

# Proposal for *Swift* Cycle 3 Observations

## Target Summary

**Principal Investigator**

DR. RITA M SAMBRUNA

**Proposal Title**

SWIFT TOO OBSERVATIONS OF FLARING BLAZARS TRIGGERED BY GLAST

Tar No	Target Name	R.A.	Dec	Obs Time (ksec)
1	PKS 0528+134	05 30 56.40	+13 31 55.0	105.00
2	OJ 287	08 54 48.80	+20 06 30.0	105.00
3	3C 279	12 56 11.20	-05 47 21.5	105.00
4	PKS 1622-297	16 26 06.00	-29 51 27.0	105.00
5	3C 454.3	22 53 57.70	+16 08 54.0	105.00
6	BL LACERTAE	22 02 43.30	+42 16 40.0	105.00
7	UNKNOWN	15 00 00.00	+42 00 00.0	105.00

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<b>Proposal Title</b> SWIFT TOO OBSERVATIONS OF FLARING BLAZARS TRIGGERED BY GLAST

<b>Target Number 1</b>	<b>Name</b> PKS 0528+134
<b>J2000</b>	<b>R.A.</b> 05 30 56.40
<b>Position</b>	<b>Dec.</b> +13 31 55.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 20.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b>  The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 2	<b>Name</b> 0J 287
<b>J2000</b>	<b>R.A.</b> 08 54 48.80
<b>Position</b>	<b>Dec.</b> +20 06 30.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 14.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 3	<b>Name</b> 3C 279
<b>J2000</b>	<b>R.A.</b> 12 56 11.20
<b>Position</b>	<b>Dec.</b> -05 47 21.5

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 17.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 4	<b>Name</b> PKS 1622-297
<b>J2000</b>	<b>R.A.</b> 16 26 06.00
<b>Position</b>	<b>Dec.</b> -29 51 27.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 20.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 5	<b>Name</b> 3C 454.3
<b>J2000</b>	<b>R.A.</b> 22 53 57.70
<b>Position</b>	<b>Dec.</b> +16 08 54.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 16.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 6	<b>Name</b> BL LACERTAE
<b>J2000</b>	<b>R.A.</b> 22 02 43.30
<b>Position</b>	<b>Dec.</b> +42 16 40.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 14.5

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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<b>Target Number</b> 7	<b>Name</b> UNKNOWN
<b>J2000</b>	<b>R.A.</b> 15 00 00.00
<b>Position</b>	<b>Dec.</b> +42 00 00.0

<b>Target Observing Time (ksec)</b> 105.00	<b>Estimated XRT Count Rate</b> 0.100
	<b>V Magnitude (if known)</b> 15.0

UVOT Filters	
U	Y
B	Y
V	Y
UVW1	Y
UVW2	Y
UVM2	Y
Ugrism	N
Vgrism	N
White	N
Blocked	N

<b>Remarks</b> The XRT count rate is arbitrary as we do not know how bright the source will be during the To0. In addition, the source will likely be variable during the requested sampling.
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