tables. However, as the host laboratory, JLab provides support for certain specific aspects of the experiment, including labor by certain JLab physicists, engineering and design oversight and coordination within Hall B, existing beamline apparatus in Hall B, and commissioning and operations of the beam delivery system and system maintenance. JLab is supported by this proposal for specific engineering and design tasks, fabrication and testing of some HPS specific hardware, and software and programming support for the TDAQ and Slow Controls systems. SLAC likewise provides support for SLAC physicists working on HPS, and is supported by this proposal for engineering, design, and fabrication of parts of the HPS apparatus.

The costs have been divided into three broad categories, Capital Equipment (CE), Infrastructure (INFRA), and Operations (OP). The categories are explained in the following table. DOE limits Capital Equipment support for small projects to \$2 M, so the HPS budget lists these expenses explicitly.

TABLE XXII: Cost categories for accounting.

Capital Equipment (CE)	Labor and Material costs of the parts constituting the experimental apparatus, which is required to perform the physics. They do not include spares and prototyping.
Infrastructures (INFRA)	Labor and Materials cost for general purpose infrastructures required by the apparatus, which can be reused for other experiments, e.g. Chiller, Power Supplies, PLC.
Operation (OP)	Labor and Material Costs incurred for the commissioning and the operation of the experimental set up, as well as Spare Parts, Prototype and R&D activities.

Beamline expenses for HPS are held to a minimum by using the 18D36 magnet currently installed in Hall B as the analyzing magnet, the two existing JLab Frascati chicane magnets and the existing Test Run vacuum chamber with the SVT vacuum box. Some overall engineering and design will be required, beam pipes fabricated, a vacuum chamber built for the downstream Frascati magnet, and a photon dump and shielding inserted behind the second chicane magnet. Total beamline expenses are about \$223k, including \$6k for infrastructure.

Three out of the five planes of the SVT Test Run will be reused after modifying their supports to provide improved mechanical stability and better cooling. Three new planes with double sensors and their supports will be designed and built from scratch. Fermilab will donate the needed silicon microstrip detectors, as it had for the HPS Test Run. The tracker/vertexer will cost about \$656k, including \$43k for infrastructure and \$75k for operations.

The SVT DAQ requires small modifications to the existing hybrid; new readout and flange board engineering design, prototyping, and production; APV25 and chip procurement; and fabrication and testing. The SVT DAQ also requires designing and prototyping the Trigger Interrupt ACTA card and new firmware for the APV25 to enable event buffering needed to accommodate higher trigger rates. ATCA crates, and standard RCE cards are also required. The expenses are dominated by engineering development, and total \$782k, including \$162k for infrastructure and \$80k for operations.

JLab had donated the  $\text{PbWO}_4$  crystals, APDs, and amplifiers for the ECal for the Test Run. All these components will be reused for full HPS. Orsay is working with JLab to replace the existing motherboards and reassemble and test the ECal. JLab will build new mounting stands for the ECal and the Ecal vacuum chamber. If supplemental funds are available in France and/or Italy, new, more sensitive APDs will replace the current ones, and an improved preamplifier will be built and installed. The total cost to DOE is \$48k, and excludes these potential improvements.

Trigger and DAQ electronics for the ECAL will use that being developed for the CLAS upgrade, so relatively little engineering and technician time will be needed for HPS except for providing special purpose firmware. Many components, including the 250 MHz FADC boards and crates are provided at no cost to HPS since they can be borrowed from the CLAS upgrade. The system test expenses will also be borne by JLab Hall B. The total cost is \$161k.

HPS plans to include a muon system as a future upgrade. The Muon system is not costed in this proposal.

The Slow Controls are needed to monitor the operations of the three sub-detectors. In addition, they will control and interlock the movements of the SVT with respect to the beamline and provide beam protection interlocks. The total cost is \$134k, which is essentially the labor required to integrate the HPS with the existing Slow Control system in Hall-B.

The infrastructure related costs are roughly one quarter of the total, or \$28k.

The offline computing resources will be provided by JLab. Local DST storage at SLAC will require purchasing tapes, for roughly \$10k.

Travel and lodging expenses for SLAC trips to JLab are also included in this proposal. During design and construction, there will be a small number of trips to solidify and review designs, and to work together to begin integration of the SLAC and JLab DAQ. Funds are also reserved for collaboration meetings to be held during calendar 2014-2016 at JLab. Travel funds for consultations and collaboration meetings total \$125k, and are considered operations. Additional travel funds are requested for integration and installation, totalling \$26k, as capital equipment funds, and for commissioning and data taking runs, totalling \$130k and treated as operations funds. The total travel expense, including both operations and capital equipment, is \$281k.

Project management for HPS is provided by our project engineer, Marco Oriunno. Project management costs total \$167K as capital equipment.

The University of California at Santa Cruz is funded through the SLAC contract to provide support for graduate student Omar Moreno, 20% of physicist Vitaliy Fadeyev, travel funds for commuting to SLAC, attending collaboration meetings and running shifts, and a little M&S. Over the three years of this proposal, the costs totals \$463K, which are accounted as operations.

**The total cost for HPS is \$2.972 M, consisting of \$1.796 M capital equipment, \$928 K operations, and \$248 K infrastructure.** The spending profile for HPS is given in Figure 75.

HPS is seeking funding from other sources for the Muon System and upgrades to the Ecal. William&Mary has submitted an MRI proposal to NSF for the Muon System, requesting  $\sim$  \$200k. IPN ORSAY (France) has submitted a proposal to a French funding agency for the various ECal upgrades, including an ECal Light Monitoring System (\$100k), new, high performance APD's to improve sensitivity (\$500k), and other expenses related to ECal fabrication and test. Note that the new APDs are not part of this proposal. If these requests are approved, the supplemental funding for any items in the present budget will be subtracted from the total cost of the HPS. The disposition of these requests should be known before the beginning of FY2014.

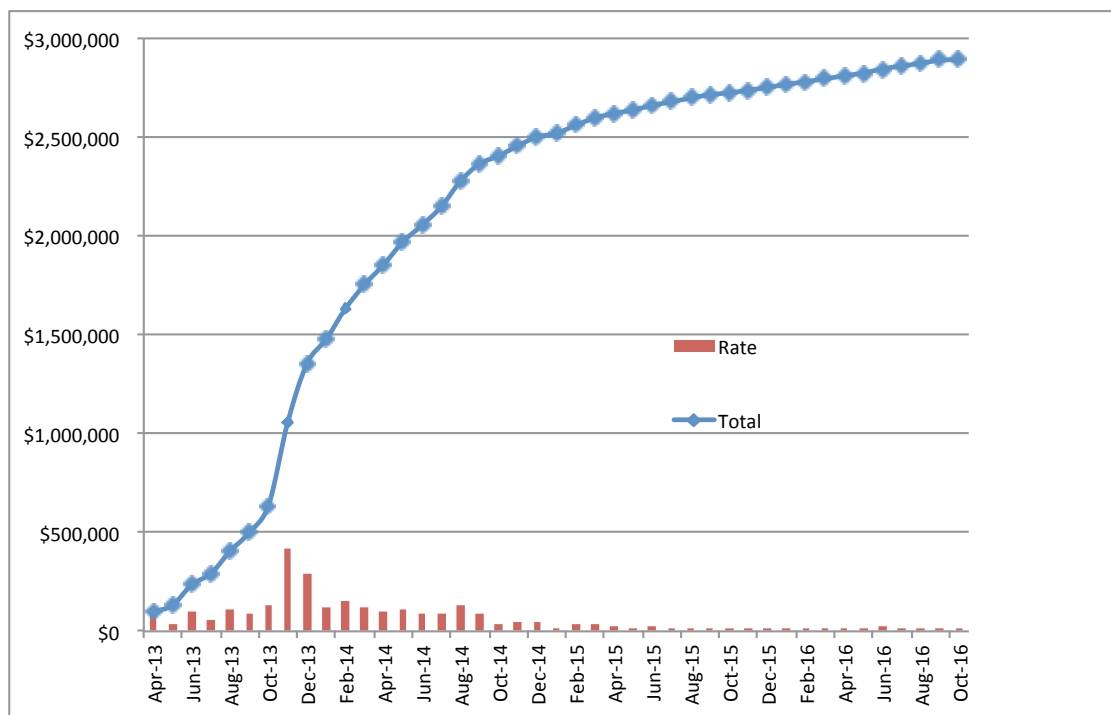


FIG. 75: Spending profile (costs after overheads and contingency).

## 9.2 Schedule

Our goal is to be ready to install the HPS at JLab by September 2014, and to proceed with commissioning on beam early in FY2015. Data taking would begin in Spring 2015 and last until Summer. Meeting this schedule has required keep-alive funding at SLAC for the period April-September, 2013, to begin critical long leadtime engineering and prototyping, and it will need approval and funding from DOE by the beginning of FY2014. Schedules for each of the major subsystems of the experiment are given in Figures 76 and 77, and summarized here. The total construction schedule extends over 16 months, assuming the funding becomes available in October 2014. The schedule contingency is about 20%.

The conceptual design of the beamline will be done during calendar 2013. Formal beamline engineering will start when funding is secured. A Beamline Engineering Design Review will be held in December 2013 to validate the concept before committing funds to construction. Final Engineering and Construction will start in Spring 2014 and will be completed well before the installation time in October 2014, providing substantial float.

Using keep-alive funds, the Test Run SVT was shipped back to SLAC in early February

2013 for continued commissioning of the SVT DAQ, reworking the modules for the first three layers of the HPS, and commissioning the motion control systems. The engineering design and prototyping of the Layers 1-2-3 and Layers 4-5-6 has already begun using keep alive funding. An Engineering Design Review of the SVT will be held late in fiscal 2013, before major construction begins.

Engineering for the new SVT DAQ has also begun. An Engineering Review will precede major construction. Integrating the SVT DAQ with the SVT and commissioning the entire system will occur in Spring 2014. The SVT will be ready for shipping in June 2014, and be ready for installation at JLab by mid-August 2014. Installing the SVT in the analyzing magnet vacuum chamber on beamline will likely occur in September, depending on the schedule at JLab. The SVT schedule has 1 month of float between the shipping and the test at JLab, which provides additional contingency for the construction work at SLAC.

The Ecal work will start when funding is received early in FY2014 and will run through June 2014. The scheduled work is relatively minor, so the ECAL will easily be ready for installation by August 2014. In the event supplemental funding is received from France or Italy, the construction work will be more ambitious, with the possible addition of new APDs and/or new preamplifiers. However, experienced teams at INFN Genova and Orsay already have familiarity with the proposed improvements, and can straightforwardly manage the construction in time for the installation at JLab.

The schedule includes a series of milestones to track the progress of each subsystem, given in Table XXIII. These milestones will help HPS management monitor the readiness of each sub-detector or system after initial assembly and testing at the respective assembly sites, and the readiness for installation at JLab. Also, ad-hoc Engineering Design Reviews, summarized in Table XXIV, will be conducted by the Project Manager for each subsystem before major costs are incurred.



TABLE XXIII: Project Milestones.

WBS	Milestones	Date
1.1	Beamline	
1.1.2	Beamline Review	20-Dec-13
1.1.12	Installation Review	18-Aug-14
1.1.14	Beamline installed	26-Sep-14
1.2	SVT	
1.2.1	SVT Design Review	20-Aug-13
1.2.3	Layer 1-3 ready	17-Feb-14
1.2.5	Layer 4-6 ready	21-Feb-14
1.2.14	SVT tested and ready for shipment	9-Jun-14
1.2.16	SVT ready for installation	15-Aug-14
1.3	SVT DAQ	
1.3.3	FE board tested w/ hybrid	28-Aug-13
1.3.5	Flange DAQ tested	10-Jan-14
1.3.7	DAQ full system Test	2-Jun-14
1.3.8.8	Single hybrid qualification	28-Aug-13
1.3.10	Flex cable w/ hybrid & FE board	10-Jan-14
1.4	ECal	
1.4.8	ECal ready for the installation	1-Aug-14
1.5	TDAQ	
1.5.8	TDAQ ready	6-Dec-13
1.7	Installation & Commissioning	
1.7.6	HPS ready for the beam	26-Sep-14

TABLE XXIV: Planned Reviews.

WBS	Engineering Reviews	Date
1.1.2	Beamline Review	20-Dec-13
1.1.11	Installation Review	18-Aug-14
1.2.1	SVT Design Review	20-Aug-13

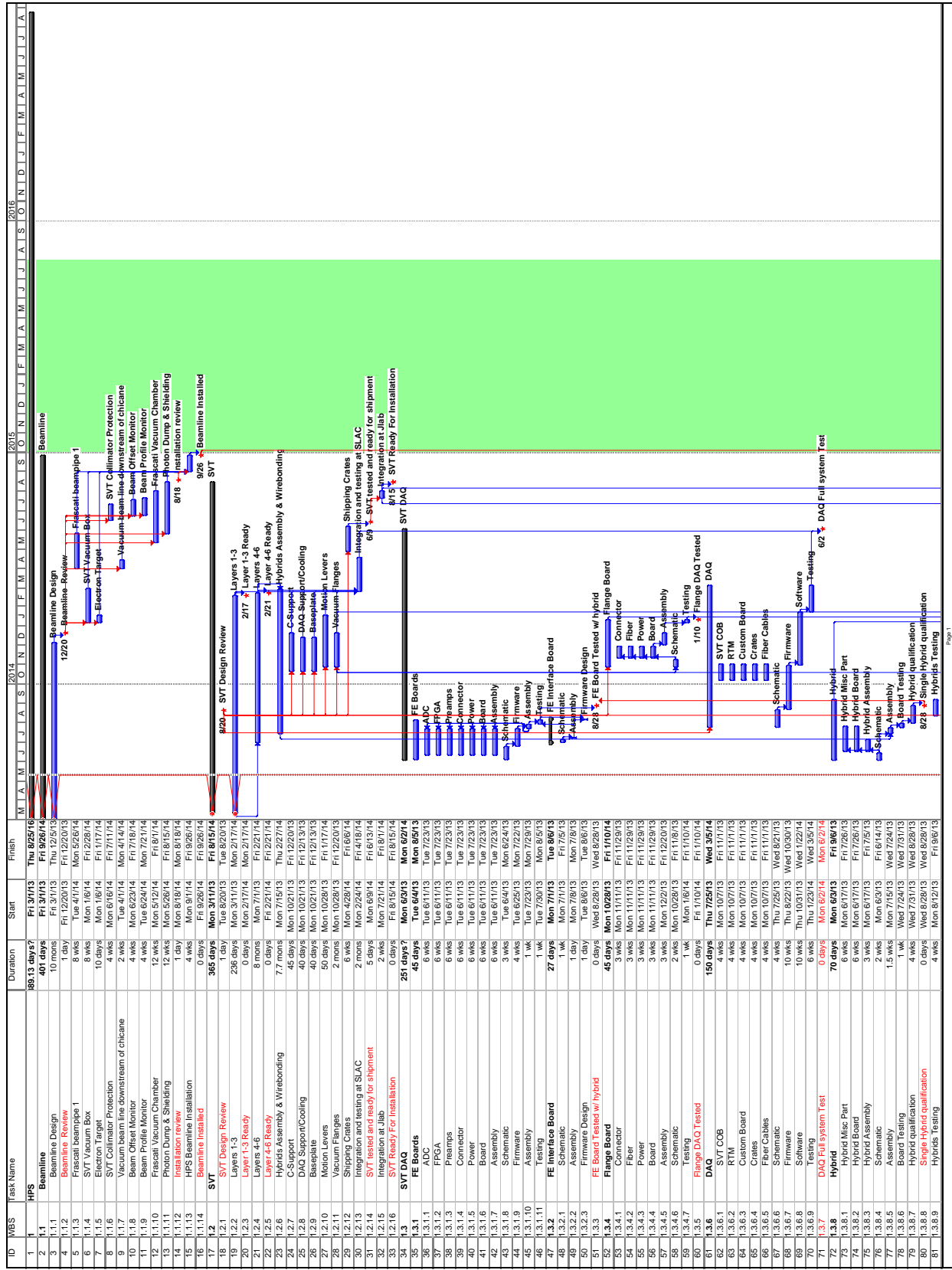


FIG. 76: HPS schedule.

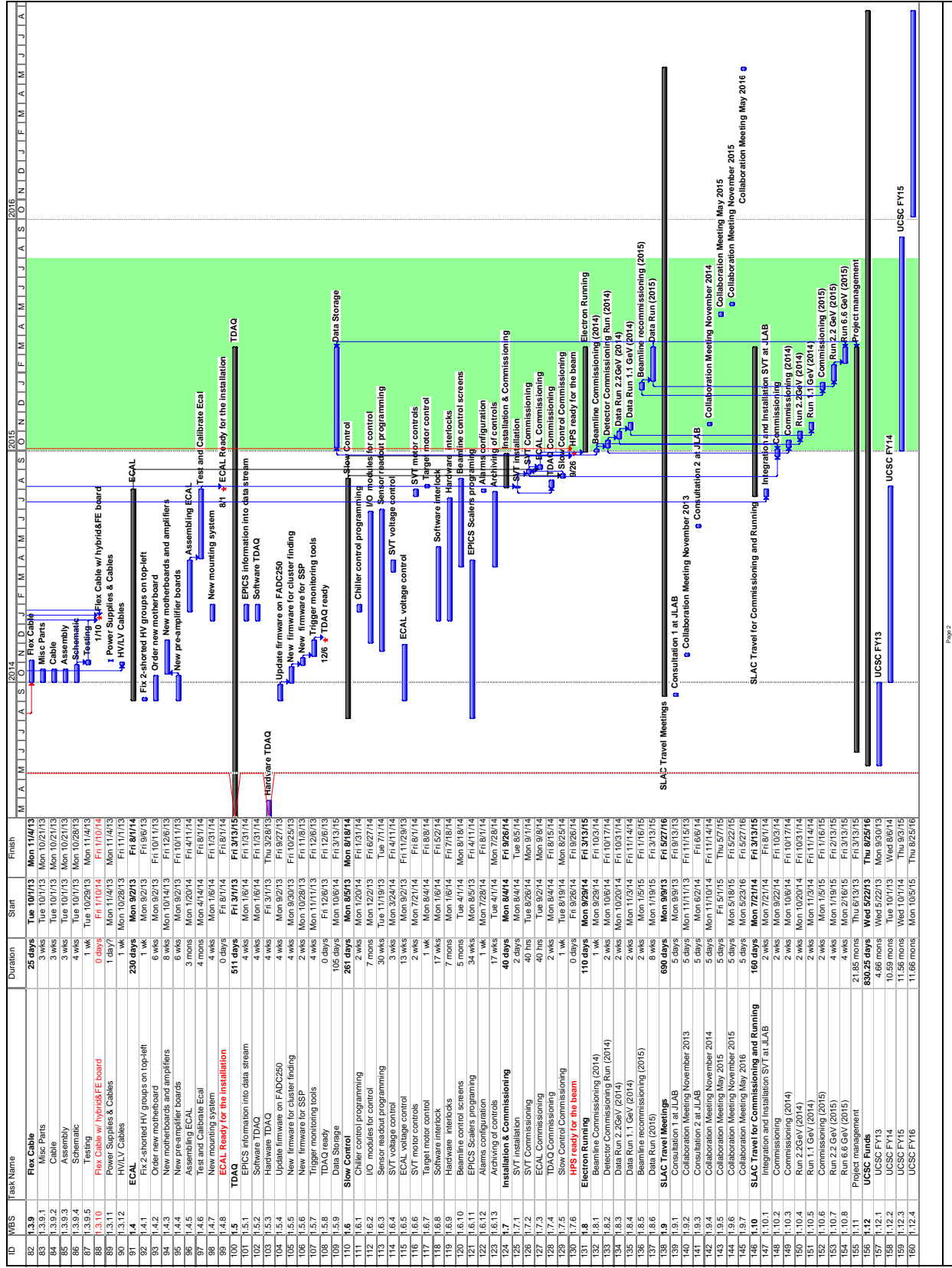


FIG. 77: HPS schedule.

### 9.3 Manpower

The manpower needed to design, fabricate, assemble, test, install, and commission the HPS is captured in the WBS tables. The HPS Collaboration successfully mounted the HPS Test Run experiment on a very aggressive schedule, and has the personnel needed to realize full HPS.

Beamline conceptual design will be done at JLab by Arne Freyberger, F-X Girod and Stepan Stepanyan and at SLAC by Ken Moffeit. Mechanical design will be done at JLab. Engineering at SLAC will be done by Marco Oriunno, Clive Field, and Takashi Maruyama. Fabrication will be done in the JLab shops, and installation by the crews at JLab.

The Tracker/Vertexer is being designed and engineered by Marco Oriunno, Matt Swift, Tim Nelson, and Per Hansson, with additional help from Vitaliy Fadeyev, Alex Grillo, and Bill Cooper, all well-experienced with silicon detector systems. Others at SLAC and UCSC will help with assembly and testing, including Matt Graham, Takashi Maruyama, John Jaros, and graduate students Sho Uemura and Omar Moreno. Matt McCulloch will serve as the technician at SLAC.

The SVT DAQ is being done by Haller's group at SLAC, including Gunther Haller, Ryan Herbst, Ben Reese, and Tung Phan, coordinated by physicist Per Hansson. SVT Physicists Alex Grillo, Vitaliy Fadeyev, and Tim Nelson will collaborate closely. Graduate students Omar Moreno and Sho Uemura will work with Per Hansson and Ryan Herbst to debug, test, and certify the SVT DAQ electronics.

The Ecal work is being coordinated by the Orsay Group, especially Philippe Rosier, Emmanuel Rindel, Emmanuel Raully, Raphael Dupre, and Michel Guidal, with participation by the JLab group, especially Stepan Stepanyan, and F.-X. Girod. Others at JLab and in the collaboration will help in assembly and test of the ECal, especially the group from INFN Genova (Italy).

The Ecal Trigger/DAQ work is done in Sergey Boyarinov's group, which supports Hall B activities at JLab, and with Chris Cuevas's group, which has designed the FADC250. R. Dupre and V. Kubarovsky will collaborate with this group in assembling and testing the electronics, programming the trigger, and integrating it with the Ecal hardware. Sho Uemura will test the trigger in simulation, and help develop diagnostics to ensure proper operations.

Slow control programming is being done by Nerses Gevorgyan (Yerevan) and Hovanes Egiyan (JLab).

The HPS collaboration is about 60 strong, so has adequate manpower for overall installation, commissioning, and data taking. Simulation work is supported by Maurik Holtrop, Matt Graham, Maurizio Ungaro, Takashi Maruyama, and students Sho Uemura and Omar Moreno. Norman Graf and Jeremy McCormick at SLAC support the lcsim simulation/reconstruction framework that is used for HPS simulation and analysis. Data management and storage and computing infrastructure will be overseen by Sergey Boyarinov, Maurik Holtrop, Homer Neal, all very experienced professionals, and graduate student Sho Uemura. Analysis and simulation studies have been initiated by Maurik Holtrop, Sarah Philips, Joey Reichert, Yuri Gershtein, Matt Graham, Per Hansson, Sho Uemura, Takashi Maruyama, and Omar Moreno. Students are actively being engaged.

The total technical labor needed for HPS, divided by category for each of SLAC and JLab, over the next three years, is given in Table XXV.

TABLE XXV: Total Labor (FTE).

SLAC					JLab				
FTE	FY13	FY14	FY15	FY16	FTE	FY13	FY14	FY15	FY16
ME	0.19	0.76	0.04	0.02	ME	0.00	0.03	0	0
MD	0.09	0.00	0.00	0.00	MD	0.05	0.06	0	0
MT	0.08	1.04	0.00	0.00	MT	0.00	0.14	0	0
EE	0.58	0.86	0.06	0.02	EE	0.12	1.19	0	0
ET	0.01	0.05	0.00	0.00	ET	0.02	0.07	0	0

## 9.4 Project Management

The HPS Collaboration formally came into existence in October, 2011 with the acceptance of Collaboration Bylaws and an initial list of members. Figure 78 shows the HPS Organization Chart. The HPS Collaboration is managed by its three spokespersons, Maurik Holtrop, John Jaros, and Stepan Stepanyan in consultation with the HPS Executive Committee, which consists of the spokespeople along with Takashi Maruyama, Matt Graham, Tim Nelson, and F-X Girod. John Jaros serves as Chair of the Executive Committee. Ten working groups, shown in Table XXVI, have been created and Chairs and Deputies

appointed with responsibilities to oversee design, schedule, and budget of each sub-system.

The overall HPS Project Manager is Marco Oriunno. The Project Management team, Oriunno, Stepanyan, and Jaros, in consultation with the Executive Committee, has the responsibility to redirect funding as needed to deal with budget and scheduling exigencies.

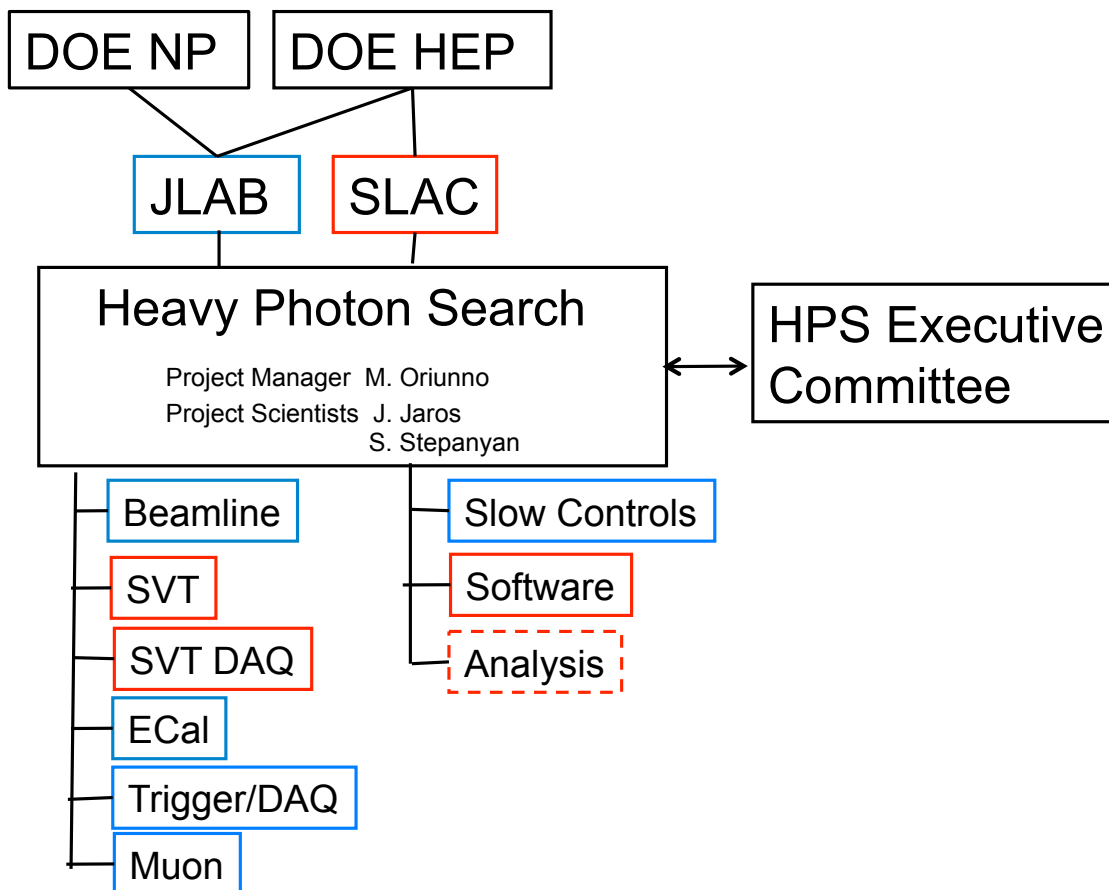


FIG. 78: HPS Organization Chart.

TABLE XXVI: Working groups.

HPS working Groups	Chair (Deputy)
Beamline	K. Moffeit (FX Girod)
SVT	T. Nelson (V. Fedayev)
ECAL	R. Dupre (S. Stepanyan)
DAQ	S. Boiarinov (P. Hansson)
Trigger	V. Kubarovsky (T. Maruyama)
Slow Control	H. Egiyan (N. Gevorgyan)
Muon	K. Griffioen (Y. Gershtein)
Software	M. Holtrop (S. Uemura, O. Moreno)
Analysis	M. Graham (S. Philips)
Project Management	M. Oriunno (S. Stepanyan, J. Jaros)

# A WBS Tables

ID	WBS	Task Name	Type	Work	#	Prototype	Labor	Total	Material	Model	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
1	1	1 HPS		33,126.6 hrs	0	0	\$1,366,880.14	\$1,761,144.08	\$1,039,083.10	\$1,210,638.48	\$2,971,722.56	\$7,915.31	\$20,834.95	\$92,715.31	\$248,127.50	\$1,796,137.20
2	1.1	1.1 Beamline		1,680 hrs	0	0	\$70,349.92	\$91,480.90	\$100,670.00	\$133,131.00	\$222,611.90	\$0.00	\$0.00	\$0.00	\$5,811.00	\$216,800.90
3	1.1.1	1.1.1 Beamline Design		440 hrs	0	0	\$8,966.40	\$11,656.32	\$0.00	\$0.00	\$11,656.32	\$0.00	\$0.00	\$0.00	\$0.00	\$11,656.32
		MD Hall-B		120 hrs	0	0	\$8,966.40	\$11,656.32	\$0.00	\$0.00	\$11,656.32	\$0.00	\$0.00	\$0.00	\$0.00	\$11,656.32
		Ken Moffat (Phys)		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Sepanyan (Phys)		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
4	1.1.2	1.1.2 Beamline Review		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
5	1.1.3	1.1.3 Frascati beampipe 1		128 hrs	0	0	\$0.00	\$0.00	\$7,450.00	\$9,685.00	\$9,685.00	\$0.00	\$0.00	\$0.00	\$0.00	\$9,685.00
		Ken Moffat (Phys)		64 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Frascati beampipe 1		64 hrs	0	0	\$0.00	\$0.00	\$7,450.00	\$9,685.00	\$9,685.00	\$0.00	\$0.00	\$0.00	\$0.00	\$9,685.00
		Stepan Sepanyan (Phys)		64 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
6	1.1.4	1.1.4 SVT Vacuum Box		248 hrs	0	0	\$18,482.40	\$24,021.12	\$8,640.00	\$11,230.00	\$35,259.12	\$0.00	\$0.00	\$0.00	\$0.00	\$35,259.12
		Metro Ormino (ME)		40 hrs	0	0	\$1,344.00	\$9,547.20	\$0.00	\$0.00	\$9,547.20	\$0.00	\$0.00	\$0.00	\$0.00	\$9,547.20
		Ken Moffat (Phys)		160 hrs	0	0	\$11,138.40	\$14,476.80	\$0.00	\$0.00	\$14,476.80	\$0.00	\$0.00	\$0.00	\$0.00	\$14,476.80
		Ken Moffat (Phys)		64 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Linear Shirts		3	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Cooling Feedthrough		2	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Electron Run Flange		1	0	0	\$0.00	\$0.00	\$5,400.00	\$7,020.00	\$7,020.00	\$0.00	\$0.00	\$0.00	\$0.00	\$7,020.00
		Vacuum Flange hardware		1	0	0	\$0.00	\$0.00	\$1,080.00	\$1,440.00	\$1,440.00	\$0.00	\$0.00	\$0.00	\$0.00	\$1,440.00
		Central Assemblies		64 hrs	0	0	\$2,116.80	\$2,819.20	\$0.00	\$0.00	\$2,819.20	\$0.00	\$0.00	\$0.00	\$0.00	\$2,819.20
		Stepan Sepanyan (Phys)		64 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
7	1.1.5	1.1.5 Electron Target		112 hrs	0	0	\$12,900.80	\$16,628.04	\$4,320.00	\$5,616.00	\$22,244.04	\$0.00	\$0.00	\$0.00	\$0.00	\$22,244.04
		40 hrs		40 hrs	0	0	\$7,344.00	\$9,547.20	\$0.00	\$0.00	\$9,547.20	\$0.00	\$0.00	\$0.00	\$0.00	\$9,547.20
		Ken Moffat (Phys)		40 hrs	0	0	\$5,446.80	\$7,088.84	\$0.00	\$0.00	\$7,088.84	\$0.00	\$0.00	\$0.00	\$0.00	\$7,088.84
		Ken Moffat (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Electron Target		1	0	0	\$0.00	\$0.00	\$4,320.00	\$5,616.00	\$5,616.00	\$0.00	\$0.00	\$0.00	\$0.00	\$5,616.00
		Stepan Sepanyan (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
8	1.1.6	1.1.6 SVT Collimator Protection		176 hrs	0	0	\$9,531.84	\$12,391.39	\$14,900.00	\$19,370.00	\$31,761.39	\$0.00	\$0.00	\$0.00	\$0.00	\$31,761.39
		Ken Moffat (Phys)		32 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Sepanyan (Phys)		32 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		SVT Collimator Protection		1	0	0	\$3,177.12	\$4,156.16	\$14,900.00	\$19,370.00	\$19,370.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,370.00
		Ken Moffat (Phys)		32 hrs	0	0	\$5,977.60	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00	\$0.00	\$0.00	\$7,770.88
		MT Accelerator JLAB		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
9	1.1.7	1.1.7 Vacuum beam line downstream of chicane		32 hrs	0	0	\$0.00	\$0.00	\$14,900.00	\$19,370.00	\$19,370.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,370.00
		Ken Moffat (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Sepanyan (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Vacuum beamline downstream chicane		1	0	0	\$0.00	\$0.00	\$14,900.00	\$19,370.00	\$19,370.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,370.00
10	1.1.8	1.1.8 Beam Offset Monitor		128 hrs	0	0	\$7,754.72	\$10,081.14	\$4,470.00	\$5,811.00	\$15,892.14	\$0.00	\$0.00	\$0.00	\$5,811.00	\$10,081.14
		ME Hall-B		16 hrs	0	0	\$1,777.12	\$2,310.26	\$0.00	\$0.00	\$2,310.26	\$0.00	\$0.00	\$0.00	\$0.00	\$2,310.26
		Stepan Sepanyan (Phys)		80 hrs	0	0	\$5,977.60	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00	\$0.00	\$0.00	\$7,770.88
11	1.1.9	1.1.9 Beam Profile Monitor	INFRA	32 hrs	0	0	\$0.00	\$0.00	\$4,470.00	\$5,811.00	\$5,811.00	\$0.00	\$0.00	\$0.00	\$0.00	\$5,811.00
		Ken Moffat (Phys)		64 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Sepanyan (Phys)		32 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Insensible YAG Viewer		1	0	0	\$0.00	\$0.00	\$10,450.00	\$13,550.00	\$13,550.00	\$0.00	\$0.00	\$0.00	\$0.00	\$13,550.00
12	1.1.10	1.1.10 Frascati Vacuum Chamber		24 hrs	0	0	\$0.00	\$0.00	\$14,900.00	\$19,370.00	\$19,370.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,370.00
		Frascati Vacuum Chamber		1	0	0	\$0.00	\$0.00	\$14,900.00	\$19,370.00	\$19,370.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,370.00
		Stepan Sepanyan (Phys)		24 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
13	1.1.11	1.1.11 Photon Dump & Shielding		176 hrs	0	0	\$5,977.60	\$7,770.88	\$20,860.00	\$27,118.00	\$34,888.88	\$0.00	\$0.00	\$0.00	\$0.00	\$34,888.88
		MD Hall-B		80 hrs	0	0	\$5,977.60	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00	\$0.00	\$0.00	\$7,770.88
		Stepan Sepanyan (Phys)		96 hrs	0	0	\$0.00	\$0.00	\$20,860.00	\$27,118.00	\$27,118.00	\$0.00	\$0.00	\$0.00	\$0.00	\$27,118.00
		Photon Dump & Shielding		0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
14	1.1.12	1.1.12 Installation review		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
15	1.1.13	1.1.13 HPS Beamline Installation		152 hrs	0	0	\$6,866.16	\$8,926.01	\$0.00	\$0.00	\$8,926.01	\$0.00	\$0.00	\$0.00	\$0.00	\$8,926.01
		ME Hall-B		8 hrs	0	0	\$888.56	\$1,155.13	\$0.00	\$0.00	\$1,155.13	\$0.00	\$0.00	\$0.00	\$0.00	\$1,155.13
		MT Hall-B		80 hrs	0	0	\$5,977.60	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00	\$0.00	\$0.00	\$7,770.88
		Ken Moffat (Phys)		32 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Sepanyan (Phys)		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
16	1.1.14	1.1.14 Beamline Installed		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00



WBS \470 Fri 5/10/13

ID	WBS	Task Name	Type	Work	#	Prototype	Spares	Material	Total	Material	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
17	1.2	1.2 SVT		5,311 hrs	0	0	0	\$346,316.76	\$461,984.75	\$160,671.00	\$204,194.20	\$656,178.95	\$10,038.00	\$74,827.44	\$42,614.00	\$538,737.51
18	1.2.1	1.2.1 SVT Design Review		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
19	1.2.2	1.2.2 Layers 1-3		1,060 hrs	0	0	0	\$53,305.20	\$66,193.92	\$26,960.00	\$37,104.00	\$103,297.92	\$0.00	\$0.00	\$0.00	\$103,297.92
		Marco Onirno (ME)		80 hrs	0	0	0	\$14,688.00	\$17,625.60	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$17,625.60
		Matt Swift (MD)		160 hrs	0	0	0	\$22,276.80	\$28,959.84	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$28,959.84
		Matt McCulloch (MT)		120 hrs	0	0	0	\$16,340.40	\$19,608.48	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$19,608.48
		Assembly Fixture		1	0	0	0	\$0.00	\$0.00	\$1,200.00	\$1,440.00	\$0.00	\$0.00	\$0.00	\$0.00	\$1,440.00
		Tensioners		2	0	0	0	\$0.00	\$0.00	\$18,144.00	\$18,144.00	\$0.00	\$0.00	\$0.00	\$0.00	\$18,144.00
		U Support		8	0	0	0	\$0.00	\$0.00	\$12,960.00	\$15,120.00	\$0.00	\$0.00	\$0.00	\$0.00	\$15,120.00
		Cables Fixtures		350 hrs	0	0	0	\$0.00	\$0.00	\$2,000.00	\$2,400.00	\$0.00	\$0.00	\$0.00	\$0.00	\$2,400.00
		Per Neilson (Phys)		350 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
20	1.2.3	1.2.3 Layer 1-3 Ready		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
21	1.2.4	1.2.4 Layers 4-6		1,400 hrs	0	0	0	\$77,785.20	\$106,671.60	\$42,470.00	\$59,058.00	\$165,729.60	\$10,038.00	\$17,934.00	\$0.00	\$147,795.60
		Marco Onirno (ME)		160 hrs	0	0	0	\$29,376.00	\$41,124.40	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$41,124.40
		Matt Swift (MD)		160 hrs	0	0	0	\$22,276.80	\$28,959.84	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$28,959.84
		Matt McCulloch (MT)		160 hrs	0	0	0	\$21,787.20	\$30,502.08	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$30,502.08
		Tung Phan (ET)		40 hrs	0	0	0	\$4,345.20	\$6,083.28	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$6,083.28
		Assembly Fixture		5	2	2	0	\$0.00	\$0.00	\$6,000.00	\$8,400.00	\$3,360.00	\$3,360.00	\$6,720.00	\$0.00	\$1,680.00
		Wirebonding Fixture		2	0	1	0	\$0.00	\$0.00	\$3,060.00	\$4,284.00	\$0.00	\$2,142.00	\$2,142.00	\$0.00	\$2,142.00
		Tensioners		8	2	2	0	\$0.00	\$0.00	\$12,960.00	\$18,144.00	\$0.00	\$4,536.00	\$9,072.00	\$0.00	\$9,072.00
		U Support		1	0	0	0	\$0.00	\$0.00	\$15,120.00	\$15,120.00	\$0.00	\$0.00	\$0.00	\$0.00	\$15,120.00
		Cables Fixtures		440 hrs	0	0	0	\$0.00	\$0.00	\$2,400.00	\$2,400.00	\$0.00	\$0.00	\$0.00	\$0.00	\$2,400.00
		Per Neilson (Phys)		440 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		FNAL Composite		1	0	0	0	\$0.00	\$0.00	\$7,650.00	\$10,710.00	\$0.00	\$0.00	\$0.00	\$0.00	\$10,710.00
		Kapton passivation		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
22	1.2.5	1.2.5 Layer 4-6 Ready		0 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
23	1.2.6	1.2.6 Hybrids Assembly & Wirebonding		0 hrs	0	0	0	\$0.00	\$0.00	\$27,000.00	\$27,000.00	\$0.00	\$0.00	\$0.00	\$0.00	\$27,000.00
24	1.2.7	1.2.7 C-Support		321 hrs	0	0	0	\$21,579.12	\$25,894.94	\$7,000.00	\$27,000.00	\$27,000.00	\$0.00	\$0.00	\$0.00	\$27,000.00
		Marco Onirno (ME)		45 hrs	0	0	0	\$8,262.00	\$9,914.40	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$9,914.40
		Matt Swift (MD)		80 hrs	0	0	0	\$13,384.80	\$13,366.08	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$13,366.08
		Matt McCulloch (MT)		16 hrs	0	0	0	\$2,178.72	\$2,614.46	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$2,614.46
		C-Support		1	0	0	0	\$0.00	\$0.00	\$2,160.00	\$2,920.00	\$0.00	\$0.00	\$0.00	\$0.00	\$2,920.00
		Per Neilson (Phys)		90 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		90 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Neilson (Phys)		248 hrs	0	0	0	\$18,360.00	\$25,704.00	\$3,725.00	\$8,215.00	\$30,919.00	\$0.00	\$0.00	\$0.00	\$30,919.00
25	1.2.8	1.2.8 DAC Support/Cooling		40 hrs	0	0	0	\$7,344.00	\$10,281.60	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$10,281.60
		Marco Onirno (ME)		40 hrs	0	0	0	\$5,569.20	\$7,996.88	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$7,996.88
		Matt Swift (MD)		40 hrs	0	0	0	\$5,446.80	\$7,625.52	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$7,625.52
		Matt McCulloch (MT)		64 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		64 hrs	0	0	0	\$0.00	\$0.00	\$3,720.00	\$5,214.00	\$0.00	\$0.00	\$0.00	\$0.00	\$5,214.00
		DAC Support		440 hrs	0	0	0	\$18,360.00	\$22,032.00	\$5,400.00	\$6,480.00	\$28,512.00	\$0.00	\$0.00	\$0.00	\$28,512.00
26	1.2.9	1.2.9 Baseplate		40 hrs	0	0	0	\$7,344.00	\$8,812.80	\$0.00	\$0.00	\$8,812.80	\$0.00	\$0.00	\$0.00	\$8,812.80
		Marco Onirno (ME)		40 hrs	0	0	0	\$5,569.20	\$6,685.04	\$0.00	\$0.00	\$6,685.04	\$0.00	\$0.00	\$0.00	\$6,685.04
		Matt Swift (MD)		40 hrs	0	0	0	\$5,446.80	\$6,536.16	\$0.00	\$0.00	\$6,536.16	\$0.00	\$0.00	\$0.00	\$6,536.16
		Baseplate		160 hrs	0	0	0	\$0.00	\$0.00	\$5,400.00	\$6,480.00	\$0.00	\$0.00	\$0.00	\$0.00	\$6,480.00
		Per Neilson (Phys)		160 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		124 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
27	1.2.10	1.2.10 Motion Levers		234 hrs	0	0	0	\$10,673.28	\$12,801.94	\$4,320.00	\$5,184.00	\$17,991.94	\$0.00	\$0.00	\$0.00	\$17,991.94
		Marco Onirno (ME)		10 hrs	0	0	0	\$1,856.00	\$2,203.20	\$0.00	\$0.00	\$2,203.20	\$0.00	\$0.00	\$0.00	\$2,203.20
		Matt Swift (MD)		40 hrs	0	0	0	\$5,569.20	\$6,685.04	\$0.00	\$0.00	\$6,685.04	\$0.00	\$0.00	\$0.00	\$6,685.04
		Matt McCulloch (MT)		24 hrs	0	0	0	\$3,288.08	\$3,921.70	\$0.00	\$0.00	\$3,921.70	\$0.00	\$0.00	\$0.00	\$3,921.70
		Flex links		80 hrs	0	0	0	\$0.00	\$0.00	\$5,184.00	\$5,184.00	\$0.00	\$0.00	\$0.00	\$0.00	\$5,184.00
		Per Neilson (Phys)		80 hrs	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		304 hrs	0	0	0	\$40,269.60	\$56,377.44	\$600.00	\$888.00	\$57,257.44	\$0.00	\$0.00	\$0.00	\$58,000.00
28	1.2.11	1.2.11 Vacuum Flanges	OP	40 hrs	0	0	0	\$7,344.00	\$10,281.60	\$0.00	\$0.00	\$10,281.60	\$0.00	\$0.00	\$0.00	\$10,281.60
		Marco Onirno (ME)	OP	80 hrs	0	0	0	\$11,138.40	\$15,992.76	\$0.00	\$0.00	\$15,992.76	\$0.00	\$0.00	\$0.00	\$15,992.76

WBS \470 Fri 5/10/13

ID	WBS	Task Name	Type	Work	# Spares	Prototype	Labor	Ltotal	Material	Mtotal	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
29	1.2.12	1.2.12 Shipping Crates	OP	160 hrs	0	0	\$21,787.20	\$30,502.08	\$0.00	\$0.00	\$30,502.08	\$0.00	\$0.00	\$30,502.08	\$0.00	\$0.00
		Matt McCulloch (MT)		12 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		12 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		8" CF Flange		2	0	0	\$0.00	\$0.00	\$880.00	\$880.00	\$880.00	\$0.00	\$0.00	\$0.00	\$0.00	\$880.00
		11" CF Flange		2	0	0	\$14,767.56	\$17,721.07	\$15,056.00	\$18,067.20	\$35,788.27	\$516.00	\$0.00	\$516.00	\$0.00	\$35,272.27
		Marco Oniano (ME)		12 hrs	0	0	\$2,203.20	\$2,643.84	\$0.00	\$0.00	\$2,643.84	\$0.00	\$0.00	\$0.00	\$0.00	\$2,643.84
		Matt Swift (MD)		12 hrs	0	0	\$1,670.76	\$2,004.91	\$0.00	\$0.00	\$2,004.91	\$0.00	\$0.00	\$0.00	\$0.00	\$2,004.91
		Matt McCulloch (MT)		80 hrs	0	0	\$10,893.60	\$13,072.32	\$0.00	\$0.00	\$13,072.32	\$0.00	\$0.00	\$0.00	\$0.00	\$13,072.32
		Tim Neilson (Phys)		60 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		60 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Crate		2	0	0	\$0.00	\$0.00	\$12,907.20	\$12,907.20	\$12,907.20	\$0.00	\$0.00	\$0.00	\$0.00	\$12,907.20
		Half-Module Shipping Box		20	2	0	\$0.00	\$0.00	\$4,300.00	\$5,160.00	\$5,160.00	\$516.00	\$0.00	\$516.00	\$0.00	\$4,644.00
30	1.2.13	1.2.13 Integration and testing at SLAC		744 hrs	0	0	\$63,187.20	\$82,143.36	\$32,780.00	\$42,614.00	\$124,757.36	\$0.00	\$0.00	\$0.00	\$42,614.00	\$19,094.40
		Marco Oniano (ME)		80 hrs	0	0	\$14,688.00	\$19,094.40	\$0.00	\$0.00	\$19,094.40	\$0.00	\$0.00	\$0.00	\$0.00	\$19,094.40
		Matt McCulloch (MT)		320 hrs	0	0	\$4,574.40	\$5,646.72	\$0.00	\$0.00	\$5,646.72	\$0.00	\$0.00	\$0.00	\$0.00	\$5,646.72
		Survey Technician SLAC		24 hrs	0	0	\$4,924.80	\$6,402.24	\$0.00	\$0.00	\$6,402.24	\$0.00	\$0.00	\$0.00	\$0.00	\$6,402.24
		Tim Neilson (Phys)		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Chiller	INFRA	1	0	0	\$0.00	\$0.00	\$29,800.00	\$38,740.00	\$38,740.00	\$0.00	\$0.00	\$0.00	\$38,740.00	\$0.00
		Equipment for chiller Control	INFRA	1	0	0	\$0.00	\$0.00	\$2,980.00	\$3,874.00	\$3,874.00	\$0.00	\$0.00	\$0.00	\$3,874.00	\$0.00
31	1.2.14	1.2.14 SVT tested and ready for shipment		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Tim Neilson (Phys)		8 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		8 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
32	1.2.15	1.2.15 Integration at Jlab		320 hrs	0	0	\$28,029.60	\$34,438.48	\$0.00	\$0.00	\$34,438.48	\$0.00	\$0.00	\$0.00	\$0.00	\$34,438.48
		Marco Oniano (ME)		80 hrs	0	0	\$14,688.00	\$19,094.40	\$0.00	\$0.00	\$19,094.40	\$0.00	\$0.00	\$0.00	\$0.00	\$19,094.40
		Ryan Herbst (EE)		80 hrs	0	0	\$13,341.60	\$17,344.08	\$0.00	\$0.00	\$17,344.08	\$0.00	\$0.00	\$0.00	\$0.00	\$17,344.08
		Tim Neilson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
33	1.2.16	1.2.16 SVT Ready For Installation		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
34	1.3	1.3 SVT DAO		3,972 hrs	0	0	\$341,554.16	\$400,545.39	\$264,406.20	\$351,672.88	\$782,218.28	\$69,503.31	\$10,796.95	\$80,300.26	\$161,765.00	\$540,153.01
35	1.3.1	1.3.1 FE Boards		504 hrs	0	0	\$60,037.20	\$84,052.08	\$45,853.80	\$64,195.32	\$148,247.40	\$33,090.50	\$3,676.72	\$36,767.22	\$0.00	\$111,480.18
		Tim Neilson (Phys)		72 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		72 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		1.3.1.1 ADC		0 hrs	0	0	\$0.00	\$0.00	\$5,760.00	\$8,064.00	\$8,064.00	\$3,628.80	\$403.20	\$4,032.00	\$0.00	\$4,032.00
		ADC		80	36	4	\$0.00	\$0.00	\$5,760.00	\$8,064.00	\$8,064.00	\$3,628.80	\$403.20	\$4,032.00	\$0.00	\$4,032.00
		1.3.1.2 FPGA		0 hrs	0	0	\$0.00	\$0.00	\$10,756.00	\$15,058.40	\$15,058.40	\$6,776.28	\$752.92	\$7,529.20	\$0.00	\$7,529.20
		FPGA		20	9	1	\$0.00	\$0.00	\$10,756.00	\$15,058.40	\$15,058.40	\$6,776.28	\$752.92	\$7,529.20	\$0.00	\$7,529.20
		1.3.1.3 Preamps		0 hrs	0	0	\$0.00	\$0.00	\$2,016.00	\$2,822.40	\$2,822.40	\$1,270.08	\$141.12	\$1,411.20	\$0.00	\$1,411.20
		Preamp		200	90	10	\$0.00	\$0.00	\$2,016.00	\$2,822.40	\$2,822.40	\$1,270.08	\$141.12	\$1,411.20	\$0.00	\$1,411.20
		1.3.1.4 Connector		0 hrs	0	0	\$0.00	\$0.00	\$430.20	\$602.28	\$602.28	\$271.03	\$30.11	\$301.14	\$0.00	\$301.14
		Connectors		20	9	1	\$0.00	\$0.00	\$430.20	\$602.28	\$602.28	\$271.03	\$30.11	\$301.14	\$0.00	\$301.14
		1.3.1.5 Power		0 hrs	0	0	\$0.00	\$0.00	\$1,075.60	\$1,505.84	\$1,505.84	\$677.63	\$75.29	\$752.92	\$0.00	\$752.92
		Power		20	9	1	\$0.00	\$0.00	\$1,075.60	\$1,505.84	\$1,505.84	\$677.63	\$75.29	\$752.92	\$0.00	\$752.92
		1.3.1.6 Board		0 hrs	0	0	\$0.00	\$0.00	\$21,512.00	\$30,116.80	\$30,116.80	\$13,552.56	\$1,505.84	\$15,058.40	\$0.00	\$15,058.40
		Board		20	9	0	\$0.00	\$0.00	\$21,512.00	\$30,116.80	\$30,116.80	\$13,552.56	\$1,505.84	\$15,058.40	\$0.00	\$15,058.40
		1.3.1.7 Assembly		0 hrs	0	0	\$0.00	\$0.00	\$4,300.00	\$5,160.00	\$5,160.00	\$516.00	\$0.00	\$516.00	\$0.00	\$516.00
		Assembly		20	9	1	\$0.00	\$0.00	\$4,300.00	\$5,160.00	\$5,160.00	\$516.00	\$0.00	\$516.00	\$0.00	\$516.00
		1.3.1.8 Schematic		120 hrs	0	0	\$20,012.40	\$28,017.36	\$0.00	\$0.00	\$28,017.36	\$0.00	\$0.00	\$0.00	\$0.00	\$28,017.36
		Ben Reese (EE)		160 hrs	0	0	\$26,683.20	\$37,356.48	\$0.00	\$0.00	\$37,356.48	\$0.00	\$0.00	\$0.00	\$0.00	\$37,356.48
		Ryan Herbst (EE)		160 hrs	0	0	\$6,670.80	\$9,339.12	\$0.00	\$0.00	\$9,339.12	\$4,202.60	\$466.96	\$4,669.56	\$0.00	\$4,669.56
		1.3.1.10 Assembly		40 hrs	0	0	\$6,670.80	\$9,339.12	\$0.00	\$0.00	\$9,339.12	\$4,202.60	\$466.96	\$4,669.56	\$0.00	\$4,669.56
		Ben Reese (EE)		40 hrs	18	2	\$6,670.80	\$9,339.12	\$0.00	\$0.00	\$9,339.12	\$4,202.60	\$466.96	\$4,669.56	\$0.00	\$4,669.56
		1.3.1.11 Testing		40 hrs	0	0	\$6,670.80	\$9,339.12	\$0.00	\$0.00	\$9,339.12	\$0.00	\$0.00	\$0.00	\$0.00	\$9,339.12
		Ryan Herbst (EE)		40 hrs	0	0	\$6,670.80	\$9,339.12	\$0.00	\$0.00	\$9,339.12	\$0.00	\$0.00	\$0.00	\$0.00	\$9,339.12
		1.3.2 FE Interface Board		48 hrs	0	0	\$8,004.96	\$9,605.95	\$0.00	\$0.00	\$9,605.95	\$9,605.95	\$0.00	\$9,605.95	\$0.00	\$9,605.95
		Interface Board		1	1	0	\$0.00	\$0.00	\$2,151.20	\$2,581.20	\$2,581.20	\$0.00	\$0.00	\$0.00	\$0.00	\$2,581.20
		1.3.2.1 Schematic		40 hrs	0	0	\$6,670.80	\$9,004.96	\$0.00	\$0.00	\$9,004.96	\$8,004.96	\$0.00	\$8,004.96	\$0.00	\$8,004.96
		Ryan Herbst (EE)		40 hrs	40	0	\$6,670.80	\$9,004.96	\$0.00	\$0.00	\$9,004.96	\$8,004.96	\$0.00	\$8,004.96	\$0.00	\$8,004.96
		1.3.2.2 Assembly		8 hrs	0	0	\$1,334.16	\$1,600.99	\$0.00	\$0.00	\$1,600.99	\$1,600.99	\$0.00	\$1,600.99	\$0.00	\$1,600.99
		1.3.2.3 Firmware Design		8 hrs	0	0	\$1,334.16	\$1,600.99	\$0.00	\$0.00	\$1,600.99	\$1,600.99	\$0.00	\$1,600.99	\$0.00	\$1,600.99
		1.3.3 FE Board tested w/ hybrid		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		1.3.3 FE Board tested w/ hybrid		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00

ID	WBS	Task Name	Type	Work	Spares	# Prototype	Labor	Ltotal	Material	Mtotal	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
52	1.3.4	1.3.4 Flange Board		264 hrs	0	0	\$20,012.40	\$25,349.04	\$18,157.44	\$25,420.42	\$50,649.46	\$6,355.10	\$3,177.55	\$9,532.66	\$0.00	\$41,236.80
		Tim Neilson (Phys)		72 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ben Reese (EE)		72 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
53	1.3.4.1	1.3.4.1 Connector		0 hrs	0	0	\$0.00	\$0.00	\$172.80	\$241.92	\$241.92	\$60.48	\$30.24	\$90.72	\$0.00	\$151.20
		Ben Reese (EE)		16	4	2	\$0.00	\$0.00	\$172.80	\$241.92	\$241.92	\$60.48	\$30.24	\$90.72	\$0.00	\$151.20
54	1.3.4.2	1.3.4.2 Fiber		0 hrs	0	0	\$0.00	\$0.00	\$7,228.00	\$10,119.20	\$10,119.20	\$2,529.80	\$1,264.90	\$3,794.70	\$0.00	\$6,324.50
		Ben Reese (EE)		16	4	2	\$0.00	\$0.00	\$7,228.00	\$10,119.20	\$10,119.20	\$2,529.80	\$1,264.90	\$3,794.70	\$0.00	\$6,324.50
55	1.3.4.3	1.3.4.3 Power		0 hrs	0	0	\$0.00	\$0.00	\$430.24	\$602.34	\$602.34	\$150.58	\$75.29	\$225.88	\$0.00	\$376.46
		Ben Reese (EE)		8	2	1	\$0.00	\$0.00	\$430.24	\$602.34	\$602.34	\$150.58	\$75.29	\$225.88	\$0.00	\$376.46
56	1.3.4.4	1.3.4.4 Board		0 hrs	0	0	\$0.00	\$0.00	\$8,604.80	\$12,046.72	\$12,046.72	\$3,011.68	\$1,505.84	\$4,517.52	\$0.00	\$7,529.20
		Ben Reese (EE)		8	2	1	\$0.00	\$0.00	\$8,604.80	\$12,046.72	\$12,046.72	\$3,011.68	\$1,505.84	\$4,517.52	\$0.00	\$7,529.20
57	1.3.4.5	1.3.4.5 Assembly		0 hrs	0	0	\$0.00	\$0.00	\$1,721.60	\$2,410.24	\$2,410.24	\$602.56	\$301.28	\$903.84	\$0.00	\$1,506.40
		Ben Reese (EE)		8	2	1	\$0.00	\$0.00	\$1,721.60	\$2,410.24	\$2,410.24	\$602.56	\$301.28	\$903.84	\$0.00	\$1,506.40
58	1.3.4.6	1.3.4.6 Schematic		80 hrs	0	0	\$13,341.60	\$18,678.24	\$0.00	\$0.00	\$18,678.24	\$0.00	\$0.00	\$0.00	\$0.00	\$18,678.24
		Ben Reese (EE)		40 hrs	0	0	\$13,341.60	\$18,678.24	\$0.00	\$0.00	\$18,678.24	\$0.00	\$0.00	\$0.00	\$0.00	\$18,678.24
59	1.3.4.7	1.3.4.7 Testing		40 hrs	0	0	\$6,670.80	\$6,670.80	\$0.00	\$0.00	\$6,670.80	\$0.00	\$0.00	\$0.00	\$0.00	\$6,670.80
		Ryan Herbst (EE)		40 hrs	0	0	\$6,670.80	\$6,670.80	\$0.00	\$0.00	\$6,670.80	\$0.00	\$0.00	\$0.00	\$0.00	\$6,670.80
60	1.3.5	1.3.5 Flange DAQ Tested		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
61	1.3.6	1.3.6 DAQ		1,640 hrs	0	0	\$173,440.80	\$233,465.60	\$86,049.60	\$120,488.64	\$333,934.20	\$7,529.20	\$0.00	\$7,529.20	\$91,855.40	\$234,949.60
		Tim Neilson (Phys)		300 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		300 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
62	1.3.6.1	1.3.6.1 SVT COB		0 hrs	0	0	\$0.00	\$0.00	\$38,721.00	\$54,209.40	\$54,209.40	\$0.00	\$0.00	\$0.00	\$0.00	\$54,209.40
		Ben Reese (EE)	INFRA	3	0	0	\$0.00	\$0.00	\$38,721.00	\$54,209.40	\$54,209.40	\$0.00	\$0.00	\$0.00	\$0.00	\$54,209.40
63	1.3.6.2	1.3.6.2 FIM		0 hrs	0	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
		Ben Reese (EE)	INFRA	3	0	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
64	1.3.6.3	1.3.6.3 Custom Board		0 hrs	0	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
		Ben Reese (EE)	INFRA	3	1	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
65	1.3.6.4	1.3.6.4 Crates		0 hrs	0	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
		Ben Reese (EE)	INFRA	2	0	0	\$0.00	\$0.00	\$16,134.00	\$22,587.60	\$22,587.60	\$0.00	\$0.00	\$0.00	\$0.00	\$22,587.60
66	1.3.6.5	1.3.6.5 Fiber Cables		0 hrs	0	0	\$0.00	\$0.00	\$4,304.00	\$6,025.60	\$6,025.60	\$0.00	\$0.00	\$0.00	\$0.00	\$6,025.60
		Ben Reese (EE)	INFRA	16	0	0	\$0.00	\$0.00	\$4,304.00	\$6,025.60	\$6,025.60	\$0.00	\$0.00	\$0.00	\$0.00	\$6,025.60
67	1.3.6.6	1.3.6.6 Schematic		160 hrs	0	0	\$26,683.20	\$37,356.48	\$0.00	\$0.00	\$37,356.48	\$0.00	\$0.00	\$0.00	\$0.00	\$37,356.48
		Ben Reese (EE)	INFRA	160	0	0	\$26,683.20	\$37,356.48	\$0.00	\$0.00	\$37,356.48	\$0.00	\$0.00	\$0.00	\$0.00	\$37,356.48
68	1.3.6.7	1.3.6.7 Firmware		320 hrs	0	0	\$53,366.40	\$64,039.68	\$0.00	\$0.00	\$64,039.68	\$0.00	\$0.00	\$0.00	\$0.00	\$64,039.68
		Ryan Herbst (EE)	INFRA	160	0	0	\$26,683.20	\$32,019.84	\$0.00	\$0.00	\$32,019.84	\$0.00	\$0.00	\$0.00	\$0.00	\$32,019.84
69	1.3.6.8	1.3.6.8 Software		320 hrs	0	0	\$53,366.40	\$64,039.68	\$0.00	\$0.00	\$64,039.68	\$0.00	\$0.00	\$0.00	\$0.00	\$64,039.68
		Ryan Herbst (EE)	INFRA	160	0	0	\$26,683.20	\$32,019.84	\$0.00	\$0.00	\$32,019.84	\$0.00	\$0.00	\$0.00	\$0.00	\$32,019.84
70	1.3.6.9	1.3.6.9 Testing		240 hrs	0	0	\$40,024.80	\$48,029.76	\$0.00	\$0.00	\$48,029.76	\$0.00	\$0.00	\$0.00	\$0.00	\$48,029.76
		Ben Reese (EE)	INFRA	120	0	0	\$20,012.40	\$24,014.88	\$0.00	\$0.00	\$24,014.88	\$0.00	\$0.00	\$0.00	\$0.00	\$24,014.88
71	1.3.7	1.3.7 DAQ Firmware Test		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
72	1.3.8	1.3.8 Hybrid		1,196.8 hrs	0	0	\$40,024.80	\$48,029.76	\$21,386.16	\$25,663.38	\$73,683.15	\$7,128.72	\$1,425.74	\$8,554.46	\$0.00	\$65,138.69
		Tim Neilson (Phys)		158.4 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		158.4 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
73	1.3.8.1	1.3.8.1 Hybrid Misc Part		0 hrs	0	0	\$0.00	\$0.00	\$8,871.12	\$10,645.34	\$10,645.34	\$2,957.04	\$591.41	\$3,548.45	\$0.00	\$7,096.90
		Ryan Herbst (EE)	INFRA	36	10	2	\$0.00	\$0.00	\$8,871.12	\$10,645.34	\$10,645.34	\$2,957.04	\$591.41	\$3,548.45	\$0.00	\$7,096.90
74	1.3.8.2	1.3.8.2 Hybrid Board		0 hrs	0	0	\$0.00	\$0.00	\$2,944.12	\$3,484.94	\$3,484.94	\$968.04	\$193.61	\$1,161.65	\$0.00	\$2,323.30
		Ben Reese (EE)	INFRA	180	50	10	\$0.00	\$0.00	\$2,944.12	\$3,484.94	\$3,484.94	\$968.04	\$193.61	\$1,161.65	\$0.00	\$2,323.30
75	1.3.8.3	1.3.8.3 Hybrid Assembly		0 hrs	0	0	\$0.00	\$0.00	\$7,744.32	\$9,293.18	\$9,293.18	\$2,581.18	\$516.29	\$3,097.47	\$0.00	\$6,195.46
		Ben Reese (EE)	INFRA	36	10	2	\$0.00	\$0.00	\$7,744.32	\$9,293.18	\$9,293.18	\$2,581.18	\$516.29	\$3,097.47	\$0.00	\$6,195.46
76	1.3.8.4	1.3.8.4 Schematic		80 hrs	0	0	\$13,341.60	\$16,009.92	\$0.00	\$0.00	\$16,009.92	\$0.00	\$0.00	\$0.00	\$0.00	\$16,009.92
		Ryan Herbst (EE)	INFRA	40	0	0	\$13,341.60	\$16,009.92	\$0.00	\$0.00	\$16,009.92	\$0.00	\$0.00	\$0.00	\$0.00	\$16,009.92
77	1.3.8.5	1.3.8.5 Assembly		120 hrs	0	0	\$20,012.40	\$24,014.88	\$0.00	\$0.00	\$24,014.88	\$0.00	\$0.00	\$0.00	\$0.00	\$24,014.88
		Ben Reese (EE)	INFRA	60	0	0	\$10,006.20	\$12,007.44	\$0.00	\$0.00	\$12,007.44	\$0.00	\$0.00	\$0.00	\$0.00	\$12,007.44
78	1.3.8.6	1.3.8.6 Board Testing		40 hrs	0	0	\$6,670.80	\$8,004.96	\$0.00	\$0.00	\$8,004.96	\$0.00	\$0.00	\$0.00	\$0.00	\$8,004.96
		Ben Reese (EE)	INFRA	40	0	0	\$6,670.80	\$8,004.96	\$0.00	\$0.00	\$8,004.96	\$0.00	\$0.00	\$0.00	\$0.00	\$8,004.96
79	1.3.8.7	1.3.8.7 Hybrid qualification		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)	INFRA	160	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
80	1.3.8.8	1.3.8.8 Single Hybrid qualification		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Sho Uemura (Phys)	INFRA	0	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
81	1.3.8.9	1.3.8.9 Hybrids testing		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ben Reese (EE)	INFRA	160	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00

WBS V470 Fri 5/10/13

ID	WBS	Task Name	Type	Work	# Spares	Prototype	Labor	Ltotal	Material	Mtotal	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments	Dt	
82	1.3.9	Per Hansson (Phys) Sho Uemura (Phys) <b>1.3.9 Flex Cable</b>		160 hrs 280 hrs 40 hrs 40 hrs	0 0 0 0	0 0 0 0	\$0.00 \$3,354.00 \$0.00 \$0.00	\$0.00 \$41,358.96 \$0.00 \$0.00	\$0.00 \$9,032.16 \$0.00 \$0.00	\$0.00 \$12,645.02 \$0.00 \$0.00	\$0.00 \$54,003.98 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	
83	1.3.9.1	Per Hansson (Phys) 1.3.9.1 Misc Parts		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
84	1.3.9.2	SVT Cable Misc parts 1.3.9.2 Cable		24 hrs	0	0	\$0.00	\$0.00	\$1,290.72	\$1,807.01	\$1,807.01	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$1,807.01
85	1.3.9.3	SVT Cable 1.3.9.3 Assembly		0 hrs	0	0	\$0.00	\$0.00	\$5,160.00	\$1,807.01	\$1,807.01	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$1,807.01
86	1.3.9.4	ST Cable Assembly 1.3.9.4 Schematic		24 hrs	0	0	\$0.00	\$0.00	\$2,581.44	\$3,614.02	\$3,614.02	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$3,614.02
87	1.3.9.5	Ryan Herbst (EE) Ben Reese (EE) 1.3.9.5 Testing		160 hrs 80 hrs 40 hrs 40 hrs	0 0 0 0	0 0 0 0	\$26,683.20 \$13,341.60 \$6,670.80 \$9,339.12	\$32,019.84 \$18,678.24 \$13,341.60 \$9,339.12	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$19,678.24 \$13,341.60 \$9,339.12	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00	
88	1.3.10	Ryan Herbst (EE) 1.3.10 Flex Cable w/ HydradFE board		0 hrs	0	0	\$0.00	\$0.00	\$59,989.20	\$69,989.20	\$69,989.20	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
89	1.3.11	SVT HV Unit SVT HV Unit	INFBA	0 hrs 4	0 0	0 0	\$0.00 \$0.00	\$0.00 \$1,216.00	\$0.00 \$17,059.20	\$0.00 \$17,059.20	\$0.00 \$17,059.20	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00
90	1.3.12	Wiener Crate 1.3.12 HV/LV Cables	INFBA	40 hrs	0	0	\$6,680.00	\$8,664.00	\$16,298.40	\$16,298.40	\$16,298.40	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$16,298.40
91	1.4	ECAL		2,924 hrs	0	0	\$27,868.48	\$30,229.02	\$8,940.00	\$11,622.00	\$47,851.02	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$47,851.02
92	1.4.1	1.4.1 Fix 2-shortied HV groups on top-left		44 hrs	0	0	\$2,988.80	\$3,885.44	\$0.00	\$0.00	\$3,885.44	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$3,885.44
93	1.4.2	ET HeilB/JLAB Phy2 ECAL/JLAB		40 hrs 4 hrs	0 0	0 0	\$2,988.80 \$0.00	\$3,885.44 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$3,885.44 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$0.00 \$0.00	\$3,885.44 \$0.00
94	1.4.3	1.4.2 Order new motherboard Phy2 ECAL/JLAB New Bottom-Left Motherboard		24 hrs 24 hrs 256 hrs	0 0 0	0 0 0	\$0.00 \$0.00 \$24,879.68	\$0.00 \$0.00 \$32,343.58	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$32,343.58	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	
95	1.4.4	1.4.3 New motherboards and amplifiers EE HeilB/JLAB Phy2 ECAL/JLAB New Amplifiers Boards		224 hrs 32 hrs 60 hrs	0 0 0	0 0 0	\$24,879.68 \$0.00 \$0.00	\$32,343.58 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$32,343.58 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	
96	1.4.5	1.4.4 New pre-amplifier boards Phy2 ECAL/JLAB Ecal Pre-amps		24 hrs 24 hrs 500 hrs	0 0 0	0 0 0	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	
97	1.4.6	1.4.5 Assembling ECAL Phy2 ECAL/JLAB 1.4.6 Test and Calibrate Ecal		160 hrs 2,080 hrs 160 hrs 40 hrs 640 hrs	0 0 0 0 0	0 0 0 0 0	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00 \$0.00 \$0.00	
98	1.4.7	1.4.7 New mounting system Phy2 ECAL/JLAB ECAL Mounting System		160 hrs 160 hrs 160 hrs	0 0 0	0 0 0	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	\$0.00 \$0.00 \$0.00	
99	1.4.8	1.4.8 ECAL Ready for the installation ME Ossay MT Ossay		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
100	1.5	1.5 DAO		1,200 hrs	0	0	\$115,512.80	\$151,055.20	\$10,000.00	\$10,000.00	\$161,055.20	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$151,055.20

WBS V470 Fri 5/10/13

ID	WBS	Task Name	Type	Work	# Spares	Prototype	Labor	Ltotal	Material	Mtotal	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
101	1.5.1	1.5.1 EPICS information into data stream		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE DAQ JLAB		160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
102	1.5.2	1.5.2 Software TDAQ		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
103	1.5.3	1.5.3 Hardware TDAQ		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
104	1.5.4	1.5.4 Update firmware on FADC250		160 hrs	0	0	\$17,771.20	\$21,325.44	\$0.00	\$0.00	\$21,325.44	\$0.00	\$0.00	\$0.00	\$0.00	\$21,325.44
		Ben Reed (EE)		160 hrs	0	0	\$17,771.20	\$21,325.44	\$0.00	\$0.00	\$21,325.44	\$0.00	\$0.00	\$0.00	\$0.00	\$21,325.44
105	1.5.5	1.5.5 New firmware for cluster finding		320 hrs	0	0	\$35,542.40	\$49,759.36	\$0.00	\$0.00	\$49,759.36	\$0.00	\$0.00	\$0.00	\$0.00	\$49,759.36
		Scott Kenata (EE)		320 hrs	0	0	\$35,542.40	\$49,759.36	\$0.00	\$0.00	\$49,759.36	\$0.00	\$0.00	\$0.00	\$0.00	\$49,759.36
106	1.5.6	1.5.6 New firmware for SSP		240 hrs	0	0	\$26,656.80	\$37,319.52	\$0.00	\$0.00	\$37,319.52	\$0.00	\$0.00	\$0.00	\$0.00	\$37,319.52
		Ben Reed (EE)		240 hrs	0	0	\$26,656.80	\$37,319.52	\$0.00	\$0.00	\$37,319.52	\$0.00	\$0.00	\$0.00	\$0.00	\$37,319.52
107	1.5.7	1.5.7 Trigger monitoring tools		80 hrs	0	0	\$8,885.60	\$12,439.84	\$0.00	\$0.00	\$12,439.84	\$0.00	\$0.00	\$0.00	\$0.00	\$12,439.84
		Scott Kenata (EE)		80 hrs	0	0	\$8,885.60	\$12,439.84	\$0.00	\$0.00	\$12,439.84	\$0.00	\$0.00	\$0.00	\$0.00	\$12,439.84
108	1.5.8	1.5.8 TDAQ ready		200 hrs	0	0	\$35,542.40	\$42,650.88	\$0.00	\$0.00	\$42,650.88	\$0.00	\$0.00	\$0.00	\$0.00	\$42,650.88
		Ben Reed (EE)		200 hrs	0	0	\$35,542.40	\$42,650.88	\$0.00	\$0.00	\$42,650.88	\$0.00	\$0.00	\$0.00	\$0.00	\$42,650.88
109	1.5.9	1.5.9 Data Storage		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Data Storage Tapes 5TB	INFRA	40 hrs	0	0	\$0.00	\$10,000.00	\$10,000.00	\$10,000.00	\$10,000.00	\$0.00	\$0.00	\$0.00	\$10,000.00	\$10,000.00
110	1.6	1.6 Slow Control		2,008 hrs	0	0	\$75,527.60	\$94,409.50	\$31,290.00	\$39,112.50	\$133,522.00	\$0.00	\$0.00	\$0.00	\$27,937.50	\$105,584.50
111	1.6.1	1.6.1 Chiller control programming		92 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
		Hovanes Heghyan (Phys)		12 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE Accelerator JLAB		80 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
112	1.6.2	1.6.2 I/O modules for control		80 hrs	0	0	\$0.00	\$0.00	\$22,350.00	\$22,350.00	\$27,937.50	\$0.00	\$0.00	\$0.00	\$0.00	\$27,937.50
		I/O Modules for controls	INFRA	1	0	0	\$0.00	\$0.00	\$22,350.00	\$22,350.00	\$27,937.50	\$0.00	\$0.00	\$0.00	\$0.00	\$27,937.50
113	1.6.3	1.6.3 Sensor readout programming		300 hrs	0	0	\$13,328.40	\$16,660.50	\$0.00	\$0.00	\$16,660.50	\$0.00	\$0.00	\$0.00	\$0.00	\$16,660.50
		Hovanes Heghyan (Phys)		180 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Heghyan (Phys)		120 hrs	0	0	\$13,328.40	\$16,660.50	\$0.00	\$0.00	\$16,660.50	\$0.00	\$0.00	\$0.00	\$0.00	\$16,660.50
114	1.6.4	1.6.4 SVT voltage control		138 hrs	0	0	\$13,328.40	\$16,660.50	\$0.00	\$0.00	\$16,660.50	\$0.00	\$0.00	\$0.00	\$0.00	\$16,660.50
		Hovanes Heghyan (Phys)		18 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Nerses Gevorgyan (EE)		120 hrs	0	0	\$13,328.40	\$16,660.50	\$0.00	\$0.00	\$16,660.50	\$0.00	\$0.00	\$0.00	\$0.00	\$16,660.50
115	1.6.5	1.6.5 ECAL voltage control		118 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Heghyan (Phys)		78 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Nerses Gevorgyan (EE)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
116	1.6.6	1.6.6 SVT motor controls		92 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
		Hovanes Heghyan (Phys)		12 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE Accelerator JLAB		80 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
117	1.6.7	1.6.7 Target motor control		46 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
		Hovanes Heghyan (Phys)		6 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE Accelerator JLAB		40 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
118	1.6.8	1.6.8 Software interlock		142 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
		Hovanes Heghyan (Phys)		102 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE Accelerator JLAB		40 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
119	1.6.9	1.6.9 Hardware interlocks		368 hrs	0	0	\$0.00	\$0.00	\$8,940.00	\$8,940.00	\$11,175.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,175.00
		EE Hall-B JLAB		200 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hardware interlock Equipments		168 hrs	0	0	\$0.00	\$0.00	\$8,940.00	\$8,940.00	\$11,175.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,175.00
120	1.6.10	1.6.10 Beamline control screens		160 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
		Hovanes Heghyan (Phys)		120 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Nerses Gevorgyan (EE)		40 hrs	0	0	\$4,442.80	\$5,553.50	\$0.00	\$0.00	\$5,553.50	\$0.00	\$0.00	\$0.00	\$0.00	\$5,553.50
121	1.6.11	1.6.11 EPICS Scalers programming		284 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Heghyan (Phys)		204 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Nerses Gevorgyan (EE)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
122	1.6.12	1.6.12 Alarms configuration		40 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
		Hovanes Heghyan (Phys)		6 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Nerses Gevorgyan (EE)		40 hrs	0	0	\$8,885.60	\$11,107.00	\$0.00	\$0.00	\$11,107.00	\$0.00	\$0.00	\$0.00	\$0.00	\$11,107.00
123	1.6.13	1.6.13 Archiving of controls		102 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Heghyan (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		EE Accelerator JLAB		62 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00

ID	WBS	Task Name	Type	Work	# Spares	Prototype	Labor	Ltotal	Material	Mtotal	Total	Spares	Prototypes	Total Operations	Total Infrastructures	Total Capital Equipments
124	1.7	1.7 Installation & Commissioning		728 hrs	0	0	\$45,244.80	\$57,936.96	\$0.00	\$0.00	\$57,936.96	\$0.00	\$0.00	\$54,999.36	\$0.00	\$2,937.60
125	1.7.1	1.7.1 SVT installation		48 hrs	0	0	\$2,937.60	\$2,937.60	\$0.00	\$0.00	\$2,937.60	\$0.00	\$0.00	\$0.00	\$0.00	\$2,937.60
		Marco Oriano (ME)		16 hrs	0	0	\$2,937.60	\$2,937.60	\$0.00	\$0.00	\$2,937.60	\$0.00	\$0.00	\$0.00	\$0.00	\$2,937.60
		Tim Neilson (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		16 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
126	1.7.2	1.7.2 SVT Commissioning	OP	120 hrs	0	0	\$6,670.80	\$6,672.04	\$0.00	\$0.00	\$6,672.04	\$0.00	\$0.00	\$8,672.04	\$0.00	\$0.00
		Ryan Herbst (EE)	OP	40 hrs	0	0	\$6,670.80	\$6,672.04	\$0.00	\$0.00	\$6,672.04	\$0.00	\$0.00	\$8,672.04	\$0.00	\$0.00
		Tim Neilson (Phys)	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
127	1.7.3	1.7.3 ECAL Commissioning	OP	120 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 2 ECAL JLAB	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		MT Orsay	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		ET Orsay	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
128	1.7.4	1.7.4 TDAQ Commissioning	OP	320 hrs	0	0	\$38,204.80	\$36,666.24	\$0.00	\$0.00	\$36,666.24	\$0.00	\$0.00	\$36,666.24	\$0.00	\$0.00
		Ryan Herbst (EE)	OP	80 hrs	0	0	\$13,341.60	\$17,344.08	\$0.00	\$0.00	\$17,344.08	\$0.00	\$0.00	\$17,344.08	\$0.00	\$0.00
		EE Hall-B JLAB	OP	80 hrs	0	0	\$9,885.60	\$11,551.28	\$0.00	\$0.00	\$11,551.28	\$0.00	\$0.00	\$11,551.28	\$0.00	\$0.00
		ET Hall-B JLAB	OP	80 hrs	0	0	\$5,977.60	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00	\$7,770.88	\$0.00	\$0.00
		Phy 1 JD	OP	80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
129	1.7.5	1.7.5 Slow Control Commissioning	OP	120 hrs	0	0	\$7,431.60	\$9,661.08	\$0.00	\$0.00	\$9,661.08	\$0.00	\$0.00	\$9,661.08	\$0.00	\$0.00
		EE Hall-B JLAB	OP	40 hrs	0	0	\$4,442.80	\$5,775.64	\$0.00	\$0.00	\$5,775.64	\$0.00	\$0.00	\$5,775.64	\$0.00	\$0.00
		ET Hall-B JLAB	OP	40 hrs	0	0	\$2,988.80	\$3,885.44	\$0.00	\$0.00	\$3,885.44	\$0.00	\$0.00	\$3,885.44	\$0.00	\$0.00
		Hovanes Hagiyan (Phys)	OP	40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
130	1.7.6	1.7.6 HPS ready for the beam		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
131	1.8	1.8 Electron Running		4,160 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
132	1.8.1	1.8.1 Beamline Commissioning (2014)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ken Mofeit (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Stepanyan (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 2 ECAL JLAB		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Tim Neilson (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 1 JM		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Hagiyan (Phys)		40 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
133	1.8.2	1.8.2 Detector Commissioning Run (2014)		640 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ken Mofeit (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Stepanyan (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 2 ECAL JLAB		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Tim Neilson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 1 JM		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Hagiyan (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
134	1.8.3	1.8.3 Data Run 2.2GeV (2014)		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
135	1.8.4	1.8.4 Data Run 1.1 GeV (2014)		0 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
136	1.8.5	1.8.5 Beamline recommissioning (2015)		640 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ken Mofeit (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Stepanyan (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 2 ECAL JLAB		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Tim Neilson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 1 JM		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Hagiyan (Phys)		80 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
137	1.8.6	1.8.6 Data Run (2015)		2,560 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Ken Mofeit (Phys)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Stepan Stepanyan (Phys)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 2 ECAL JLAB		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Tim Neilson (Phys)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Per Hansson (Phys)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Phy 1 JM		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00
		Hovanes Hagiyan (Phys)		320 hrs	0	0	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00	\$0.00







## B Additions and improvements to the HPS setup using non-DOE sources of funding

While the proposed baseline equipment will be sufficient to carry out the proposed measurements, the HPS collaboration is seeking funding from non-DOE sources to improve and enhance capabilities of the HPS detector. Plans for improvements include a light monitoring system, new preamplifiers, and large area APDs for the ECal. The HPS collaborators from IPN Orsay have applied for a number of European grants, including European Research Council (ERC) Advanced Grant 2013 to purchase APDs, for manpower costs to replace the old ones, to design and build new preamplifier boards, and to assemble and test the ECal with the new modules. The total cost of replacing all ECal APDs is about 500K\$. These grants also include the light monitoring system. If successful, European funding will cover most of the ECal modifications, and will significantly reduce the support needed from DOE for the more modest upgrades we have proposed above.

Besides ECal improvements, the collaboration intends to add a muon detector to the HPS setup. Collaborators from the College of William&Mary (PI Prof. Keith Griffioen) together with collaborators from Rutgers University (PI Prof. Yuri Gerstein) and Old Dominion University (PI Prof. Lary Weinstein) submitted an MRI proposal to NSF for the Muon System, requesting  $\sim 300k$ .

The details of ECal improvements, and the motivation and description of the muon system are presented below.

### B.1 Improvements to ECal

#### 1. Light monitoring system

For an experiment like HPS, where backgrounds must be well understood and need to be strongly suppressed, the trigger bias must be minimized. In particular, having stable, known, and uniform thresholds at the trigger readout is necessary in order to avoid bias in the event selection. Such uniformity and stability can be achieved with the installation of an online gain monitoring system. This system will introduce short light pulses into the front face of the crystals. The crystals already have fiber

holders attached, allowing implementation of this system without having to modify the crystals or wrapping.

Optical fibers will be used to transmit light from a calibration source to the crystals to test the response of the APDs. The response of the system could change in time because of losses in crystal transparency due to radiation damage or because of gain variations of the APDs. Such a calibration system has been developed for several experiments (CMS at CERN for instance) with various light sources. The system for the ECal will be developed at IPN Orsay during 2013 and in the first half of 2014, and will be ready for installation at JLAB for the commissioning run in the fall of 2014. Each module will have a red and blue mono-color LED light source for monitoring purposes. Blue light transmission, corresponding to the domain of the crystal's emission spectrum, is very sensitive to the presence of color centers which are produced by radiation damage. So the blue light source will monitor variations of the response in the main domain of the spectrum. The response to red light is less sensitive to the color centers, and so permits monitoring the APD gains more directly. Thus the use of two colors separates gain variations due to the APD and electronics from those due to changes in the crystal transparency, and provides clear information on the state of the electronics. The main challenge for the system is to guarantee stability at a level of 2%. The test of the system will be carried-out at IPN Orsay, in order to study its efficacy and also to test the radiation resistance of the various materials.

2. **Modifications to the side brackets to accommodate fiber bundles for the light monitoring system** - A light monitoring system was not used during the test run. While the design of the ECal enclosure was done in such a way that it can accommodate optical fibers attached to the front face of crystals, the side plates that hold the crystal frames do not have inlets for the accompanying fiber bundles. Space is available on the side plates for a straight-forward modification which will allow the addition of a light monitoring system.
3. **New low-voltage power supply** - The existing low voltage power supply is a manually controlled, single output power supply that feeds the four ECal motherboards through a custom-made patch panel. The present system cannot control the voltage

supplied to preamplifiers at different parts of the ECal, and controlling or resetting them remotely has proved to be very inconvenient, requiring frequent access to the Hall, especially during commissioning. Newly available low voltage power supplies are much more flexible. The one that is the most suitable for the ECal APD preamplifiers is the WIENER MPV 8008LD. This power supply is being used at JLAB and the control software exists, so it will be easy to incorporate it into HPS.

4. **New preamplifiers** - A low noise, low gain preamplifiers will be needed to take advantage of increased signal on the input of FADC after removing the splitter. The impact of the lower noise/threshold system is twofold: first it will improve the ECal's energy resolution; and second it will make the ECal sensitive to minimum ionizing particles which pass through the crystals transversely. With sensitivity to cosmic ray muons, which will pass through the ECal transversely when it is installed in HPS, the Ecal crystals can be calibrated for MIPS, and their effective gains balanced. HPS collaborators from INFN Genova have shown that with such a low noise, low threshold system, the ECal can distinguish the MIP energy deposition from noise, see left plot on Figure 79.
5. **Possibilities with new APDs** - Installing new APDs on the existing crystals will significantly improve the ECal performance, but doing so is expensive, so replacement is only being considered if a funding source beyond DOE HEP is secured. Replacing the old  $5 \times 5 \text{ mm}^2$  Hamamatsu S8664-55 APDs with  $10 \times 10 \text{ mm}^2$ , Hamamatsu S8664-1010 will improve two critical characteristics of the present calorimeter modules. First, the new APDs from Hamamatsu have much better performance than the ones which are currently installed. Data from Hamamatsu shows that APDs made from the same wafer have excellent gain uniformity. With  $\pm 10\%$  known uniformity at the gain of 100, the required variations in bias voltage are only  $\sim 4.5 \text{ V}$ . Even for large samples of APDs ( 1300), the required bias voltage differences are  $\sim 50 \text{ V}$ , which is half that of the current APDS. The ECal supplies bias voltage to groups of APDs, so with new APDs, with their smaller voltage-gain variations, it will be possible to achieve much better uniformity in the response of the calorimeter modules, and consequently better uniformity in trigger thresholds.

Secondly, the new APDs have a 4 times larger readout area, ensuring 4 times more

light collection and therefore 4 times larger signals. This will allow the use of lower gain amplifier modules which in turn will decrease electronic noise. Tests performed for another calorimeter, now in production at INFN Genova for JLAB Hall-B, showed that amplifier boards with a factor 2 lower gains have a noise level  $< 5$  MeV. The energy deposition in the HPS  $\text{PbWO}_4$  crystals from minimum ionizing cosmic muons passing transversely to the crystal axis is  $\sim 15$  MeV. Moving the energy thresholds close to 5 MeV will allow the MIP peak to be clearly distinguished, and will let the calorimeter be calibrated and monitored with cosmic muons. The HPS collaborators from the INFN group have performed the first tests of the Hamamatsu  $10 \times 10$  mm<sup>2</sup> APDs and a new amplifier board on HPS crystals. In Figure 79, the charge distribution of a single crystal system is shown for  $5 \times 5$  mm<sup>2</sup> (left) and  $10 \times 10$  mm<sup>2</sup> (right) APDs. A coincidence signal from scintillator telescopes positioned above and below the module provides the trigger. The crystal was positioned horizontally as it would be in HPS, so the cosmic muons would pass through it perpendicular to the crystal axis. The red line histogram is for all events triggered by the scintillation telescope and corresponds to the noise. The black line histogram corresponds to the charge detected within 100 ns of the trigger time. The MIP peak is clearly visible and well isolated from the noise for the S8664-1010 APD readout. For the S8664-55 APD, the MIP signal is also seen, but its charge distribution is under the noise peak and it does not have well defined peak position. Using MIP calibration in conjunction with the light monitoring system will ensure stable and reliable performance of the ECal and the trigger system. As a bonus, the lower noise will allow the use of lower thresholds and improve the calorimeter's energy resolution. The new amplifier boards have to be designed to work with new APDs.

## B.2 Muon system

The muon detector will match the geometrical acceptances of the tracker and ECal, and will be about a cubic meter in size. With such geometrical coverage, the efficiency of detecting high mass  $A'$ s in the  $\mu^+\mu^-$  decay channel will be higher than for  $e^+e^-$  decays, see Figure 80. The invariant mass resolution with di-muons is also improved, see Figure 71. The di-muon decay channel of the  $A'$  has the advantage of having greatly reduced electromagnetic

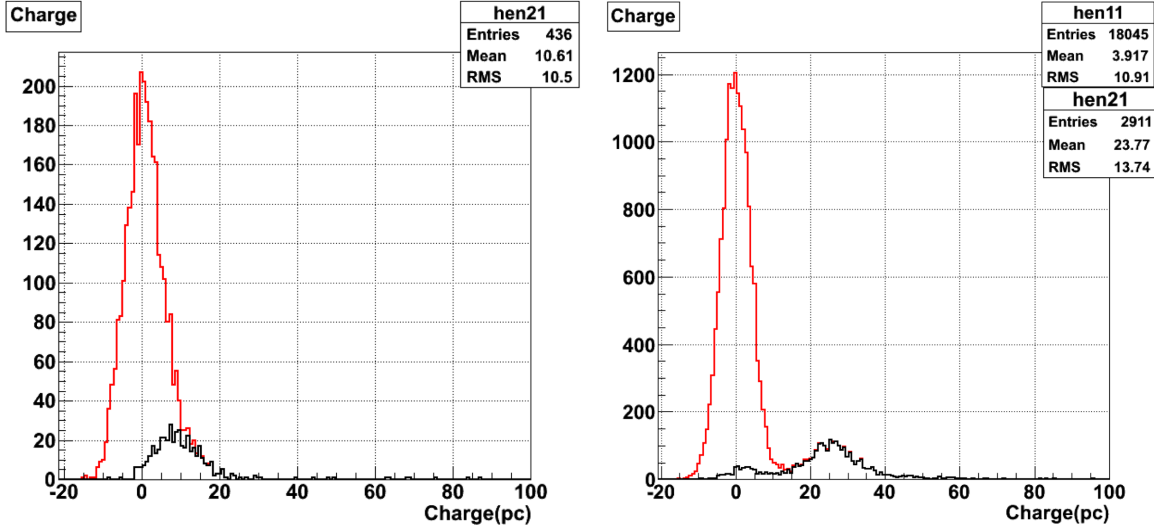


FIG. 79: Charge distribution from readout of the HPS calorimeter crystal with Hamamatsu S8664-55 (left) and S8664-1010 (right) APDs, and the new low noise amplifier board. The red line histogram corresponds to the charge distribution for all triggers coming from the scintillators positioned above and below the crystal. The black line shows the distribution for hits in the crystal within 100 ns of the trigger signal.

backgrounds. In this case, the only particle background in a muon counter would come from the photoproduction of  $\pi^+$  and  $\pi^-$  pairs that are not fully stopped in the ECal or absorber. The expected low background and the high detection efficiency makes the di-muon final state an attractive complement to the  $e^+e^-$  final state. It will add substantial territory in the mass and coupling parameter space as show in Figure 81 and will offer valuable cross-checks. With the addition of a muon system, HPS will be the only fixed target experiment proposed to date to search for heavy photons in an alternative to the  $e^+e^-$  decay mode.

The muon system can easily be constructed with layers of scintillator hodoscopes sandwiched between iron absorbers, and can easily be added downstream of the rest of the HPS apparatus. The number of layers and the thickness of absorbers is defined by the  $\pi/\mu$  rejection factor. The schematic design of the muon detector was optimized using the GEANT-3 model for the ECal with added layers of iron and scintillators. In the simulation, muons and pions in the momentum range of 1 to 4 GeV/c first passed through the 16 cm of lead tungstate in the ECal and then entered a muon counter with various total absorber thicknesses (see [5] for details). Detection efficiencies for pions ( $\epsilon_\pi$ ) and muons ( $\epsilon_\mu$ ) were then calculated as a function of absorber thickness and particle momentum. For low-energy particles ( $< 1.7$  GeV) detection in all four layers of scintillator hodoscopes was not considered.

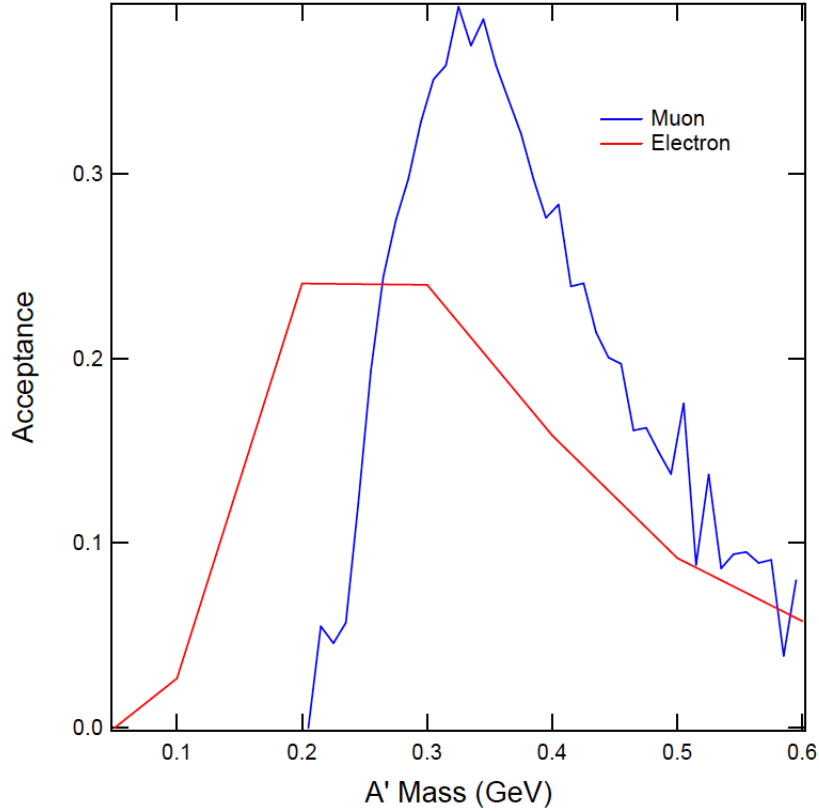


FIG. 80:  $A'$  detection efficiency through  $\mu^+\mu^-$  (blue) and  $e^+e^-$  (red) decay channels as a function of mass for 6.6 GeV beam energy.

Depending on the momentum, particles were not traced behind the third, fourth or fifth absorber. Figure 82 shows the resulting rejection factor  $\epsilon_\pi/\epsilon_\mu$ . The right-hand plot shows the dependence of  $\epsilon_\pi/\epsilon_\mu$  on the total thickness of the iron absorber, with the best rejection at about 75 cm. The right-hand plot shows  $\epsilon_\pi/\epsilon_\mu$  for a 75 cm absorber as a function of muon momentum. The suppression of individual pions by two orders of magnitude will suppress pion pairs by 4 orders of magnitude.

### B.2.1 Conceptual Design

On the basis of these simulations, we have designed a muon detector composed of four iron absorbers (total length of  $30 + 15 + 15 + 15 = 75$  cm) with a double-layer scintillator hodoscope positioned after each absorber. The muon detector will be mounted behind the ECal. The front face of the first absorber will be at  $\sim 180$  cm from the target. Similar to the Ecal, the muon detector will consist of two halves, one above and one below the beam plane.

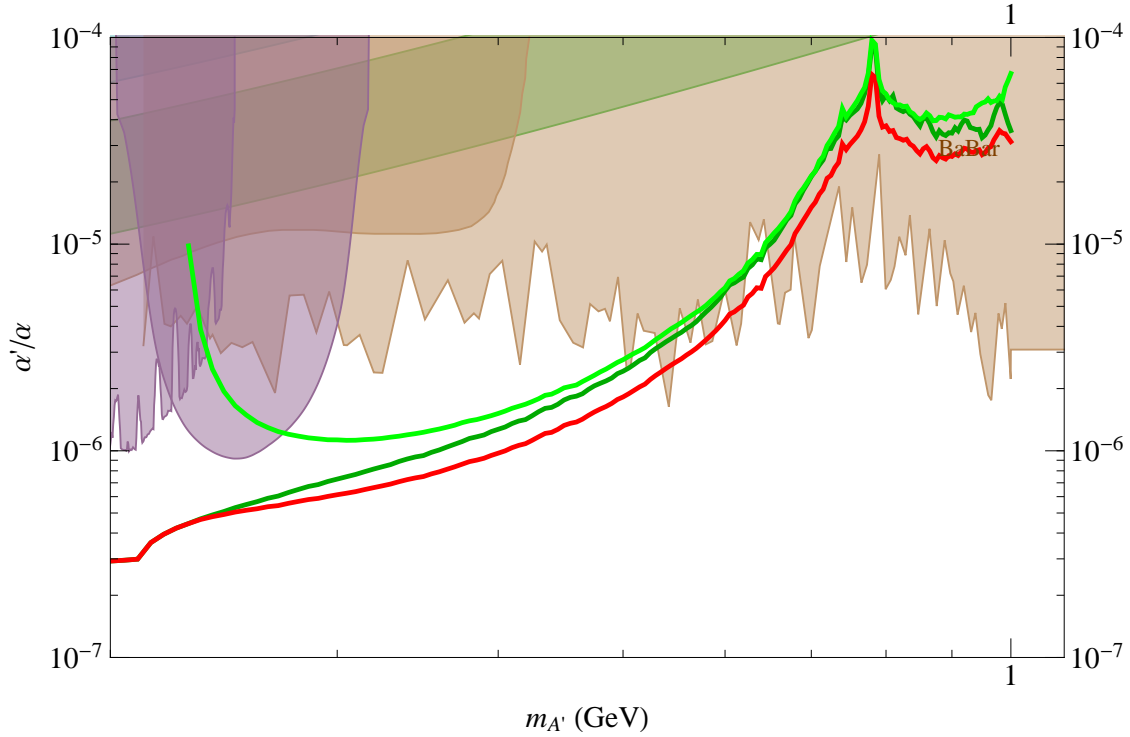


FIG. 81: Experimental reach for 2 weeks of beam time at 6.6 GeV for  $\mu^+\mu^-$  (light green) and  $e^+e^-$  (dark green) decay channels. The red line is the combined reach.

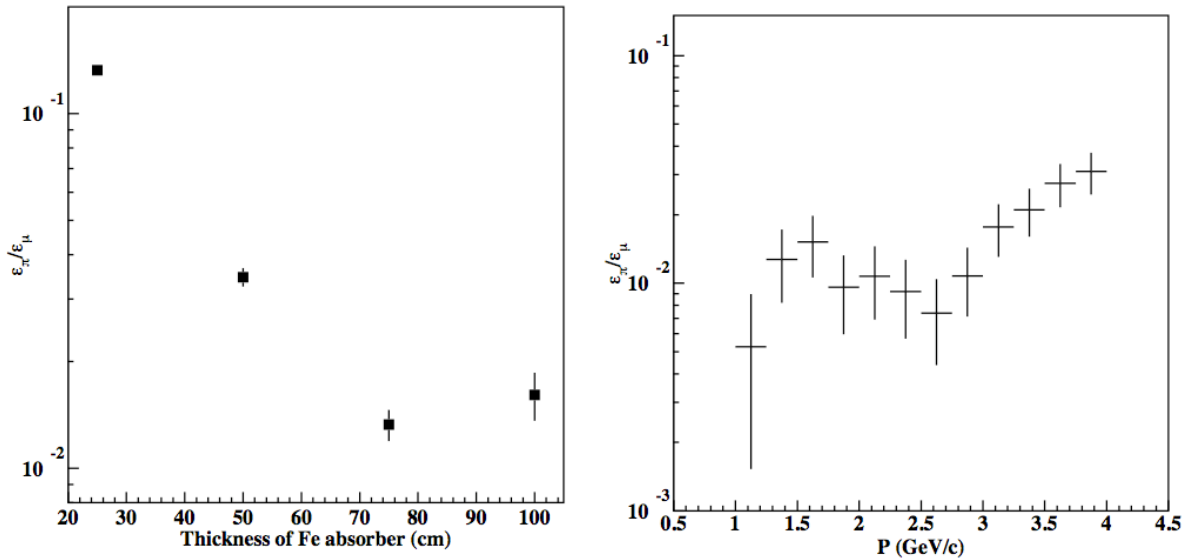


FIG. 82: Pion-muon rejection factor  $\epsilon_\pi/\epsilon_\mu$  versus total iron absorber thickness (left) and versus particle momentum for a 75 cm absorber (right).

This segmentation is necessary in order to minimize the effects of the “sheet-of-flame,” the multitude of low-energy particles in the horizontal plane, swept into the detector acceptance by the dipole analyzing magnet.

The dimensions of the hodoscopes and absorbers are defined using simulations of  $A' \rightarrow \mu^+\mu^-$ . In Figure 83 the points where muon pairs from  $A'$  decays of different masses intercept the  $(XY)$  plane located at 210 cm from the target are shown. Both muons are required to be detected in the ECal. Overall dimensions of the hodoscopes and absorbers are shown in Table XXVII. Figure 84 shows a CAD drawing of the HPS detector with the muon system on the right, which includes the 4 absorbers (gray), the vacuum box (light gray) between the upper and lower sections, and the final set of scintillator paddles (red). The ECal is directly upstream from the muon detector, with its crystals shown in yellow. In front of the ECal is a large gray vacuum flange. The silicon tracker is represented by red and gray rectangles and the red point on the left is the target position. The vertical gap between the first hodoscope layers of the two halves is about 5 cm and increases to about 7 cm for the last layer of hodoscopes.

TABLE XXVII: Dimensions (in cm) of the muon system scintillation hodoscopes (H) and iron absorbers (A).

	H1	H2	H3	H4
Distance from target	212	232	252	272
Width	112	125	138.5	152
Height	10.5	11.5	12.5	13.5
	A1	A2	A3	A4
Distance from target	207	227	247	267
Width	108.5	122	135	148.5
Height	10	11	12	13
Thickness	30	15	15	15

For the hodoscopes we plan to use the same extruded scintillator strips with embedded wavelength-shifting fiber and phototube readout that was developed for the CLAS12 Preshower Calorimeter. These scintillator strips are 45 mm x 10 mm in cross section, and can be cut to any length. Widths can be reduced as needed for the muon counter. Each strip contains two long tunnels, created in the original extrusion process, into which wavelength shifting fibers can be inserted. Each hodoscope will consist of one  $X$  and one  $Y$  plane. The vertical strips of the last hodoscope are shown in two colors in Figure 84. The horizontal counters of the last hodoscope plane is shown in Figure 85. The horizontally aligned strips will extend over half the width of the detector and will be read out from the outer ends. The upper and lower hodoscopes in each plane will have their own vertically aligned strips,



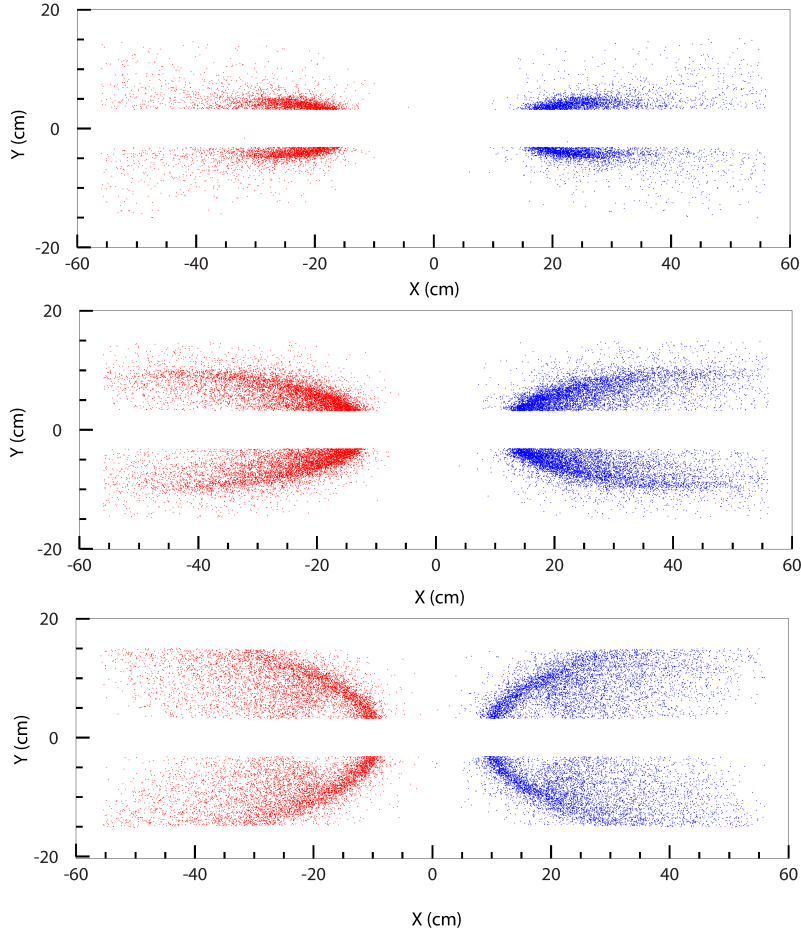


FIG. 83: Points where  $\mu^+$  (red) and  $\mu^-$  (blue) pairs from  $A'$  decays intercept the  $(XY)$  plane located 210 cm from the target for  $A'$  masses 250 MeV (top), 300 MeV (middle), and 400 MeV (bottom).

which will be read out only on their outer ends. The inner end is inaccessible because of the vacuum box, and there is no particular advantage to having a double readout on these short (135 mm) strips.

The system will be instrumented with less than 256 readout channels so that the requisite electronics will fit into a single VME/VXS crate. The signals from each channel (PMT) will be sent to a FADC. We intend to borrow the CLAS12 Preshower Calorimeter electronics and HV system. Just as with the ECal, the FADCs will be used to construct a muon trigger for the experiment. In the current design there will be 3 pairs (left-right) of horizontal strips in each of 8 hodoscope planes (48 total) and a total of 208 vertical strips in 8 hodoscope planes. The number of vertical strips per plane increases slightly with distance from the target to keep a constant angular coverage. The maximum is 33 per hodoscope in the back

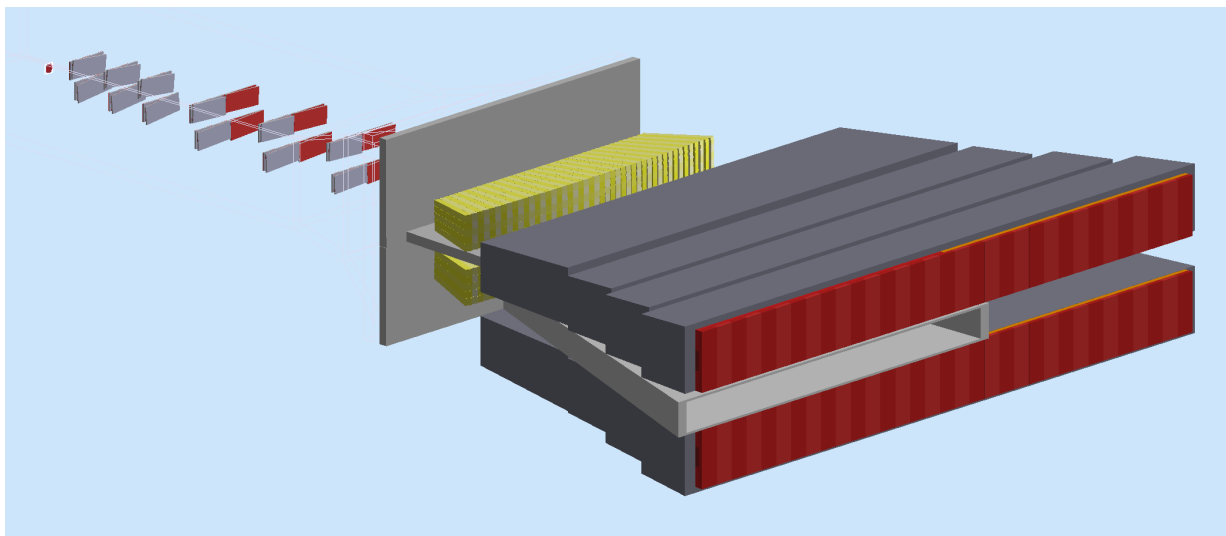


FIG. 84: CAD drawing of the HPS detector setup. From left to right this consists of the target (red dot), the silicon tracker (gray and red rectangles), the large shielding wall (gray), the ECal lead tungstate crystals (yellow, two shades), the muon counter absorbers (gray), and the final muon counter scintillators (red, two shades).

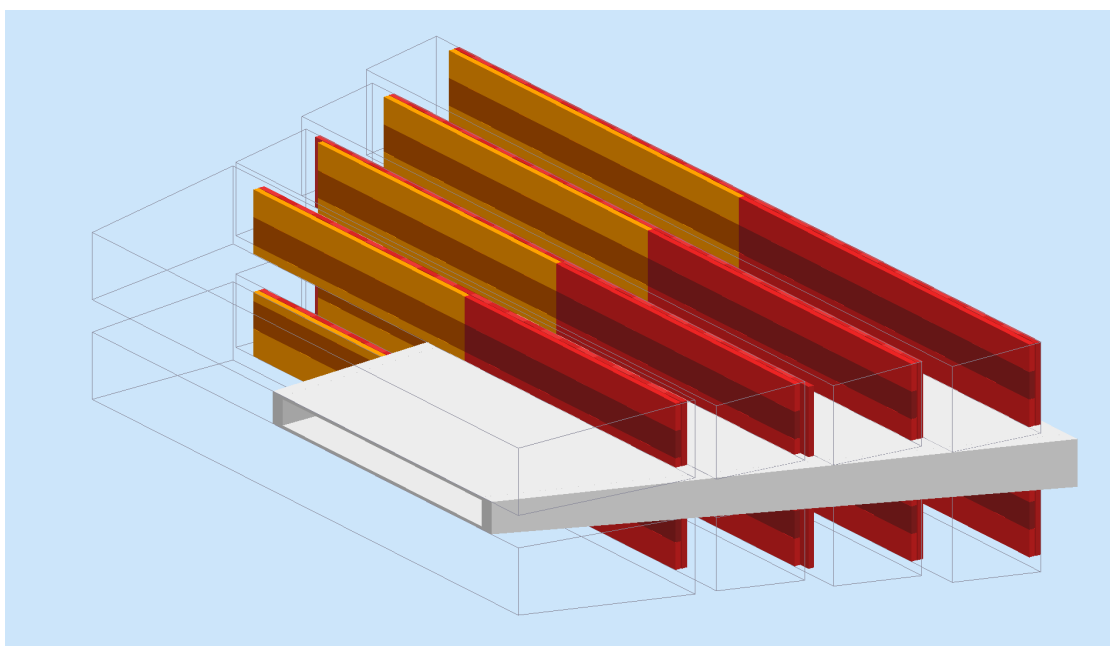


FIG. 85: The horizontal scintillator configuration for the muon counter. Scintillators are shown in red and yellow/brown. The white/gray structure is the vacuum box. Each hodoscope layer (top and bottom) contains six strips (three left and three right), read out from the outer ends.

plane.

Full Monte Carlo simulations with realistic event rates are currently under study in order to finalize the design details of the muon counter. The crucial issues are the event rates in

the scintillators near the beamline (which already has initiated a redesign of the vacuum chamber to reduce background), the target-to-muon-counter tracking resolution and the detection efficiency, and the achievable trigger rates. Any changes to the detector as a result of these studies are expected to be minor and will not alter the conceptual design presented here.

### B.2.2 $\mu^+\mu^-$ Trigger

The muon trigger will look for  $\mu^+\mu^-$  pairs by finding energy depositions consistent with those expected from minimum ionizing particles in the layers of the muon system. The trigger algorithm in the CTP of the muon system VXS crate will produce a muon pair trigger in four steps:

- search for MIP hits using energy selections on the hits reported by the FADCs which satisfy  $E_{MIP}^{thr} < E_{\mu.hodo}$
- use the time information of the reported MIP hits to select coincidences between the two planes of each hodoscope layer, quadrant by quadrant.
- look for coincidences in successive quadrants of at least the first three layers of the muon hodoscopes
- select pairs of triple coincidences in opposite quadrants of the detector and report the times and positions of coincident triple hits to the SSP

If it is necessary to reduce the rate further, the SSP can in addition look for time and position coincidences of MIP hits in the ECal, defined in the ECal crate CTPs as hits with 1 or 2 crystals and energy within predefined thresholds:  $E_{MIP}^{min} < E_{ECal.channel} < E_{MIP}^{max}$ . The SSP will send the final decision regarding the  $\mu^+\mu^-$  trigger to the Trigger Supervisor board.

### B.2.3 Muon system trigger rates

Like the ECal, the muon system trigger rates are dominated by beam backgrounds. A GEANT4 model of the HPS detector was used to estimate the rates, following the conceptual design for the muon system presented in Section B.2. Figure 84 shows the layout of the

system. Each of eight hodoscope layers (four layers in the top part of the detector and four in the bottom) consists of two planes of scintillator strips. One plane, called the Y-plane, has strips oriented horizontally. The other, called the X-plane, has them oriented vertically. The Y-strips are segmented exactly in the middle and outside ends, left or right, are read out. Six Y-strips make up each half plane, so the total number of Y-strips is 48. X-planes are divided into 4.5 cm wide segments, with 240 strips in total. It should be noted that the number of vertical strips in the conceptual design is only 208 to limit the total number of readout channels to 256, one crate’s worth. Since the rates in the vertical strips at the edges of the hodoscope are very low, eight strips in each plane can be paired to make four readout channels without negatively impacting the detector occupancy. In Table XXVIII, the lengths and widths of the detector readout segments (counters) used in the simulation are presented. There is a 14 cm wide and 3.5 cm high gap introduced into the model, centered on the point where the electron beam passes through the muon vacuum chamber, to avoid the high rate region. As shown in Figure 83 this gap has a negligible effect on the detection efficiency of muon pairs from  $A'$  decays.

TABLE XXVIII: Lengths and widths of the hodoscope strips. Dimensions are centimeters.

Readout plane	Layer 1	Layer 2	Layer 3	Layer 4
X-plane width	4.5	4.5	4.5	4.5
X-plane length	10.5	11.5	12.5	13.5
Y-plane width	3.5 all three	3.5, 4, 4	3.5, 4.5, 4.5	4.5 all three
Y-plane length	56	62.5	70	76

Events generated in EGS5 were used as an input to the GEANT4 simulation. The CEBAF beam bunch structure was simulated by sending one bunch equivalent of electrons, 5,625  $e^-$ ’s (6.6 GeV), through the target to generate secondaries and scattered beam particles. The secondaries were followed through the apparatus to simulate the detector response. As expected, the highest background rates are seen in the Layer 1 hodoscope and are  $\sim 0.7$  MHz in both the X-strips near the electron beam location and the beam-left Y-strip closest to the beam plane, see Figure 86. Rates in the vertical strips far from the beam position are very low, allowing multiple strips to be combined into a single readout channel in order to reduce the number of PMTs and electronic channels.

The coincidence rates between hodoscope planes in a given layer and between different

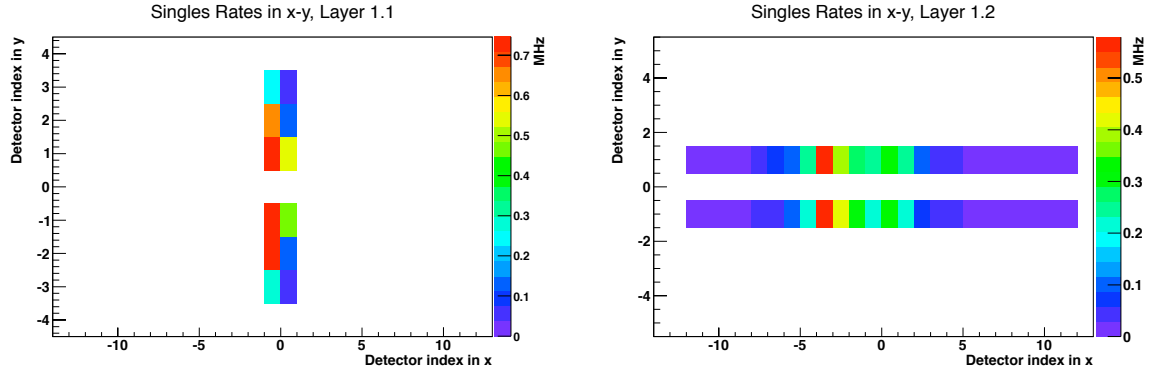


FIG. 86: Rates in Y- (left) and X-planes (right) of the Layer 1 hodoscopes for hits with energy deposition  $> 0.5$  MeV.

layers have been studied using a 16 ns coincidence time window. On the left of Figure 87, the coincidence rates between X- and Y-quadrants, top-left (TL), top-right (TR), bottom-left (BL), and bottom-right (BR) of the Layer 1 hodoscope are shown. On the right, the figure shows coincidence rates of respective quadrants of Layers 1 and 2. The fact that there is a significant reduction of the rates from 2-plane ( 1.2 MHz) to 2-layers (0.07MHz) coincidences indicates that hits are mostly from uncorrelated background. For the muon trigger, a coincidence of two opposite quadrants (TLxBR) or (TRxBL) is required along with triple coincidences of the first three layers of hodoscopes in each quadrant. The rates of the triple quadrant coincidences within 16 ns are shown in Figure 88. The maximum trigger rates are in the beam-right (electron side) quadrants and are on order of 7 kHz. In the beam-left quadrants (positron side), the tripple coincidence rates are  $< 1$  kHz. Since an overall trigger requires hits in two opposite quadrants, the maximum rate will be  $< 1$  kHz. While a further reduction of rates will be possible with the inclusion of MIP hits in the ECal, the  $< 1$  kHz is a small addition to the total trigger rate from the Ecal trigger (see discussions in Section 6.2) and will keep overall trigger rate well within the limit of allowed rates for the HPS DAQ.

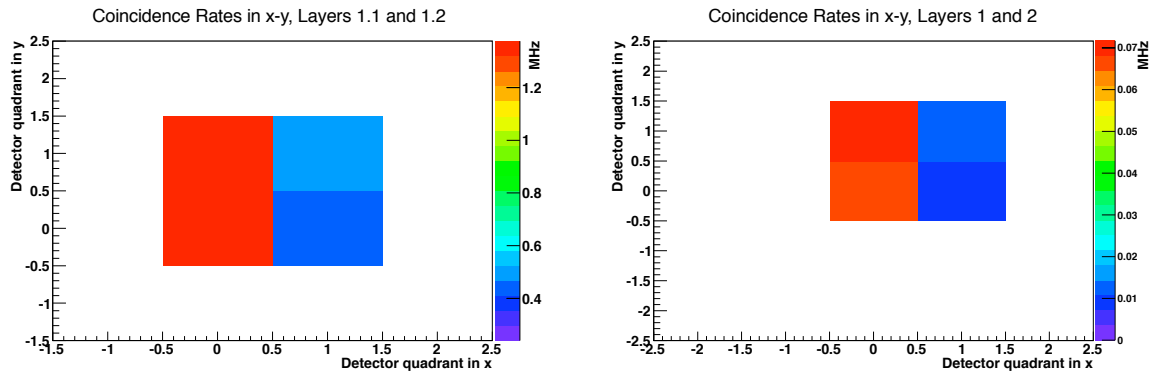


FIG. 87: Coincidence rates between X- and Y-quadrants of the Layer 1 hodoscope (left graph) and coincidence rates between Layer 1 and 2 (right graph).

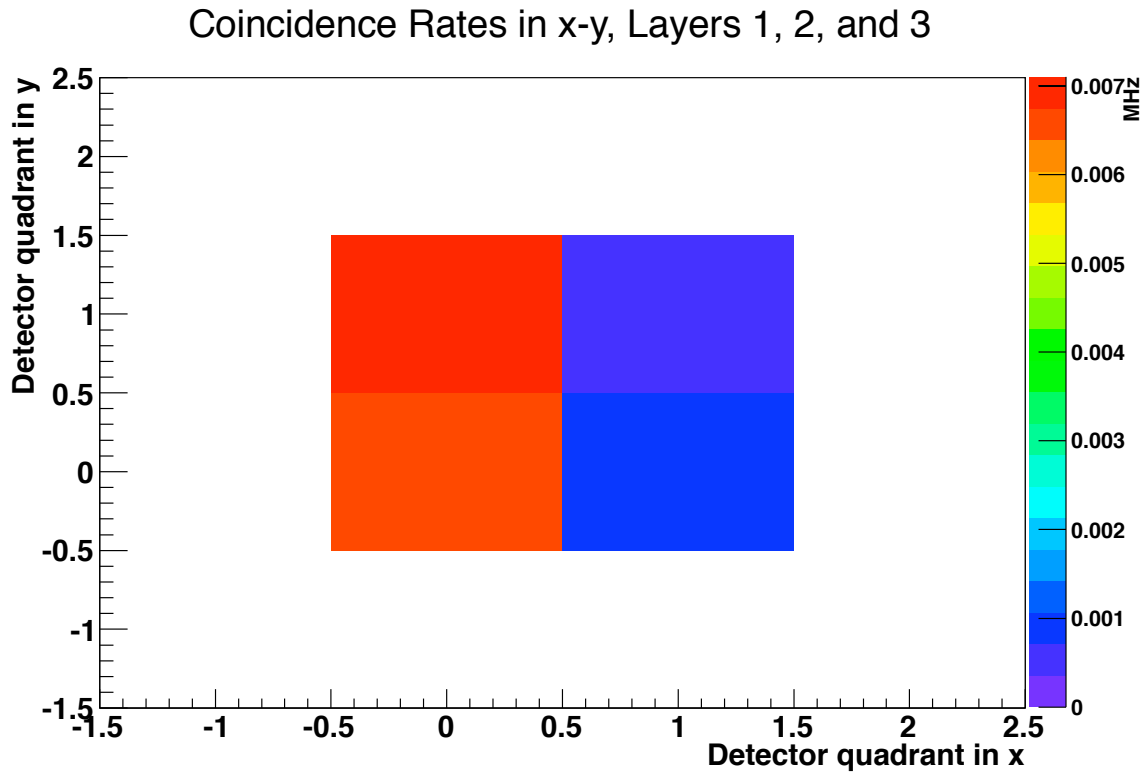


FIG. 88: Coincidence rates in first three layers of muon hodoscopes. The coincidence time window is set to 16 ns.

## C Simulation Tools

The simulation tools play a critical role in simulating the background environment, optimizing the detector setup, and developing the trigger and reconstruction strategies. We use GEANT4 and EGS5 to simulate electromagnetic interactions. There is generally good agreement between these two codes. In particular, no inconsistencies have been found on secondary particle yields or energy spectra. However, we have found significant disagreements on the angular distributions in the multiple scattering, bremsstrahlung and pair production processes.

### Multiple Scattering Simulation

EGS5 simulates the electron elastic scattering using the Molière theory [110] as formulated by Bethe. [111] It is based on a small angle approximation ( $\theta \ll 1$  radian), and the angular distribution approaches asymptotically to Gaussian at small angles, and to Rutherford's Coulomb scattering function at large angles given by,

$$F(\theta) \sim \frac{1}{\left(1 - \cos\theta + \frac{\chi^2}{2}\right)^2}. \quad (\text{C1})$$

Instead of using the complex and time consuming Molière's formula, GEANT4 uses two functions explicitly, Gaussian at small angles and the Rutherford function Eq. C1 at large angles with a requirement that these two functions and their first derivatives are joined continuously. GEANT4, however, uses a different power in the denominator in Eq. C1 which is close to 2 but not exactly equal to 2 and is dependent on the target material and thickness.

Several comparisons have been made in the angular distribution  $F(\theta)$  in the differential cross section  $d\sigma = F(\theta)d(\cos\theta)d\phi$  for 2.2 GeV electron scattering from 0.125%  $X_0$  Tungsten target. The EGS5 simulation is compared with Molière's analytical formula in Figure 89(a), demonstrating a good agreement between EGS5 and the Molière theory. While the Molière theory is based on a small angle approximation, the multiple scattering theory developed by Gaudsmit and Saunderson is valid for any angle by means of an expansion in Legendre

polynomials. [112] The validity of the small angle approximation is checked by comparing the Molière integral with the Goudsmit-Saunders theory as shown in Figure 89(b), demonstrating that the Molière theory is accurate in the angular region of the HPS detector.

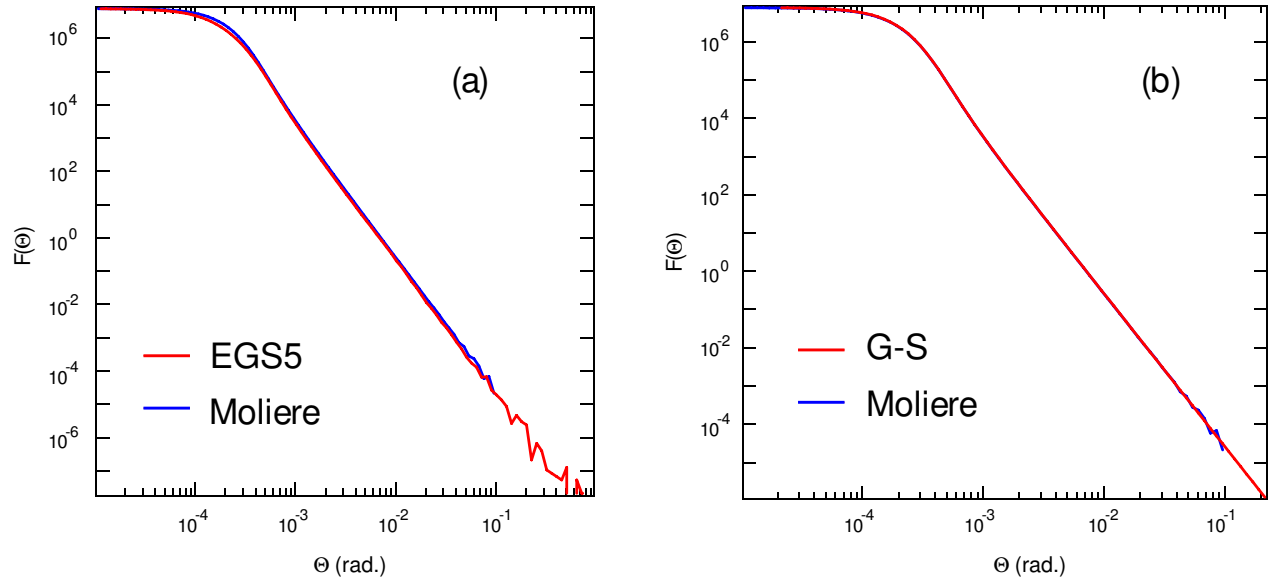


FIG. 89: (a) Molière vs. EGS5 (b) Molière vs. Goudsmit-Saunders

Figure 90 shows the angular distribution comparison between the GEANT4 simulation and the Molière integral. GEANT4 is in good agreement with the Molière integral up to about 1 mrad, then it deviates at larger angles, predicting roughly twice the cross section at 15 mrad, where the HPS tracker sensor edge is located.

D. Attwood et al. measured 170 MeV muon angular distributions and compared with GEANT4 simulations and the Molière theory. [113] They concluded that GEANT4 simulation over-estimated the scattering tail by about a factor of two, and the data were consistent with the Molière theory. G. Shen et al. [114] and B. Gottschalk et al. [115] also showed that the Molière theory was consistent with the measurements on a wide variety of target materials.

## Angular Distributions

While GEANT4 and EGS5 are in good agreement in the production rates and the secondary particle energy spectra, there are significant differences in the angular distribution in the secondary particles. In EGS5, the angular distributions are sampled from the following



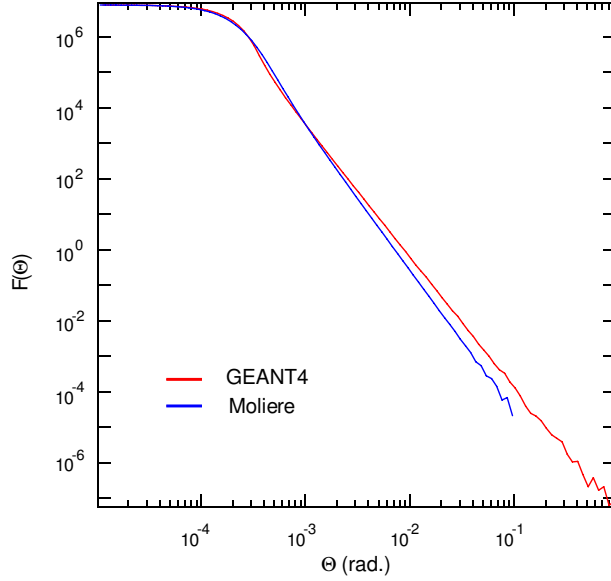


FIG. 90: Molière vs. GEANT4

differential cross section for the bremsstrahlung process, [116]

$$d\sigma(k, \theta_\gamma) = \frac{4Z^2 r_0^2 dk}{137 k} y dy \left\{ \frac{16y^2 E}{(y^2 + 1)^4 E_0} - \frac{(E_0 + E)^2}{(y^2 + 1)^2 E_0^2} + \left\{ \frac{E_0^2 + E^2}{(y^2 + 1)^2 E_0^2} - \frac{4y^2 E}{(y^2 + 1)^4 E_0} \right\} \ln M(y) \right\},$$

where  $k$  is photon energy,  $\theta_\gamma$  is photon polar angle,  $E_0$  and  $E$  are initial and final electron energy, and

$$y = E_0 \theta_\gamma; \frac{1}{M(y)} = \left( \frac{k}{2E_0 E} \right)^2 + \left( \frac{Z^{1/3}}{111(y^2 + 1)} \right)^2,$$

and for the pair production process, [117]

$$\frac{d\sigma}{dE_\pm d\Omega_\pm} = \frac{2\alpha Z^2 r_0^2 E_\pm^2}{\pi k^3} \left\{ -\frac{(E_+ - E_-)^2}{(u^2 + 1)^2} - \frac{16u^2 E_+ E_-}{(u^2 + 1)^4} + \left\{ \frac{E_+^2 + E_-^2}{(u^2 + 1)^2} + \frac{4u^2 E_+ E_-}{(u^2 + 1)^4} \right\} \ln M(u) \right\},$$

where  $k$  photon energy,  $E_\pm$   $e^\pm$  energy,  $\theta_\pm$   $e^\pm$  polar angle, and

$$u = E_\pm \theta_\pm; \frac{1}{M(u)} = \left( \frac{k}{2E_+ E_-} \right)^2 + \left( \frac{Z^{1/3}}{111(u^2 + 1)} \right)^2.$$

GEANT4 uses an approximate function to simulate the angular distributions in the bremsstrahlung and pair production processes given by

$$f(u) = C[ue^{-au} + due^{-3au}],$$

with  $u = E_0\theta_\gamma$  for incident electron energy  $E_0$  and the polar angle  $\theta_\gamma$  of the bremsstrahlung photon, and  $u = E_\pm\theta_\pm$  for the pair energy  $E_\pm$  and polar angle  $\theta_\pm$  in the pair production. Since the production angle is typically  $1/\gamma$ , GEANT4's approximations are acceptable for most of the high energy detector simulations. However, GEANT4 simulations are inconsistent with the data in the following two cases in the HPS Test Run:

- The bremsstrahlung photon angular distribution is too narrow, resulting in too few scatters in the collimator.
- The prediction on the pair angular distribution is too narrow, resulting in too few Ecal trigger rates.

## Simulation Tools Setup in HPS

Because of the inaccuracies in GEANT4 described above the electromagnetic interactions in the target are simulated by EGS5, and all the particles that come out of the target are passed on to the HPS detector simulation system based on GEANT4.

## D Test Run ECal Calibration

The noise and pedestal of the readout chain are calibrated by sampling the preamplifier output out-of-time with the trigger.

We calibrate gain of the individual ECal channels using the SVT measurement of track momentum and comparison to Monte Carlo simulation. We disable all SVT and ECal channels in the simulation that were inoperable or noisy in the test run, so any efficiency or bias effects that affect the real data should be reflected in the simulation as well; then we use a formula to compute the “weighted E/p” for a crystal, representing the average E/p for clusters that include the crystal:  $\frac{\sum_j w_{j,i}}{\sum_j \frac{P_j}{E_j} w_{j,i}}$ , and iteratively adjust the gains until the weighted E/p is equal for test run data and simulation.

These gains can then be used to convert from ADC counts in a channel to the energy

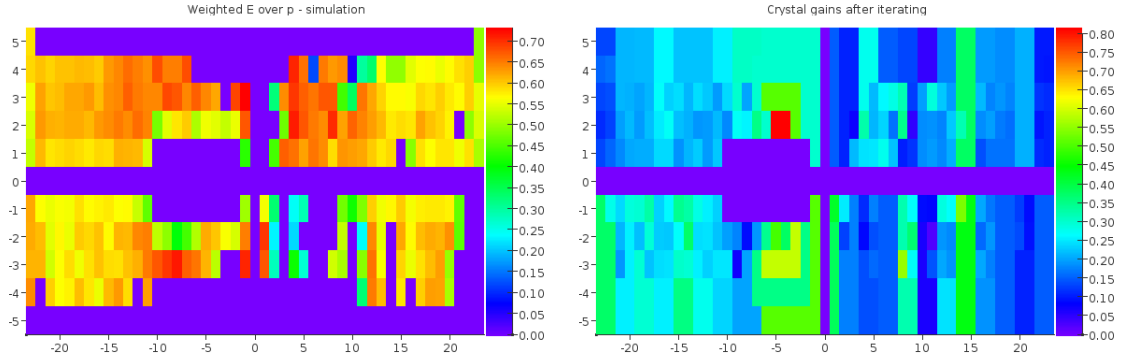


FIG. 91: Weighted  $E/p$  from Monte Carlo simulation (left), calibrated values of gain in units of MeV per ADC count (right).

deposited into that ECal crystal. The other information needed to find the energy of an incident particle is the sampling fraction—the ratio of energy read out from crystals to energy of an incident particle. The conventional sampling fraction—the fraction of incident energy that is deposited in crystals—is approximately 0.9 for our ECal, and less at edges. For our readout, there is additional energy lost because crystals under the readout threshold are not read out. The weighted  $E/p$  used in calibration (see Figure 91) is an approximate measurement of sampling fraction, but the sampling fraction is energy-dependent because of the effect of readout threshold. A full computation of sampling fraction can be done using simulation.

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