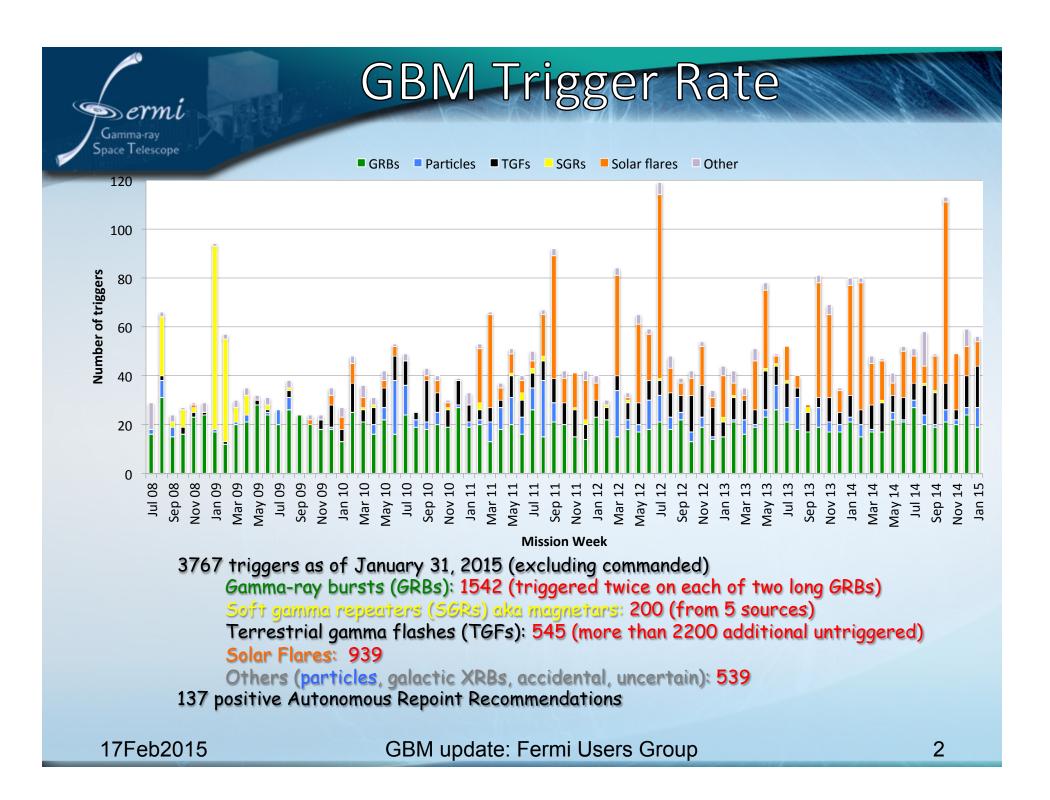
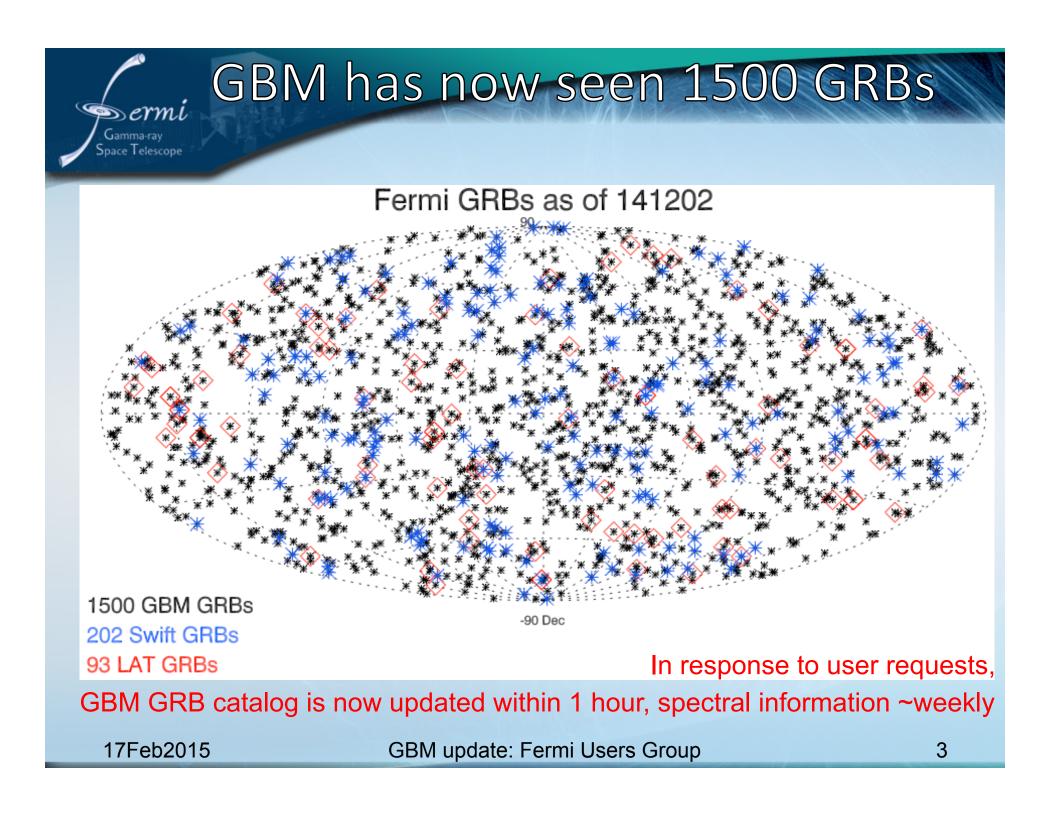


Fermi GBM Status, Results, Plans Linda Sparke NASA (HQ, on detail to MSFC)

Fermi Users Group 17 February 2015





Operational Changes & Improvements

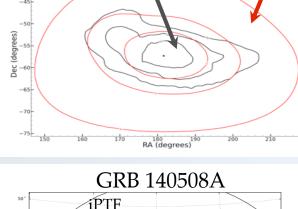
- We continue to disable soft energy (22–50 keV) trigger algorithms at weekends & periods of high solar activity
- Continuous Time Tagged Event (TTE) data are available since 2012.11.26. In August 2014, the TTE pipeline code was updated to fix occasional missing TTE data files and a timing glitch; a clock rollover ("Y2K" problem) was handled pre-emptively without causing problems. Documentation is ready for the next rollover in 12.5 years...
- Capabilities to search continuous TTE data off-line for short GRBs that did not trigger GBM:

-Work in progress (next talk by M. Briggs)

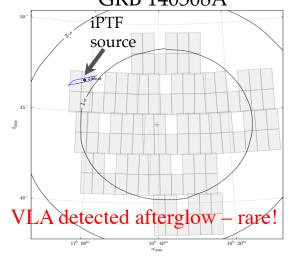
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Gamma-ray pace Telescope

GRB Localization & Follow-up



Total Uncertainty



- GBM proposal: GBM locates bursts to 15°
- Main error sources are systematic: analysis in V Connaughton et al. 2015 ApJS 216, 32 (arXiv:1411.2685)
- GBM ground automated processing (<1min) yielded location to 7-8° for 68% of bursts
- Since Jan 2014, improved location to $\sim 6^{\circ}$
- In ~ a month: BAP will be modified to include FITS maps of ground-automated probability contours (red curves, top plot)
- Human being (~1 hour) locates 68% of bursts to ~5°, 90% of bright bursts to ~6° Error ellipse files are uploaded when final position GCN is sent
- 8 successful follow-ups (~41 attempts) with iPTF using GBM contour files for location; GRBs 140801A (auto loc), 150210A (human loc) found by MASTER based on GBM data
- Collaborations with iPTF, IPN, FIGARO, RAPTOR, MASTER, aLIGO, IceCube, Swift

GBM update: Fermi Users Group

Samma-rav

Statistical Uncertainty

Space Telescope

Gamma Ray Burst spectroscopy

4-year catalog of time-integrated Gamma Ray Burst spectra is now published: D Gruber et al 2014 ApJS 211,12

4-year catalog of time-resolved spectroscopy is in progress for 81 bursts with high fluence, peak flux, signal-to-noise: H-F Yu et al. 2015

Examples of science enabled by GRB spectroscopy:

"Synchrotron cooling in energetic gamma-ray bursts observed by the Fermi Gamma-Ray Burst Monitor": time-resolved spectroscopy of 8 bright bursts, H-F Yu et al 2015 A&Ap 573, 81

"The width of gamma-ray burst spectra": width in EF_E distinguishes long and short bursts, M Axelsson & L Borgonovo, MNRAS, 447, 3150, 2015 "Are GRB blackbodies an artefact of spectral evolution?" JM Burgess, & F Ryde, MNRAS, 447, 3087, 2015

Starquakes in Magnetar Storm

Press release 21 Oct 2014

SGR J1550–5418 is a magnetar, a highly magnetized neutron star with B>10¹⁴G. In 2008-9 it had a series of soft gamma-ray outbursts.

Huppenkothen et al (ApJ 1 Jun 2014) analyzed GBM observations of 286 bursts in a 'storm', Jan 22-29 2009. They found quasi-periodic oscillations at 260Hz.



Just as earthquakes travel deep inside the Earth, and solar oscillations probe the Sun's interior, these oscillations sample the equation of state of matter (plus a strong magnetic field) at neutron star densities.

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GBM update: Fermi Users Group

Gamma rays from thunderstorms

Press release 15 Dec 2014

With continuous time-tagged event data, GBM now detects 5 times as many gamma-ray flashes from thunderstorms.

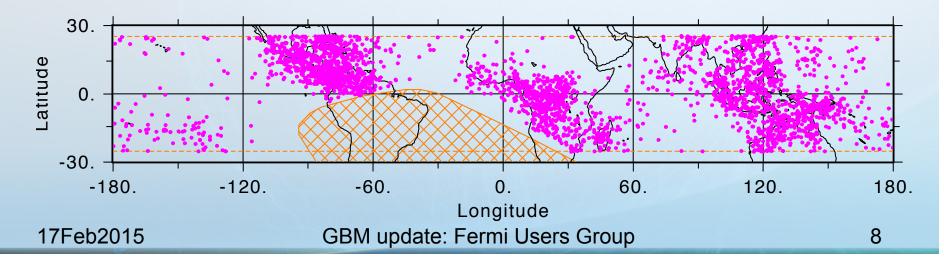
M. Briggs and collaborators tracked simultaneous strong bursts of radio waves to locate the gamma-ray flashes to a specific storm.

All thunderstorm types, strong and weak, appear to produce gamma rays. GBM sees flashes from the upper regions of thunderstorms – gamma rays from lower down may be absorbed in air before reaching Fermi.

The first GBM TGF catalog was released Jan2015, at http://gammaray.nsstc.nasa.gov/gbm/science/terr_grf.html

Gamma-ray

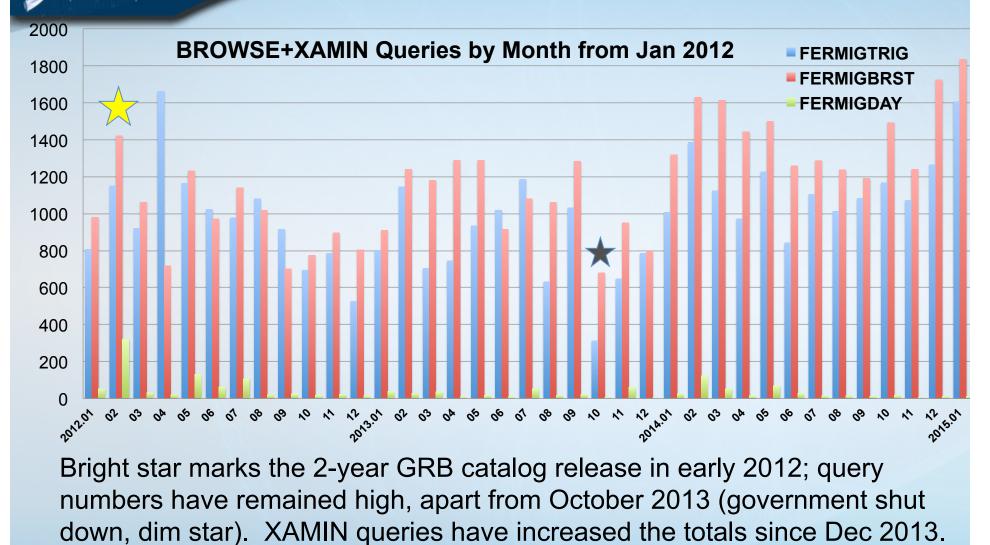
Space Telescope



Usage of GBM archived data

Gamma-ray Space Telescope

Dermi

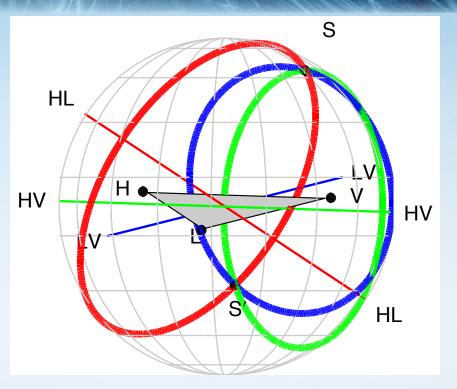


GBM update: Fermi Users Group

The Future

Advanced LIGO begins observing late in 2015.

Short GRB (<1s) are likely mergers of compact stellar-mass objects – these are the main expected LIGO sources! GBM triggers on ~40/year. For an electromagnetic signal, to identify the source for follow-up, GBM is the best bet!



Timing measurement on each baseline localizes a source on an annulus in the sky. A-LIGO alone will locate sources to 100-1000 deg², but we'd have to get lucky: expect only ~0.01/year close-enough short GRB (more from untriggered TTE search). With Advanced Virgo, after 1-2 years, sensitivity improves and sources will be located to tens of deg².

Second joint LIGO-Fermi workshop in Pasadena, CA: 2015 March 14-15 includes Swift: http://www.ligo.caltech.edu/~jkanner/ligo-fermi/

Gamma-rav

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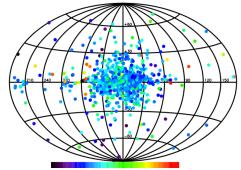
Summary & Near-term Plans

- GBM operations and performance are nominal
 - Full-orbit untriggered Time Tagged Event data collection is proceeding smoothly
- No significant operational changes planned for next few months
- Work on searching TTE data for short GRBs continues: M Briggs talk
- Will add FITS files distribution for GRB localization contours
 - Planned work to improve localization algorithms: summer student may work on atmospheric scattering
- Science catalogs

Dermi Gamma-ray

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- GBM Burst Catalog is now continuously updated on-line at FSSC; next catalog paper likely when data doubles (8 years) rather than at 6 years
- GRB 4-year spectroscopic catalog published; 4-year catalog of time-resolved spectroscopy in preparation
- Terrestrial Gamma-Ray Flash catalog released January 2015
- Earth Occultation Light Curves and Spin Histories for accreting pulsars being updated: access via http://fermi.gsfc.nasa.gov/ssc/data/access/gbm/
- 3-year catalog of Type 1 X-ray bursts in preparation (see image)



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GBM update: Fermi Users Group