

Variability of VHE γ -ray sources

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Method

- presence of source in a given region:
on-off method (Li&Ma, ApJ 272 (1983) 317)

Citing Articles: 451
(from All Databases)

For: ANALYSIS-METHODS FOR RESULTS IN GAMMA-RAY ASTRONOMY

Binomial significance

$$S_{Bi} = \frac{N_{on} - \alpha N_{off}}{\sqrt{\alpha(N_{on} + N_{off})}}$$

← normally distributed



hypothesis:

$$\mu_{on} = \alpha \mu_{off}$$

$$\mu_S = 0$$

Li–Ma significance

$$S_{LM} = s\sqrt{2} \left\{ N_{on} \ln \left[\frac{1 + \alpha}{\alpha} \frac{N_{on}}{N_{on} + N_{off}} \right] + N_{off} \ln \left[(1 + \alpha) \frac{N_{off}}{N_{on} + N_{off}} \right] \right\}^{\frac{1}{2}}$$

Method

- presence of source in a given region:
on-off method (Li&Ma, ApJ 272 (1983) 317)
- modification: test of constant intensity
source parameter β

Citing Articles: 451
(from All Databases)

For: ANALYSIS-METHODS FOR RESULTS IN GAMMA-RAY ASTRONOMY

Binomial significance

$$S_{Bi} = \frac{N_{on} - \beta\alpha N_{off}}{\sqrt{\beta\alpha(N_{on} + N_{off})}}$$

hypothesis:

$$\mu_{on} = \beta\alpha\mu_{off}$$

~~$\mu_{on} = \mu_{off}$~~

$$\mu_S = (\beta - 1)\alpha\mu_{off}$$

← normally distributed



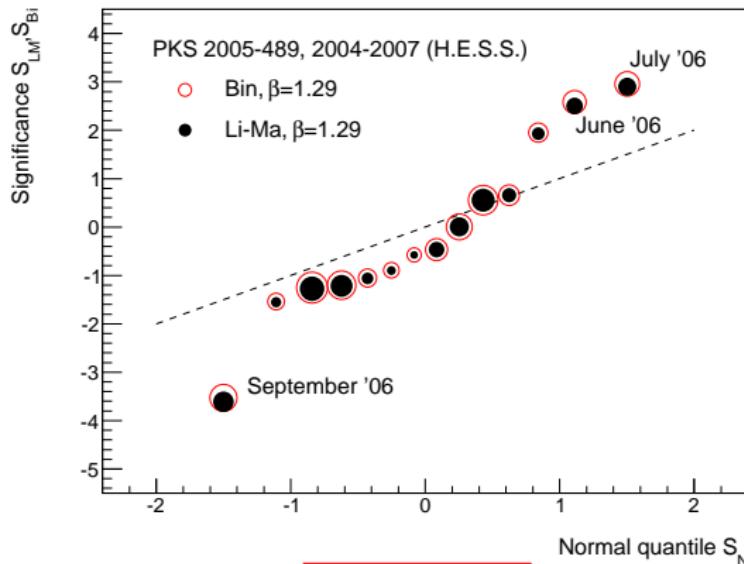
Li-Ma significance

$$S_{LM} = s\sqrt{2} \left\{ N_{on} \ln \left[\frac{1 + \beta\alpha}{\beta\alpha} \frac{N_{on}}{N_{on} + N_{off}} \right] + N_{off} \ln \left[(1 + \beta\alpha) \frac{N_{off}}{N_{on} + N_{off}} \right] \right\}^{\frac{1}{2}}$$

Nosek et al., ICRC 2013, arXiv:1309.6476

Applications: PKS 2005–489

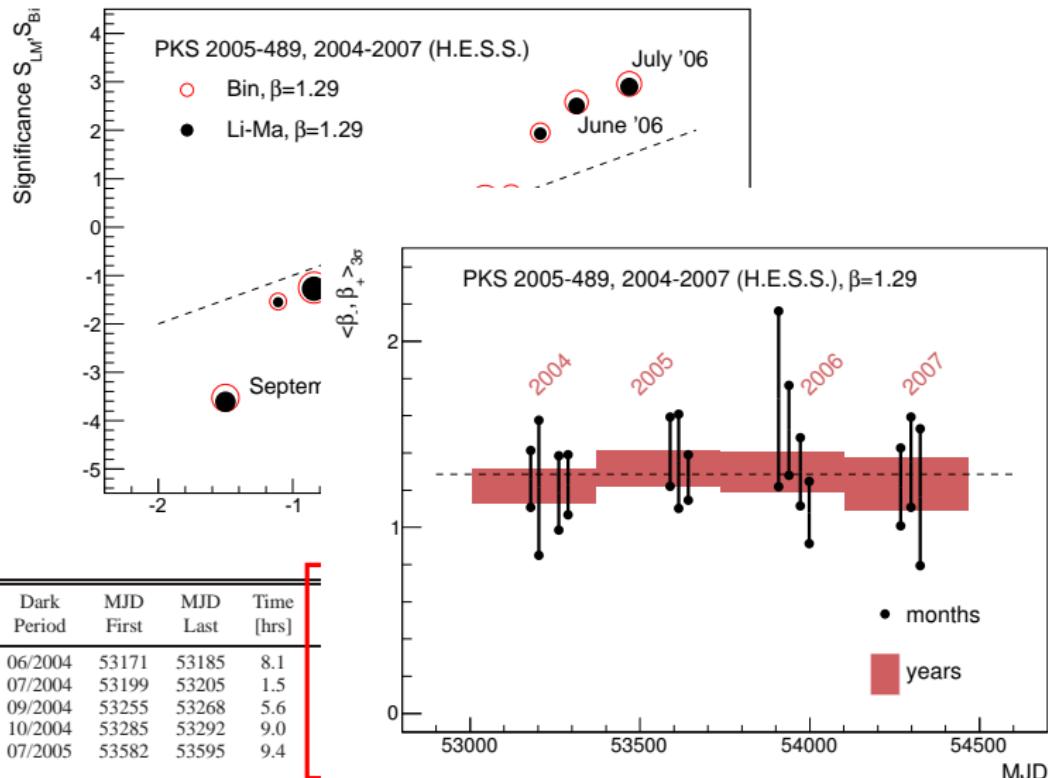
- F.Acero et al. (H.E.S.S. Collaboration), A&A 511 (2010) A52



| Dark Period | MJD First | MJD Last | Time [hrs] | On | Off | α | Excess | Sig [σ] | $I(>400 \text{ GeV})^a$ [$10^{-12} \text{ cm}^{-2} \text{ s}^{-1}$] | Crab ^b % | χ^2 (NDF ^c) | $P(\chi^2)^c$ |
|-------------|-----------|----------|------------|-----|------|----------|--------|------------------|---|---------------------|------------------------------|---------------|
| 06/2004 | 53171 | 53185 | 8.1 | 678 | 5877 | 0.0919 | 138 | 5.4 | 2.64 ± 0.50 | 3.0 | 7.2 (10) | 0.71 |
| 07/2004 | 53199 | 53205 | 1.5 | 105 | 985 | 0.0909 | 15 | 1.5 | 1.06 ± 1.18 | 1.2 | 1.4 (1) | 0.75 |
| 09/2004 | 53255 | 53268 | 5.6 | 342 | 3171 | 0.0916 | 52 | 2.8 | 1.89 ± 0.78 | 2.1 | 2.1 (5) | 0.84 |
| 10/2004 | 53285 | 53292 | 9.0 | 569 | 5065 | 0.0916 | 105 | 4.5 | 2.58 ± 0.62 | 2.9 | 1.5 (4) | 0.83 |
| 07/2005 | 53582 | 53595 | 9.4 | 573 | 4504 | 0.0908 | 164 | 7.3 | 3.56 ± 0.67 | 4.0 | 9.2 (9) | 0.42 |

Applications: PKS 2005–489

- F.Acero et al. (H.E.S.S. Collaboration), A&A 511 (2010) A52

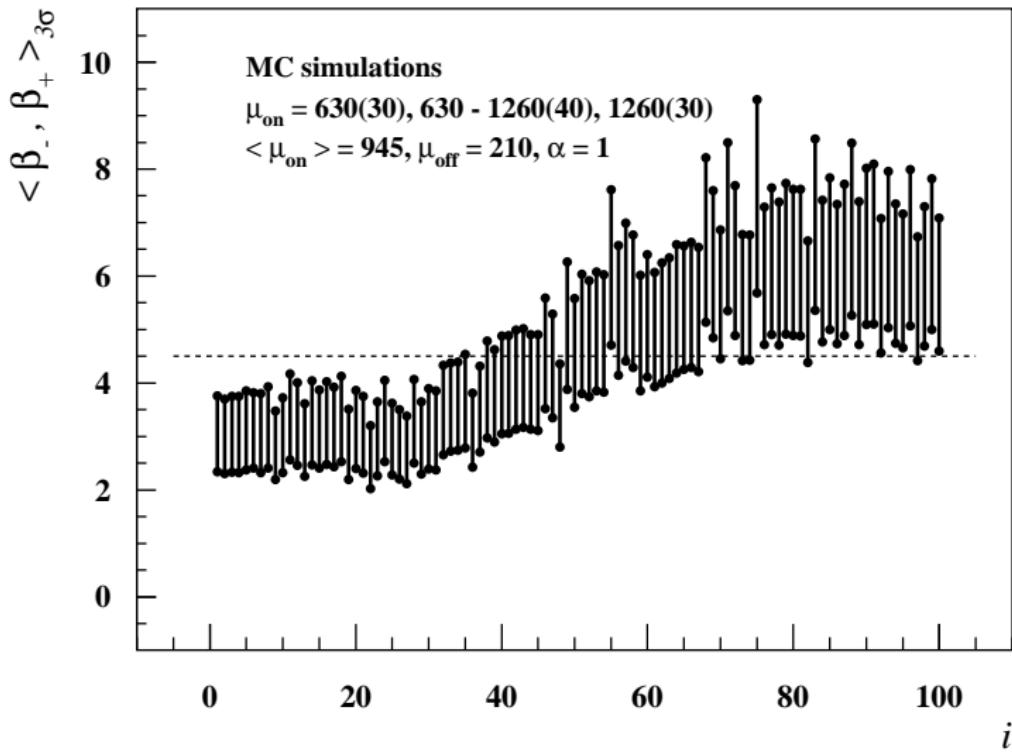


Thank you!

Back-up

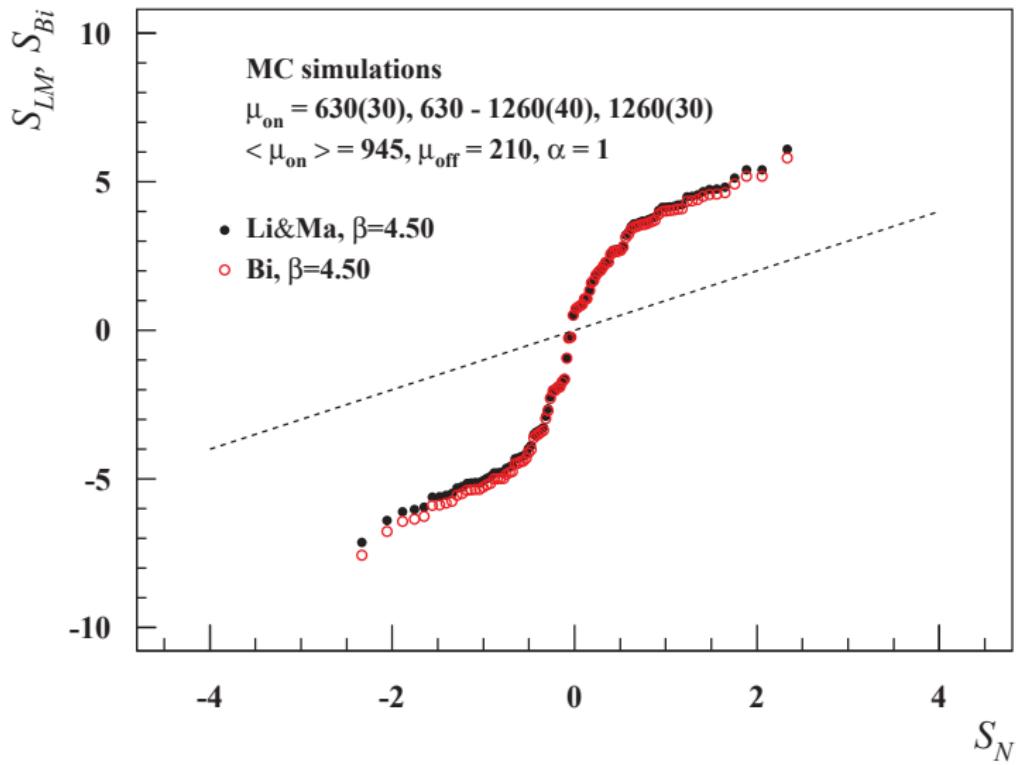
MC example

- evolution of the source intensity $|S_{LM}(\beta)| < 3$

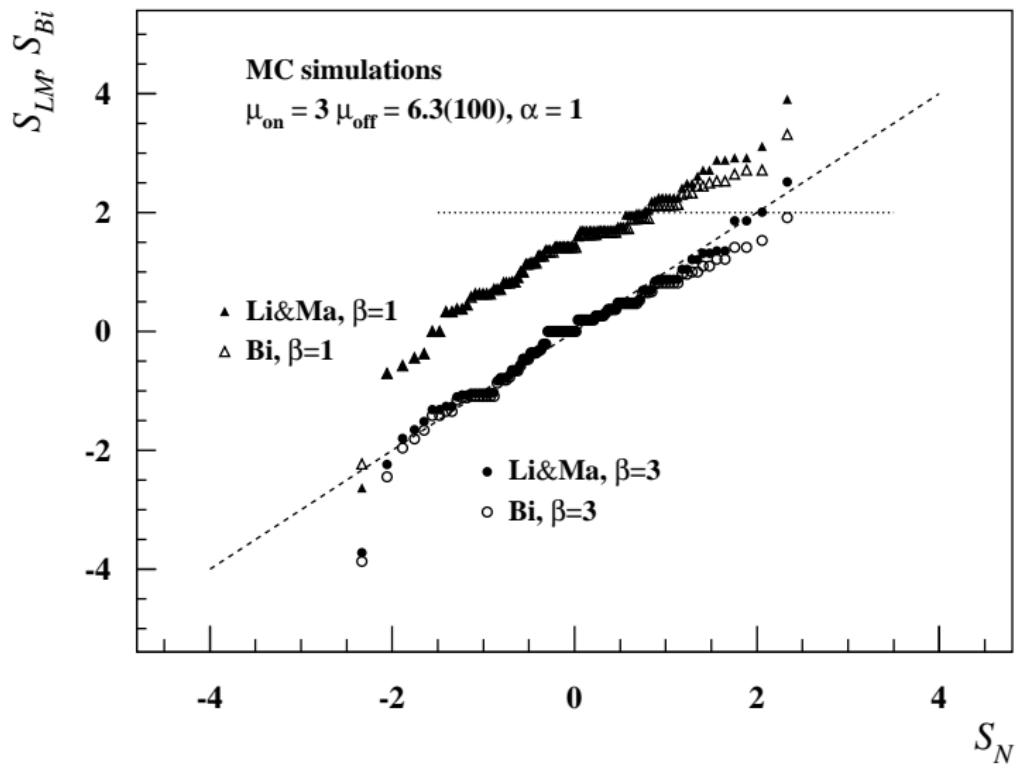


MC example

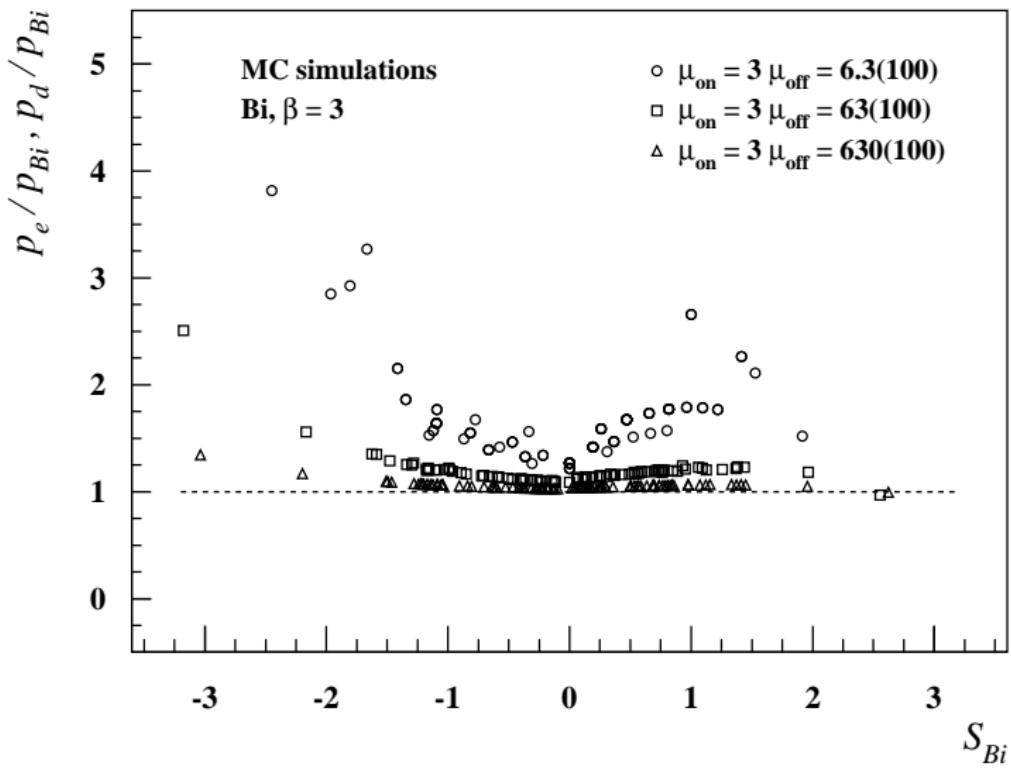
- quantile–quantile plots



Hypotheses tests

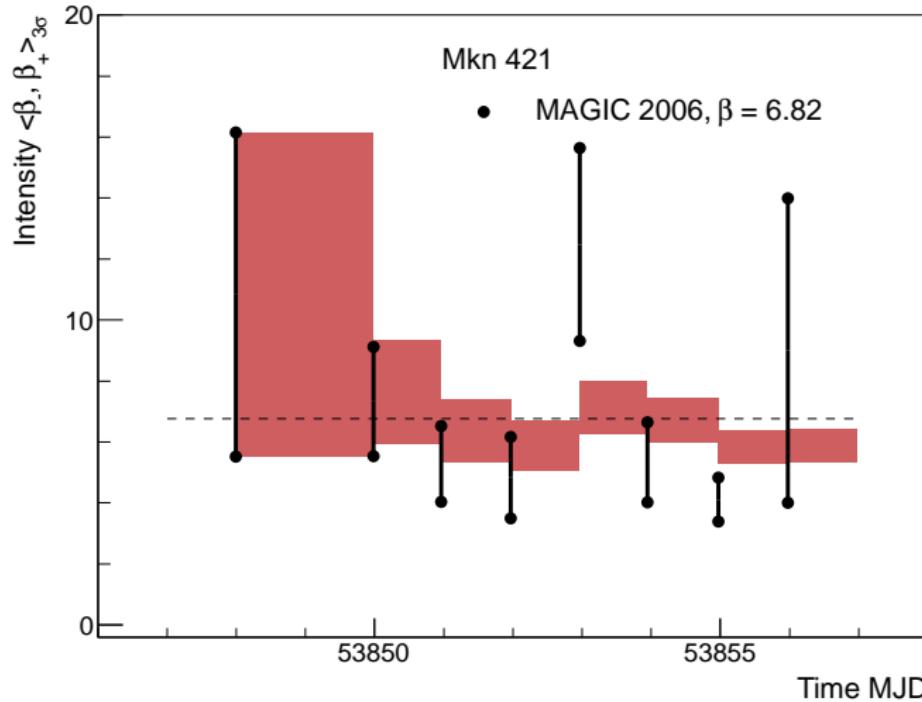


Asymptotic properties



Mkn 421

- J.Aleksić et al. (MAGIC Collaboration), A&A 519 (2010) A32



Other sources

| Epoch | N_{on} | N_{off} | α | N_s | S_{LM}^0 | S_{Bi}^0 | S_{LM} | S_{Bi} | p_e or p_d | $\langle \beta_-, \beta_+ \rangle_{3\sigma}$ |
|---|-----------------|------------------|----------|-------|-------------------|-------------------|-----------------|-----------------|----------------------|--|
| PSR B1259-63/LS 2883 (HESS Collaboration 2013), $\beta = 4.08$ | | | | | | | | | | |
| Pre-flare | 44 | 133 | 0.0760 | 34 | 7.4 | 9.2 | 0.4 | 0.4 | 0.38 | $\langle 2.5, 7.2 \rangle$ |
| Flare | 68 | 232 | 0.0760 | 50 | 8.5 | 10.4 | -0.5 | -0.5 | 0.34 | $\langle 2.5, 5.7 \rangle$ |
| PG 1553+113 (Aharonian et al. 2008), $\beta = 1.16$ | | | | | | | | | | |
| 04/2005 | 1210 | 8154 | 0.1250 | 191 | 5.5 | 5.6 | 0.8 | 0.8 | 0.23 | $\langle 1.1, 1.3 \rangle$ |
| 08/2005 | 491 | 3462 | 0.1250 | 58 | 2.5 | 2.6 | -0.5 | -0.5 | 0.33 | $\langle 1.0, 1.3 \rangle$ |
| 04/2006 | 1811 | 12742 | 0.1250 | 218 | 5.0 | 5.1 | -0.8 | -0.8 | 0.22 | $\langle 1.1, 1.2 \rangle$ |
| 07/2006 | 2236 | 15341 | 0.1250 | 318 | 6.7 | 6.8 | 0.2 | 0.2 | 0.41 | $\langle 1.1, 1.2 \rangle$ |
| B2 1215+30 (Aliu et al. 2013b), $\beta = 1.39$ | | | | | | | | | | |
| 2008–09 | 304 | 2288 | 0.1243 | 20 | 1.1 | 1.1 | -4.4 | -4.3 | $5.1 \cdot 10^{-6}$ | $\langle 0.9, 1.3 \rangle$ |
| 2011 | 472 | 2325 | 0.1161 | 202 | 10.4 | 11.2 | 4.4 | 4.6 | $5.1 \cdot 10^{-6}$ | $\langle 1.5, 2.0 \rangle$ |
| 2012 | 443 | 898 | 0.1177 | 37 | 3.2 | 3.4 | -0.3 | -0.3 | 0.40 | $\langle 1.0, 1.8 \rangle$ |
| 1ES 0229+200 (Aharonian et al. 2007; Aliu et al. 2014) ^a , $\beta = 1.20$ | | | | | | | | | | |
| 2005 | 246 | 2238 | 0.0916 | 41 | 2.7 | 2.7 | -0.0 | -0.0 | 0.52 | $\langle 1.0, 1.5 \rangle$ |
| 2006 | 1344 | 12304 | 0.0914 | 220 | 6.1 | 6.2 | -0.1 | -0.1 | 0.46 | $\langle 1.1, 1.3 \rangle$ |
| 2009–10 | 1054 | 7601 | 0.0909 | 363 | 12.2 | 12.9 | 7.1 | 7.3 | $7.8 \cdot 10^{-13}$ | $\langle 1.4, 1.7 \rangle$ |
| 2010–11 | 614 | 5862 | 0.0909 | 81 | 3.3 | 3.3 | -1.0 | -1.0 | 0.17 | $\langle 1.0, 1.3 \rangle$ |
| 2011–12 | 249 | 2241 | 0.0909 | 45 | 2.9 | 3.0 | 0.3 | 0.3 | 0.40 | $\langle 1.0, 1.5 \rangle$ |

1ES 0229+200 data

Table 1. The MJD of the first and last night of HESS observations of 1ES 0229+200, the live time of the observations, the number of on- and off-source events measured, the on/off normalization (α), the excess, and the significance of the excess are given. In addition, the integral flux above 580 GeV (assuming $\Gamma = 2.50$), and the corresponding percentage of the Crab Nebula flux above 580 GeV are shown. The χ^2 , degrees of freedom (NDF), and χ^2 probability, $P(\chi^2)$, for a fit of a constant to the flux binned by dark period within each year, or yearly within the total, are also given.

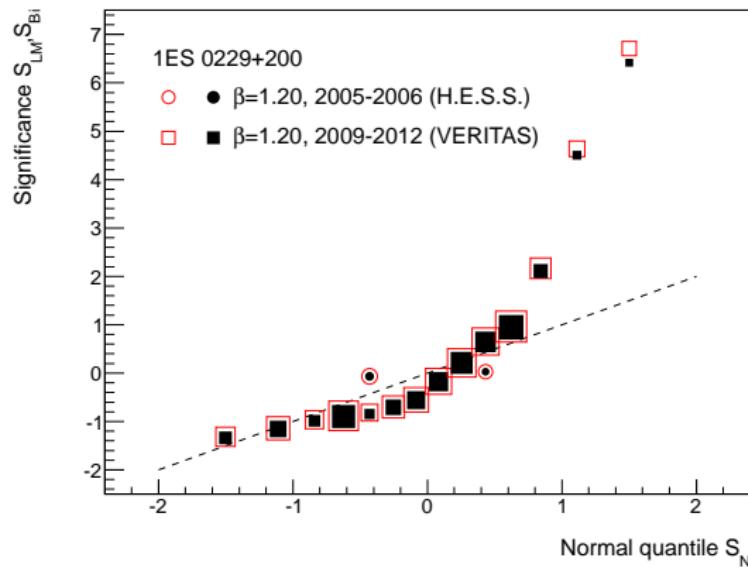
| Epoch | MJD First | MJD Last | Time [h] | On | Off | α | Excess | Sig [σ] | I(>580 GeV) [$10^{-13} \text{ cm}^{-2} \text{ s}^{-1}$] | Crab % | χ^2 , NDF | $P(\chi^2)$ |
|-------|-----------|----------|----------|------|-------|----------|--------|------------------|---|--------|----------------|-------------|
| 2005 | 53614 | 53649 | 6.8 | 246 | 2238 | 0.09160 | 41 | 2.7 | $6.8 \pm 3.1_{\text{stat}} \pm 1.4_{\text{syst}}$ | 1.3 | 1.6, 1 | 0.21 |
| 2006 | 53967 | 54088 | 35.0 | 1344 | 12304 | 0.09136 | 220 | 6.1 | $10.0 \pm 1.7_{\text{stat}} \pm 2.0_{\text{syst}}$ | 1.9 | 1.5, 3 | 0.68 |
| Total | 53614 | 54088 | 41.8 | 1590 | 14542 | 0.09140 | 261 | 6.6 | $9.4 \pm 1.5_{\text{stat}} \pm 1.9_{\text{syst}}$ | 1.8 | 0.8, 1 | 0.37 |

Table 1: The VERITAS 1ES 0229+200 observation details. α (the ratio of the area \times livetime of the *on* source and *off* source regions) is 1/11. The integral flux is calculated assuming an overall spectral index of 2.59. Upper limits at the 99% confidence level using the Rolke method (Rolle & López 2001) are presented when the significance is less than two standard deviations. The horizontal lines delineate the results for the full time period, the data divided by season, and the data divided by observing period (dictated by the lunar cycle and indicated by 'P. 1' through 'P. 5' in each season).

| Period | Dates [MJD] | Live Time [minutes] | On [events] | Off [events] | Significance [σ] | Flux ($> 300 \text{ GeV}$) [$10^{-9} \text{ m}^{-2} \text{ s}^{-1}$] | UL ($> 300 \text{ GeV}$) [$10^{-9} \text{ m}^{-2} \text{ s}^{-1}$] |
|----------------|---------------|---------------------|-------------|--------------|---------------------------|--|--|
| 2009-2012 | 55118 - 55951 | 3260 | 1917 | 15704 | 11.7 | $23.3 \pm 2.8_{\text{stat}} \pm 5.8_{\text{syst}}$ | N/A |
| 2009-2010 | 55118 - 55212 | 1674 | 1054 | 7601 | 12.2 | $30.3 \pm 3.9_{\text{stat}} \pm 7.6_{\text{syst}}$ | N/A |
| 2010-2011 | 55476 - 55587 | 1079 | 614 | 5862 | 3.3 | $18.7 \pm 5.1_{\text{stat}} \pm 5.7_{\text{syst}}$ | N/A |
| 2011-2012 | 55828 - 55951 | 507 | 249 | 2241 | 2.9 | $9.9 \pm 6.4_{\text{stat}} \pm 2.5_{\text{syst}}$ | N/A |
| 2009-2010 P. 1 | 55118 - 55131 | 715 | 484 | 3210 | 9.7 | $41.8 \pm 6.4_{\text{stat}} \pm 10.5_{\text{syst}}$ | N/A |
| 2009-2010 P. 2 | 55144 - 55159 | 844 | 524 | 3880 | 8.1 | $24.2 \pm 5.4_{\text{stat}} \pm 6.1_{\text{syst}}$ | N/A |
| 2009-2010 P. 3 | 55183 - 55183 | 24 | 10 | 120 | -0.3 | $1 \pm 26_{\text{stat}} \pm 1_{\text{syst}}$ | 100 |
| 2009-2010 P. 4 | 55200 - 55212 | 91 | 36 | 391 | 0.1 | $3 \pm 10_{\text{stat}} \pm 1_{\text{syst}}$ | 51 |
| 2010-2011 P. 1 | 55476 - 55482 | 319 | 187 | 1900 | 1.0 | $15 \pm 9_{\text{stat}} \pm 4_{\text{syst}}$ | 41 |
| 2010-2011 P. 2 | 55501 - 55513 | 162 | 121 | 901 | 3.8 | $39 \pm 14_{\text{stat}} \pm 10_{\text{syst}}$ | N/A |
| 2010-2011 P. 3 | 55526 - 55538 | 127 | 69 | 692 | 0.7 | $1 \pm 14_{\text{stat}} \pm 1_{\text{syst}}$ | 60 |
| 2010-2011 P. 4 | 55555 - 55570 | 297 | 147 | 1490 | 1.0 | $15 \pm 10_{\text{stat}} \pm 4_{\text{syst}}$ | 40 |
| 2010-2011 P. 5 | 55583 - 55587 | 174 | 90 | 879 | 1.1 | $26 \pm 13_{\text{stat}} \pm 7_{\text{syst}}$ | 54 |
| 2011-2012 P. 1 | 55828 - 55840 | 101 | 46 | 434 | 1.0 | $13 \pm 14_{\text{stat}} \pm 3_{\text{syst}}$ | 66 |
| 2011-2012 P. 2 | 55855 - 55861 | 111 | 55 | 460 | 1.9 | $15 \pm 14_{\text{stat}} \pm 4_{\text{syst}}$ | 78 |
| 2011-2012 P. 3 | 55886 - 55895 | 119 | 68 | 608 | 1.6 | $13 \pm 14_{\text{stat}} \pm 3_{\text{syst}}$ | 77 |
| 2011-2012 P. 4 | 55916 - 55922 | 103 | 41 | 435 | 0.2 | $-6 \pm 13_{\text{stat}} \pm 2_{\text{syst}}$ | 51 |
| 2011-2012 P. 5 | 55940 - 55951 | 73 | 39 | 304 | 1.9 | $16 \pm 18_{\text{stat}} \pm 4_{\text{syst}}$ | 100 |

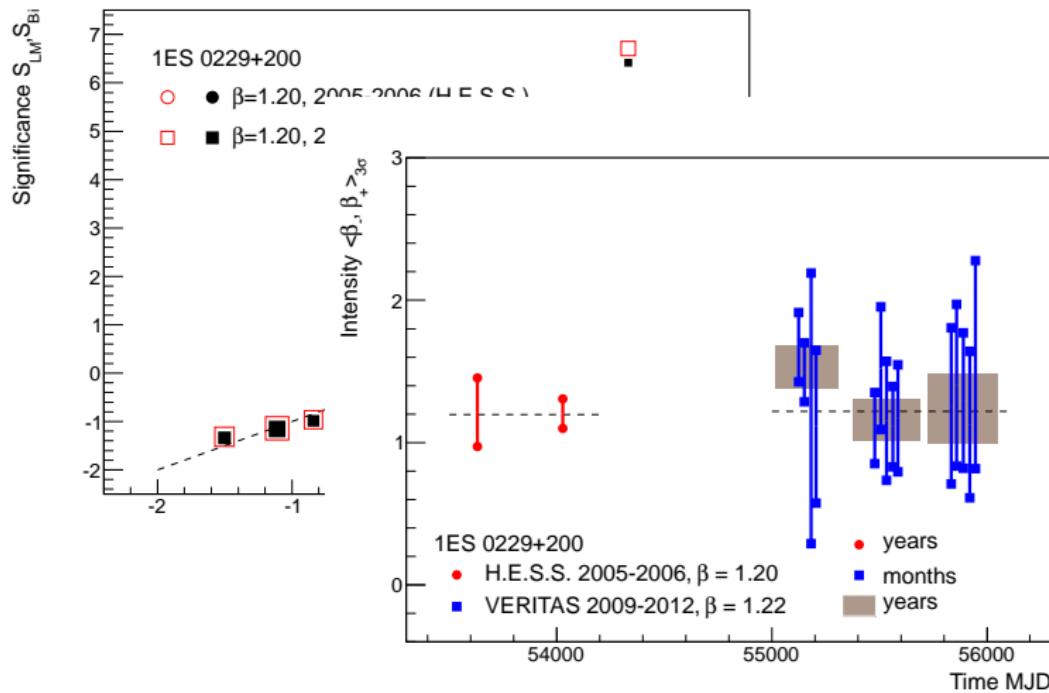
Applications: 1ES 0229+200

- F.Aharonian et al. (H.E.S.S. Collaboration), A&A 475 (2007) L9
- E.Aliu et al. (VERITAS Collaboration), ApJ 782 (2014) 13



Applications: 1ES 0229+200

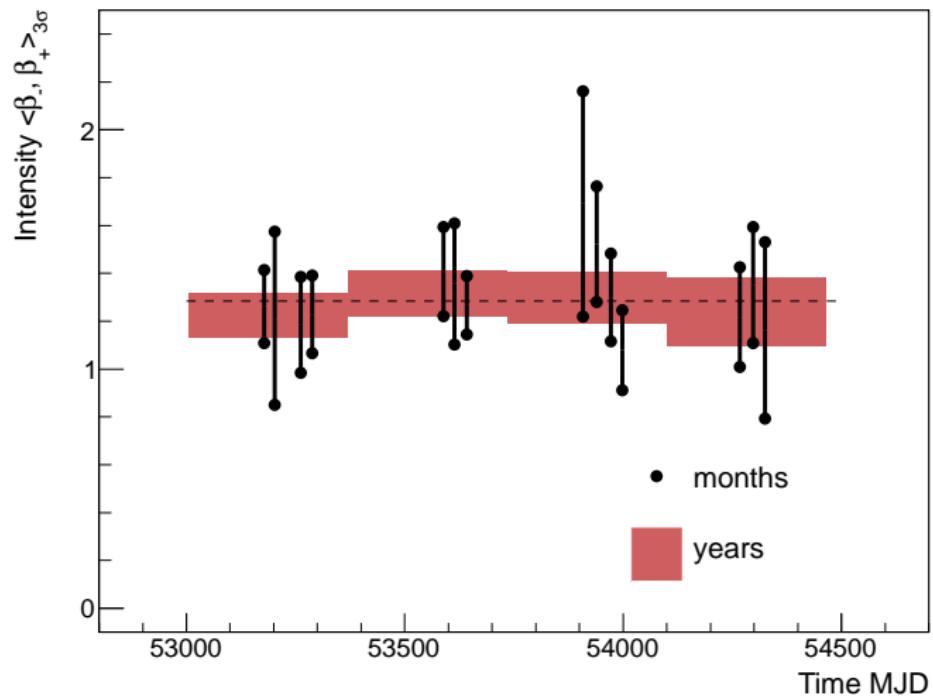
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- E.Aliu et al. (VERITAS Collaboration), ApJ 782 (2014) 13



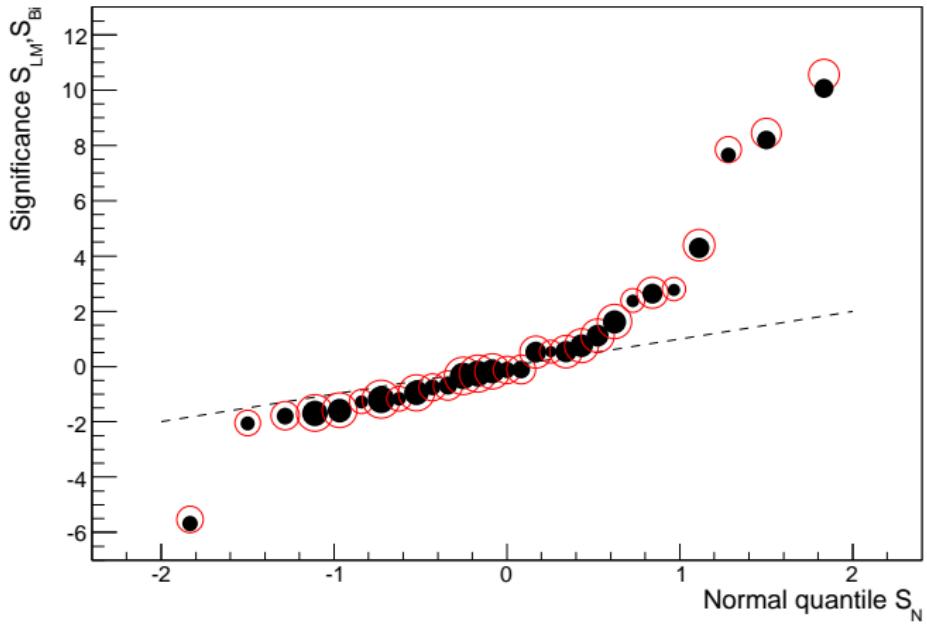
PKS 2005–489 data

Table 1. Results from long-term HESS observations of PKS 2005–489.

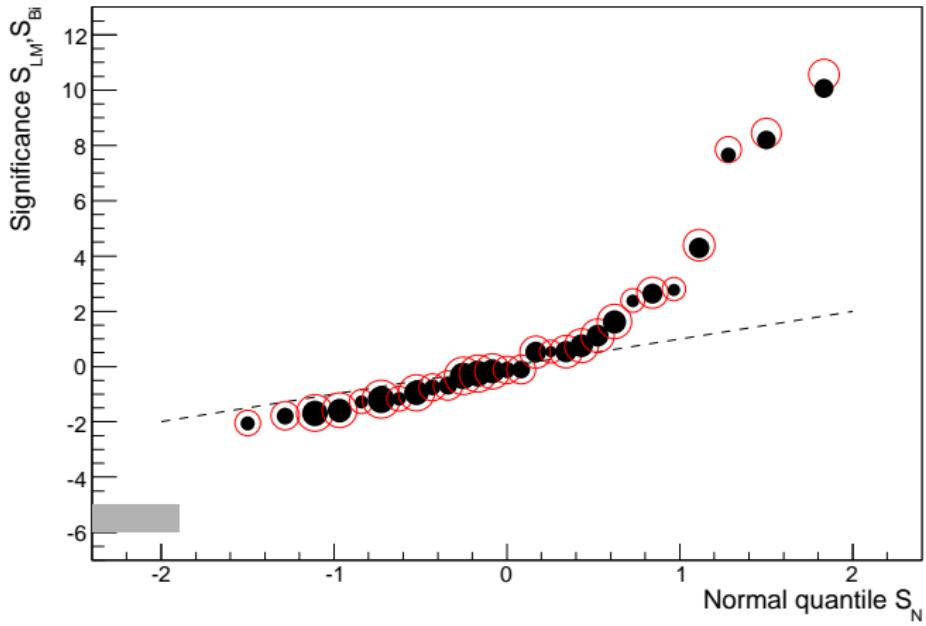
| Dark Period | MJD First | MJD Last | Time [hrs] | On | Off | α | Excess | Sig [σ] | $I(>400 \text{ GeV})^a$ [$10^{-12} \text{ cm}^{-2} \text{s}^{-1}$] | Crab ^b % | χ^2 (NDF ^c) | P(χ^2) ^c |
|-------------|-----------|----------|------------|------|-------|----------|--------|------------------|--|---------------------|------------------------------|----------------------------|
| 06/2004 | 53171 | 53185 | 8.1 | 678 | 5877 | 0.0919 | 138 | 5.4 | 2.64 ± 0.50 | 3.0 | 7.2 (10) | 0.71 |
| 07/2004 | 53199 | 53205 | 1.5 | 105 | 985 | 0.0909 | 15 | 1.5 | 1.06 ± 1.18 | 1.2 | 1.4 (1) | 0.75 |
| 09/2004 | 53255 | 53268 | 5.6 | 342 | 3171 | 0.0916 | 52 | 2.8 | 1.89 ± 0.78 | 2.1 | 2.1 (5) | 0.84 |
| 10/2004 | 53285 | 53292 | 9.0 | 569 | 5065 | 0.0916 | 105 | 4.5 | 2.58 ± 0.62 | 2.9 | 1.5 (4) | 0.83 |
| 07/2005 | 53582 | 53595 | 9.4 | 573 | 4504 | 0.0908 | 164 | 7.3 | 3.56 ± 0.67 | 4.0 | 9.2 (9) | 0.42 |
| 08/2005 | 53609 | 53618 | 5.1 | 286 | 2159 | 0.0989 | 73 | 4.5 | 2.86 ± 0.93 | 3.2 | 4.6 (4) | 0.33 |
| 09/2005 | 53639 | 53646 | 18.1 | 1072 | 9125 | 0.0928 | 225 | 7.1 | 2.93 ± 0.46 | 3.3 | 13.1 (6) | 0.041 |
| 06/2006 | 53908 | 53909 | 1.3 | 127 | 824 | 0.0938 | 50 | 4.9 | 4.72 ± 1.32 | 5.3 | 0.0 (1) | 0.97 |
| 07/2006 | 53938 | 53940 | 4.4 | 397 | 2927 | 0.0901 | 133 | 7.3 | 4.33 ± 0.70 | 4.8 | 1.3 (2) | 0.52 |
| 08/2006 | 53967 | 53977 | 8.4 | 500 | 4261 | 0.0913 | 111 | 5.1 | 2.74 ± 0.57 | 3.1 | 15.1 (7) | 0.035 |
| 09/2006 | 53995 | 54002 | 7.4 | 405 | 4116 | 0.0920 | 26 | 1.3 | 1.31 ± 0.58 | 1.5 | 6.0 (5) | 0.30 |
| 06/2007 | 54264 | 54270 | 5.3 | 333 | 3006 | 0.0924 | 55 | 3.1 | 1.02 ± 0.60 | 1.1 | 1.9 (6) | 0.93 |
| 07/2007 | 54291 | 54304 | 4.4 | 309 | 2412 | 0.0964 | 76 | 4.5 | 2.51 ± 0.70 | 2.8 | 15.2 (8) | 0.056 |
| 08/2007 | 54321 | 54329 | 1.8 | 93 | 849 | 0.0977 | 10 | 1.0 | 0.52 ± 0.92 | 0.6 | 2.9 (3) | 0.41 |
| 09/2007 | 54345 | 54345 | 0.5 | 11 | 114 | 0.1000 | 0 | -0.1 | < 6.60 ^d | <7.4 | — | — |
| 2004 | 53171 | 53292 | 24.2 | 1694 | 15098 | 0.0917 | 310 | 7.7 | 2.37 ± 0.33 | 2.6 | 2.0 (3) | 0.57 |
| 2005 | 53582 | 53646 | 32.6 | 1931 | 15785 | 0.0930 | 462 | 11.0 | 3.09 ± 0.35 | 3.5 | 0.6 (2) | 0.73 |
| 2006 | 53908 | 54002 | 21.5 | 1429 | 12128 | 0.0914 | 320 | 8.8 | 2.84 ± 0.34 | 3.2 | 13.5 (3) | 0.0037 |
| 2007 | 54264 | 54345 | 12.0 | 746 | 6381 | 0.0947 | 141 | 5.3 | 1.50 ± 0.40 | 1.7 | 3.8 (3) | 0.28 |
| Total | 53171 | 54345 | 90.3 | 5800 | 49392 | 0.0924 | 1233 | 16.7 | 2.57 ± 0.18 | 2.9 | 10.2 (3) | 0.016 |



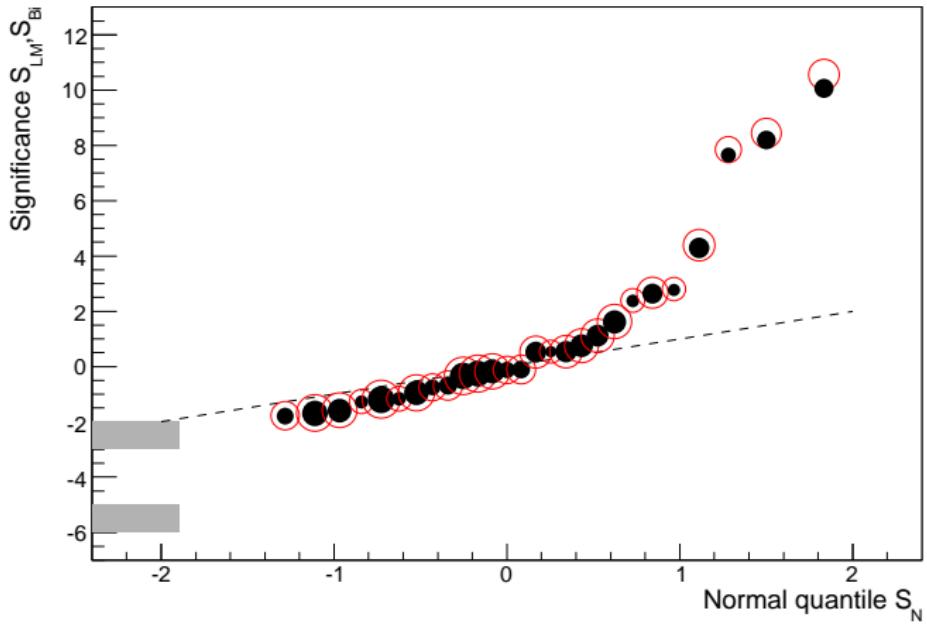
PKS 2155–304



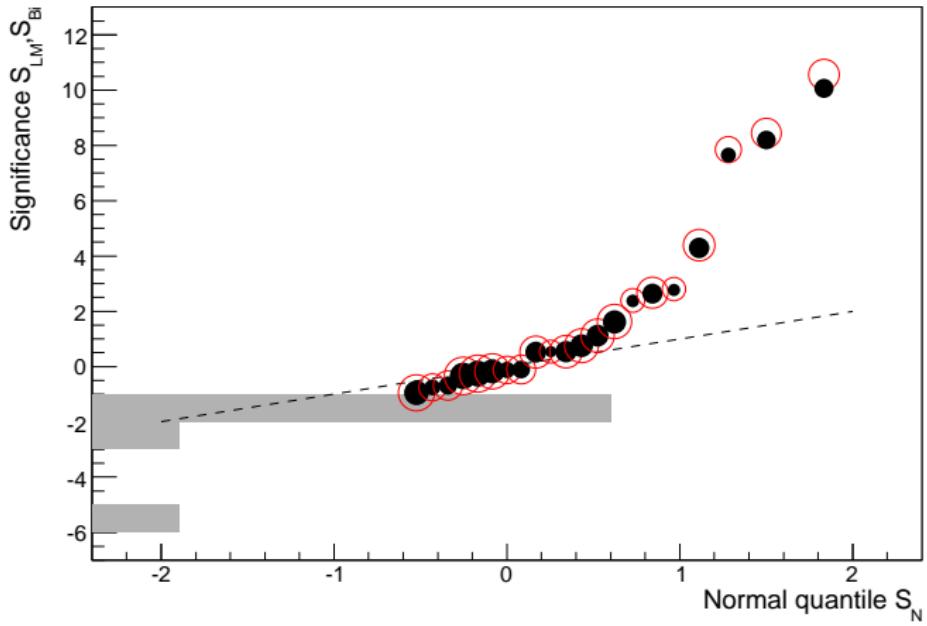
PKS 2155–304



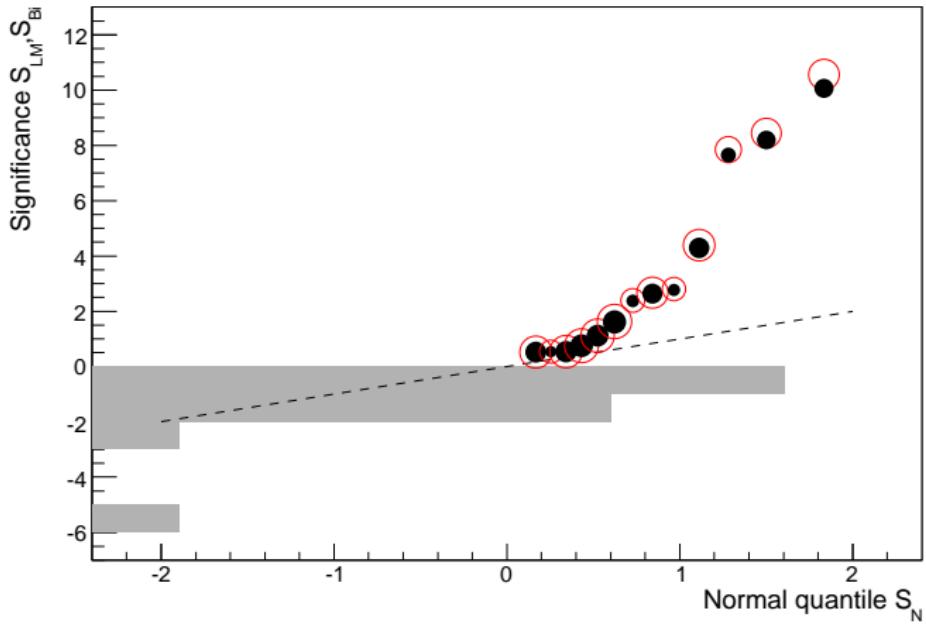
PKS 2155–304



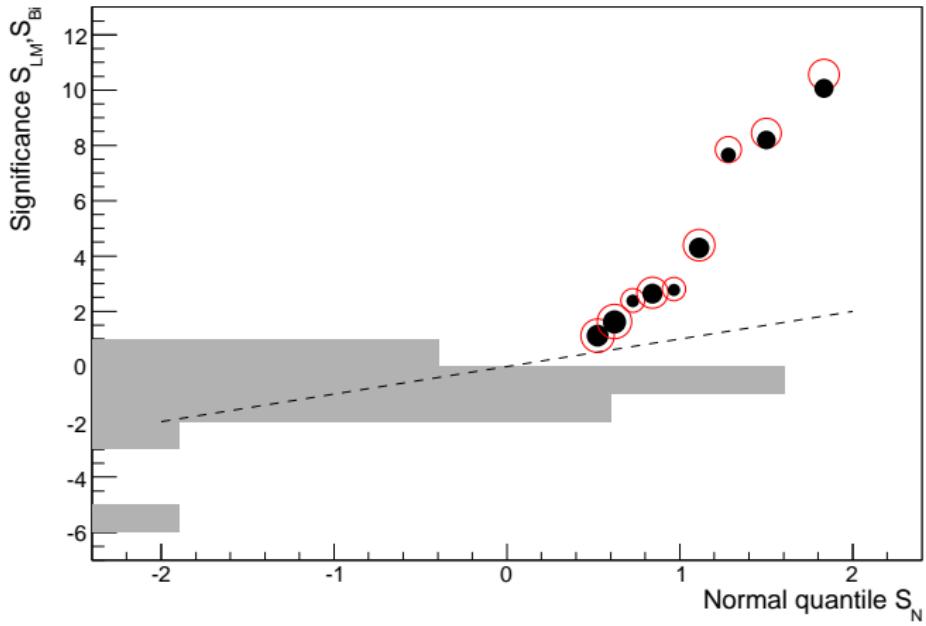
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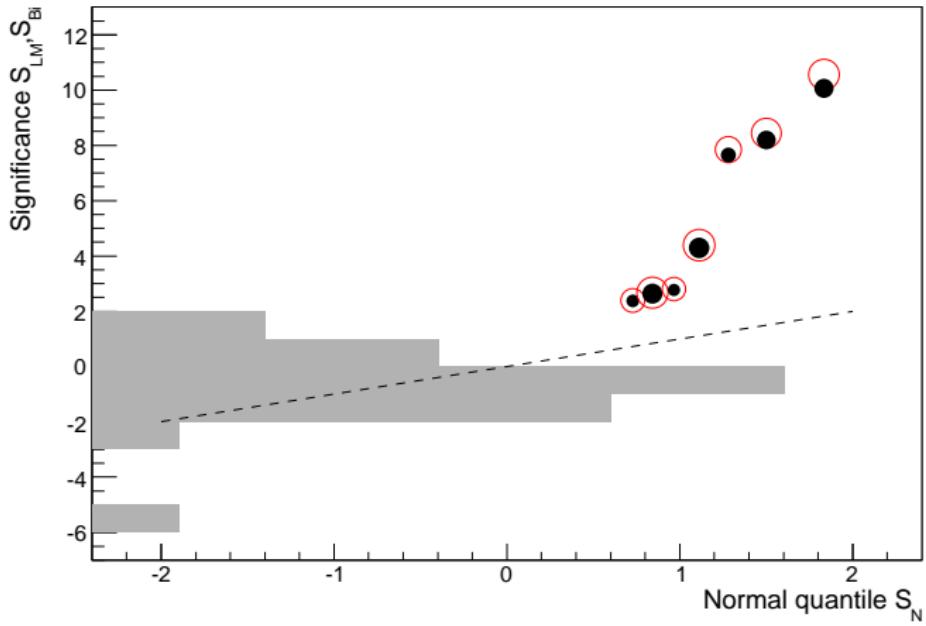
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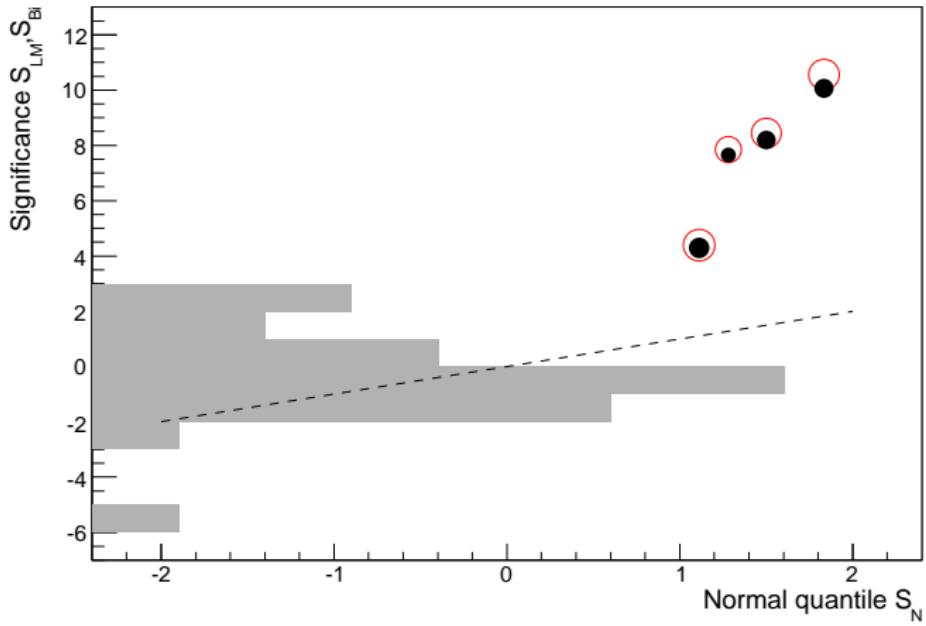
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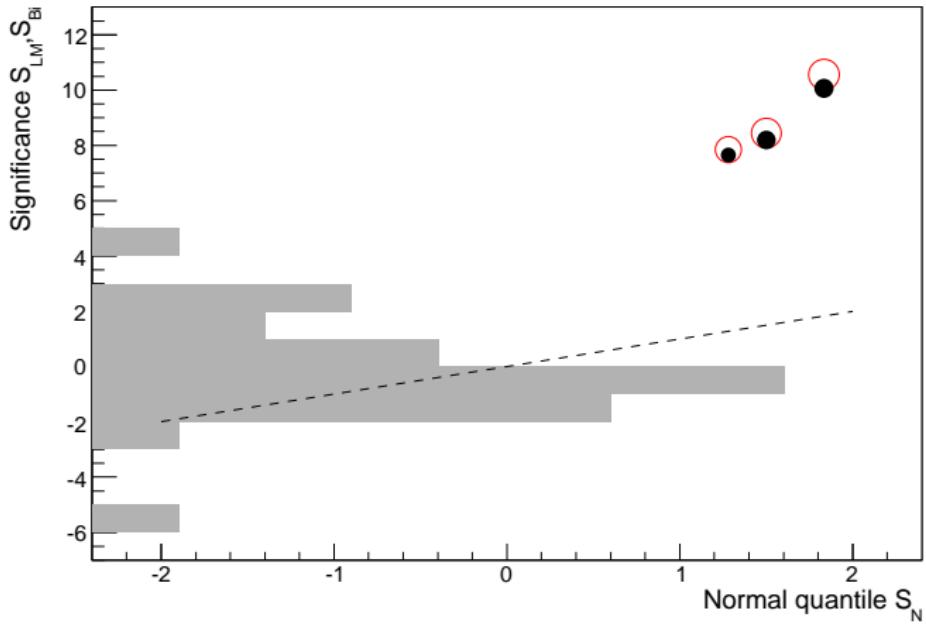
PKS 2155–304



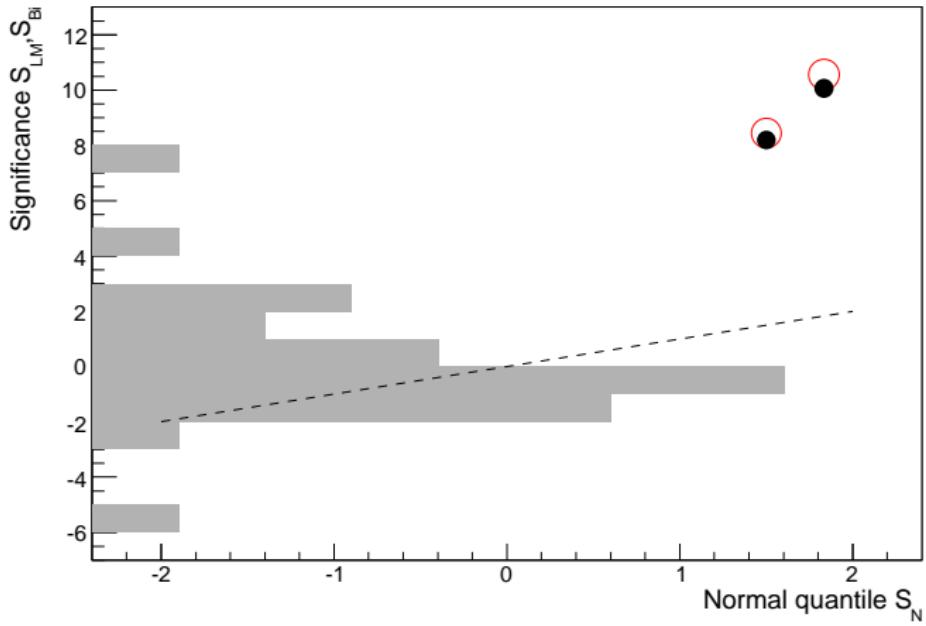
PKS 2155–304

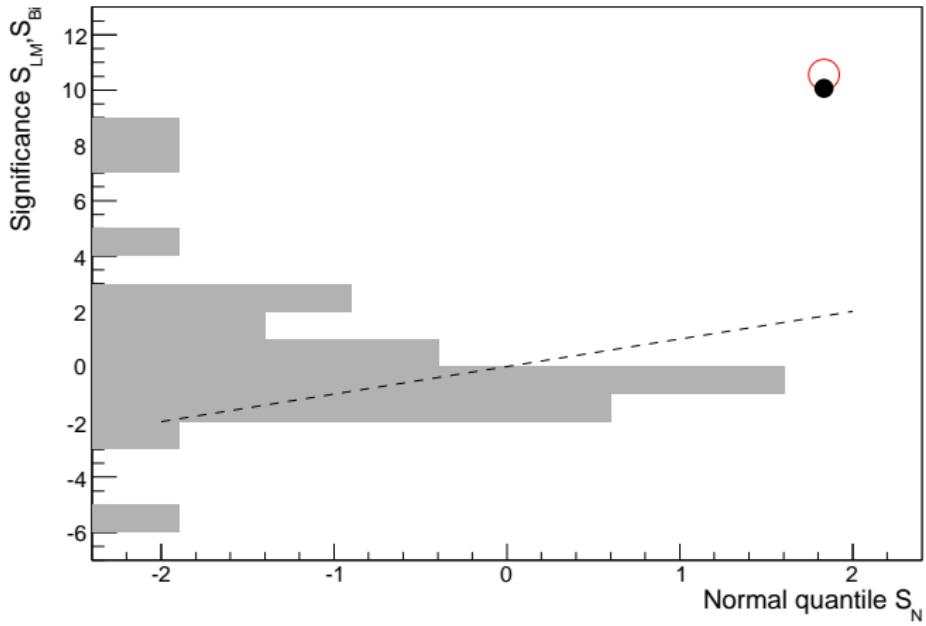


PKS 2155–304

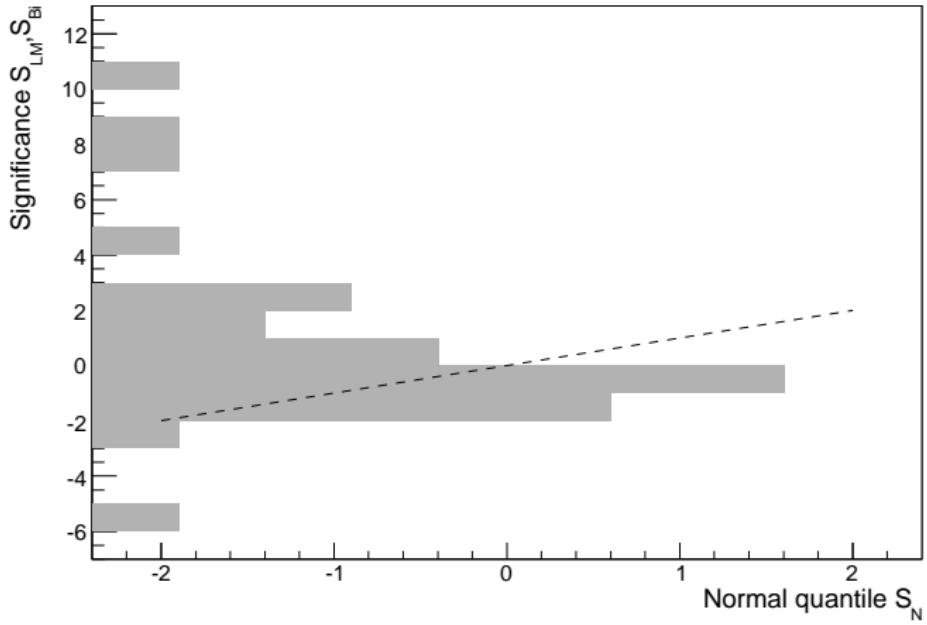


PKS 2155–304





PKS 2155–304



PKS 2155–304

