



Routine Science Processing of Fermi/LAT Data for Monitoring X-ray Binary Systems

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on behalf of the Fermi Large Area Telescope Collaboration

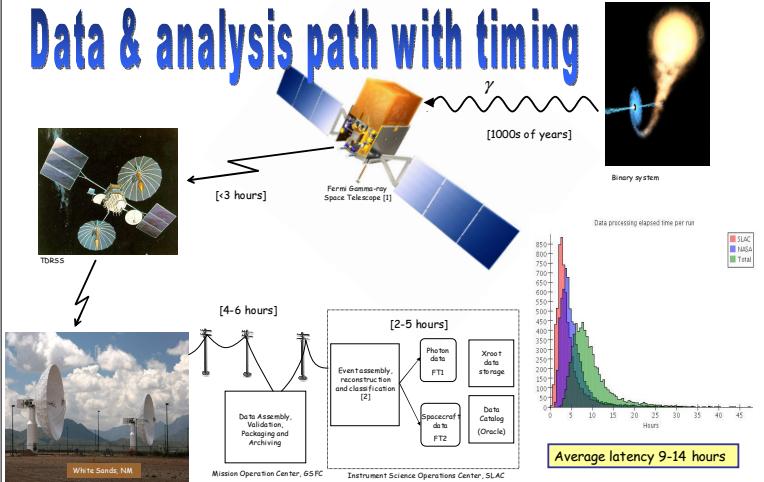


Summary: Part of the FGST routine data analysis includes monitoring of X-ray binary systems. This processing is described with illustrations from the LSI +61 303 system.

Abstract

The Fermi Gamma-ray Space Telescope normally transmits science data to earth every two orbits (~180 minutes); basic event reconstruction is typically complete 24 hours later. An Automated Science Processing (ASP) task then takes a quick look at all data, recording, for example, blazar activity. This is followed by a number of semi-automated Routine Science Processing (RSP) tasks. One such RSP task monitors pulsars. Another RSP task, and the focus of this work, analyzes and monitors X-ray binary systems, including source detection, spectral analysis, and activity trending for each source. The structure and features of this task, plans for future enhancements, and its application to the LSI +61 303 system are described.

Data & analysis path with timing



Routine Science Processing

Objectives:

- Continuously improved detections and fits due to steadily increasing statistics
- Early notification of flares and other activity changes
- Rapid analysis of selected binary systems (leveraging SLAC batch farm to extract ~300 CPU-hours per processing cycle)
- Routine full processing cycles, ~1 per week, automatically and on demand

Analysis Requirements:

- Standard analysis for all sources
- Extensive, custom analysis for better-known sources
- Source detection
- Spectral fitting with flexible source modeling
- Monitor for flaring activity
- Monitor for baseline activity shifts
- Periodicity determination
- Maintain library of web-accessible standard plots
- Maintain tables of web-accessible trending data

Current list of monitored sources

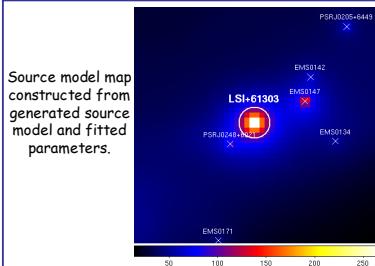
LSI +61 303

Pipeline Processing

(current) code outline

- Setup**
 - Define analysis time period (MET)
 - Extract spacecraft data (FT2)
 - Submit jobs - 1 per source
- mgAnalysis (parallel jobs, one per source)**
 - Extract photon data (FT1)
 - Prepare source model
 - (Basic) analysis**
 - Count maps, live time cubes, exposure maps, etc.
 - Event selection (e.g., ROI & zenith angle cuts)
 - Spectral fitting
 - Aperture photometry (light curve)
- finisup**
 - Register output data in Data Catalog
 - Load trending data
 - Other bookkeeping

- Fermi Pipeline [3,4]:** a system to organize and manage a large automated processing task, including an interface to the SLAC and Lyon batch farms, a comprehensive Data Catalog, self-monitoring, and reporting.
- Processing task driven by python scripts, including general-purpose analysis classes, optimized for automated, parallel pipeline operation.



Source Modeling

- General purpose python classes for source model generation, editing, and xml production for use with maximum likelihood tools
- Candidate Sources from special database table (at right)
 - Choice of spectral model behavior (e.g., power law, broken power law, power law with exponential cutoff, etc.)
 - Ad hoc candidates (not in DB table) may be invented dynamically
- Nearby Sources from LAT catalog
 - Select observed nearby sources, e.g., within desired ROI
 - Initial spectral parameters (from LAT catalog) may be fixed or left floating
 - After spectral fit, parameters may be adjusted, nearby sources removed, etc. for a refit

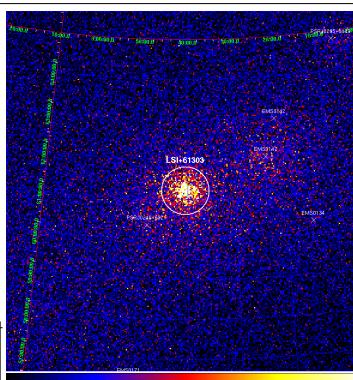
Source	RA	DEC
1A 0858-26	04:47:39.8	+26:31:57.6
1A 0803-00	05:58:45.6	-03:04:08.5
1A 1740-27-2942	23:17:46.6	-27:46:17.1
1E 2259-58	04:52:56.6	-58:57:00.1
3EG J2021-01-18	03:08:41	-01:18:32
4U 015-04	01:15:29.2	-04:33:17.4
4U 014-07	01:14:29.1	-07:04:48.4
4U 1545-47	23:15:45.9	-47:09:44.4
4U 1649-52	23:16:49.7	-52:42:09.6
4U 1630-47	23:16:50.8	-47:39:17.7
4U 1636-53B	23:16:51.2	-53:37:53.9
4U 1702-429	23:16:58.8	-43:03:57.5
4U 1705-40	23:17:06.9	-40:11:22.8
4U 1755-444	23:17:42.6	-44:44.5
4U 1820-303	23:18:59.5	-30:36:11.1
All Air	31:08:18	-08:37:05
Ari X-1	00:01:48.7	-00:51:06.5
Cam X-1	17:03:35.7	-40:22:09.7
Cir X-1	23:01:33.9	-57:16:27.4
Cyg X-1	23:59:50.1	-35:20:18.6
Cyg X-2	23:58:17.9	-38:32:07.7
Cyg X-3	23:58:10.4	-40:05:57.1
GRO J0422-32	04:42:32.1	-32:30:42.5
GRO J1655-40	16:55:06.8	-40:06:48.1
GRO J1705-28	17:05:28.5	-28:45:35.3
GRO J1744-24	17:44:24.5	-24:03:09.9
GRO J1805-20	18:05:20.6	-20:48:55.8
GRO J2014-08	20:14:08.1	-08:47:06.4
GRO J2000-251	20:00:25.1	-25:06:59.8
GRO J2022-33B	20:22:33.8	-33:08:59.6
GX 13-1	27:03:46	-17:15:42
GX 17-2	27:04:08.8	-14:09:03.9
GX 3-1	26:08:33.3	-26:56:08.1
GX 30-4	16:52:51.7	-41:01:09.4
GX 33-4	23:55:26.5	-45:38:07.2
GX 340-4	23:44:48.7	-45:51:11.1
GX 34-4	23:44:48.7	-45:51:11.1
GX 34-5	23:44:48.7	-45:52:00.7
GX 5-1	23:08:07	-05:20:27.1
GX 9-1	27:03:46	-17:15:42
GX 9-9	26:03:17	-16:56:01
H1705-250	23:05:08	-25:05:19.3
H1743-322	23:48:53.3	-32:22:06.7
HESS J0930+057	09:30:42.9	5.8057
Her X-1	19:27:45.6	-35:53:42.9
IGR J17535-0201	17:53:50.7	-02:54:54.7
LMX X-4	17:03:04.6	-44:48:41.1
LMX X-4	17:03:04.6	-44:48:41.1
LSI +61 303	04:13:1	-61:29
PSR J0205-6449	02:05:34.7	-64:59:59
SAX J0905-2053	09:05:22.5	-20:53:58.8
SLX J1745-331	17:45:32.8	-33:20:55.6
SMC X-1	19:27:45.6	-35:44:53.4
SS 433	28:07:56.5	4.98267
SWIFT J1753.5-0207	17:53:53.5	-02:54:54.5
Suz X-1	24:07:49.6	-15:54:02.2
TELE J-4	27:04:08.8	-14:09:03.9
TELE J-4	27:04:08.8	-14:09:03.9
Vela X-1	15:58:59.8	-40:54:47
XTE J118-330	11:18:45.2	-33:40:00.7
XTE J1550-564	15:50:56.4	-56:40:00.4
XTE J1749-276	17:49:27.6	-27:46:39.7
XTE J1806-500	18:06:50.4	-50:50:17
XTE J1811-202	18:11:20.2	-20:18:57
XTE J1901-028	19:01:02.8	-02:54:05.8
XTE J1923-318	19:23:31.8	-31:56:58.3
XTE J1926-288	19:26:28.8	-28:47:08.3
XTE J1917-330	19:17:33.0	-33:01:58.3
XTE J1928-245	19:28:24.5	-24:54:34.3
XTE J1930-248	19:30:24.8	-24:54:34.3
XTE J1936-054	19:36:05.4	-05:27:17
XTE J1946-024	19:46:02.4	-02:54:06.7
XTE J2012-081	20:12:08.1	-08:46:48.9
XTE J2012-081	20:12:08.1	-08:46:48.9

Analysis

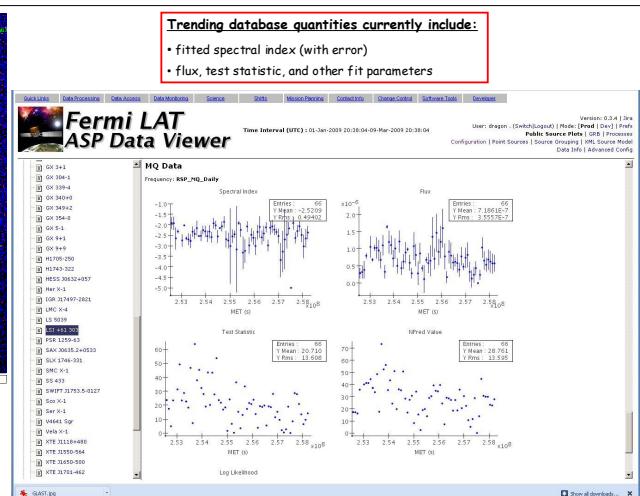
All analysis is performed with Fermi ScienceTools [5] and the HEASoft library [6]

Automatically produced data products:

- Source model (.xml)
- Ancillary files (livetime cubes, exposure maps, etc.)
- Counts map (image and FITS)
- Fit results: summary (.txt) and counts spectra (FITS)
- Light curve (image and FITS)



- LSI +61 303 [7,8] examples:
 - Counts map (actual data in 0.03-deg bins, above)
 - Light curve (1-day binning, left)
 - Trending database web application (right)



References

- Atwood, W.B., et al. 2009 ApJ 697 1071-1102
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