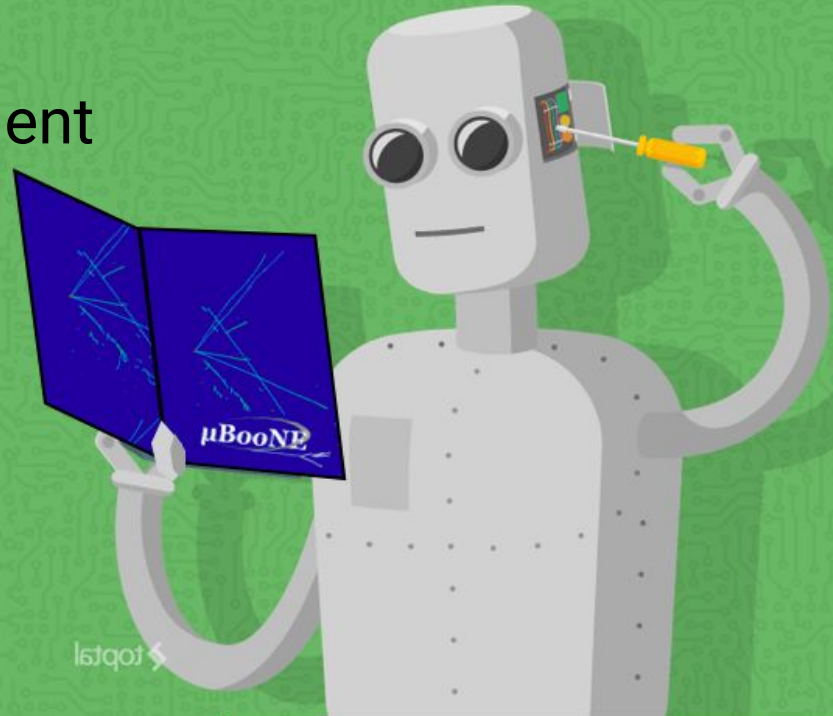


From Pixels to Neutrinos

Convolution Neural Networks
Applied to the MicroBooNE experiment

Taritree Wongjirad (Tufts U.)
SLAC Seminar
June 6, 2019



Outline

Discuss efforts on the MicroBooNE experiment to use convolutional neural networks to improve physics analysis

- What are we trying to solve with CNNs? -- and why?
- From physics problem to ML problem
- Current Analysis Effort
- Developments



“Eyes” of the MicroBooNE detector -- a time-projection chamber,-- before installation

The MicroBooNE Experiment in a nutshell

“An accelerator neutrino oscillation experiment”

Neutrinos

“An accelerator neutrino oscillation experiment”

Neutrinos

Neutrinos are a type of the fundamental particle

Key stats:

- No electric charge
- Interacts only via weak force and gravity
- Very, very small mass: ~ 8 orders smaller than next heaviest particle!
- Come in three “flavors” -- based on what other particle they make during certain interactions

The Three “FLAVORS”



electron



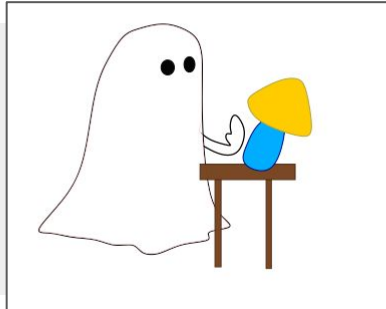
muon



tau

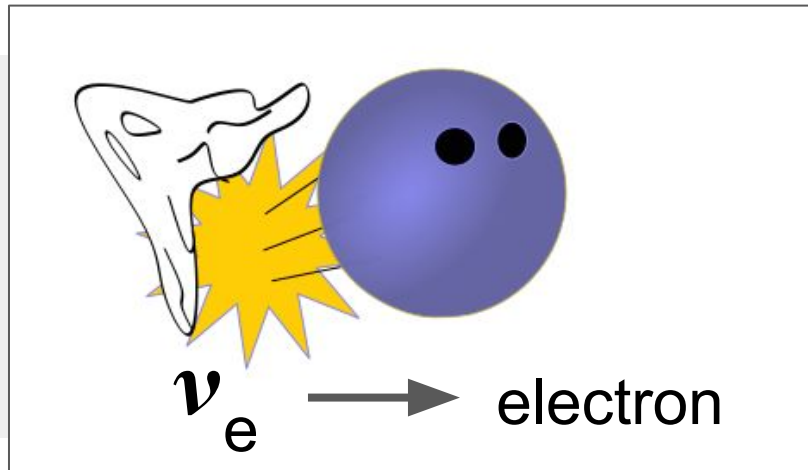
Life as Neutrino Physicist

No electric charge
=> ***observe indirectly***



Only Weak Force
=> ***rare process***

Flavor related to particle produced
during certain interaction
=> ***key to identifying type***



Neutrino Source

“An accelerator neutrino oscillation experiment”

MicroBooNE uses a high-intensity beam of neutrinos made at Fermi National Accelerator Lab



Neutrino Oscillations

“An accelerator neutrino oscillation experiment”

Goal is to look for evidence of Neutrino Oscillations

Flavor detected oscillations over distance traveled

Neutrino created in certain flavor ...



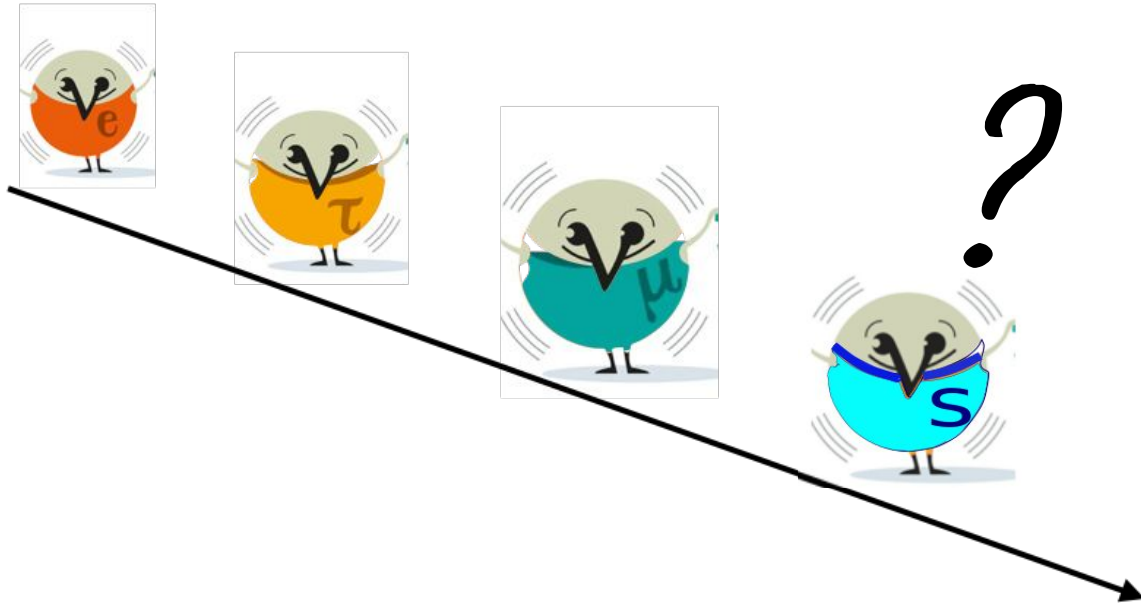
can later be detected in other flavor ...

or later in original ...

Neutrino Oscillations

Anomalies in other past experiments can be interpreted as oscillations occurring because of new flavor or neutrino

Exciting if true!



MicroBooNE setup

Beam created as almost entirely muon neutrinos

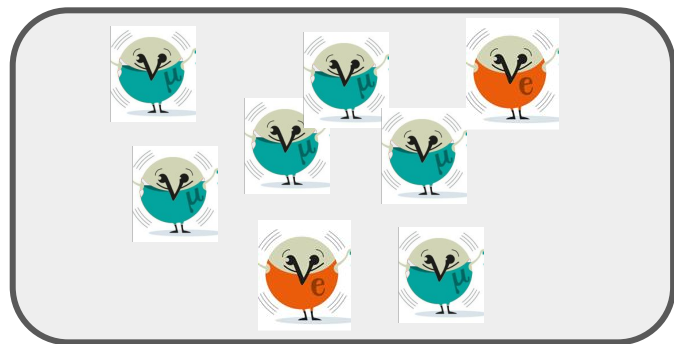
Given distance traveled and energy of neutrinos,
Should be the case @ detector



MicroBooNE setup

But if consistent with past
Anomaly (measured in same
beam line)

We will see **excess** of
electron neutrinos
(at low energies)



MicroBooNE setup

Our target measurement, then, is the energy of neutrinos and the counts of the different flavors

The detector, **a liquid argon time projection chamber**, provides high resolution images with which we can extract these measurements

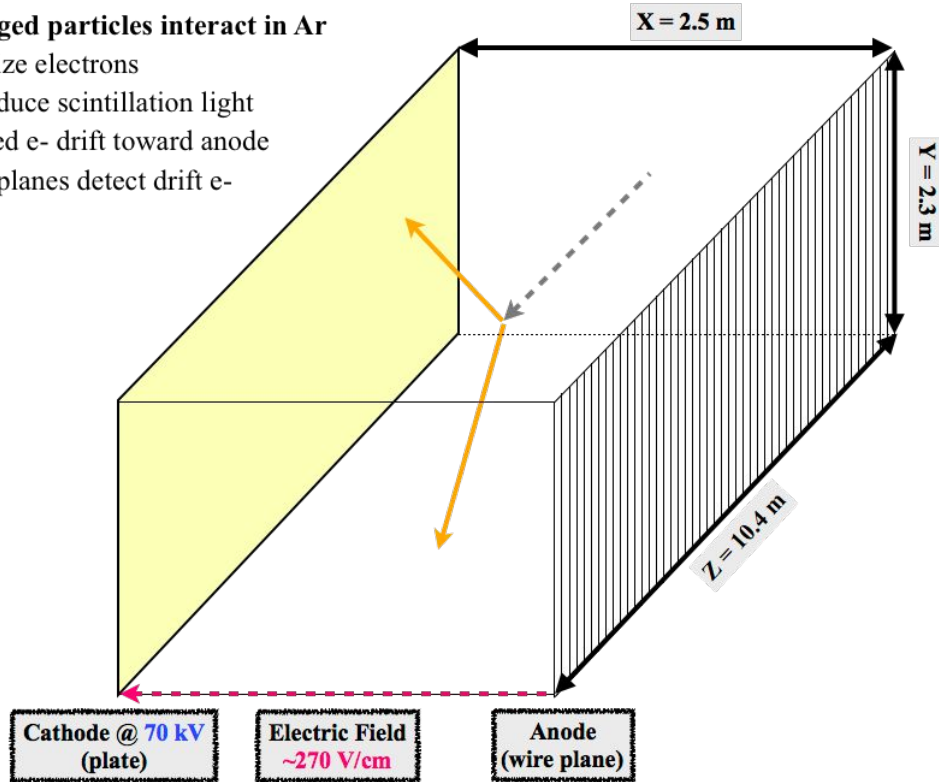
Capturing Images of Neutrino Interactions

1. Charged particles interact in Ar

- Ionize electrons
- Produce scintillation light

2. Ionized e- drift toward anode

3. Wire planes detect drift e-



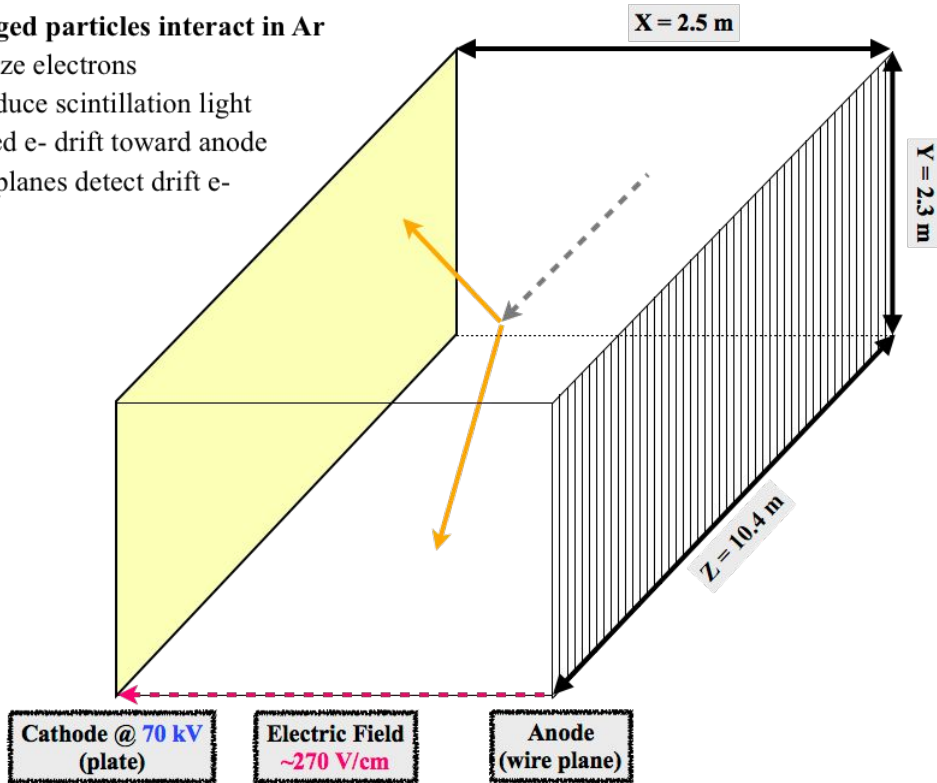
Capturing Images of Neutrino Interactions

1. Charged particles interact in Ar

- Ionize electrons
- Produce scintillation light

2. Ionized e- drift toward anode

3. Wire planes detect drift e-



A neutrino (dashed grey) passes into the detector and interacts producing charged particles (solid yellow)

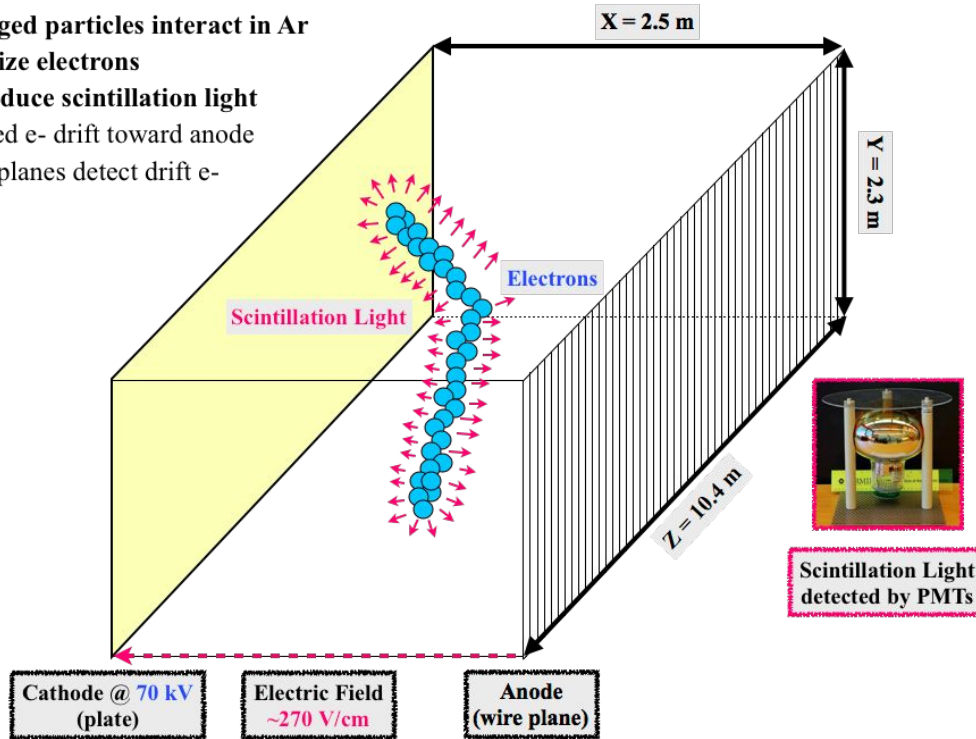
Capturing Images of Neutrino Interactions

1. Charged particles interact in Ar

- Ionize electrons
- Produce scintillation light

2. Ionized e- drift toward anode

3. Wire planes detect drift e-

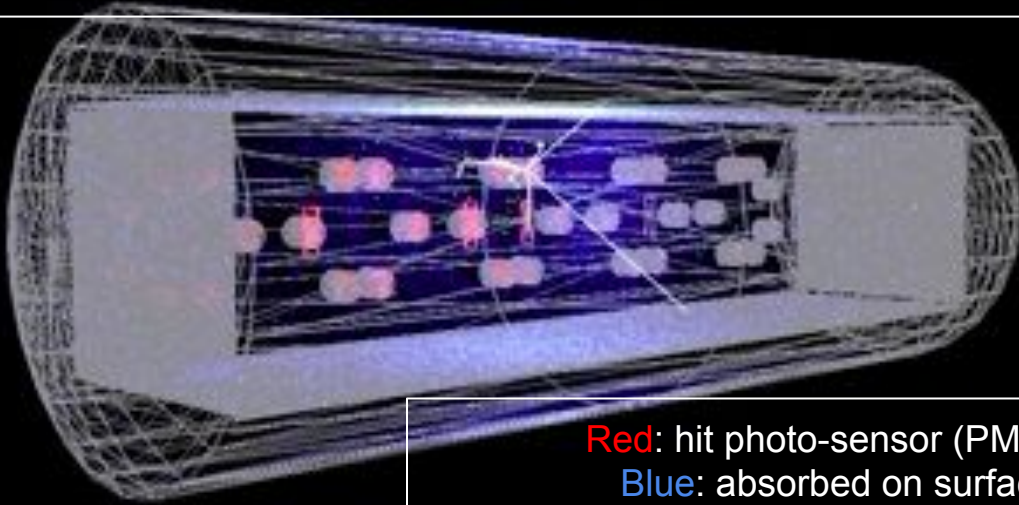


Charged particles produce ionization electrons along path

(neutrino neutral and leaves no direct signature)

Capturing Images of Neutrino Interactions

GPU-accelerated photon simulation, showing final location of photons



Red: hit photo-sensor (PMT)
Blue: absorbed on surface
Green: absorbed in bulk
White-lines: particle trajectories

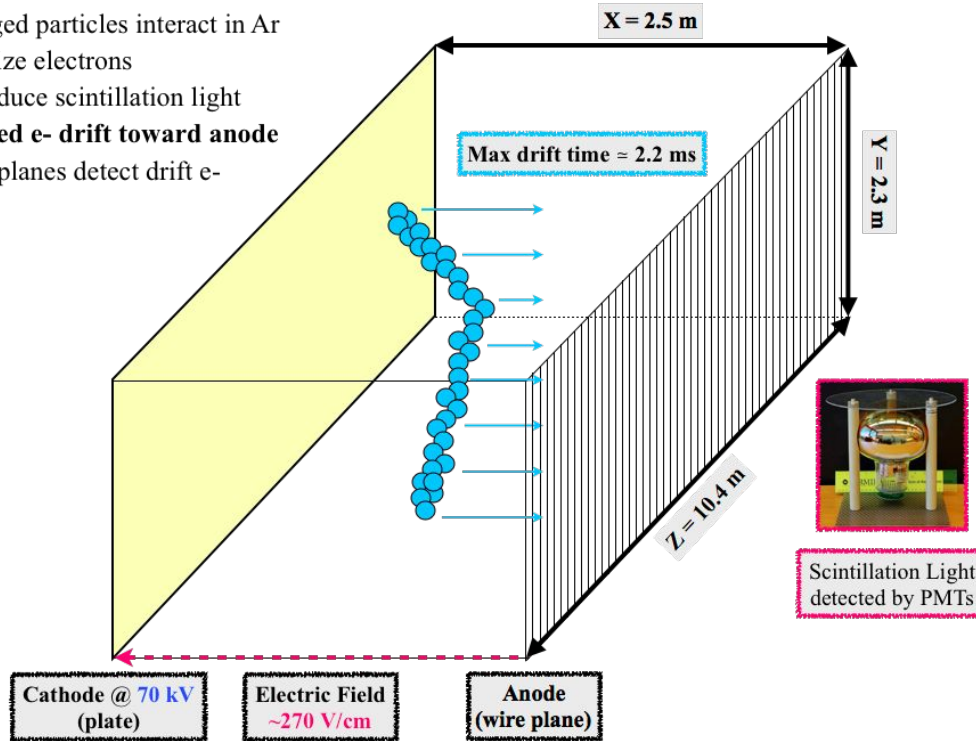
Light also produced by charged particles.

Travels to sensors on short (ns) timescales

Light provides timing for event -- and course position info.

Capturing Images of Neutrino Interactions

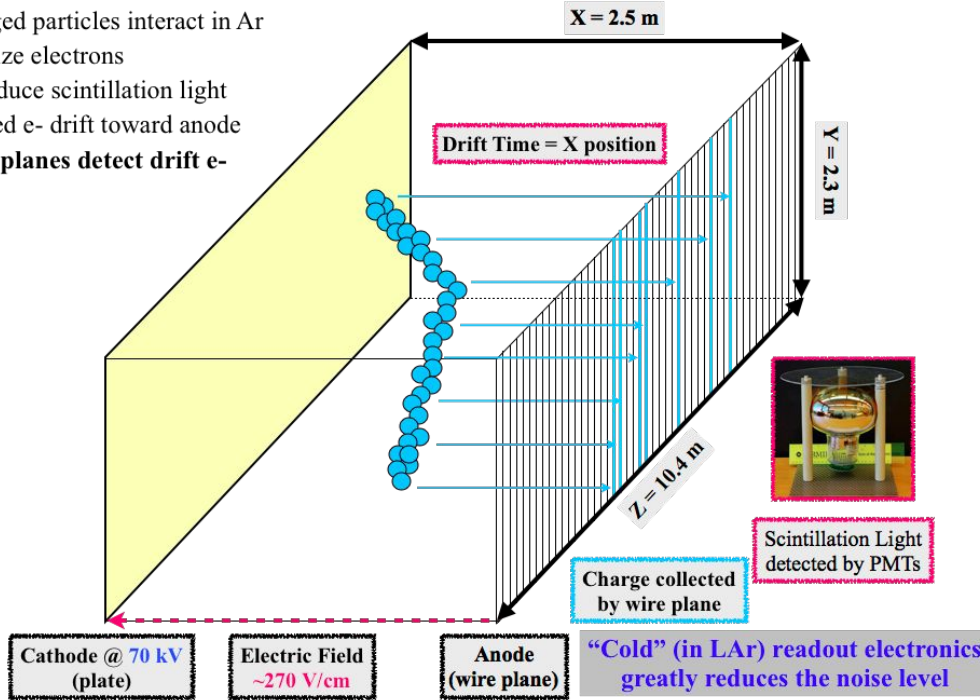
1. Charged particles interact in Ar
 - Ionize electrons
 - Produce scintillation light
2. Ionized e- drift toward anode
3. Wire planes detect drift e-



Ionization electrons
drift towards
wireplanes

Capturing Images of Neutrino Interactions

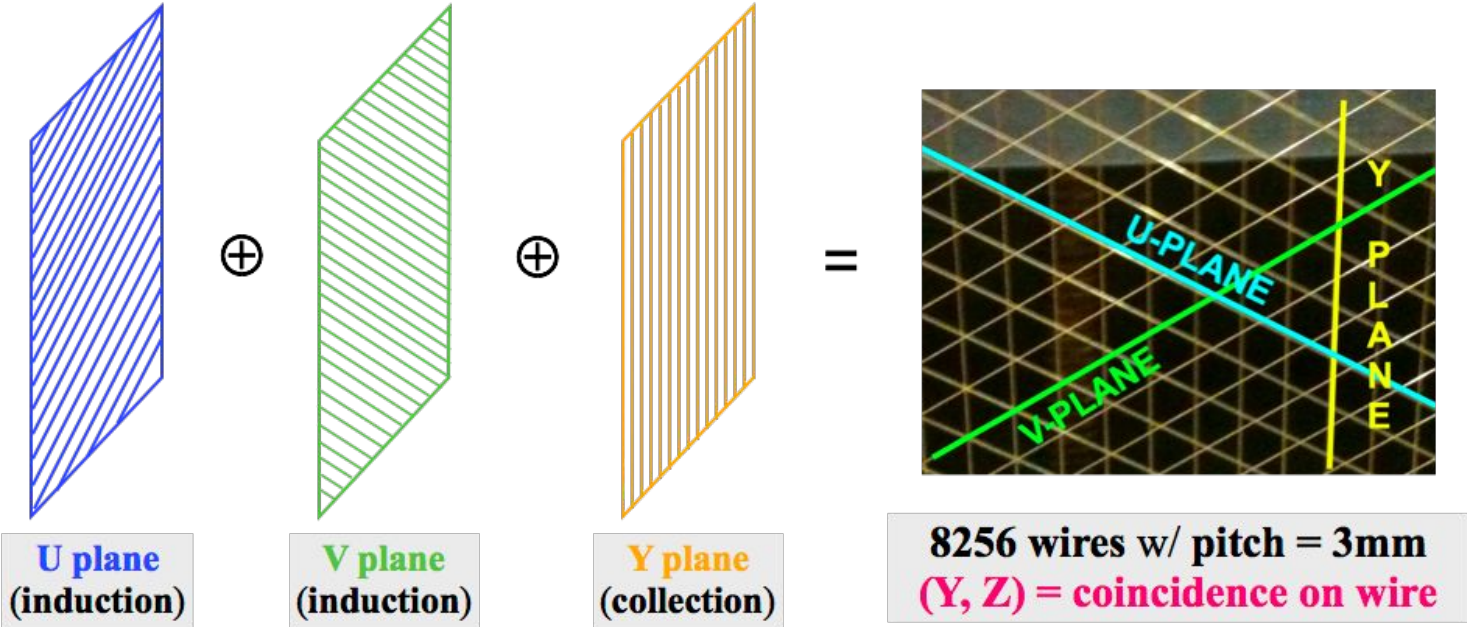
1. Charged particles interact in Ar
 - Ionize electrons
 - Produce scintillation light
2. Ionized e- drift toward anode
3. Wire planes detect drift e-



Ionization induce detectable signals on nearby wires

Capturing Images of Neutrino Interactions

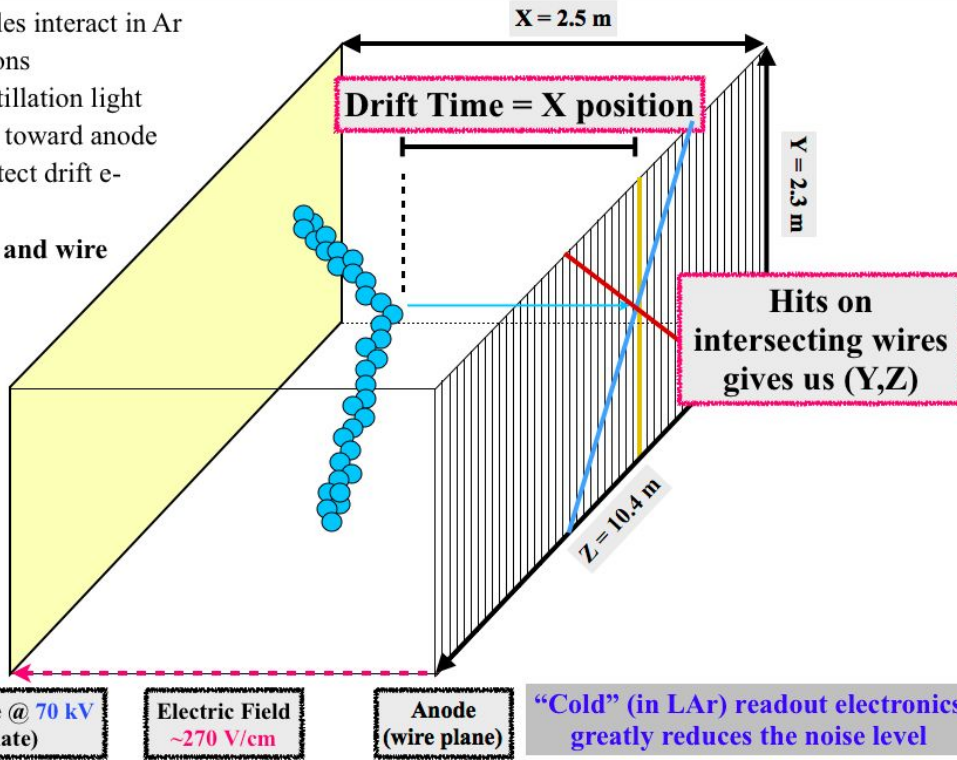
Three Wire Planes



Capturing Images of Neutrino Interactions

1. Charged particles interact in Ar
 - Ionize electrons
 - Produce scintillation light
2. Ionized e- drift toward anode
3. Wire planes detect drift e-

Combine timing and wire info. to get 3D reconstruction.

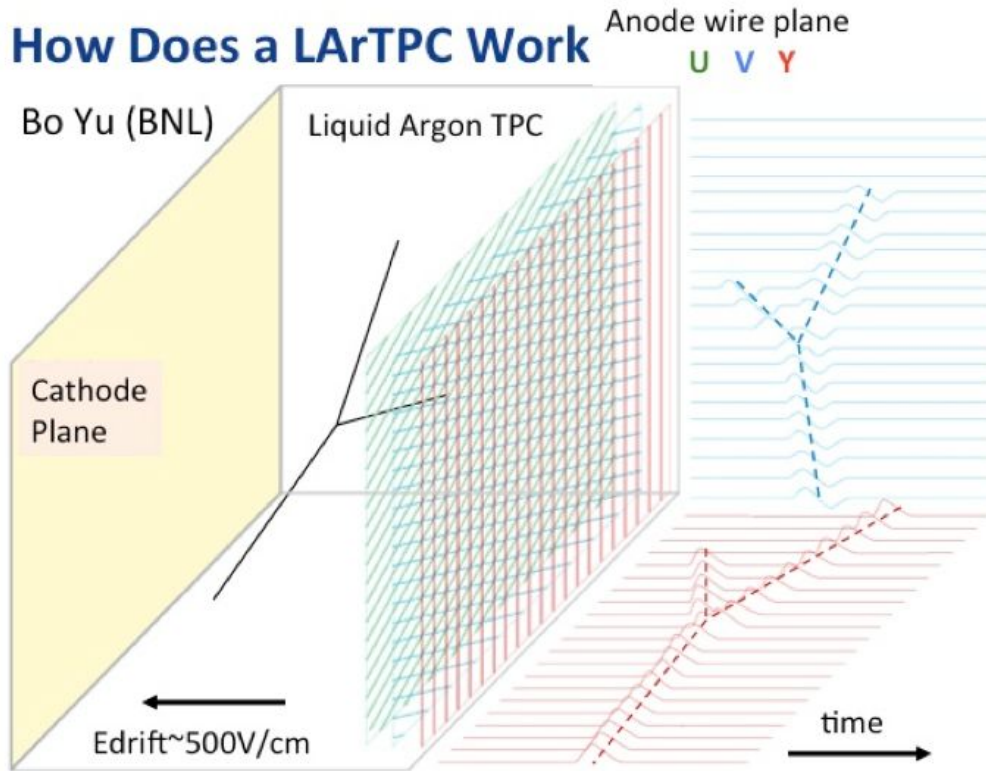


In principle: enough information for 3D reconstruction

(Y,Z) position of ionization recorded through coincident signals on different wire planes

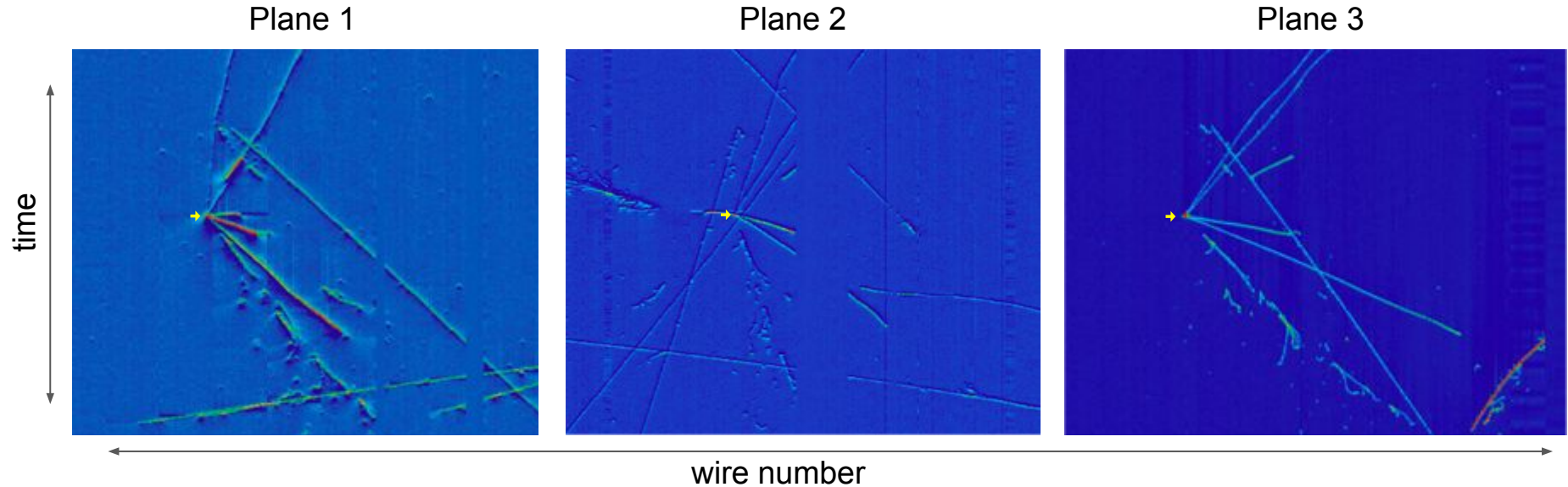
X position give by time delay from light signal

Capturing Images of Neutrino Interactions



Recording wire signals over time, detector produces image-like data

Capturing Images of Neutrino Interactions



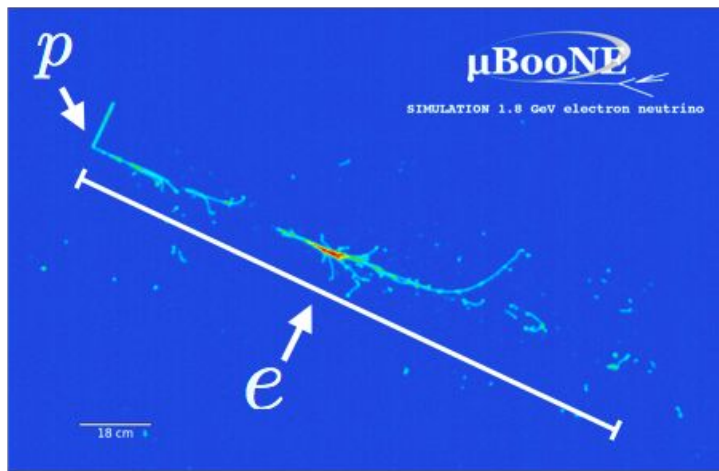
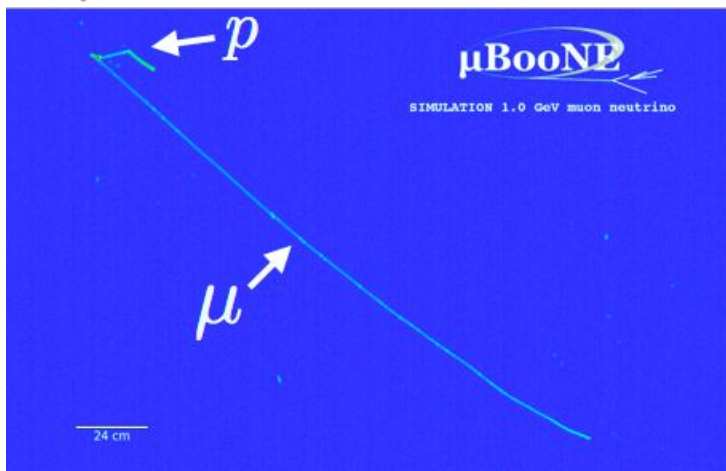
Example of data event in MicroBooNE. View of same event for each projection.

Color scale indicates amount of ionization electrons seen on wire at given time

Measurables in the image

Flavor determined from finding partner lepton (muon, electron) produced in interaction

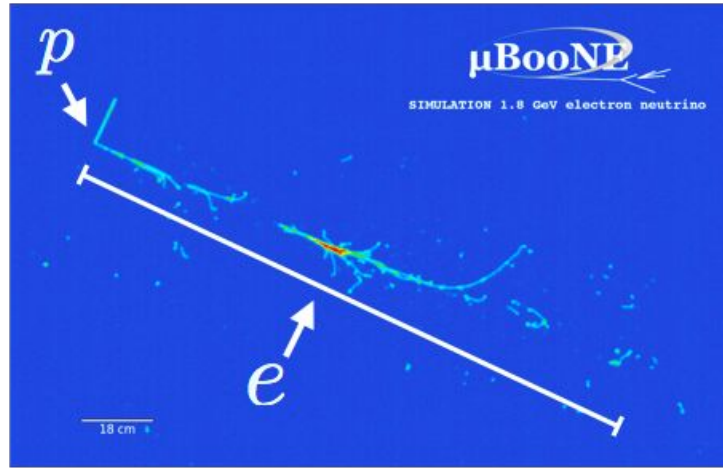
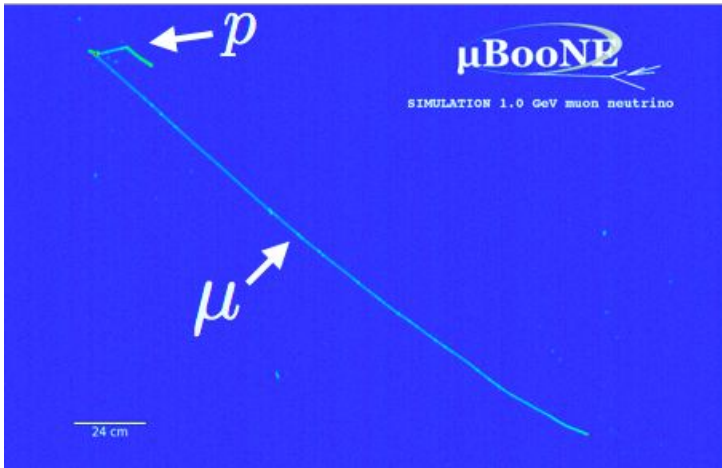
Neutrino energy inferred from momenta of resulting particles



Measurables in the image

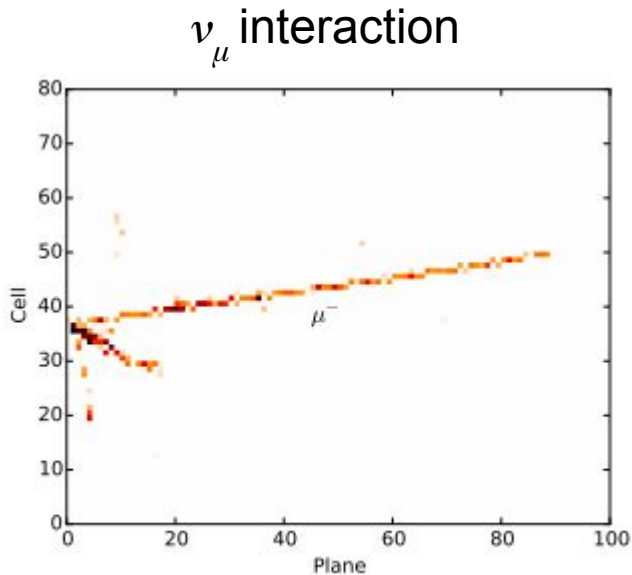
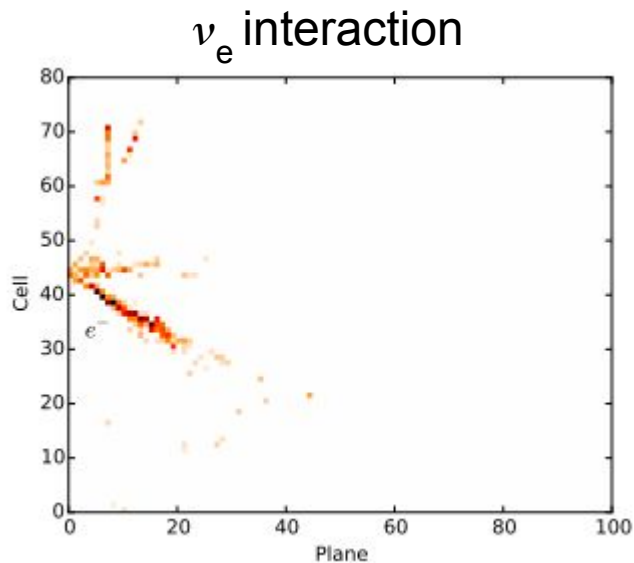
From ML perspective: **the problem is predicting class and energy from images**

We have simulations that can produce training examples with various labels



One approach: directly predict values

Example: CNNs have been used (another experiment) to predict the neutrino flavor



[Aurisano, A. et al. JINST 11 \(2016\) no.09](#)

One approach: directly predict values

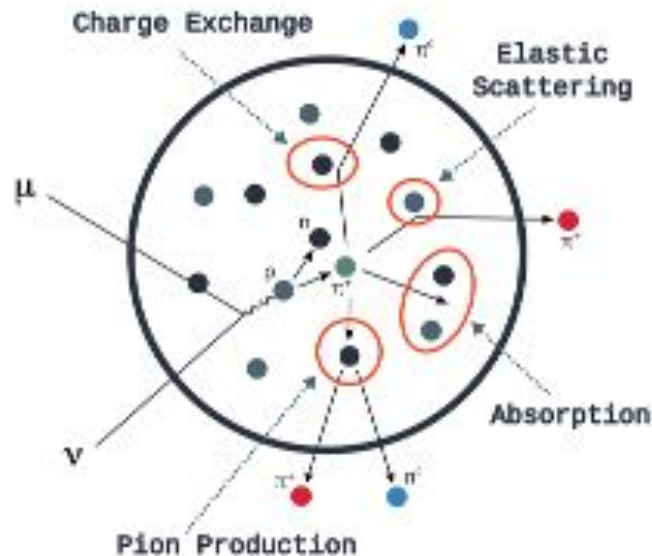
Issue with training network to target final observable: modeling uncertainties

Neutrinos hit constituents of the nucleus -- an extremely complicated system to model

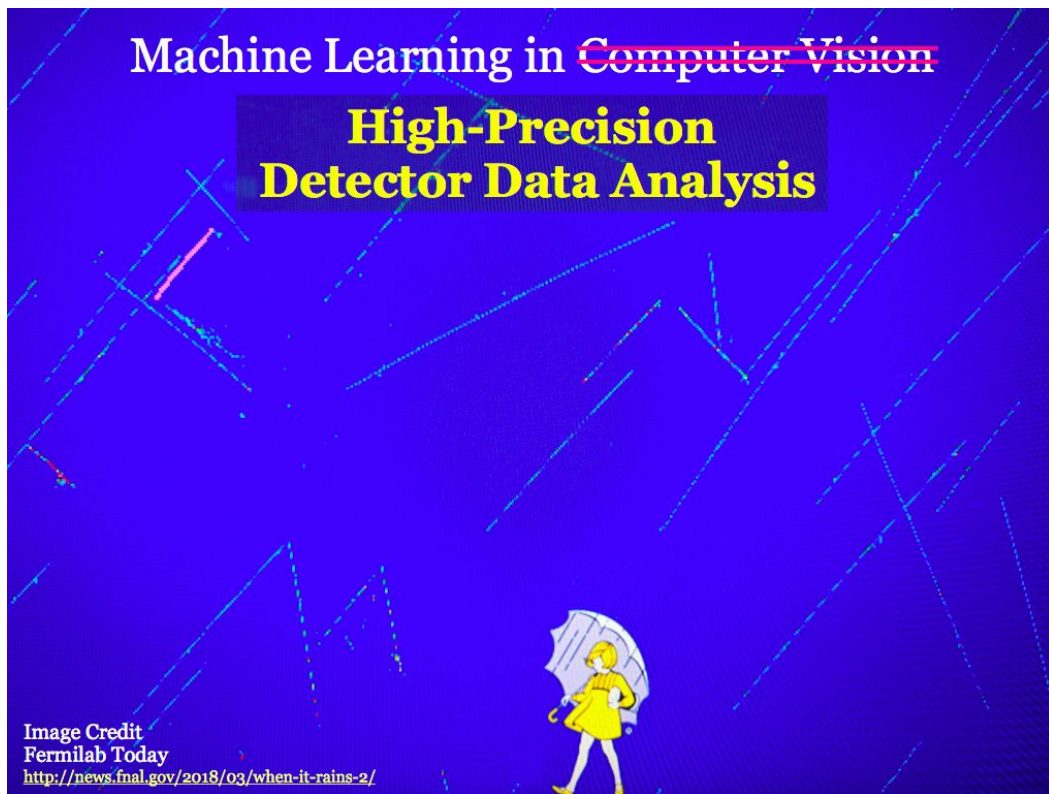
Neutrino energy inference influenced by knowing “type” of interaction

Can produce particles that cause patterns which fake the signal (primary e.g. photons from interaction can look like electrons)

Risk of model errors being trained into the networks



MicroBooNE specific issue: backgrounds



For MicroBooNE, lots of backgrounds from cosmic rays since detector is on the surface

Requires parsing of image to find neutrinos

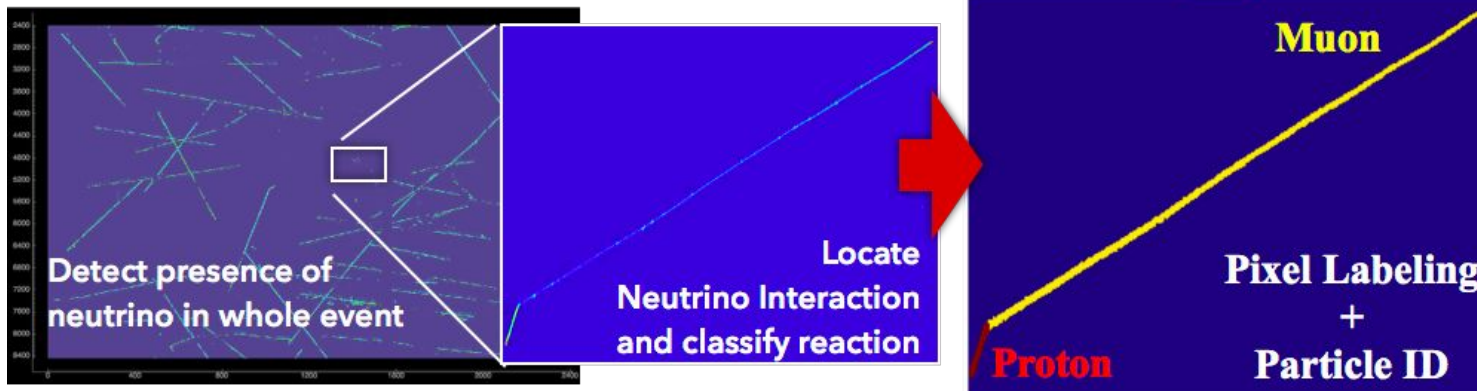
Images with target events only
1 in 1000-10,000

DL Reconstruction

Our goal is to produce constituent particles. Many techniques available

- Need object detection
- Classification

Evaluation with respect to analysis performance important

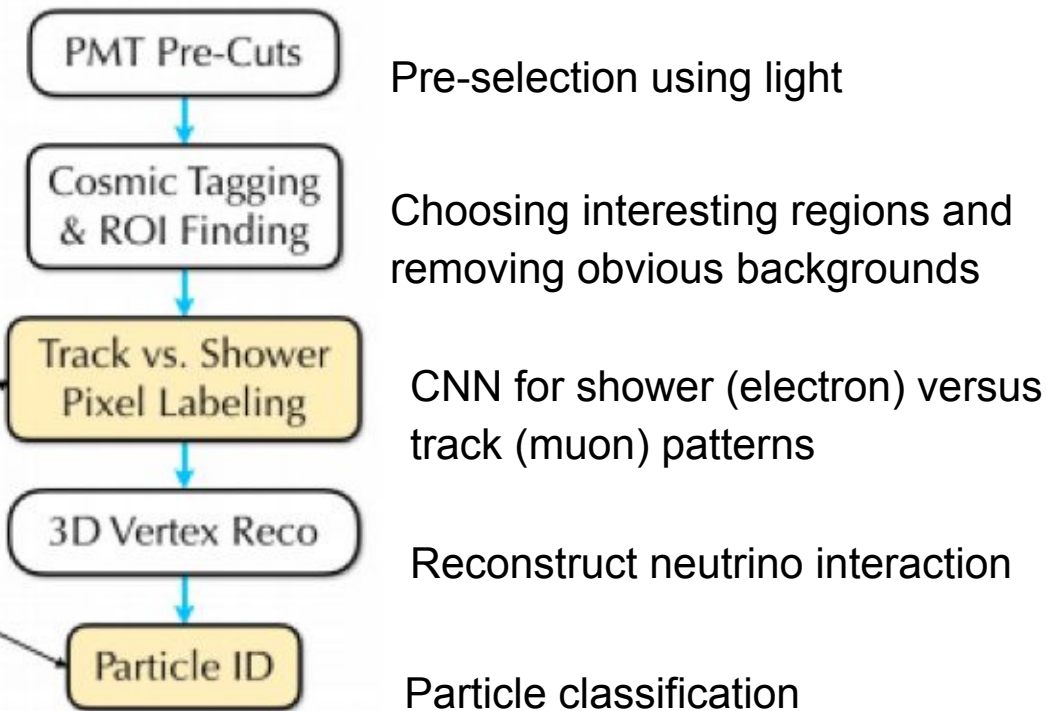


Current Analysis

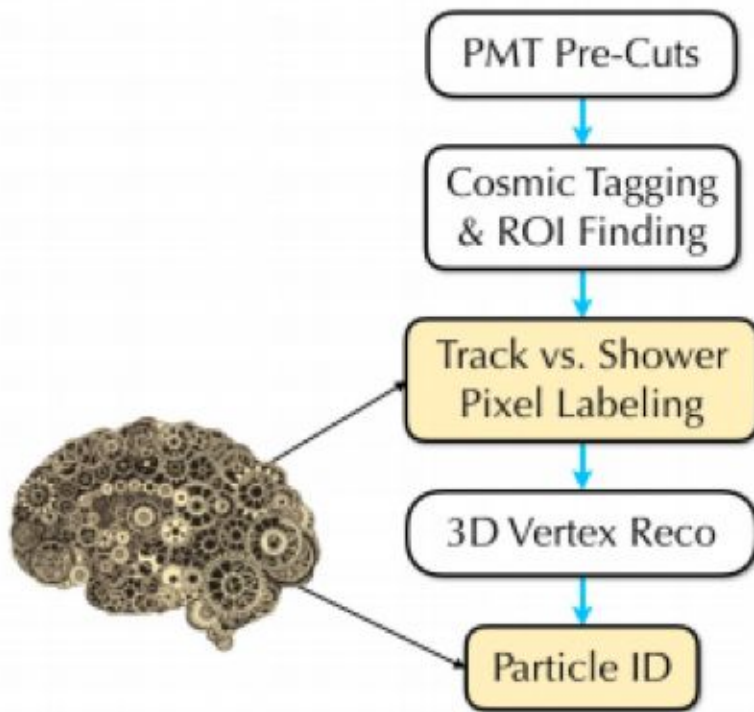
Have built a full reconstruction/analysis chain to search for the oscillation signal

A mixture of CNNs and traditional algorithms

Working analysis benchmarks improvements from CNN techniques



Current Analysis



First applications chosen as they are techniques where we could use non-signal data to evaluate network behavior on real data

Preparing the data

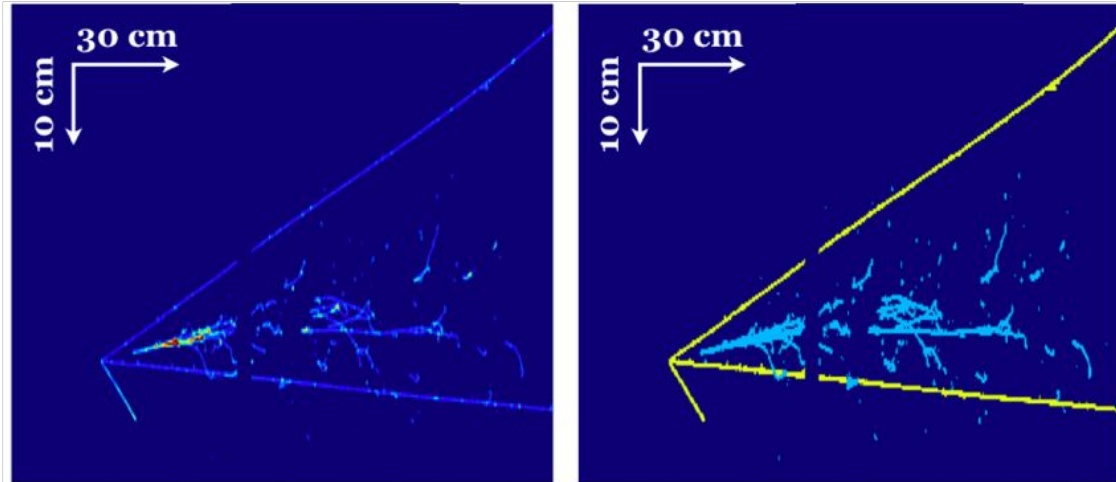
- Images preparation:
 - Noise filtering
 - pulse finding + zero suppression
 - Deconvolve wire response
 - Accounting for electronics response + expected induced signal
 - Downsample in time (summed) by factor of 6
- 3D consistent cropping
 - Full size: 3456 (wire) x 6448 (ticks)
 - Downsampled size: 3456 x 1008 -- both dimensions about 3 mm
 - Cropped into 832 wire x 512 ticks (24 images per plane)

Pixel labeling

In reconstructing events, useful to be able to separate two types of patterns: tracks and showers

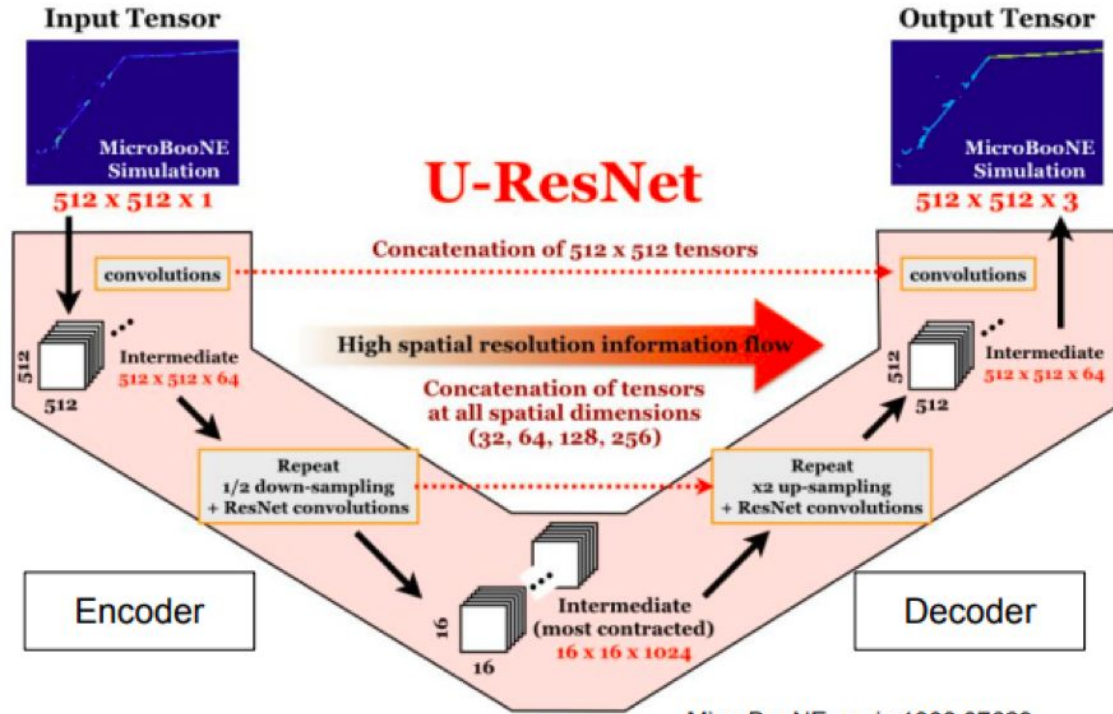
The Goal

yellow: track
cyan: shower



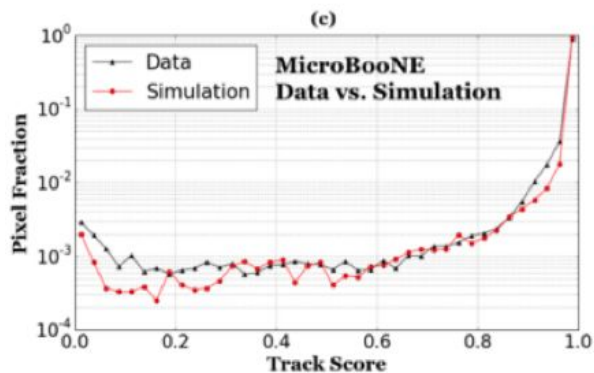
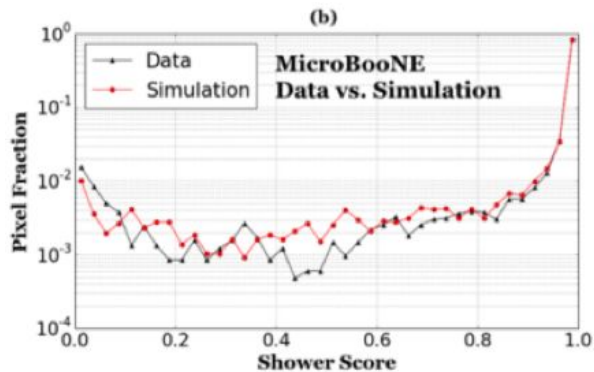
Pixel labeling

We use a U-Net for this problem

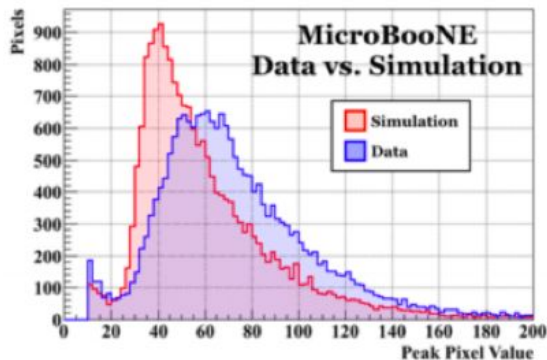


MicroBooNE: arxiv:1808.07629

Pixel labeling: behavior on real versus sim. images



- Sample: stopping muons
- Score distributions similar
- Robust to moderate difference in images as shown by peak pixel distributions

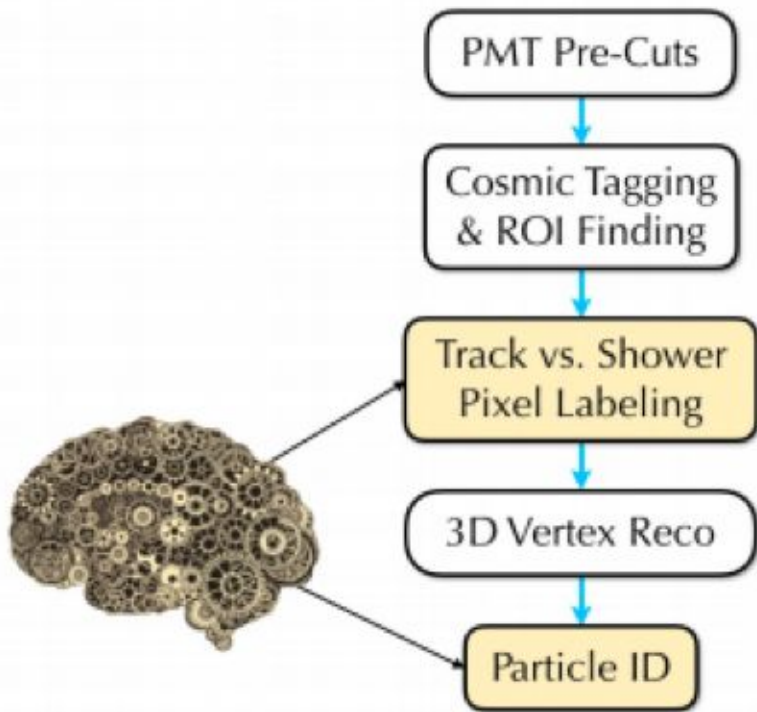


These are cosmic particles that come to rest in the detector

Mostly muons, many of which decay into electrons

Use to check track and shower labeling

Analysis Status



Analysis components complete

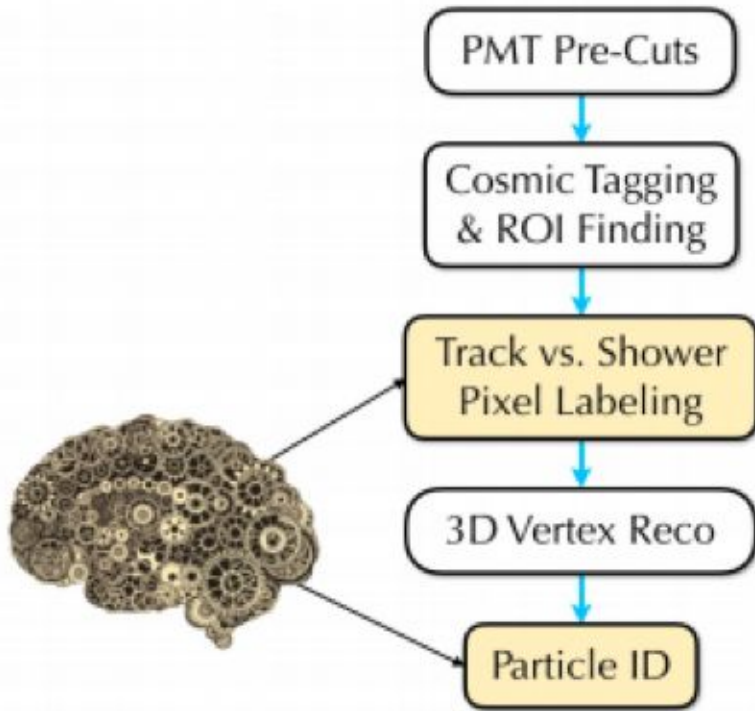
Evaluating:

- Data versus sim. differences through distributions of particle kinematics
- Sensitivity of analysis to see anomalous signal (or excluding it)

Not yet ready

- Hitting various performance milestone with simulation dataset
- Aim is to release result within year

Further CNN techniques in the works



Also tackling more parts of the reconstruction chain

Finding and removing non-neutrino tracks

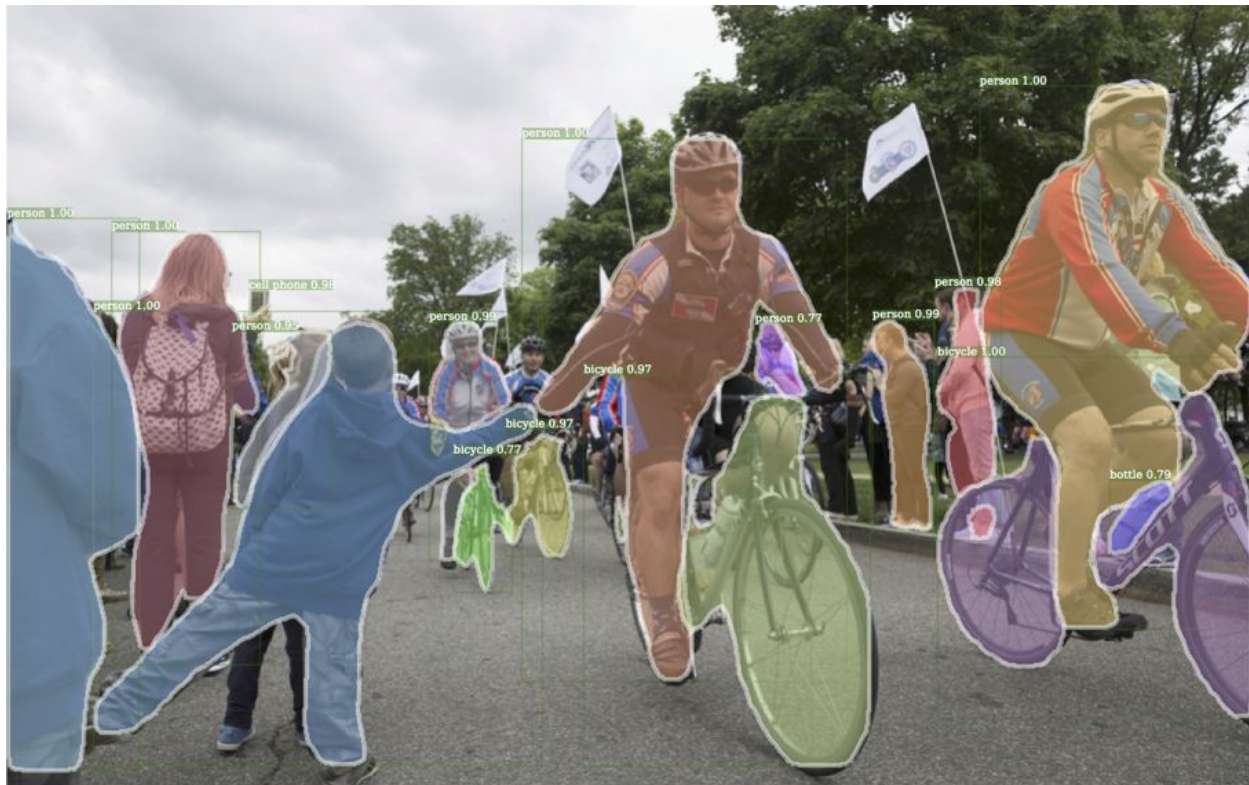
Providing 3D spacepoints to perform 3D reconstruction at earlier stage

“Repairing” images to assist track reconstruction

Applying Instance Aware Segmentation

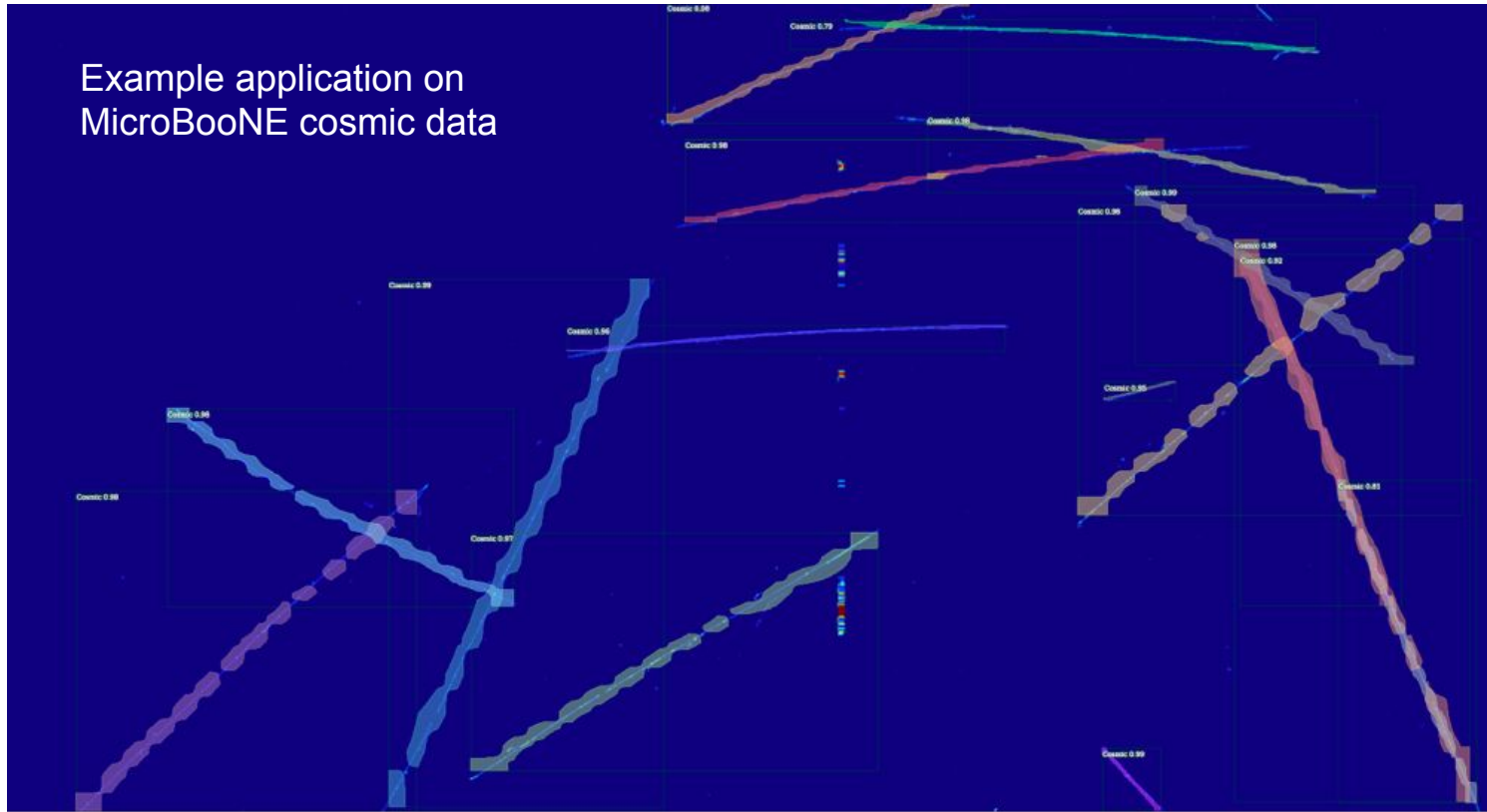
Currently
adapting
Detectron

Mask-RCNN
network



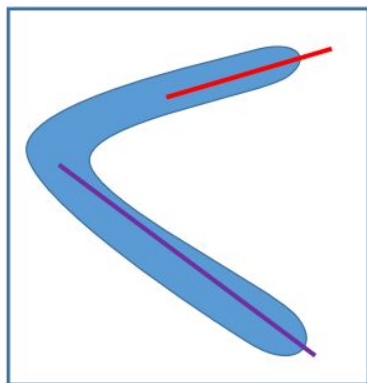
Mask R-CNN for cosmic detection and rejection

Example application on
MicroBooNE cosmic data



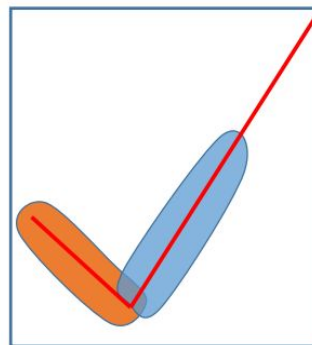
Mask R-CNN: evaluating performance

$$\text{Purity} = \frac{\text{Sum}[(\text{Prediction}) \times (\text{Ground Truth}) \times (\text{ADC Binary})]}{\text{Sum}[(\text{Prediction}) \times (\text{ADC Binary})]}$$



- Ground Truth 1
- Ground Truth 2
- Prediction Mask

$$\text{Efficiency} = \frac{\text{Sum}[(\text{Prediction } \mathbf{Union}) \times (\text{Ground Truth}) \times (\text{ADC Binary})]}{\text{Sum}[(\text{Ground Truth}) \times (\text{ADC Binary})]}$$



- Ground Truth
- Prediction 1
- Prediction 2

Mask R-CNN: evaluating performance

- MCC 8 Simulation
- y – Plane
- Log Z Axis
- Specialized Epochs correspond to training on a dataset with more overlapping clusters



Eff vs Pur 1.75 Epochs, 7.976 Specialized Epochs

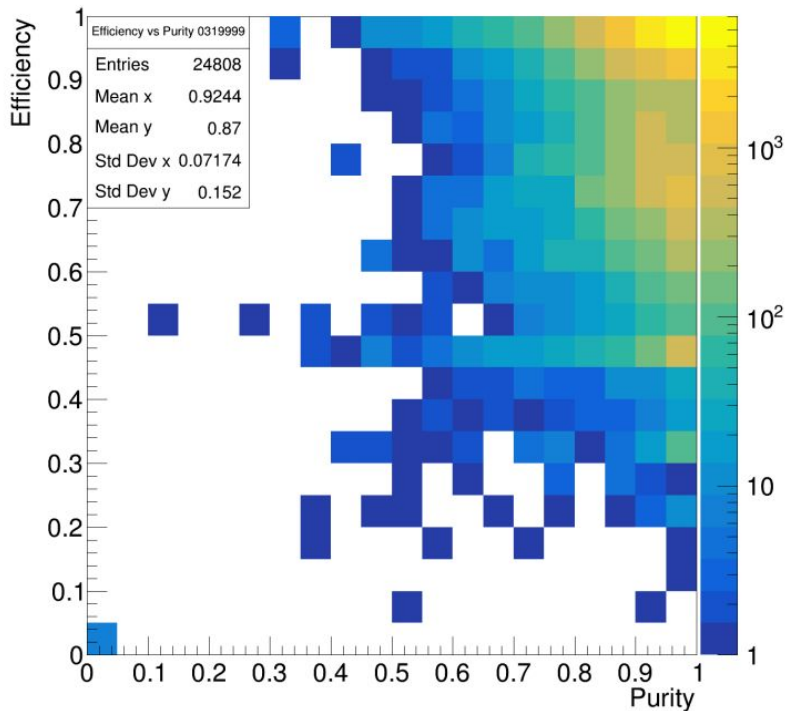


Image Repair and Tracking

- In industry: filling in blanked-out regions in images
- Using a similar idea to fill in missing parts of track in MicroBooNE
- Useful for 3D track reco (trajectory only, not calorimetry)

<http://arxiv.org/pdf/1804.07723.pdf> NVIDIA Corporation

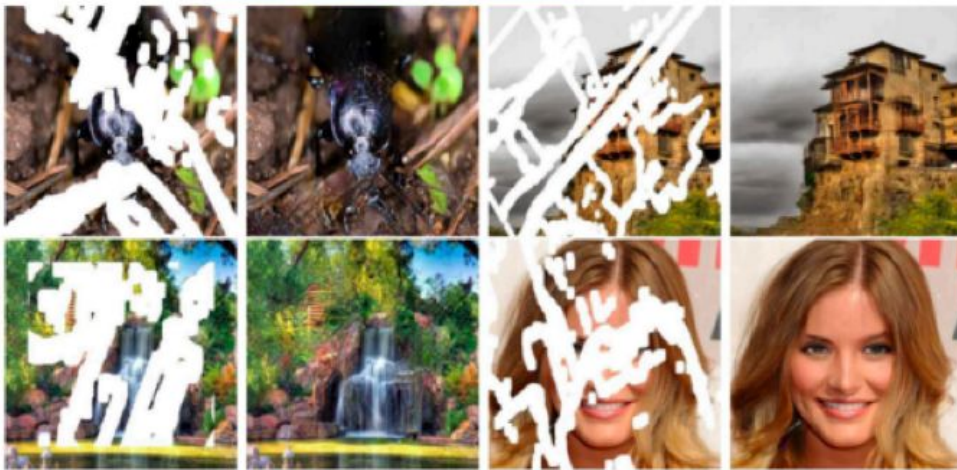
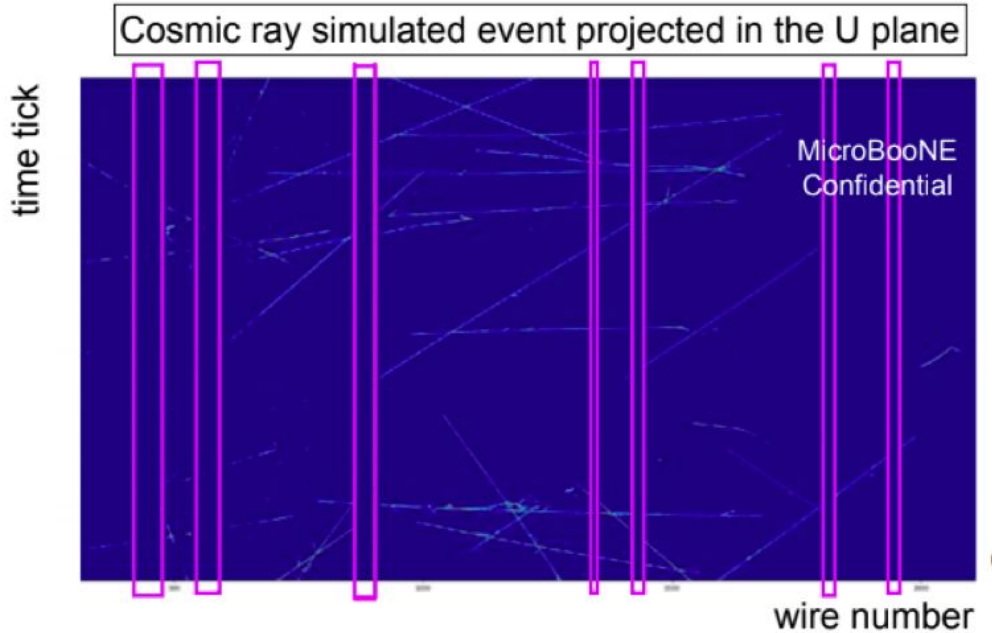


Image Repair and Tracking

- Tracking: clustering of continuous clusters of 3D points
 - difficult in regions w/ dead readout channels: track gaps

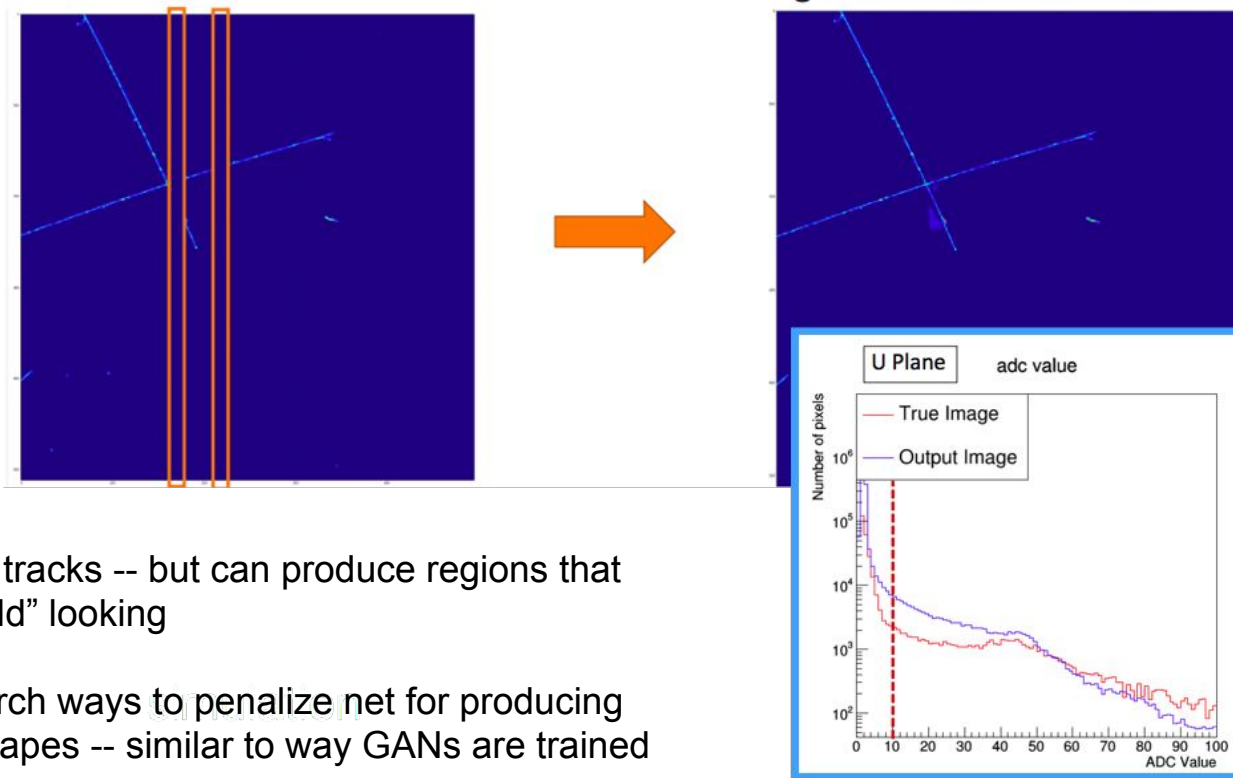


Gaps impair ability to track reconstruction to accurately get momentum

Currently try to detect when track ends in gap and remove events

Overcoming this can also help with efficiency

Image Repair and Tracking



Fills in tracks -- but can produce regions that are "odd" looking

Research ways to penalize net for producing odd shapes -- similar to way GANs are trained

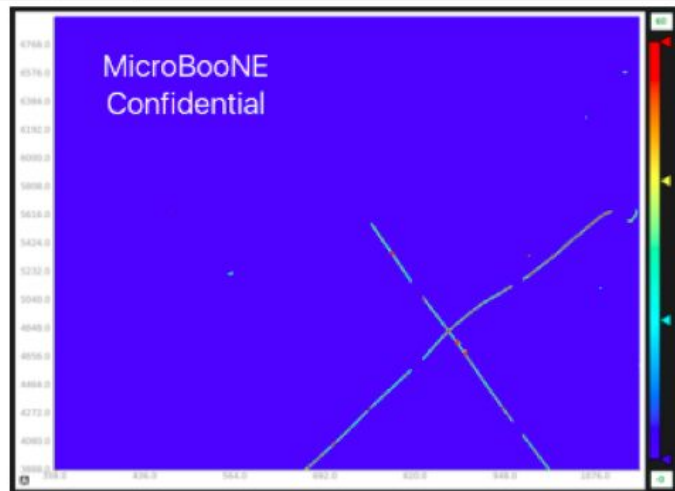
3D Space Points

- To reconstruct 3D position of a charge deposit: need to match charges in same time window on at least 2 wire planes
- 3D position from wire intersection

Charge depositions in U plane



Charge depositions in same time window in Y

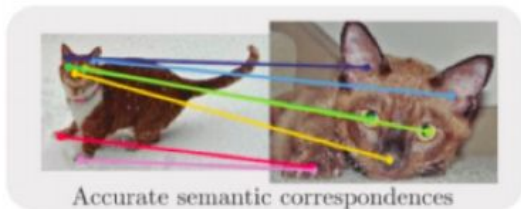


wire number

wire number

3D Space Points

- Goal of dense pixel correspondence: match regions of one image to another, connecting semantically similar items



Choy et al. "Universal Correspondence Network" NIPS 2016



Zhou, Krähenbühl et al. "Learning Dense Correspondence via 3D-guided Cycle Consistency" CPVR 2016

colors indicate what should be matched

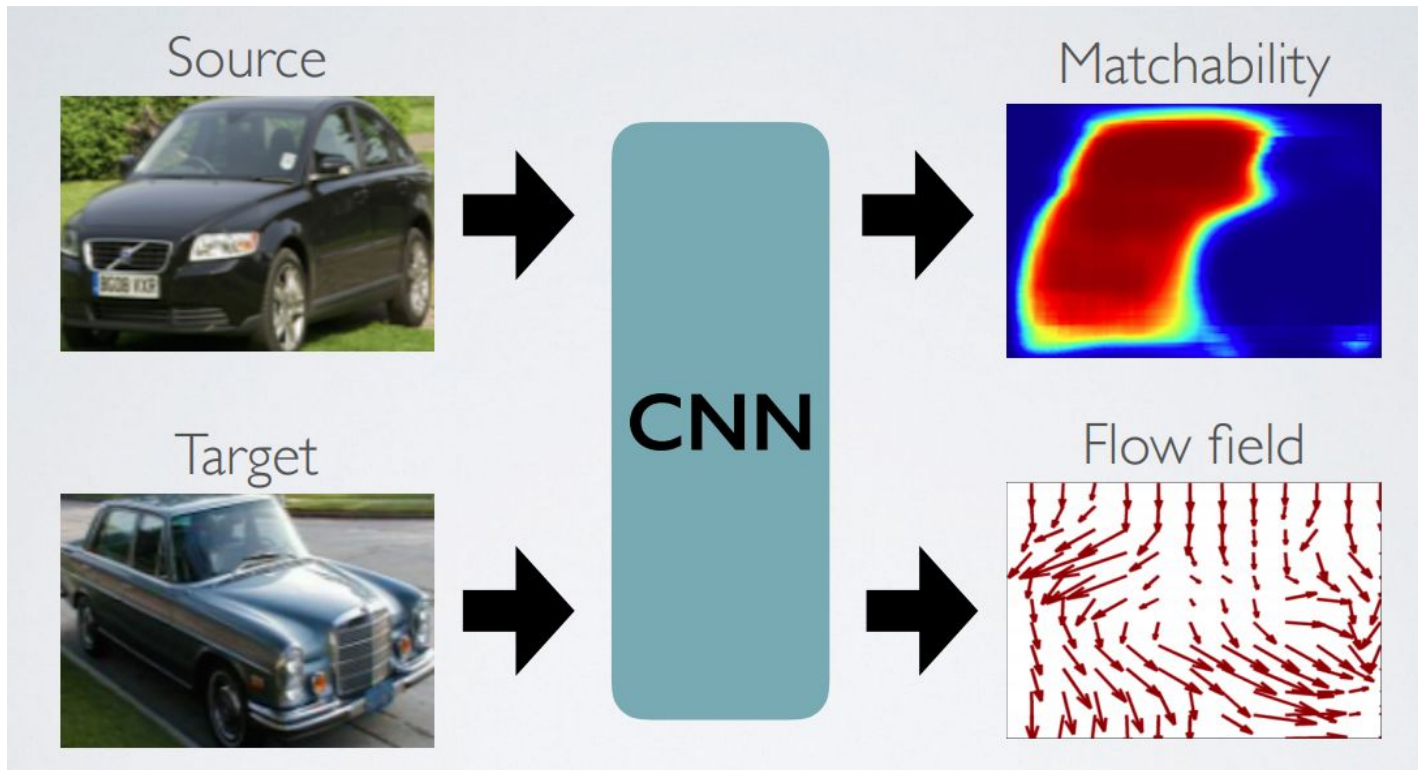
SCAPE



Wei et al. "Dense Human Body Correspondences Using Convolutional Networks" CPVR 2016

37

LArFlow



in LArTPC
context

matchability = 0
when true target pixel in
dead wires, below
thresh, etc.

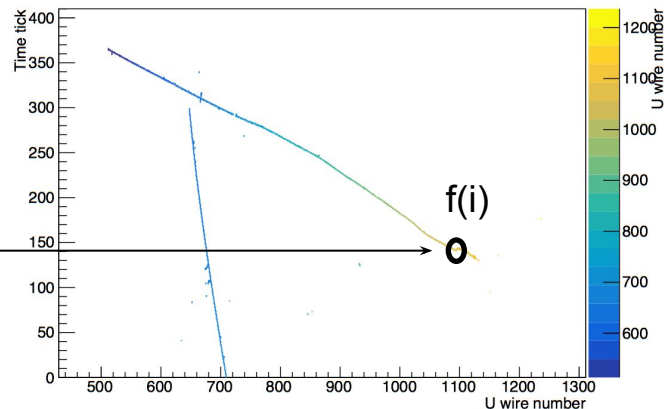
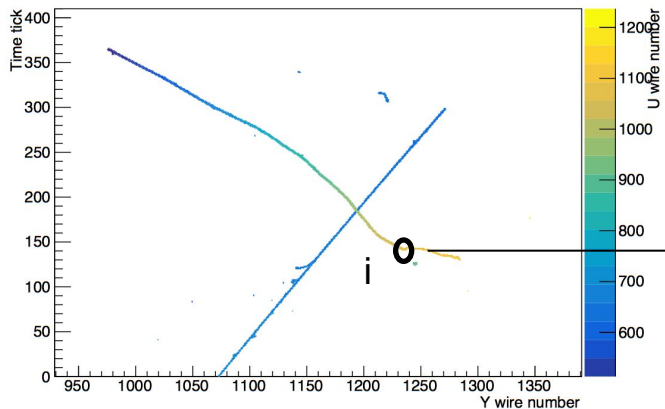
enforce same-time tick,
so only wire-direction
flow predicted

LArFlow

Network predicts correspondence between pixels (charges) in Y, U, V ADC images

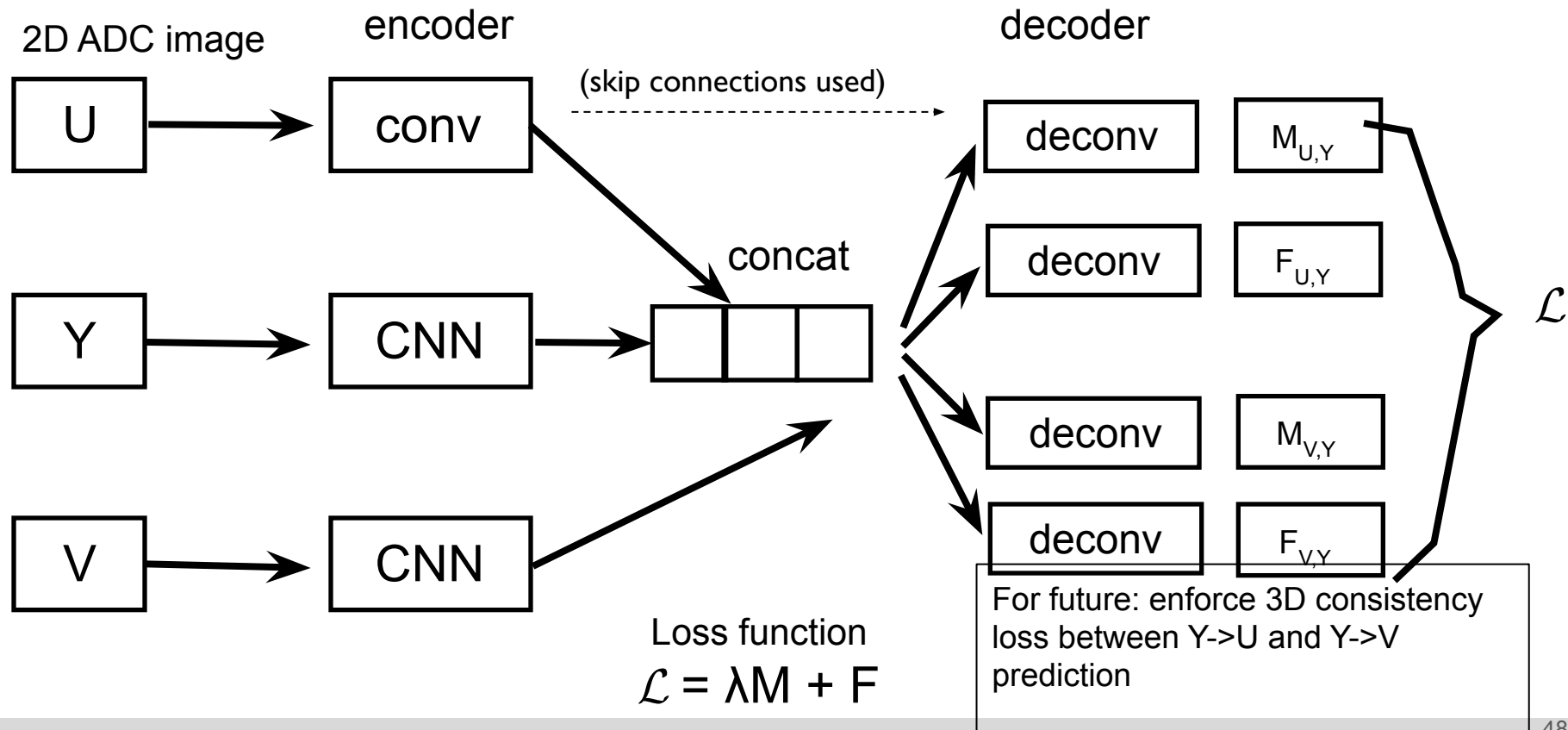
For pixel i in Y plane: the CNN is asked to predict **shift** needed to move to pixel 1175

Which is where the corresponding pixel is in U plane

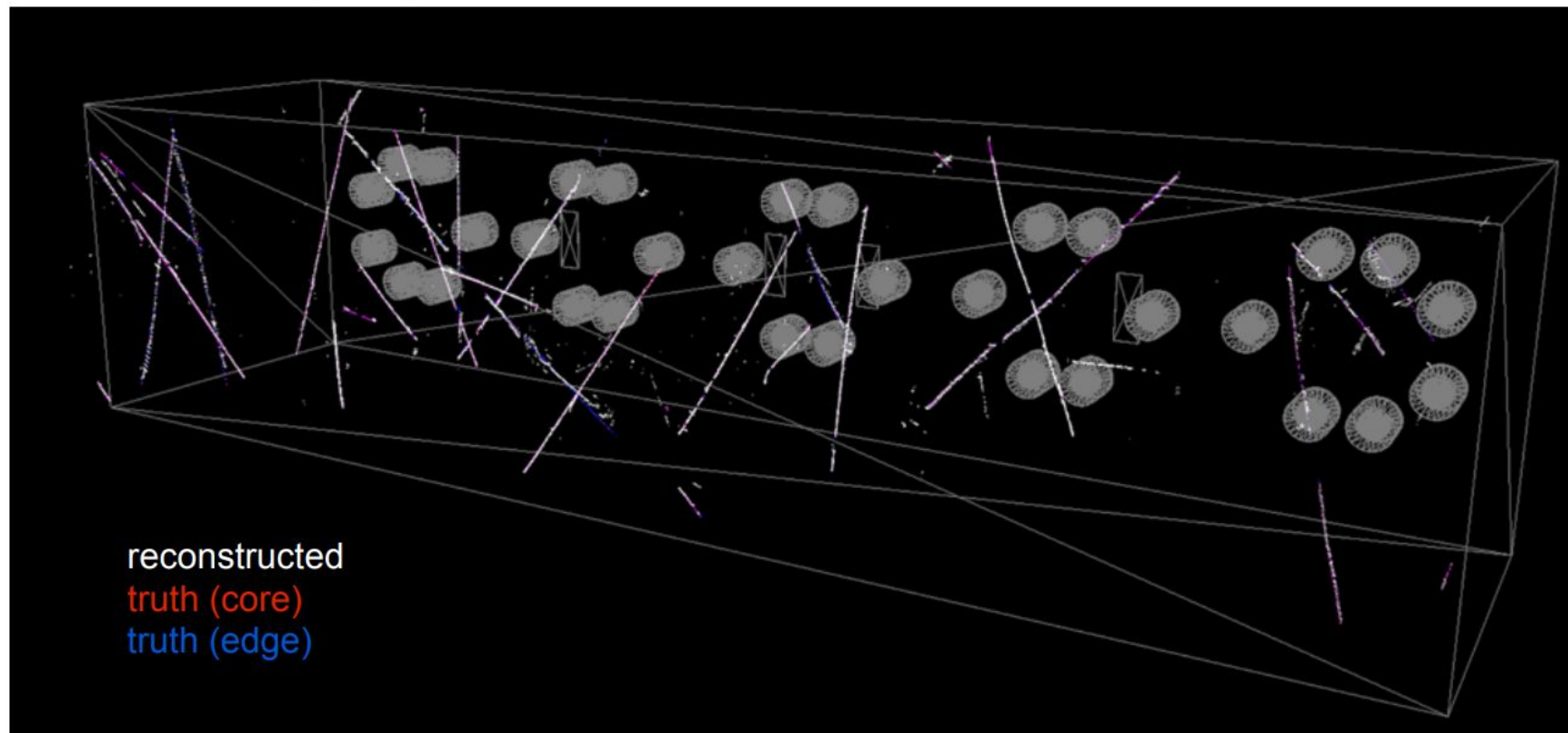


Correspondence prediction gives 3D space-point for that charge

LArFlow: Network

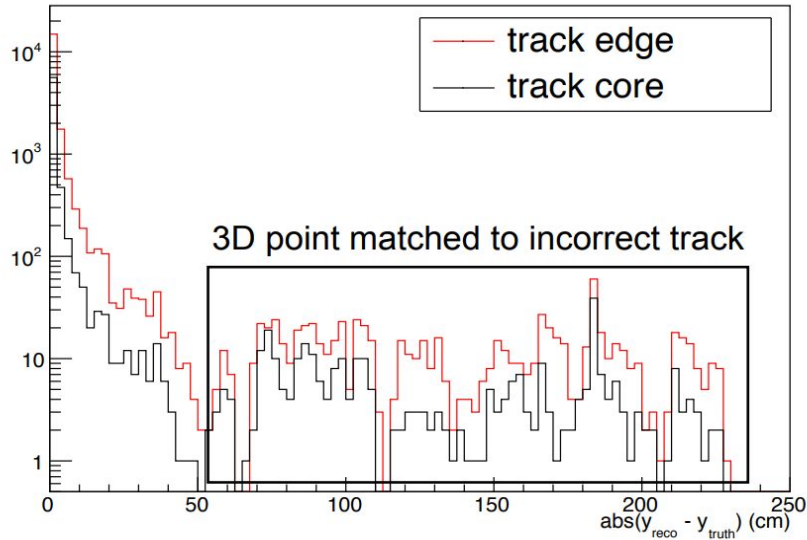


LArFlow: Loss



LArFlow: Initial Performance

Absolute distance in y (cm) between reco and truth



Within 10cm for 92% of hits

Within 50cm for 95% of hits

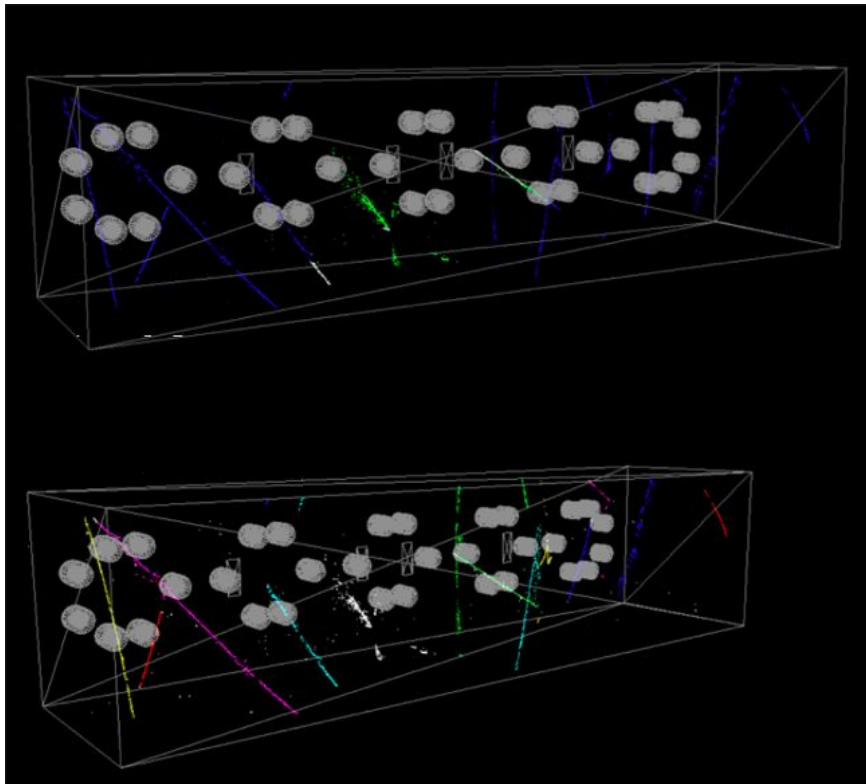
If flow prediction (U or V wire) is wrong, we shift to incorrect y

Have plans to use cosmic muon data to evaluate similar metrics

Good enough for cosmic rejection

Improvements in precision needed for neutrino reconstruction

LArFlow: Initial Performance

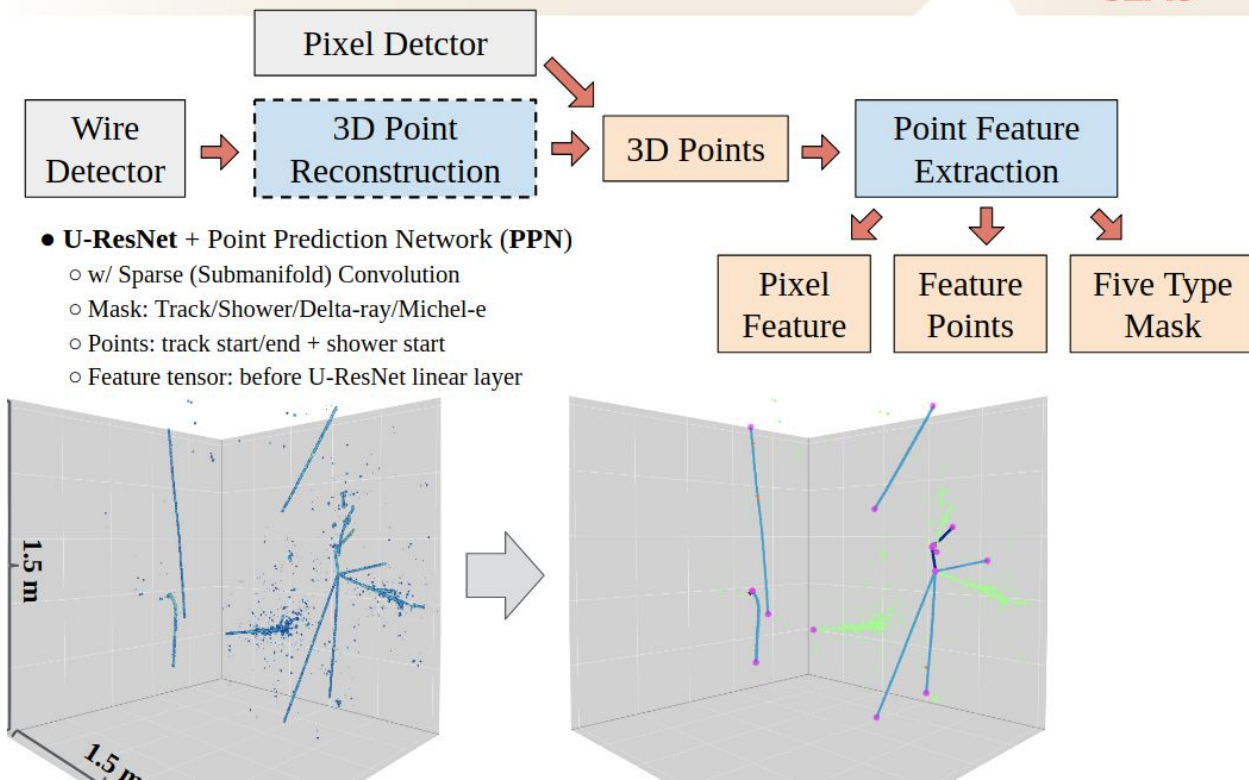


Top: combining 3D points with track/shower labeling

Bottom: using Mask-RCNN network to cluster cosmic muon candidates

Towards 3D space-point reconstruction

Our work is in collaboration with DL-based reco on space-points done here at SLAC



Summary

CNNs well-suited to analysis of LArTPC images

Applications developed in conjunction with physics analyses -- important for knowing effect on ultimate goal

Moving towards an end-to-end reconstruction chain using networks

Stay tuned for analysis result!

Results from work by:



Katie Mason
(grad)



Joshua Mills
(grad)



Ralitsa Sharanova
(post-doc)