

MAGNETIC MIRROR

(H_p)

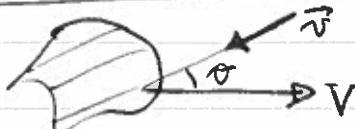
CONSERVATION OF MAGNETIC MOMENT

If the B field varies adiabatically, then $\frac{p_L^2}{B} = \text{const}$

Since B does no work $p^2 = p_L^2 + p_\parallel^2 = \text{const}$

In a region where B increases, $p_L^2 \uparrow \Rightarrow p_\parallel^2 \downarrow$, and the particle may end up being reflected

FERMI ACCELERATION



LAB FRAME: $E_i, p_i, [p_c = \beta E$
 $p_\parallel \approx \beta E]$

In the MIRROR FRAME

$$\begin{cases} E_i' = \gamma_v (\beta_v p_{x,i} + E_i) \\ p_{x,i}' = \gamma_v (p_{x,i} + \beta_v E) \end{cases} \xrightarrow{\text{REFLECTION}} \begin{cases} E_f' = E_i' \\ p_{x,f}' = -p_{x,i}' \end{cases}$$

Back in the LAB FRAME:

$$E_f = \gamma_v (\beta_{-v} p_{x,f}' + E_f') \quad [\beta_{-v} = -\beta_v]$$

$$= \gamma_v (\beta_v p_{x,i}' + E_i') \quad [p_{x,i}' = \beta E \text{ const}]$$

$$E_f = \gamma_v^2 E_i [\gamma_v^2 \beta_v \beta \cos\theta + 1 + \beta_v^2]$$

$$\Delta E = E_f - E_i = E_i [\gamma_v^2 (2 \beta_v \beta \cos\theta + \beta_v^2) + \underbrace{\gamma_v^2 - 1}_{\beta_v^2 \gamma_v^2}]$$

$$\frac{\Delta E}{E} = 2 \beta_v \beta \cos\theta + 2 \beta_v^2$$

$$\boxed{\frac{\Delta E}{E} = 2 \beta_v (\beta \cos\theta + \beta_v)}$$

$$V_{\text{tot}} = \frac{u + V_{\text{rel}}}{1 + \frac{u V_{\text{rel}}}{c^2}}$$

$$V_{\text{rel}} = \frac{u_{\text{tot}} - u}{1 - \frac{u_{\text{tot}} u}{c^2}}$$

In general the # of collisions per unit time is proportional to the relative velocity between the particle and the mirror:

$$dP(\omega) \propto \frac{V_{\text{tot}} + v}{1 + v V_{\text{tot}}/c^2} \sin \omega d\phi d\theta \propto \frac{V_{\text{tot}} + v}{1 + v V_{\text{tot}}/c^2} d\omega$$

(HP): $v \sim c$ and $v \gg V$ (at the $O(1)$ in $\frac{v}{c}$)

$$P(\omega) d\omega = \frac{1}{2} \left(1 + \frac{V_{\text{tot}}}{c} \right) d\omega \propto \left[1 + \beta_v \cos \omega \right]$$

$$\left. \langle \omega \rangle_p \right|_x = \frac{\int_{-1}^1 x (1 + \beta_v x) dx}{\int_{-1}^1 (1 + \beta_v x) dx} = \frac{\frac{x^3}{3} \Big|_{-1}^1}{\frac{x^2}{2} \Big|_{-1}^1} = \beta_v \frac{\frac{2}{3}}{2} = \frac{\beta_v}{3}$$

And finally

$$\left. \langle \frac{\Delta E}{E} \rangle_p \right|_x = \frac{2 \beta_v^2}{3} + 2 \beta_v^2 = \frac{8}{3} \beta_v^2$$

$$\boxed{\left. \langle \frac{\Delta E}{E} \rangle_p \right|_x = \frac{8}{3} \beta_v^2}$$

II order Fermi acceleration

(I) If we are at a shock, assuming isotropy, $P(\omega)$ $\propto \omega$, which implies:

$$\left. \langle \omega \rangle_p \right|_x = \frac{\int_0^1 x \cdot x dx}{\int_0^1 x dx} = \frac{\frac{x^3}{3} \Big|_0^1}{\frac{x^2}{2} \Big|_0^1} = \frac{2}{3}, \text{ so that}$$

$$\left. \langle \frac{\Delta E}{E} \rangle_p \right|_x = \frac{4}{3} \beta_v \beta + 2 \beta_v^2$$

I order Fermi acceleration

CAVEAT: If $v \gg V$ there may be 2 averages involved [still assuming specular reflection!]

Shock HYDRODYNAMICS

Let us start from the conservation eqs for an IDEAL FLUID; described by $\{\vec{v}(x, t); \rho(x, t); p(x, t)\}$

$$\text{MASS: } \frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) = 0$$

$$\text{MOMENTUM: } \frac{\partial \vec{p}}{\partial t} + (\vec{p} \cdot \vec{\nabla}) \vec{v} = - \frac{\vec{\nabla} p}{\rho} (+ \frac{\vec{p} \cdot \vec{\nabla} t}{\rho}) = - \vec{\nabla} w$$

$$\text{ENERGY: } \frac{\partial}{\partial t} \left(\rho \frac{v^2}{2} + pE \right) + \vec{\nabla} \cdot \left[\rho \vec{v} \left(\frac{v^2}{2} + w \right) \right] = 0$$

Assume: 1D + stationarity ($\vec{v} \rightarrow u$)

$$\begin{cases} w = \text{enthalpy} \\ \text{unit mass} \\ = E + pV; V = \frac{1}{\rho} \end{cases}$$

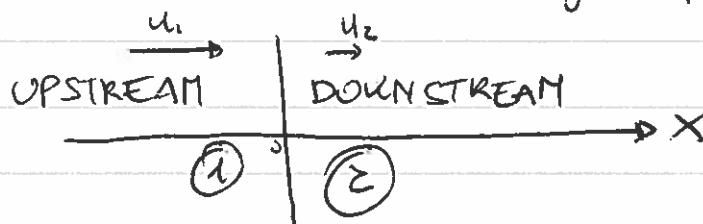
$$\begin{cases} w = \frac{\gamma}{\gamma-1} \frac{p}{\rho} = \frac{c_s^2}{\gamma-1} \\ E = \frac{1}{\gamma-1} \frac{p}{\rho} = \frac{c_s^2}{(\gamma-1)} \end{cases}$$

$$\text{MASS: } \rho u = \text{const}$$

$$\text{MOMENTUM: } \rho u^2 + p = \text{const}$$

$$\text{ENERGY: } \cancel{\rho u^3 + p \left(\frac{u^2}{2} + w \right)} = \text{const} \Rightarrow \frac{1}{2} \rho u^3 + \frac{\gamma}{\gamma-1} u p = \text{const}$$

Let's consider a solution with a jump (and $\delta s > 0$)



RANKINE-HUGONIOT
CONDITIONS

$$\frac{u_2}{u_1} = \frac{u_1}{u_2} = \frac{\gamma+1}{\gamma-1 + \frac{2}{M_1^2}}$$

$$M_1^2 = \frac{u_1^2}{c_{s1}^2} \quad \left[\frac{P_1}{\rho u_1^2} = \frac{1}{\gamma M_1^2} \right]$$

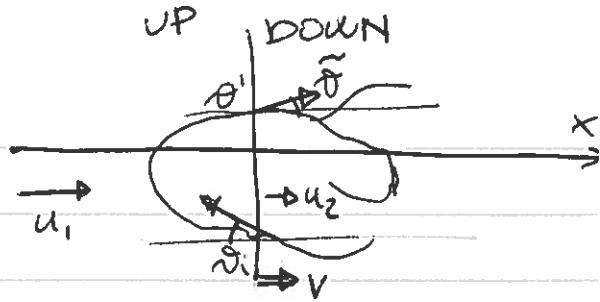
$$\frac{P_2}{P_1} = \frac{2\gamma M_1^2 - (\gamma-1)}{\gamma+1} ; \quad \frac{T_2}{T_1} = \frac{P_2}{P_1} \cdot \frac{P_1}{P_2}$$

$$P_{pr} = \text{const}$$

$$M_1^2 \gg 1 : \quad \frac{P_2}{P_1} = \frac{\gamma+1}{\gamma-1} ; \quad \frac{P_2}{P_1} = \frac{2\gamma M_1^2}{\gamma+1} ; \quad \frac{T_2}{T_1} = \frac{2\gamma(\gamma-1)M_1^2}{(\gamma+1)^2} ; \quad M_2^2 = \frac{\gamma-1}{2\gamma}$$

$$\frac{\gamma+3}{\gamma-1} = 4 \quad 2M_1^2 \quad \frac{\gamma-3}{\gamma-1} = \frac{1}{5}$$

SHOCK ACCELERATION



Consider a TEST-PARTICLE

in the DS FRAME, with velocity $\beta \gg u_1, u_2$.

(Fp) Let us assume that there is TURBULENCE to scatter the part. in pitch angle θ .

Let the flight cosine be $\mu = \cos\theta$

and be $v = u_1 - u_2$ the relative velocity between UP/DS

$$[DS] \quad E = E_i \quad p_{xi} = \beta \mu_i E_i \quad p_c = \beta E \quad p_{xi} = p \mu_i$$

$$DS \rightarrow US \quad E' = \Gamma E_i (1 + \beta_v \beta \mu_i) \quad \Gamma = \gamma_v$$

The part. is isotropized by turbulence, and may re-encounter the shock with a different $\theta = \theta'$ and with $\beta = \beta'$

$$\hookrightarrow E'_f = E'_i \quad \& \quad p'_{xf} = E'_f \beta' \mu'$$

Back in the DS.

$$[DS_f] \quad E_f = \Gamma^2 E_i (1 + \beta_v \beta \mu_i) (1 - \beta_v \beta' \mu')$$

$$[P_v] = -\beta_v$$

Now, I have to also convert μ' into $\tilde{\mu}_f$

$$\mu'_f = -\frac{\beta_v - \mu' \beta}{1 - \mu' \beta_v}$$

$$\mu' = \frac{\mu_f + \beta_v}{\beta + \mu_f \beta_v}$$

$$[\beta, \beta' \approx 1]$$

$$\text{Finally: } \frac{E_f}{E_i} = \Gamma^2 (1 + \beta_v \mu_i) \left[1 - \beta_v \frac{\mu_f + \beta_v}{\beta + \mu_f \beta_v} \right]$$

$$= \Gamma^2 (1 + \beta_v \mu_i) \frac{1 - \beta_v^2}{1 + \mu_f \beta_v} = \frac{1 + \beta_v \mu_i}{1 + \beta_v \mu_f}$$

$$\boxed{\frac{E_f}{E_i} \approx (1 + \beta_v \mu_i) (1 - \beta_v \mu_f)} \quad \text{for } \beta_v \ll 1$$

AVERAGE OVER μ_i, μ_f

(DS)

We want to calculate the flux of particle in direction μ : $\phi \propto n v_x \propto \mu$; therefore

$$\left\langle \frac{E_F}{E_i} \right\rangle_{\mu_f} = \frac{\int d\mu_f \mu_f (1 + \beta_v \mu_f) (1 - \mu_f \beta_v)}{\int d\mu_f \mu_f}$$

$$P(\mu) \propto \mu$$

(DS)

Extremes of integration.

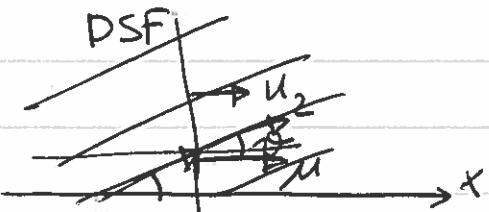
The rel velocity between the port and the shock is

- if the port distribution is isotropic in the DSF -
is $\mu_f + u_2/c$

UP \rightarrow DS

For the port to enter downstream:

$$\mu_f + u_2 \geq 0 \Rightarrow -u_2 \leq \mu_f \leq 1$$



DS \rightarrow UP

For the port to enter upstream:

$$\mu_i + u_2 \leq 0 \Rightarrow -1 \leq \mu_i \leq -u_2$$

Performing the average:

$$\left\langle \frac{E_F}{E_i} \right\rangle_{\mu_f} = \frac{1 + V \mu_i}{\frac{1}{2}(1 - u_2^2)} \int_{-u_2}^1 d\mu_f (1 - \mu_f V) \mu_f$$

$$\int_{-u_2}^1 d\mu_f \mu_f = \frac{1}{2}(1 - u_2^2)$$

$$= \frac{1 + V \mu_i}{\frac{1}{2}(1 - u_2^2)} \left[\frac{1}{2}(1 - u_2^2) - \frac{V}{3}(1 - u_2^3) \right]$$

$$\int_{-u_2}^1 d\mu_f \mu_f^2 = \frac{1 + u_2^3}{3}$$

$$\left\langle \frac{E_F}{E_i} \right\rangle_{\mu_f, \mu_i} = \left\{ 1 - \frac{2}{3} V \left(\frac{1 + u_2^3}{1 - u_2^2} \right) \right\} \frac{1}{u_2^2 - 1} \int_{-1}^{-u_2} d\mu_i \mu_i (1 + V \mu_i)$$

$$\int_{-1}^{-u_2} d\mu_i \mu_i = \frac{u_2^2 - 1}{2}$$

$$= \left[1 - \frac{2}{3} V \left(\frac{1 + u_2^3}{1 - u_2^2} \right) \right] \left[1 + \frac{2}{3} V \left(\frac{1 - u_2^3}{u_2^2 - 1} \right) \right] \approx$$

$$\int_{-1}^{-u_2} d\mu_i \mu_i^2 = \frac{1 - u_2^3}{3}$$

$$\approx 1 + \frac{2}{3} V \left(\frac{1 + u_2^3}{u_2^2 - 1} \right) - \frac{2}{3} V \left(\frac{1 - u_2^3}{1 - u_2^2} \right) + O(V^2) \approx 1 + \frac{4}{3} V$$

$$\left\langle \frac{\Delta E}{E} \right\rangle_{\mu_i, \mu_f} = \frac{4}{3} (u_1 - u_2)$$

BELL'S APPROACH

Let's start with N_0 particles with energy E_0

bc: - G : the energy gain per step: $\bar{E}_K G^k E_0$

- P : the probability of staying in the accelerator

After K cycles: $N_K = P^K N_0$

[G, P indep of E]

Since:

$$\frac{N_K}{N_0} = P^K \quad \log\left(\frac{N_K}{N_0}\right) = K \log P$$

$$\frac{\bar{E}_K}{E_0} = G^K \quad \log\left(\frac{\bar{E}_K}{E_0}\right) = K \log G \Rightarrow$$

$$\log\left(\frac{N_K}{N_0}\right) = \frac{\log P}{\log G} \log\left(\frac{\bar{E}_K}{E_0}\right) \Rightarrow \log\left(\frac{N_K}{N_0}\right)$$

$$\frac{N_K}{N_0} = \left(\frac{\bar{E}_K}{E_0}\right)^{\alpha}, \quad \alpha = -\frac{\log P}{\log G}$$

RETURN PROBABILITY

The ret. prob. P is the ratio of FLUXES of particles returning to the shock to entering part

$$P = \frac{|\text{J}_{\text{ret}}|}{\text{J}_{\text{in}}} = \frac{- \int d\mu(u+u_2) \Big|_{u+u_2 \leq 0}}{\int d\mu(u+u_2) \Big|_{u+u_2 \geq 0}} = \frac{- \int_{-u_2}^{u_2} d\mu(u+u_2)}{\int_{-u_2}^1 d\mu(u+u_2)}$$

$$= \frac{- \left[\mu u_2 + \frac{\mu^2}{2} \right]_{-1}^{u_2}}{\left[\mu u_2 + \frac{\mu^2}{2} \right]_{-u_2}^1} = \frac{- \left[-u_2^2 + \frac{u_2^2}{2} + \mu u_2 - \frac{1}{2} \right]}{u_2^2 + \frac{u_2^2}{2} + \mu u_2 - \frac{u_2^2}{2}} = \frac{- \left[-u_2^2 + 2u_2 - 1 \right]}{u_2^2 + 2u_2 + 1} = \frac{(1-u_2)^2}{1+u_2}$$

$$P \approx 1 - 4u_2 \quad \text{for } u_2 \ll 1$$

$$v = \frac{u_1}{u_2}$$

$$\alpha = -\frac{\log P}{\log G} = -\frac{\log(1-4u_2)}{\log(1+\frac{4}{3}(u_1-u_2))} \approx \frac{4u_2}{\frac{4}{3}(u_1-u_2)} = \frac{3u_2}{u_1-u_2} = \frac{3}{v-1}$$

$$\frac{N_K}{N_0} = \left(\frac{\bar{E}_K}{E_0}\right)^{\frac{3}{v-1}} \Rightarrow N(E) = N_0 \left(\frac{E}{E_0}\right)^{-\frac{3}{v-1}} \Rightarrow \frac{dN}{dE} \propto \left(\frac{E}{E_0}\right)^{-\frac{3}{v-1}} \propto \left(\frac{E}{E_0}\right)^{\frac{3}{v-1}}$$