

# The First GRBAlpha and VZLUSAT-2 catalogue: gamma-ray transients and detector sensitivity

Marianna Dafcikova, Masaryk University, Czech republic

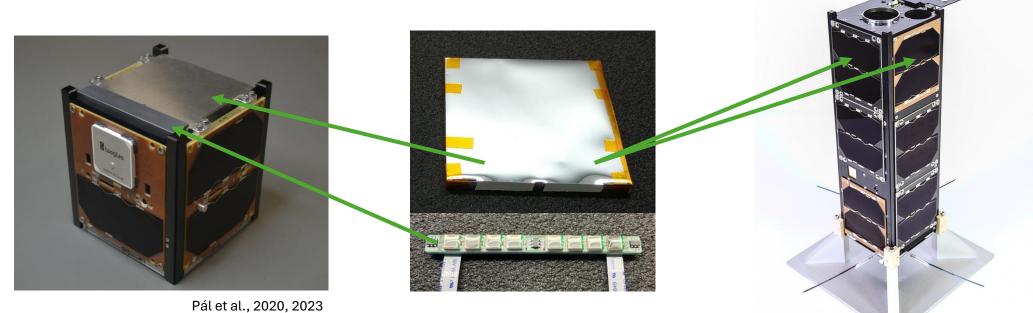
Collaborators: J. Ripa, A. Pal, N. Werner, F. Munz, M. Kolar, L. Szakszonova, M. Duriskova, N. Husarikova, et al.

# GRBAlpha

- 1U CubeSat
- Launched in March 2021
- 550 km polar orbit
- CsI(Tl) scintillator read-out by 2x4 SiPMs
- Technological experiment for CAMELOT

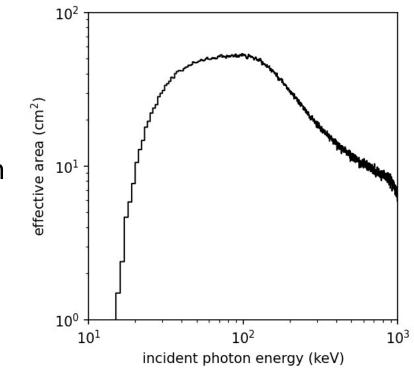
# VZLUSAT-2

- 3U CubeSat
- Launched in January 2022
- 530 km polar orbit
- Secondary payload: 2 GRB detectors



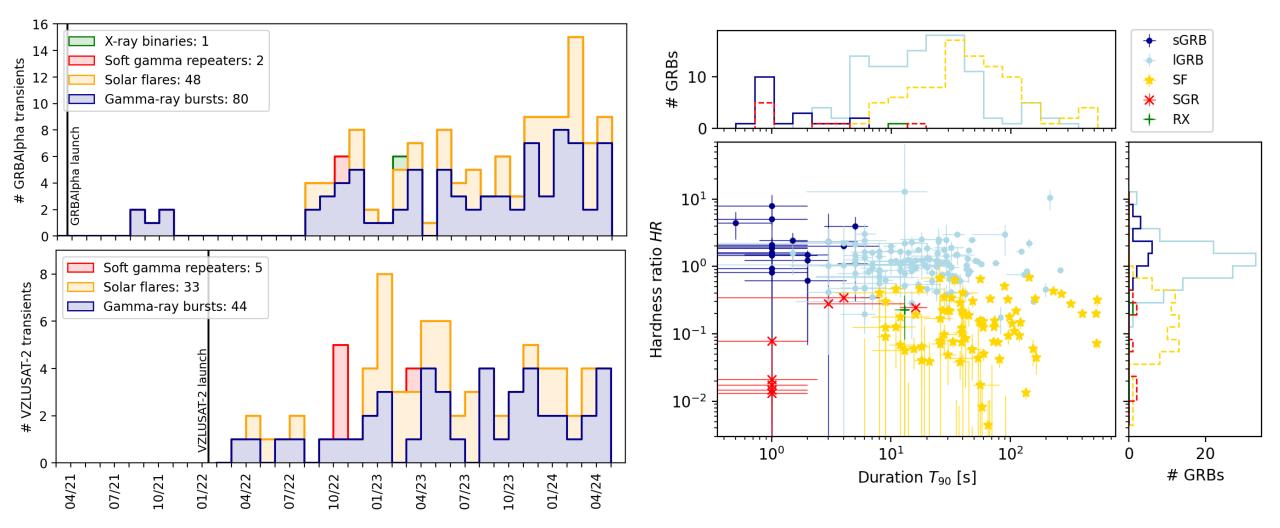
## Observations

- GRBAlpha operations done by students, nonstop measurements
- VZLUSAT-2 measurements less frequent
- 0.5 and 1 s exposure time, 4 energy bands (70 950 keV)
- No trigger algorithm yet, correlation with other missions (Fermi, Swift, INTEGRAL, Konus, AGILE, CALET, GECAM, AstroSat, FRB + GW triggers)
- 2 detections/week -> 250 detections/year for a constellation



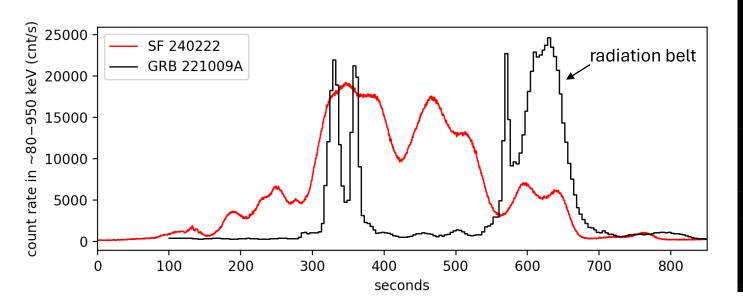
## **Confirmed detections**

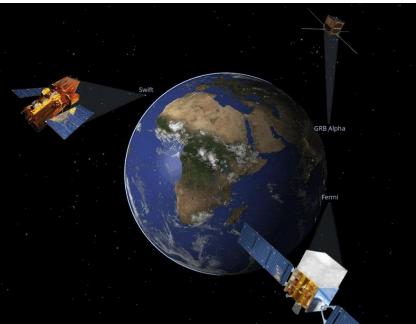
https://monoceros.physics.muni.cz/hea/GRBAlpha/ https://monoceros.physics.muni.cz/hea/VZLUSAT-2/



#### Notable detections

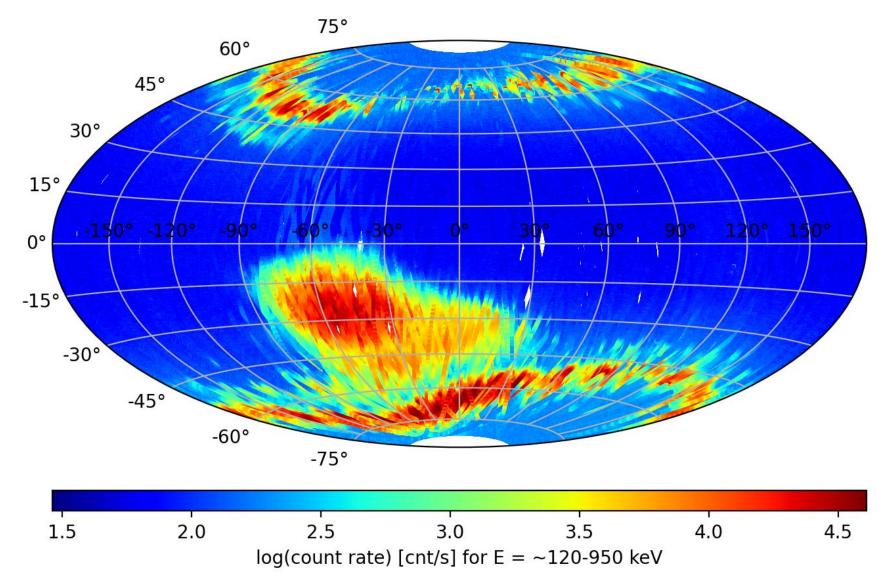
- GRB 230709B and GRB 230709C: 42 minutes apart
- GRB 231215A: most distant GRB at z = 2.305 (10.8 Gyr)
- GRB 221009A (+ GRB 230307A): peak flux without saturation effects
- X6 class solar flare on 2024-02-22: nearly as bright as the BOAT





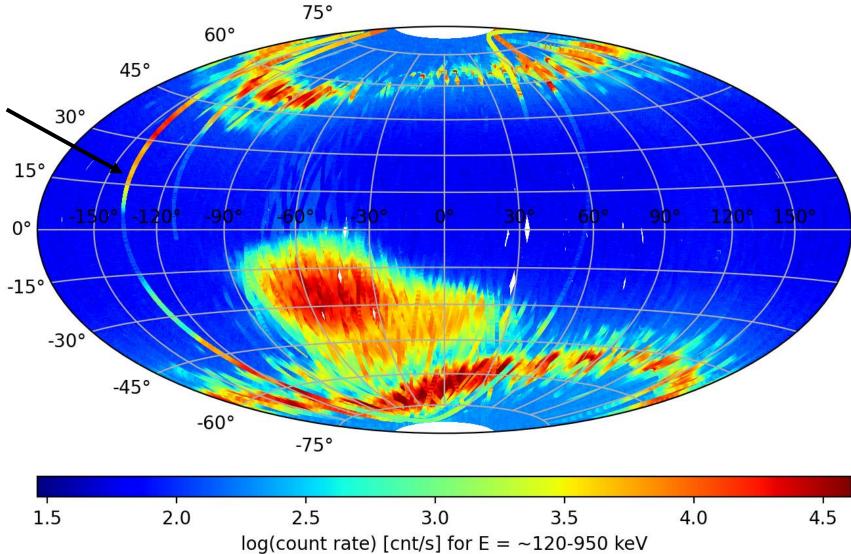
Credit: NASA

#### LEO environment

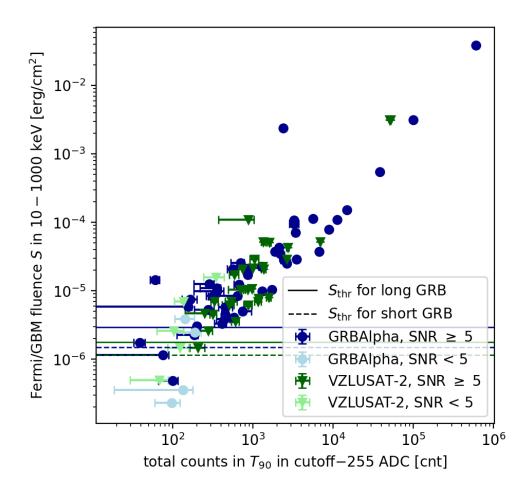


#### LEO environment

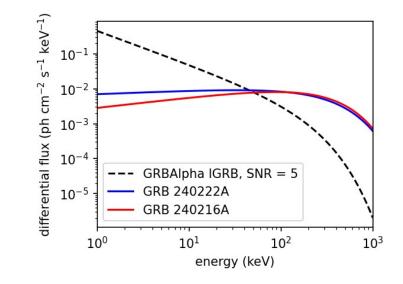




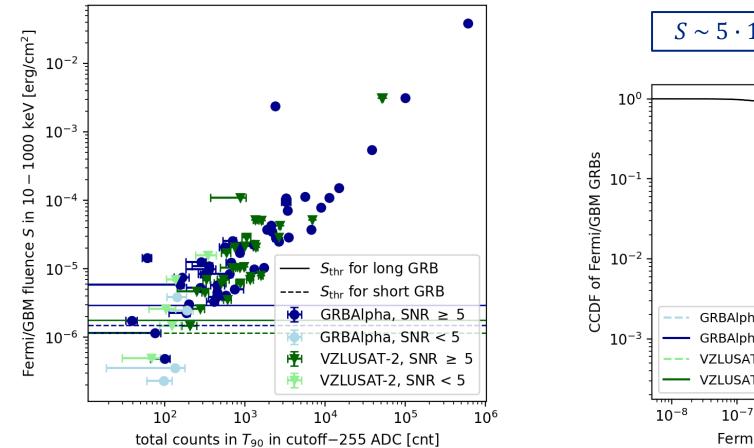
## Cross-correlation with Fermi/GBM

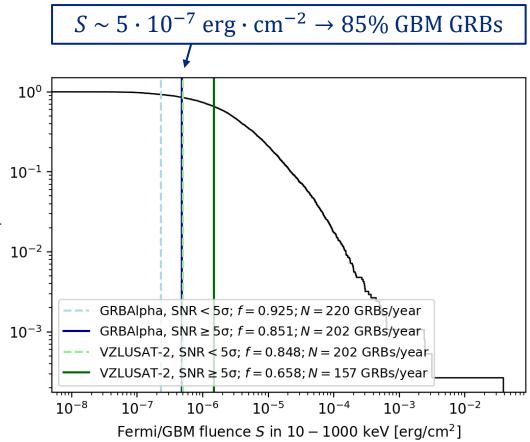


- Unknown GRBAlpha orientation
- No detector response matrix for VZLUSAT-2
- $S_{thr} = 5\sigma$  detection of a typical GRB
- Significant GRBs with  $S < S_{thr}$  were harder

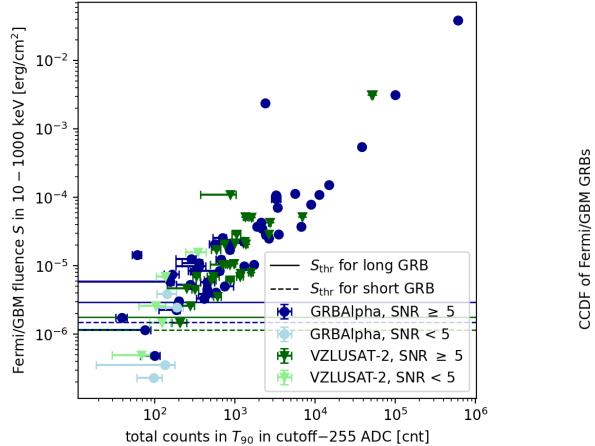


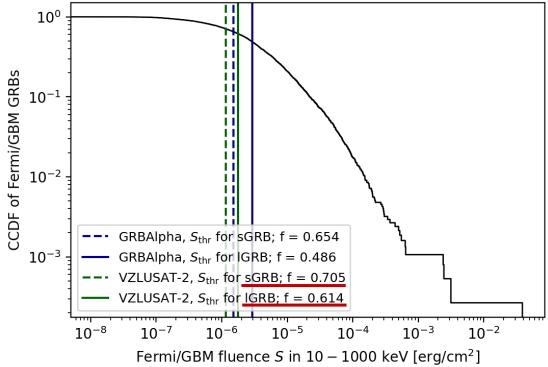
#### **Empirical sensitivity**





#### Theoretical sensitivity





# Summary & future plans

- GRBAlpha: 131+ detections in 3 years
- VZLUSAT-2: 83+ detections in 2 years
- Nonstop measurements: 2 detections/week
- Feasible detection of 60 85% of all Fermi/GBM GRBs
- Sensitivity higher for short GRBs –> advantage in search for GW counterparts
- LEO background monitoring, SiPM degradation, ...
- Trigger algorithm
- GRBBeta (2U CubeSat) launch in mid 2024