



# Fermi Science: Theoretical Overview

27 May 2014

Fermi Science School, Lewes, Delaware

“Not only is the devil in the details,  
but he lurks in the background.”

Charles D. Dermer and Govind Menon

## HIGH ENERGY RADIATION FROM BLACK HOLES

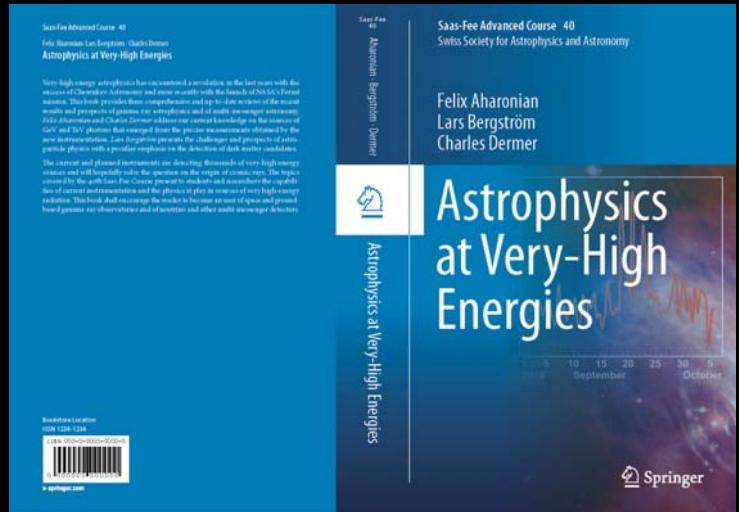
Gamma Rays,  
Cosmic Rays,  
and Neutrinos.

PRINCETON SERIES IN ASTROPHYSICS

## Outline

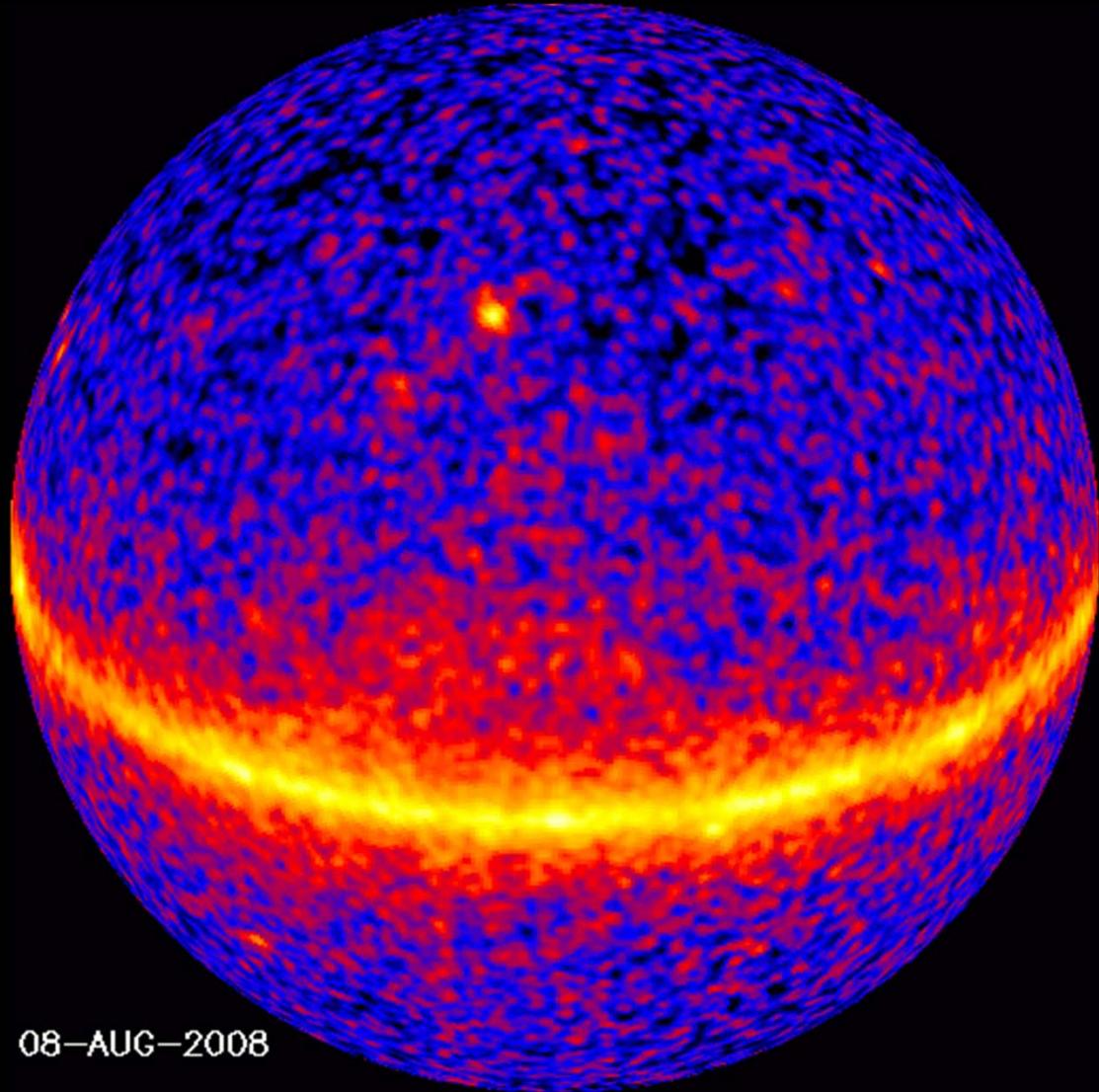
1. Sources and Subjects
2. Processes
3. Themes

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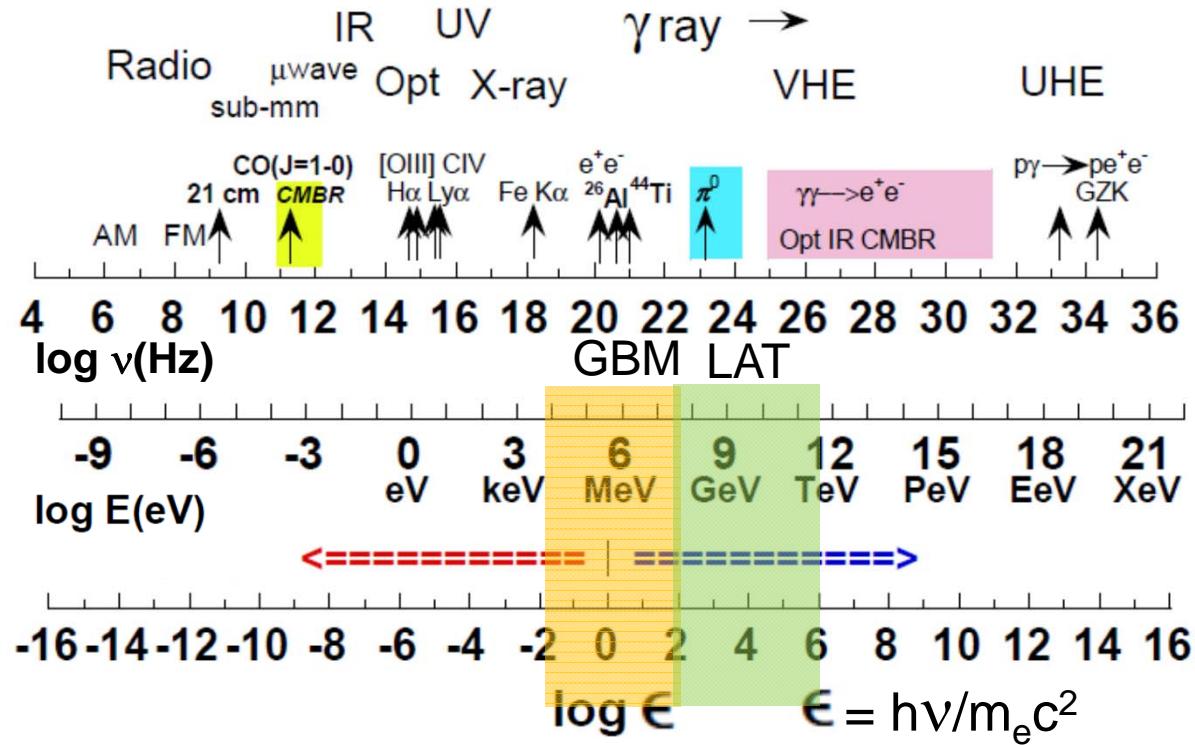
# Fermi $\gamma$ -ray sky



08-AUG-2008

- >100 MeV, 36 months
- Galactic  $\gamma$ -ray glow: accelerated particles meet target gas and photons
- ~80% of the emission is diffuse
- Transient and flaring sources
- Normal and ms pulsars
- Blazars
- GRBs
- Other  $\gamma$ -ray galaxies
- Unidentified sources

# Multiwavelength Astronomy



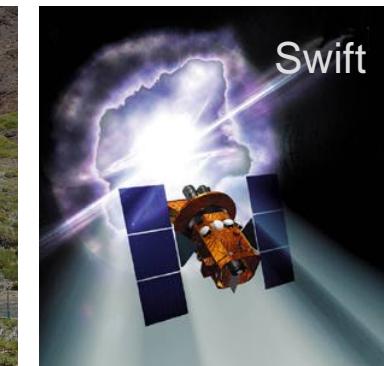
thermal vs. nonthermal

$\gamma$  rays: particle acceleration or dark matter signatures

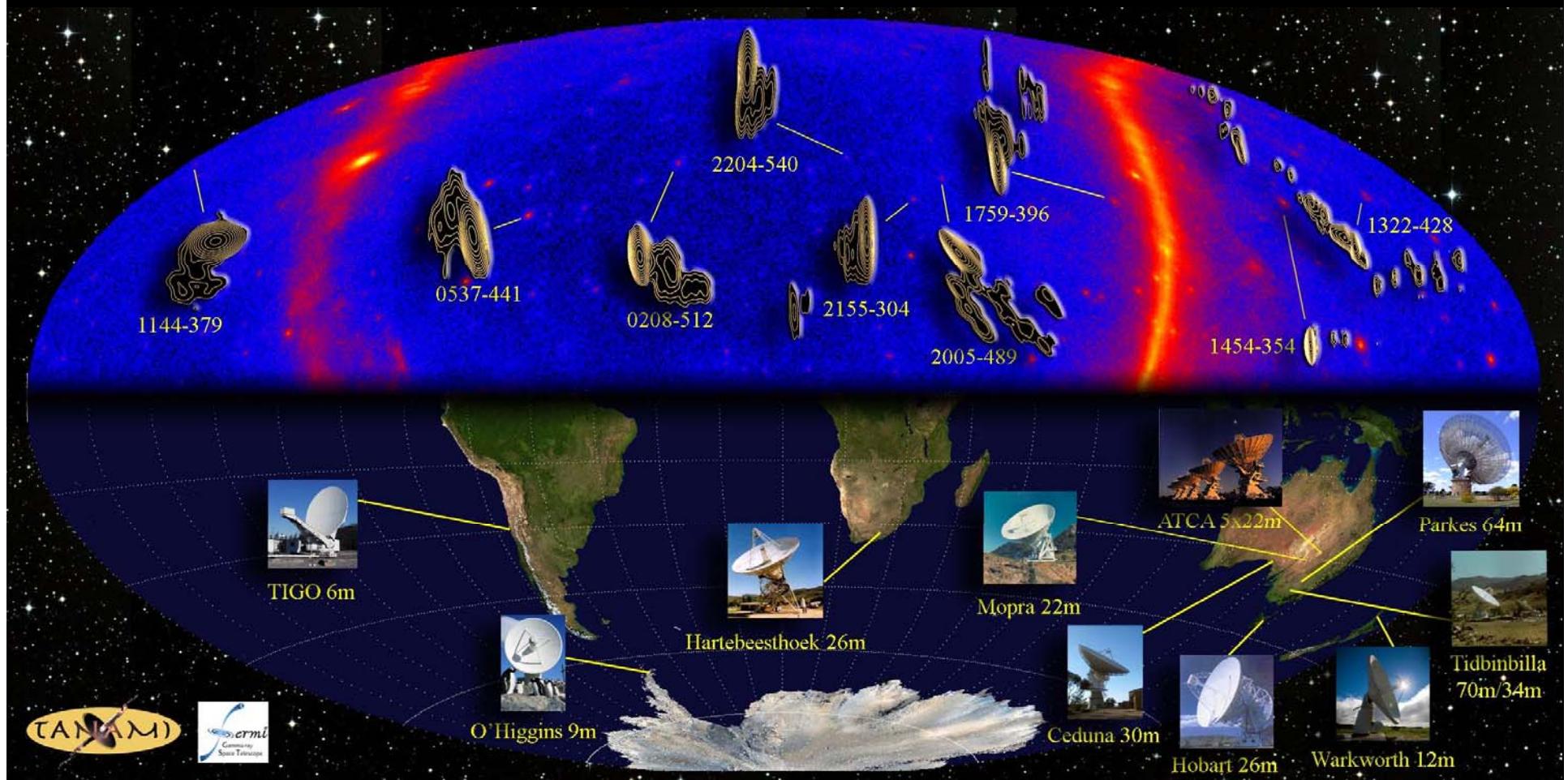
GeV vs. TeV astronomy

# High-Energy Observatories

- 0. Compton Gamma-Ray Observatory: Pioneering  $\gamma$ -ray space observatory (1991-2000)
  - 1. Swift Gamma-ray Burst Explorer (NASA 2004 MidEx)
  - 2. High Energy Stereoscopic Observatory (HESS)  
(Ground-based  $\gamma$ -ray telescope; Namibia, 2004)
  - 3. Very Energetic Radiation Imaging Telescope Array System  
(VERITAS) (Arizona; 2007)
  - 4. Major Atmospheric Gamma-ray Imaging Cherenkov Telescopes (MAGIC)  
(Canary Islands; 2004)
  - 5. Fermi Gamma-ray Space Telescope (2008)
- High Altitude Water Cherenkov (HAWC)  
➤ Cherenkov Telescope Array (CTA)  
Future ground-based  $\gamma$ -ray telescope arrays



# Worldwide Radio Networks of Observers



TANAMI: Tracking AGN with Austral Milliarcsecond Interferometry  
MOJAVE: Monitoring of Jets of AGN with VLBA Experiments

# Optical Networks: WEBT-- GASP



Whole Earth Blazar Telescope

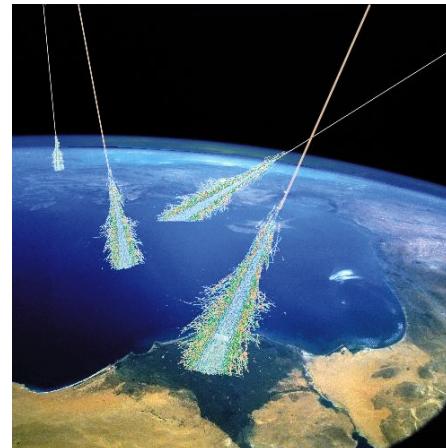
GLAST-AGILE Support Program



# Multimessenger Astronomy

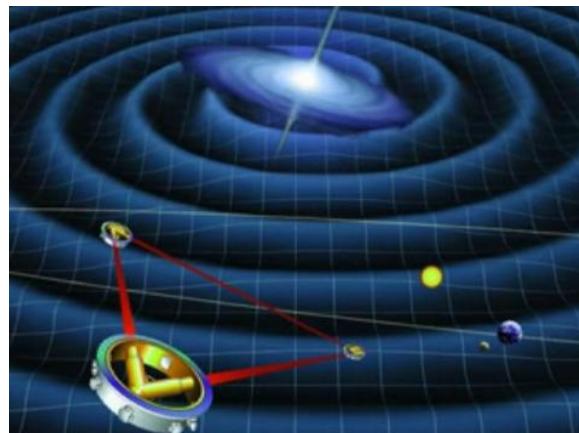


Photons

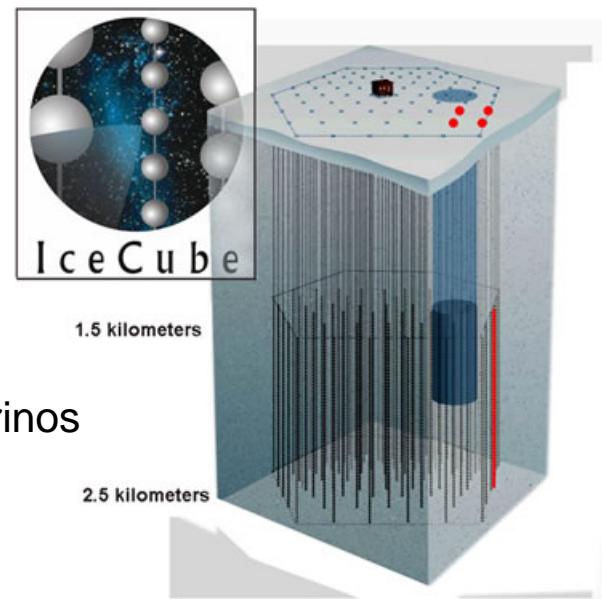


Cosmic rays

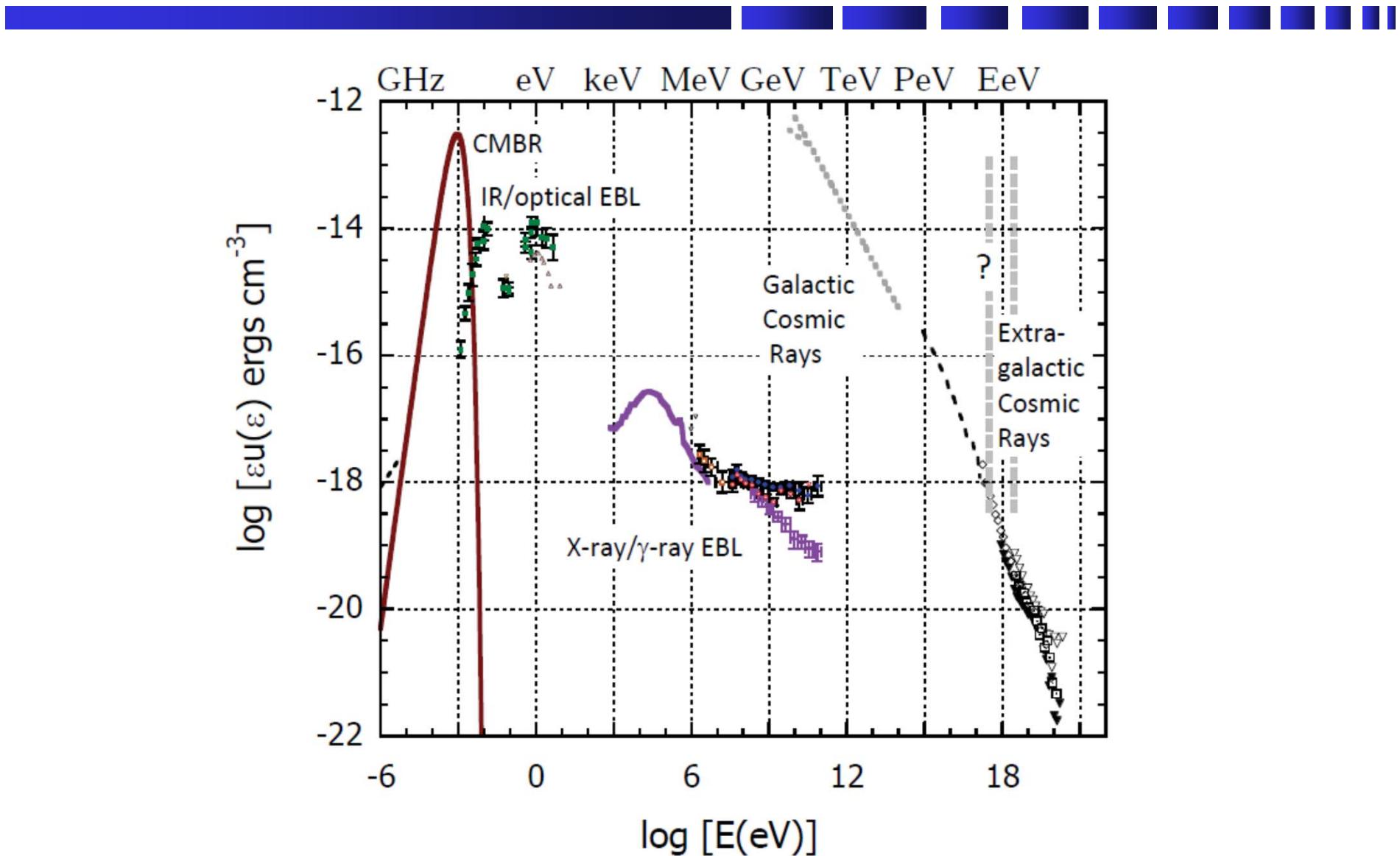
Gravitational waves



Neutrinos



# Diffuse Radiations



# Fermi LAT: Subjects and Sources



## Calibration and Analysis

Energy/Effective Area; FoV; Imaging  
LLE/Transient/Source/Diffuse Classes

## Catalogs

205    132    1500    700    1900    1017    2100    1450  
BSL/LBAS/1FGL/1LAC/2FGL/2LAC/3FGL/3LAC  
1<sup>st</sup> PSR/2<sup>nd</sup> PSR/GRB/SNR/1FHL/FAVA

45    114    7/28    62/6

## Galactic Sources

Pulsars  
    normal (radio-loud and radio quiet), ms  
Pulsar Wind Nebulae/Crab/flares  
Globular Clusters  
SNRs  
High-mass X-ray/gamma-ray binaries  
Novae

## Diffuse, Molecular Clouds, and Other Galaxies

Diffuse Galactic Emission  
Galactic Center/Fermi Bubbles  
LMC, SMC, M31  
Starburst Galaxies  
Cosmic-ray electrons/positrons  
EBL  
DIGB/EGB

## Sources in the Solar System

Earth/CRs  
Sun/ Quiescent and Solar Flares  
Moon  
Jupiter/Zodiacal Light  
TGFs

## Blazars and other AGNs

Blazars/ FSRQs and BL Lacs  
Radio Galaxies  
RLNLSy1

## Gamma Ray Bursts

Long Soft  
Short Hard  
SGRs

## Dark Matter and New Physics

DM lines and features  
Axions  
Lorentz Invariance Violation  
Primordial Black Hole Annihilation  
Magnetogenesis

# Fermi: Processes



## Particle Interactions and Radiation

Leptonic

e

e

Hadronic

Photons

p (ion)

p (ion)

Magnetic Fields

Cascades

$\gamma$

$\gamma$

## Acceleration Physics

B

B

Fermi Processes

Electrodynamic Acceleration

Magnetic Reconnection

## Relativistic Flows

## Black Hole Physics

Accretion

Blandford-Znajek Process

Jet formation

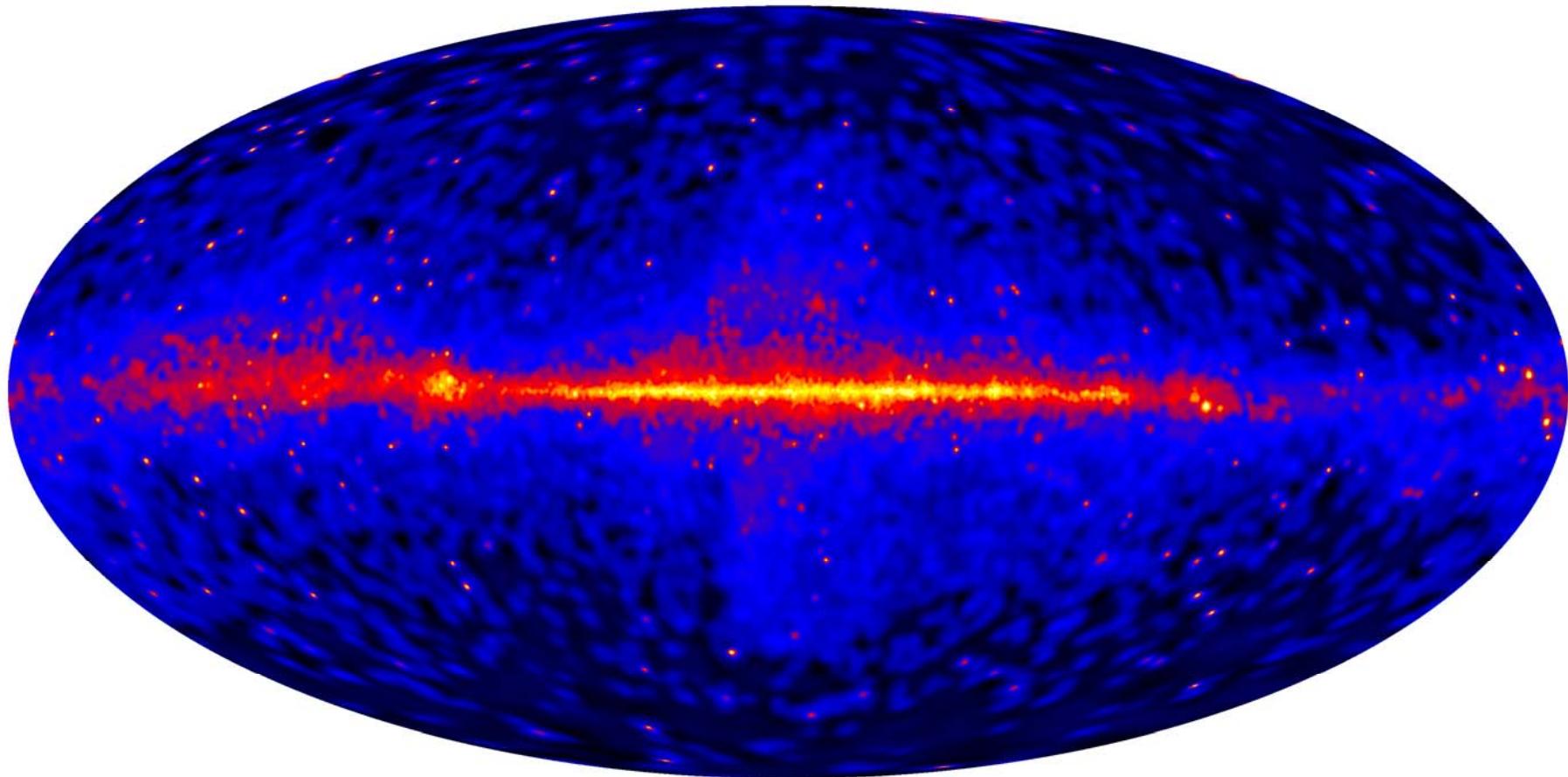
## Dark Matter Decay and Annihilation

## Fermi: Themes



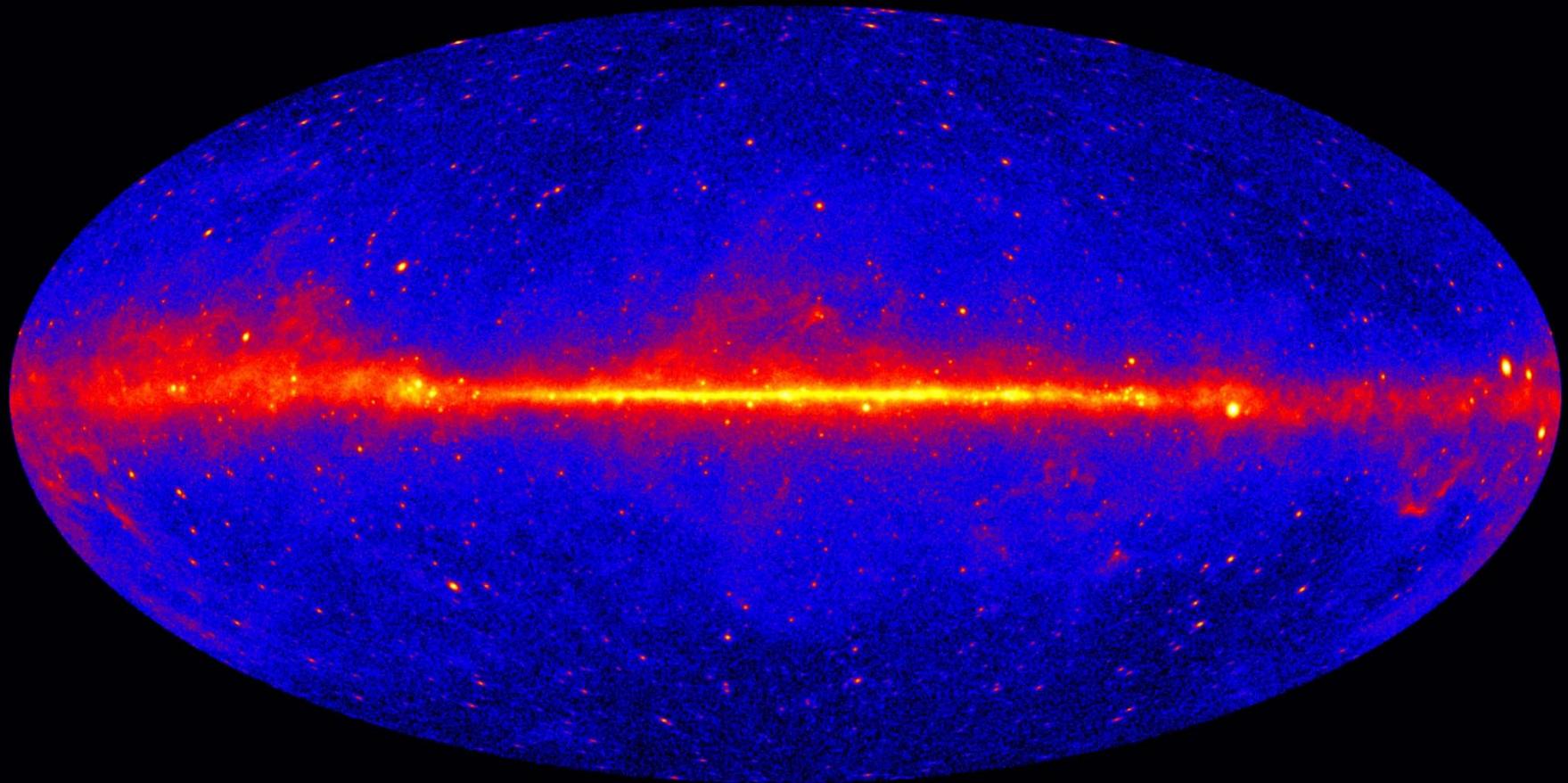
- Cosmic Ray and Ultra-High Energy Cosmic Ray Origin**
- The Search for Dark Matter**
- Extreme Stars and the Endpoints of Stellar Evolution**
- Structure and Content of the Galaxy in  $\gamma$  rays**
- Nonthermal Particle Acceleration in Nature**
- The Black-Hole Universe**
- Radiation and Field Content of the Universe**

# Fermi: >10 GeV Skymap





# Fermi >1 GeV 5 Year Sky Map



Dermer

Fermi Summer Science School 2014

Lewes, DE