



Fermi
Gamma-ray Space Telescope

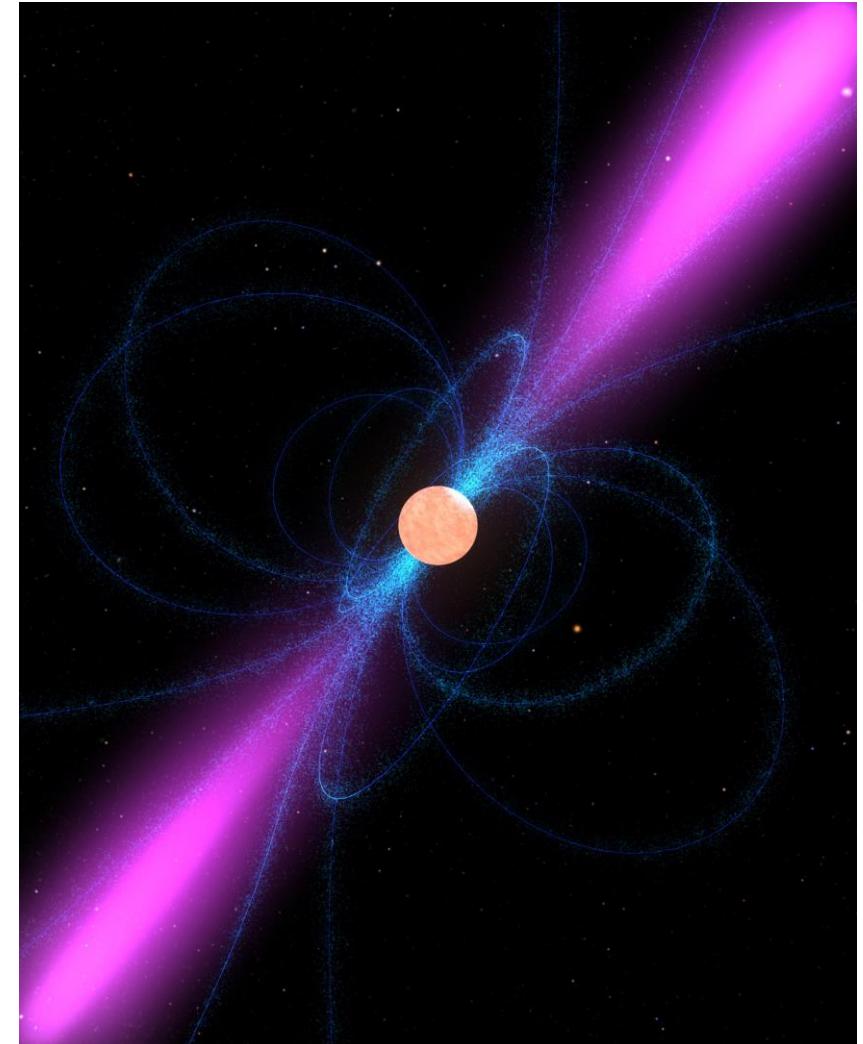
Analyzing Pulsar “Glitches” Using the Fermi LAT

Brent Limyansky
1st Year Graduate Student
University of California, Santa Cruz
Santa Cruz Institute for Particle Physics



Common Characteristics

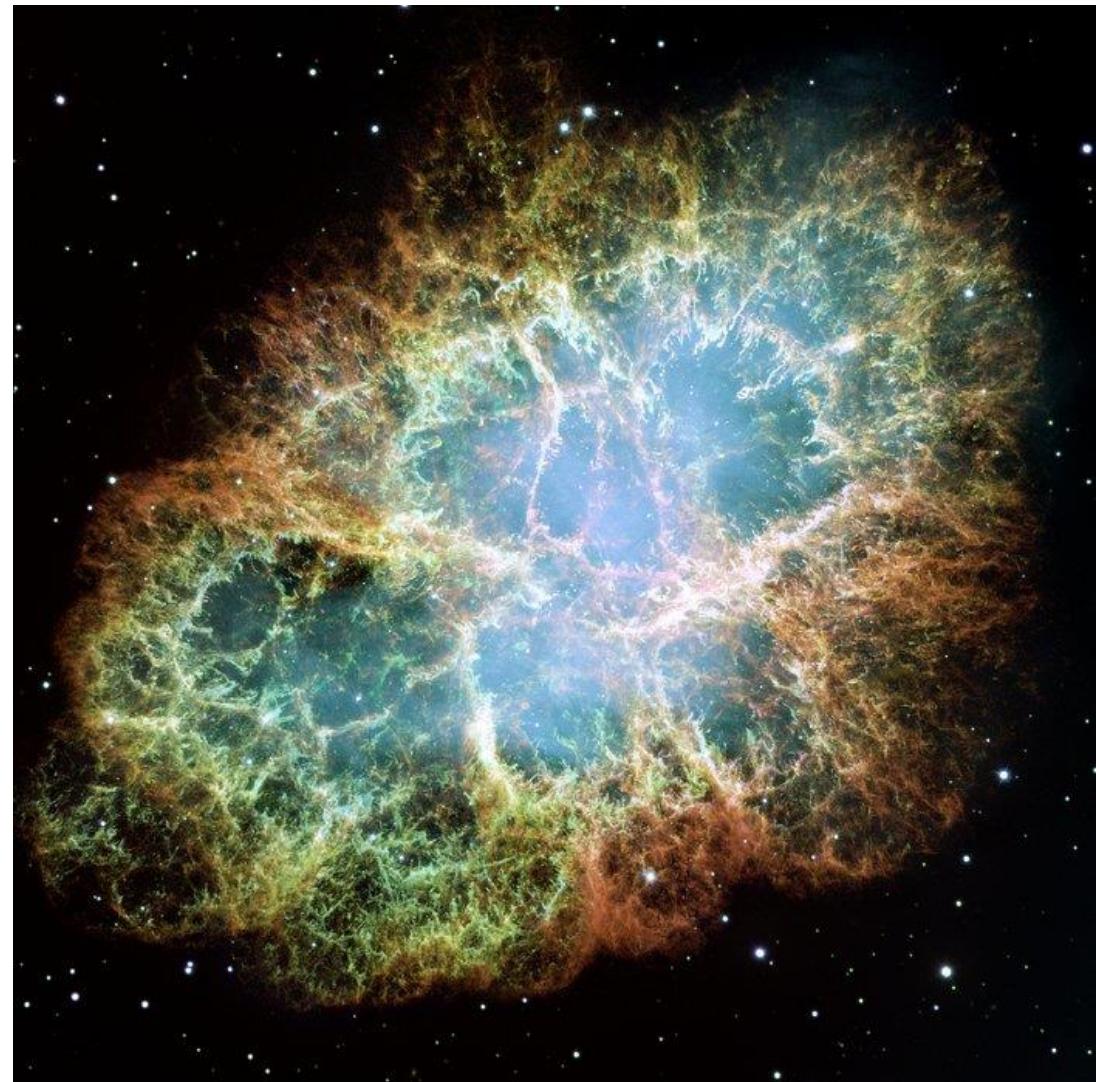
- Neutron Stars
- Variable Sources
 - **Milliseconds to Minutes**
- Radius ~ 10 km
- Mass ~ $1.4 M_{\odot}$
- Strong Magnetic Fields
- High rates of Spin
- Focused Radiation Beams



NASA

Stellar Origins

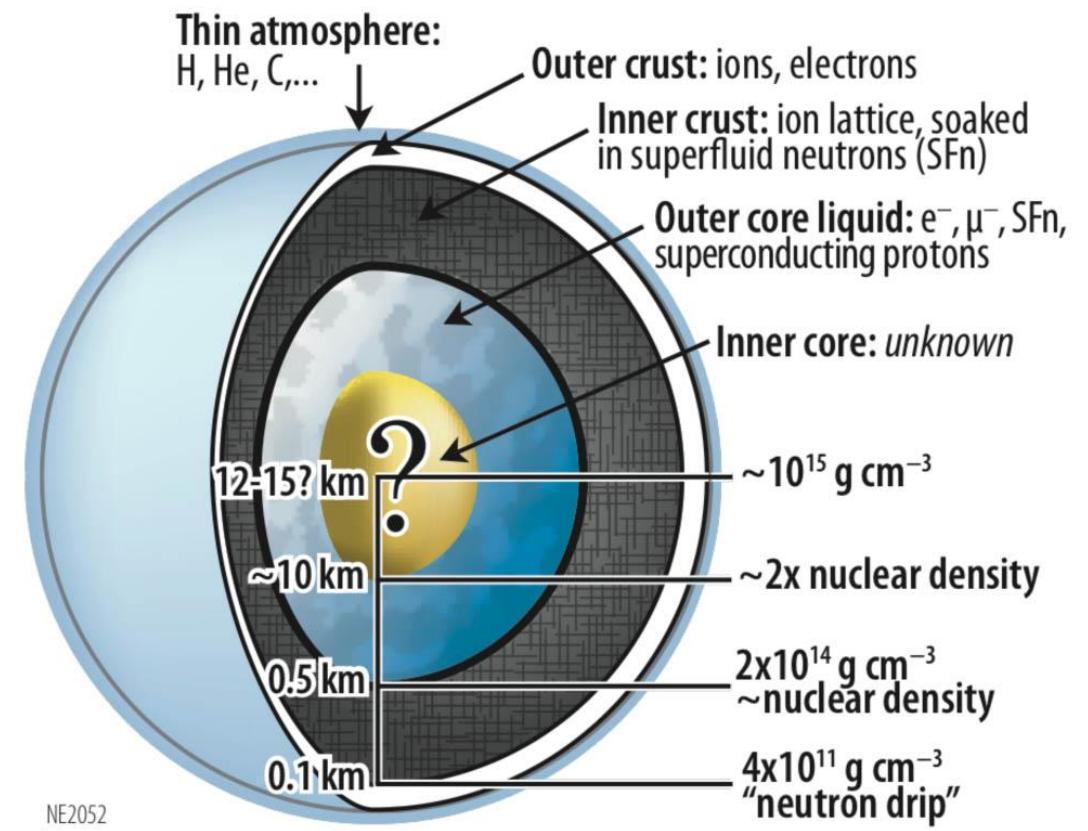
- Formed During Supernovas
 - Commonly Associated with SNRs
 - Core Collapse
- Progenitor Mass $8\text{-}25 M_{\odot}$
 - Possibly Greater in Binary Systems
- Between White Dwarf and Black Hole



NASA, ESA and Allison Loll/Jeff Hester

Internal Structure

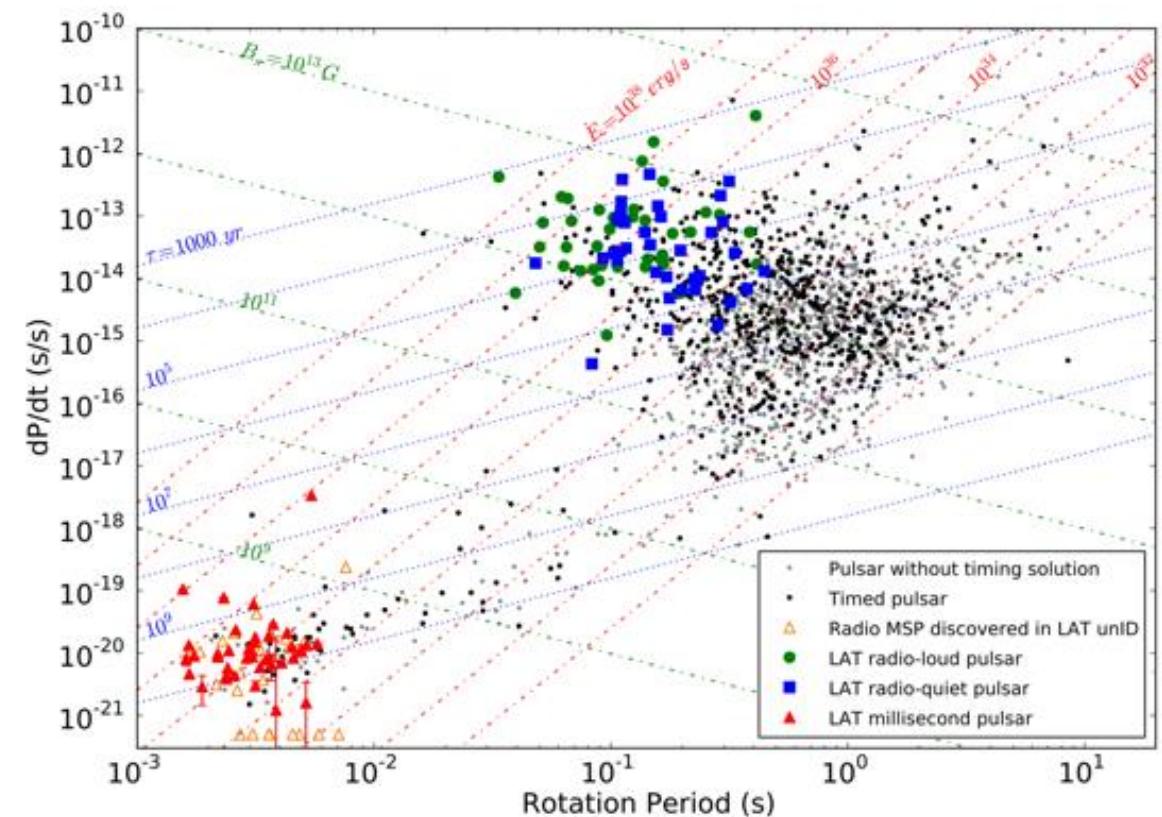
- Magnetosphere
 - Coupled to Crust
 - Radiation Generation
- Inner Superfluid Component
 - Differential Rotation
- Core
 - Unknown Composition
 - Determines Maximum Mass



Gendreau et al. 2012

Classification

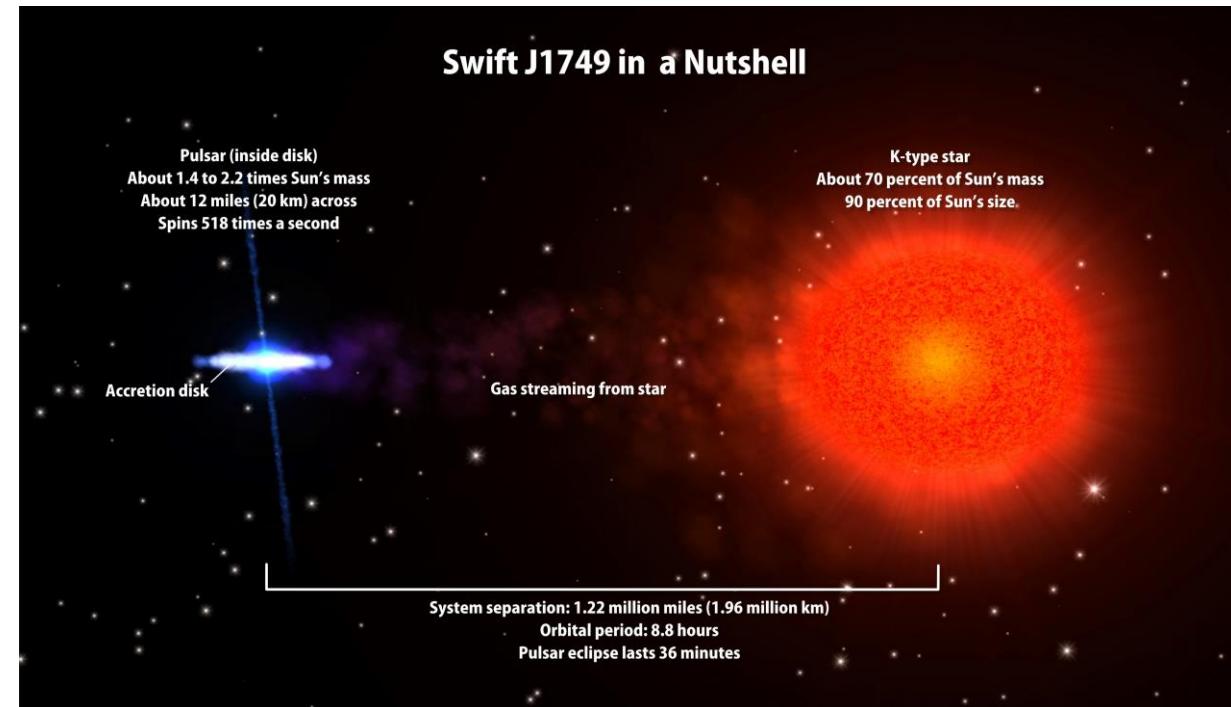
- Main Factors
 - Period
 - Period Derivative
 - Power Source
- Correlated Factors
 - Magnetic Field Strength
 - Characteristic Age



Fermi-LAT Collaboration

Notable Evolutions

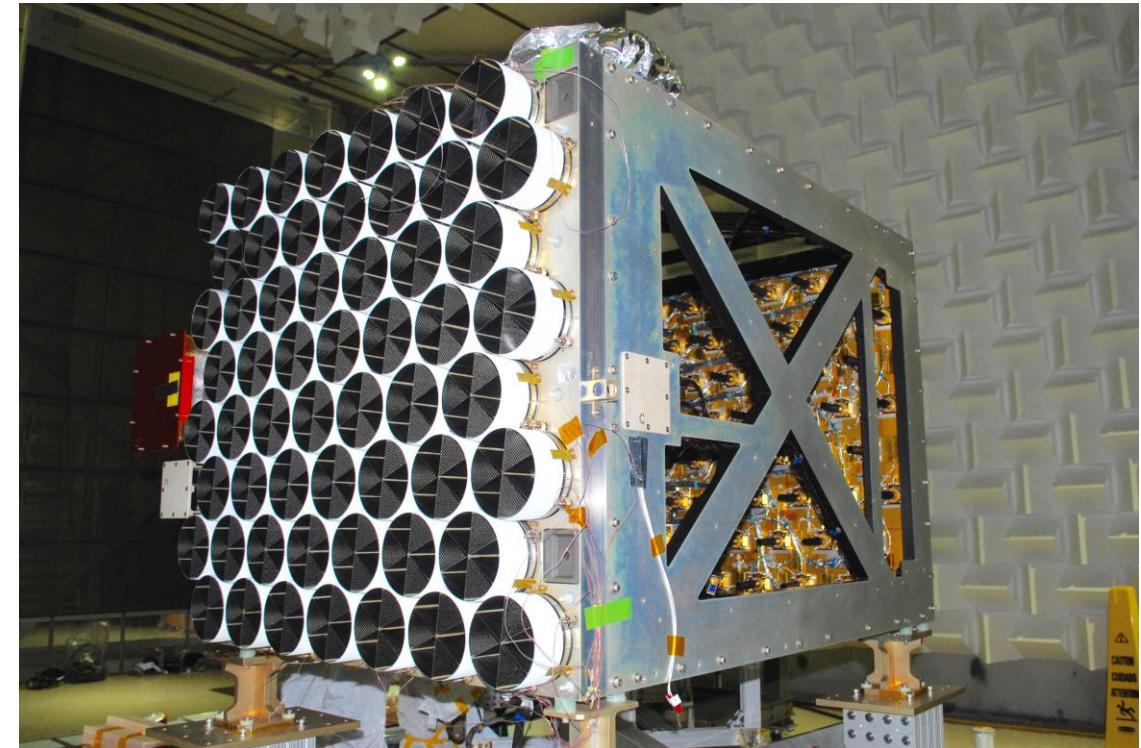
- Binary Systems
- Accreting Neutron Stars
 - Mass Transfer from Companion
 - Bright in X-Ray (GBM)
- Recycled Pulsars
 - Redback System
 - Black Widow System



NASA

Millisecond Pulsars

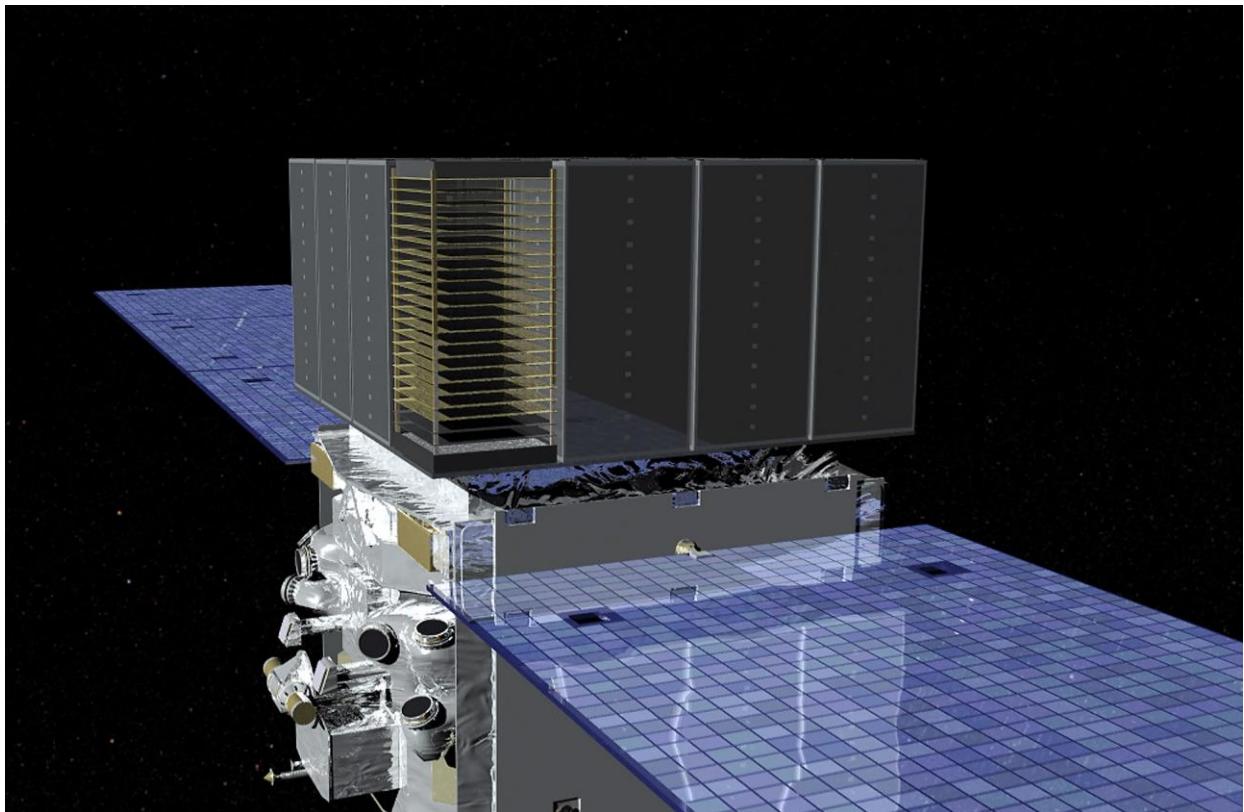
- Precise
 - Rival Early Atomic Clocks
 - Don't Glitch
- Possible Gravitational Wave Detectors
 - NANOGrav
- Possible Navigational Applications
 - NICER



NASA

Fermi and Pulsars

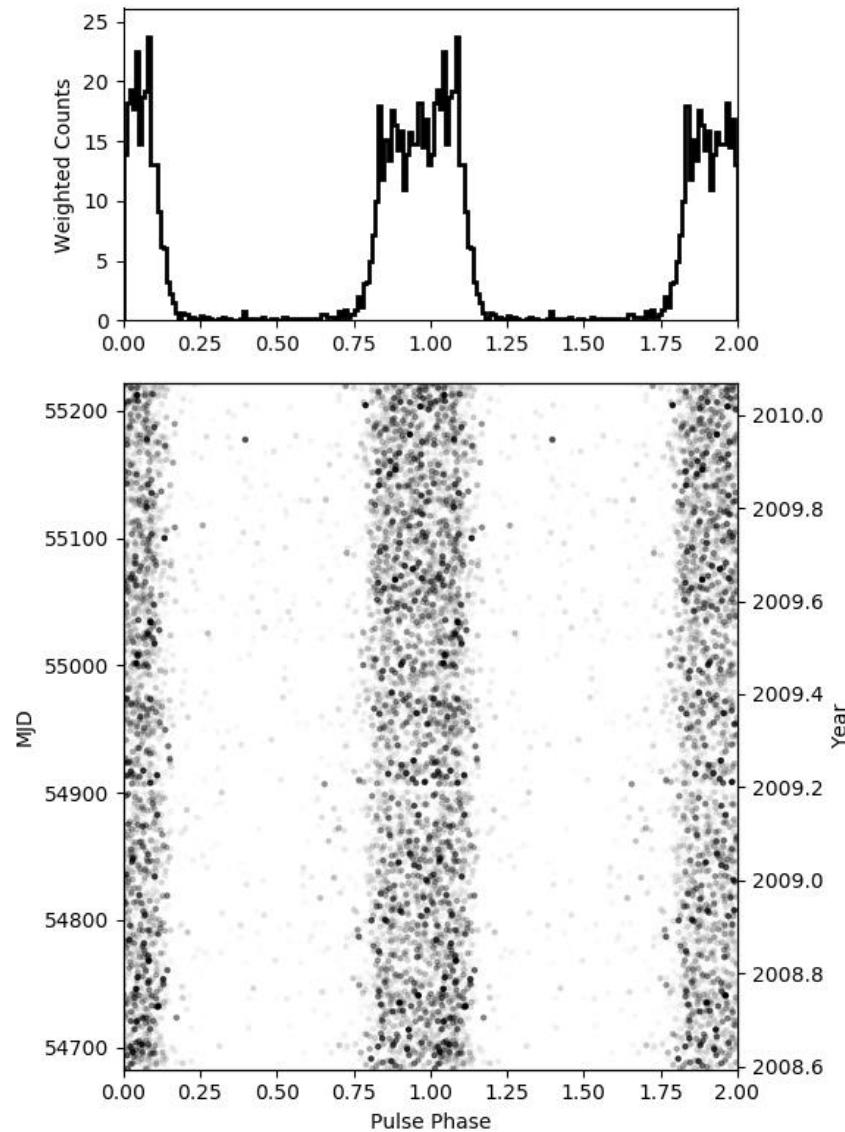
- Detection of over 200 Gamma-Ray Pulsars
- Discovery of Many Millisecond Pulsars
 - Spectral Searches
- This Project:
 - Build a Catalog of Glitches Observed by the Fermi Spacecraft



NASA

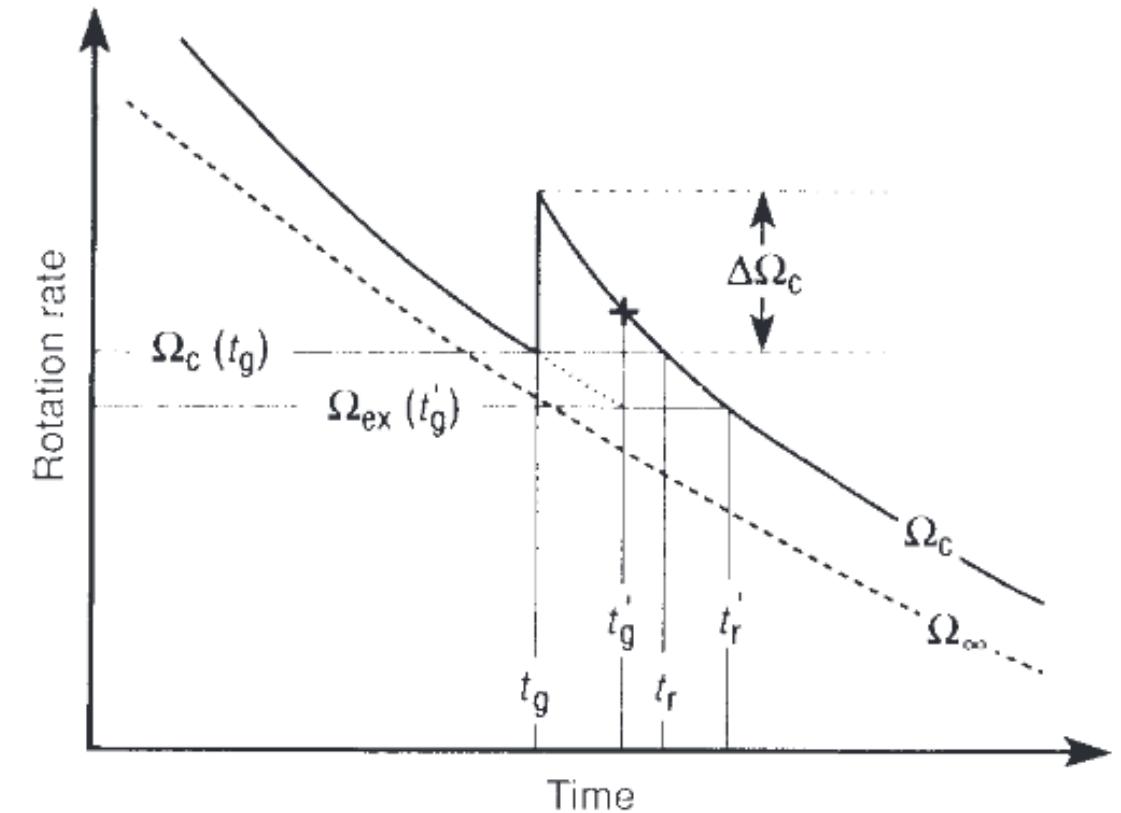
Pulsar Challenges

- One Photon for Many Revolutions
 - Timing Noise
- Barycentering
 - Position of Spacecraft
 - Position of Pulsar
- Techniques
 - Existing Solutions
 - Blind Searches
 - Spectral Searches



Pulsar “Glitches”

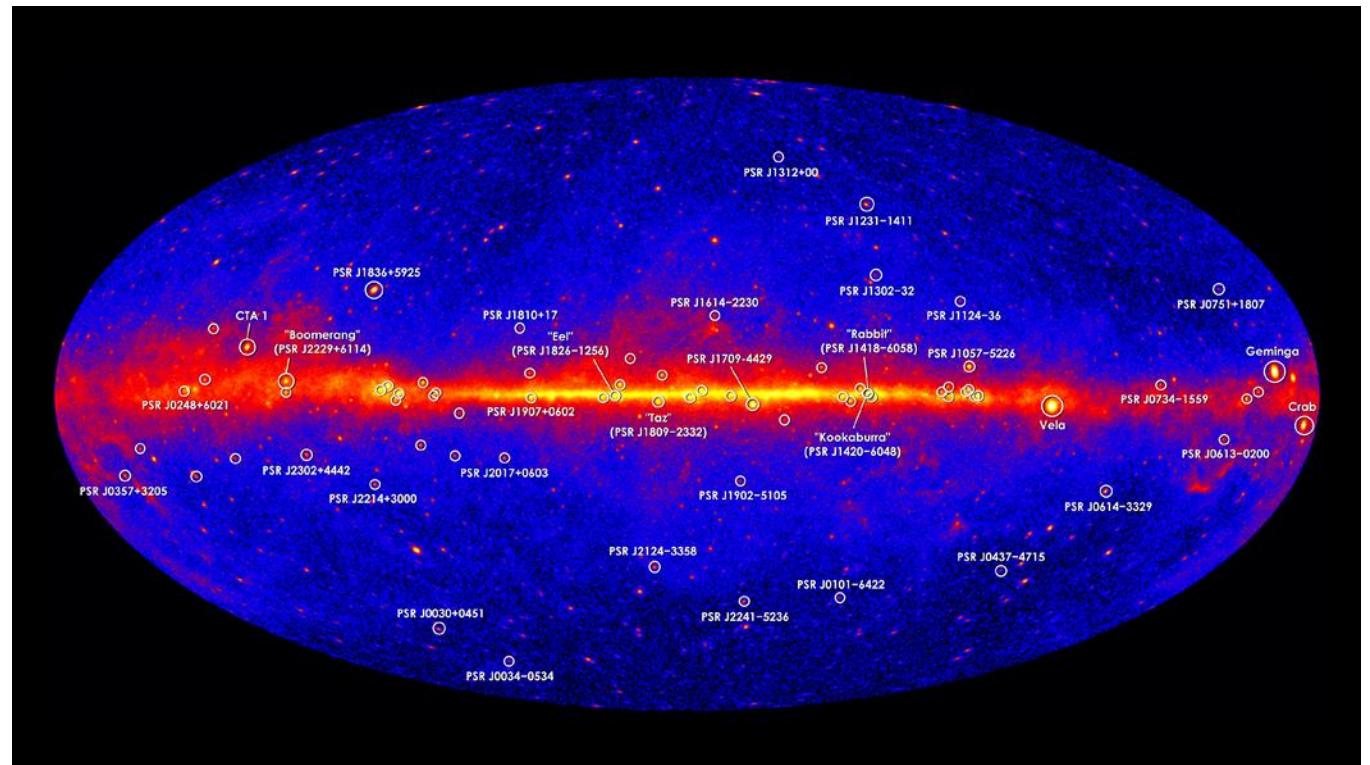
- Sudden Changes in Rotation Rate
- Mechanisms
 - Coupling of Crust and Superfluid
 - Starquakes and Crust Rearrangement
- Investigate Equation of State
- Gravitational Wave Source



Link et al., Nature, 1992

Fermi and Glitches

- Continual Monitoring
 - Prompt Observation
- All Sky
 - Large Population
- Previous Searches
 - 128 Glitches
 - 63/700+ Since 1978
- Fermi Subset
 - 44 Glitches
 - 24/81 Monitored
 - ~ 5 years



NASA/DOE/Fermi LAT Collaboration

Works Cited

Elan Stopnitzky and Stefano Profumo. Gravitational waves from gamma-ray pulsar glitches. *The Astrophysical Journal*, 787(2):114, 2014.

C. M. Espinoza, A. G. Lyne, B. W. Stappers, and M. Kramer. A study of 315 glitches in the rotation of 102 pulsars. *Monthly Notices of the Royal Astronomical Society*, 414(2):1679, 2011.

Alice K. Harding. The neutron star zoo. *Frontiers of Physics*, 8(6):679–692, 2013.

Keith C Gendreau, Zaven Arzoumanian, and Takashi Okajima. The neutron star interior composition explorer (nicer): an explorer mission of opportunity for soft x-ray timing spectroscopy. In *SPIE Astronomical Telescopes+ Instrumentation*, pages 844313–844313. International Society for Optics and Photonics, 2012.

Bennett Link, Richard I. Epstein, and Kenneth A. Van Riper. Pulsar glitches as probes of neutron star interiors. *Nature*, 359(6396):616–618, Oct 1992.