At the end of July 2006, the blazar PKS 2155-304 ($z=0.116$) literally "exploded" into an exceptional TeV activity observed by HESS (George et al., 1997); HESS Collaboration).

**History of the TeV event**

(Benbow et al., 2006, Aharonian et al., in preparation. We present here a summary of the X-ray/UV/Optical follow-up performed with the Swift satellite. More details can be found in Benbow et al., astro-ph/0703869.)

- **27 July 2006 (afternoon):** Release of ATel 867 (Benbow et al.) about the increase of TeV activity of PKS 2155-304.
- **27/28 July 2006:** First outburst, with night average of 8 Crab (E>200 GeV) and flares up to 17 Crab.
- **29/30 July 2006:** Second outburst with night average of 5 Crab and flares up to 15 Crab.

**Swift/BAT:** No detection, upper limit (20-40 keV; 3σ): 42 mCrab on 29-31 July.

**Swift/NRT** energy band 0.3-10 keV.

**Swift/UVOT** Optical fluxes in units $10^{-14}$ erg cm$^{-2}$ s$^{-1}$ Å$^{-1}$ No data in U filter are shown, since they are saturated for most of the observing time.

Each mark in the abscissa corresponds to the 0:00:00 UTC of the day. For example, the mark at day 2063 corresponds to 0:00:00 UTC. The night between 29 and 30 July, when the second TeV outburst occurred, is approximately emphasized by the yellow rectangle.

**CONCLUSIONS:**

We presented the observations of the blazar PKS 2155-304 performed by the Swift satellite immediately after the giant TeV flare observed by HESS at the end of July 2006 (Raue et al. 2006; Aharonian et al., in prep.). The most important result appears to be that, in correspondence with the dramatic TeV activity, the X-ray intensity changed by a factor 5 but without large spectral changes. In particular the frequency of the synchrotron peak occurred at the peak of 3.8 Crab occurred on 16 April, with an average of 1.6 Crab and no hourly timescale variability (Gatster et al. 1997, Ran et al. 1998).

**A less extreme**, though analogous, behaviour was observed in Mkn 421 in 1998-2000. The X-ray and TeV activity were correlated also on short timescale (Me先sha et al. 1999, Takahashi et al. 2000) with larger amplitude variations in the TeV band. The synchrotron peak appeared to shift to higher energies but not as dramatically as for Mkn 501.

**The behaviour of PKS 2155-304 appears less striking in X-rays than for the previous sources but more extreme in the TeV variability.**