

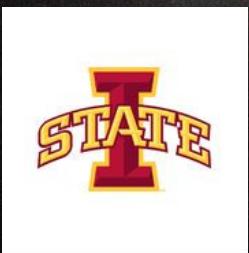
The Starburst Galaxy M82 at Very High Energy (100GeV>)

Use of the FERMI LAT high energy data with
VERITAS Imaging Atmospheric Cherenkov Telescope
observations of M82 to constrain the broad band spectrum

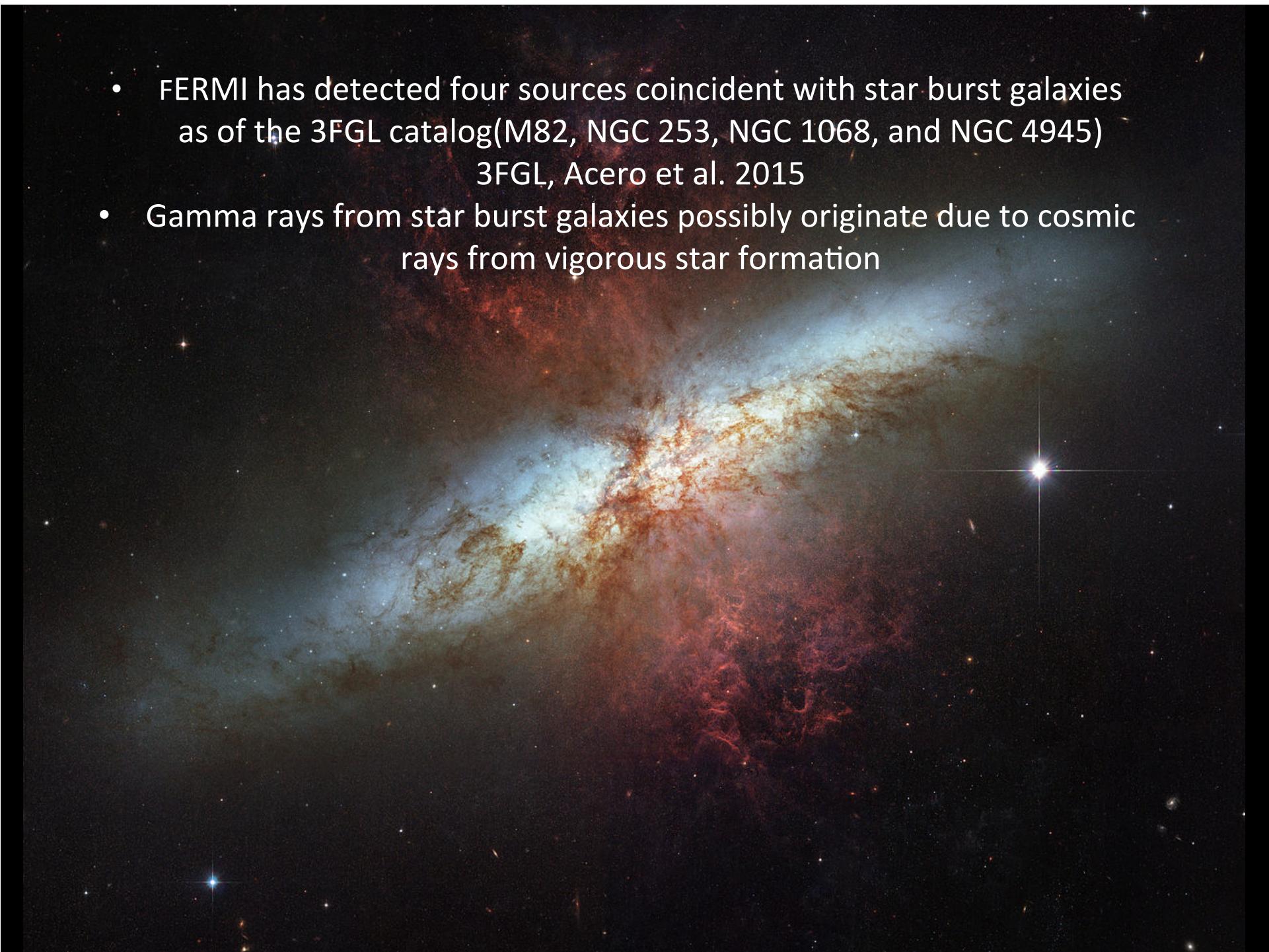
Analysis performed with FERMI Science Tools Binned
Likelihood Method

Alisha Chromeley for FERMI Summer School 2015

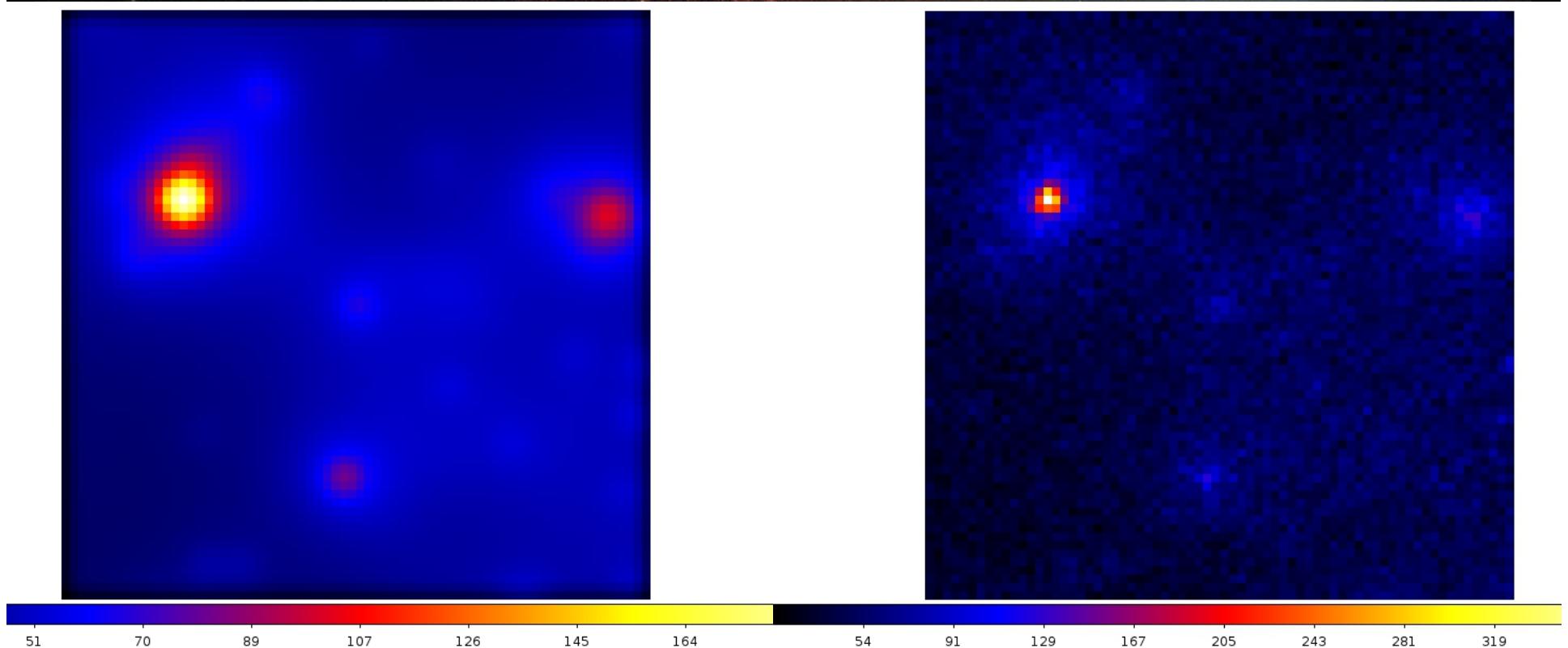
NASA, ESA, and The Hubble Heritage Team (STScI/AURA)



- FERMI has detected four sources coincident with star burst galaxies as of the 3FGL catalog(M82, NGC 253, NGC 1068, and NGC 4945)
3FGL, Acero et al. 2015
- Gamma rays from star burst galaxies possibly originate due to cosmic rays from vigorous star formation

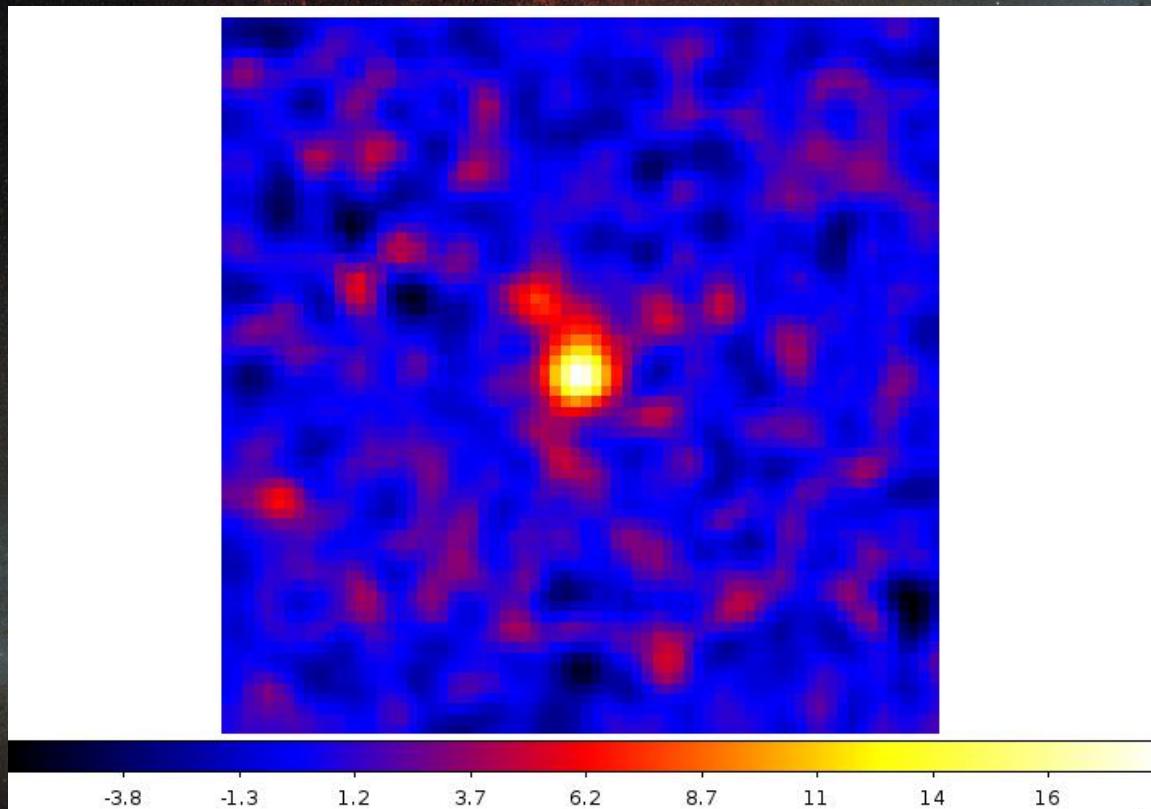


- FERMI has detected four sources coincident with star burst galaxies as of the 3FGL catalog(M82, NGC 253, NGC 1068, and NGC 4945) 3FGL, Acero et al. 2015
- Gamma rays from star burst galaxies possibly originate due to cosmic rays from vigorous star formation



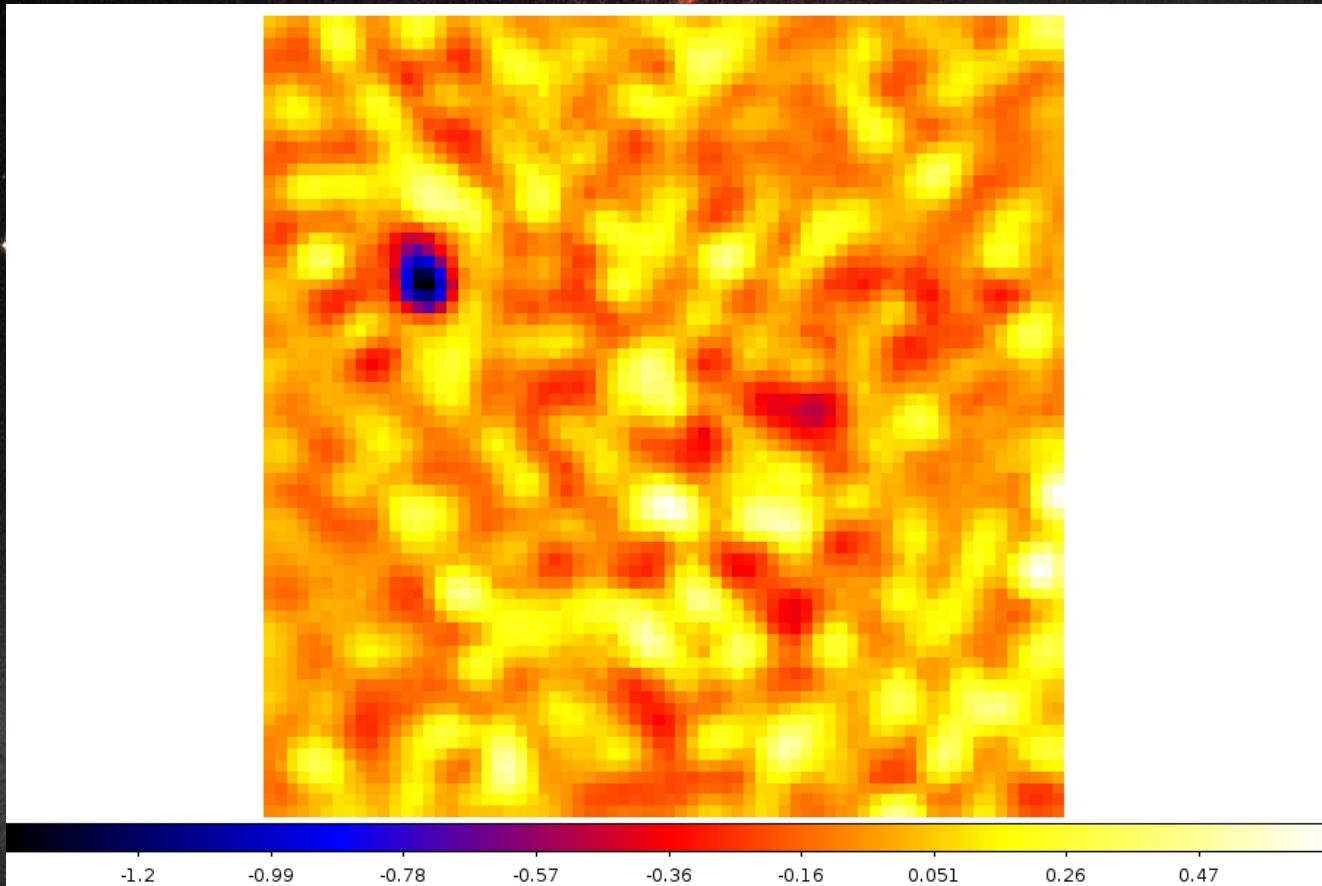
Background Source Model vs. Counts Map
(100MeV – 300GeV) (08/14/2008 – 01/19/2015)

- FERMI has detected four sources coincident with star burst galaxies as of the 3FGL catalog(M82, NGC 253, NGC 1068, and NGC 4945) 3FGL, Acero et al. 2015
- Gamma rays from star burst galaxies possibly originate due to cosmic rays from vigorous star formation



Residual of Counts Map vs. Background Source Model
(100MeV – 300GeV) (08/14/2008 – 01/19/2015)

Creating Energy Bins Residual Map for E>3GeV



- Use different spectral models in the likelihood analysis
- Separate data between time eras

3FGL Catalog J0955.4+6940

- Spectral Index: 2.21003
- Spectral Index Error: 0.05857
- Flux: **1-100 GeV** 9.34374e-10 (photon/cm²/s)
- Flux Error: 8.27854e-11
- Flux: **100-300 MeV** 7.79228e-09 (photon/cm²/s)
- Flux Error: 2.16458e-09

Fermi Analysis(6 ½ years)

3FGL J0955.4+6940

Spectrum: PowerLaw

Index: 2.189e+00 9.731e-04 0.000e+00 1.000e+01 (-1.000e+00)

Flux = 1.03872320302e-08 [s⁻¹ cm⁻²] **100-300 MeV**

Flux Error = 2.30384331379e-11

Flux= 9.19453146415e-10 [s⁻¹ cm⁻²] **1-100 GeV**

Flux Error = 1.35496497742e-12

Flux= 1.4e-08 +/- 2.7e-11 cm⁻² s⁻¹ **0.1-300.0 GeV**

