



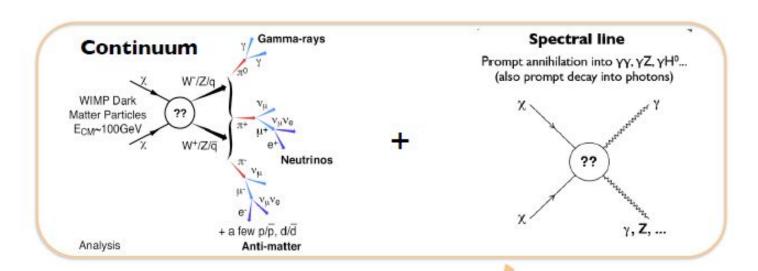
Search for Dark Matter Signal from Galaxy Clusters

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Dark Matter Indirect Detection



Gamma-ray Flux

Particle Physics

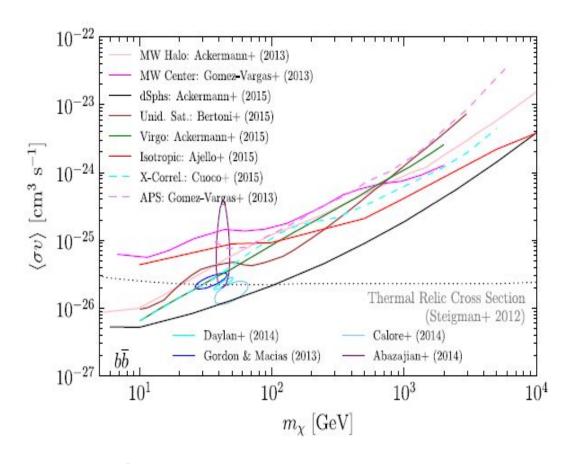
$$\frac{d\Phi_{\gamma}}{dE_{\gamma}}(E_{\gamma}, \phi, \theta) = \frac{1}{4\pi} \frac{\langle \sigma_{ann} v \rangle}{2m_{WIMP}^2} \sum_{f} \frac{dN_{\gamma}^f}{dE_{\gamma}} B_f$$

$$\times \int_{\Delta\Omega(\phi, \theta)} d\Omega' \int_{los} \rho^2(r(l, \phi')) dl(r, \phi')$$

DM Distribution (J-Factor)

[Matthew Wood, Fermi Summer School 2014]

Dark Matter Indirect Detection

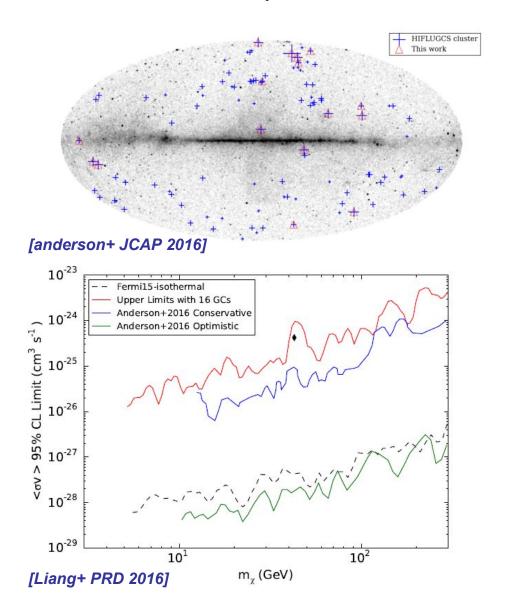


DM search targets:

- Galactic Center
- Milky Way Halo
- Galaxy Clusters
- Dwarf Satellites
- Isotropic Backgroud
- and so on

[Charles+ PhysRep 2016]

Search for a gamma-ray line feature from a group of nearby Galaxy clusters with Fermi LAT Pass 8 data



- search for line-like signals
 from 16 nearby galaxy clusters
- stacking analysis method is used
- improvements may be due to Pass 8 data and stacking method we used
- there is a weak excess around 43 GeV
- more details can be found in Liang+ PRD 2016 (arXiv:1602.06527)

Thanks for your attention!