



Fermi  
Gamma-ray Space Telescope



# A tale of cosmic rays narrated in $\gamma$ rays by *Fermi*

Luigi Tibaldo  
[ltibaldo@slac.stanford.edu](mailto:ltibaldo@slac.stanford.edu)



on behalf of  
the *Fermi*-LAT collaboration

ICRC 2013  
July 8 2013

# Outline

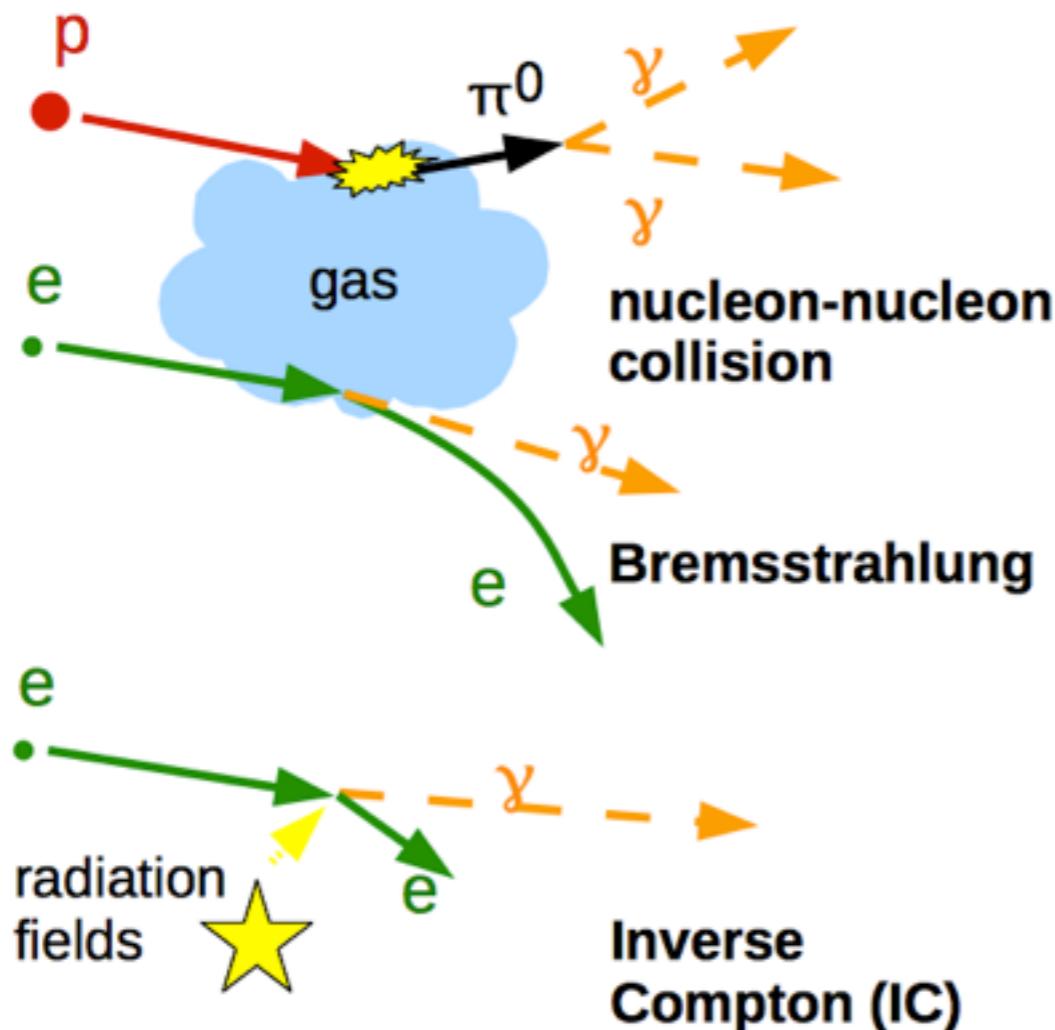
- **$\gamma$  rays as a cosmic-ray tracer**
- Cosmic-ray acceleration in supernova remnants
- Cosmic rays in massive star-forming regions
- Large-scale propagation of cosmic rays in the Milky Way
- Cosmic rays in external galaxies
- Challenges and summary

# Chasing cosmic rays



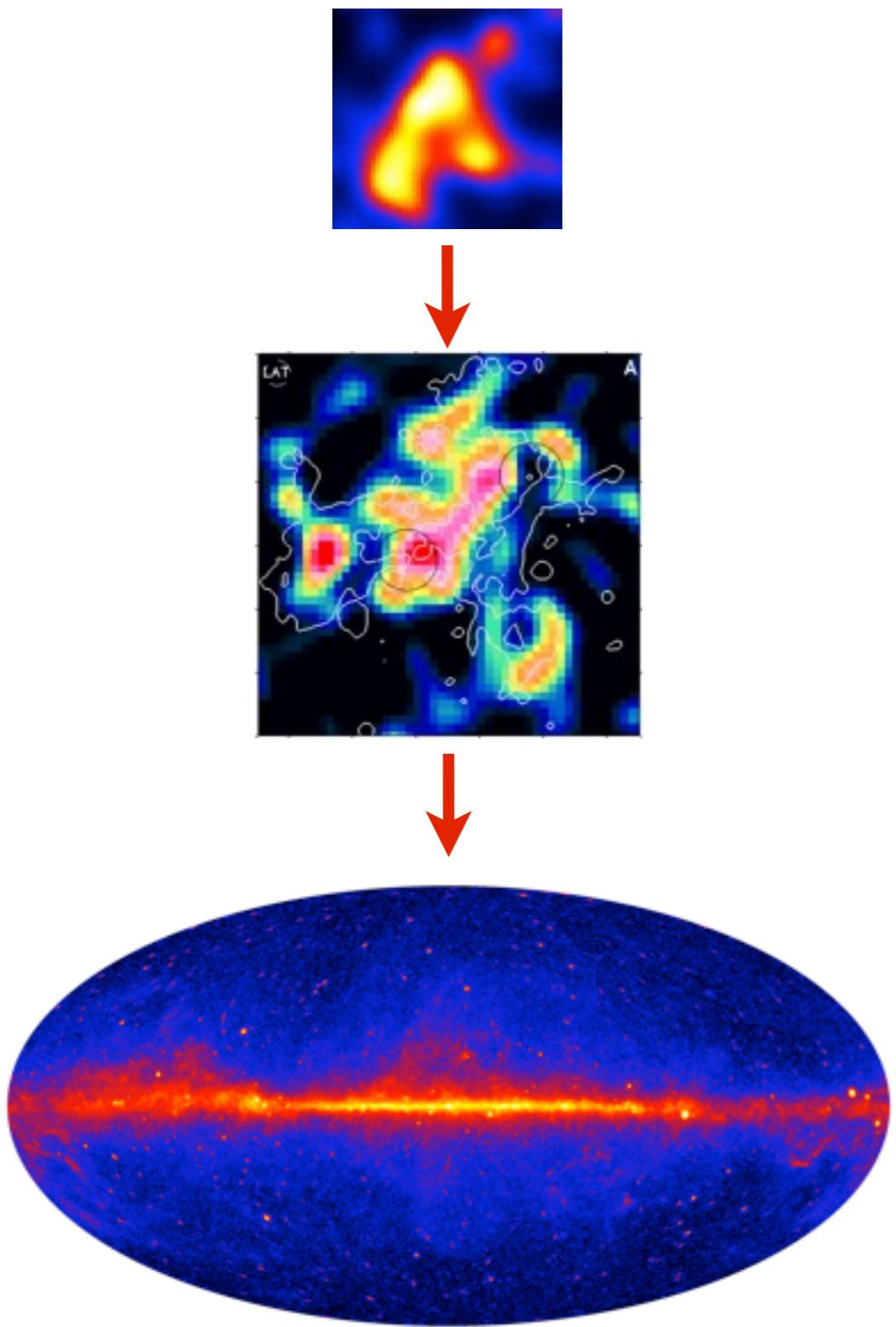
CR are charged + B fields  
→ do not track back to  
sources ( $< 10^{18}$  eV)

# $\gamma$ rays as a charged particle tracer



- neutral secondaries → complement direct observations
- $\gamma$  rays → neutral and easy to detect

# Fermi tells us the story of cosmic rays



focus on CRs below the knee,  $<10^{15}$  eV

- acceleration in supernova remnants
- link with massive-star forming regions/early propagation
- large-scale propagation
- external galaxies



# Outline

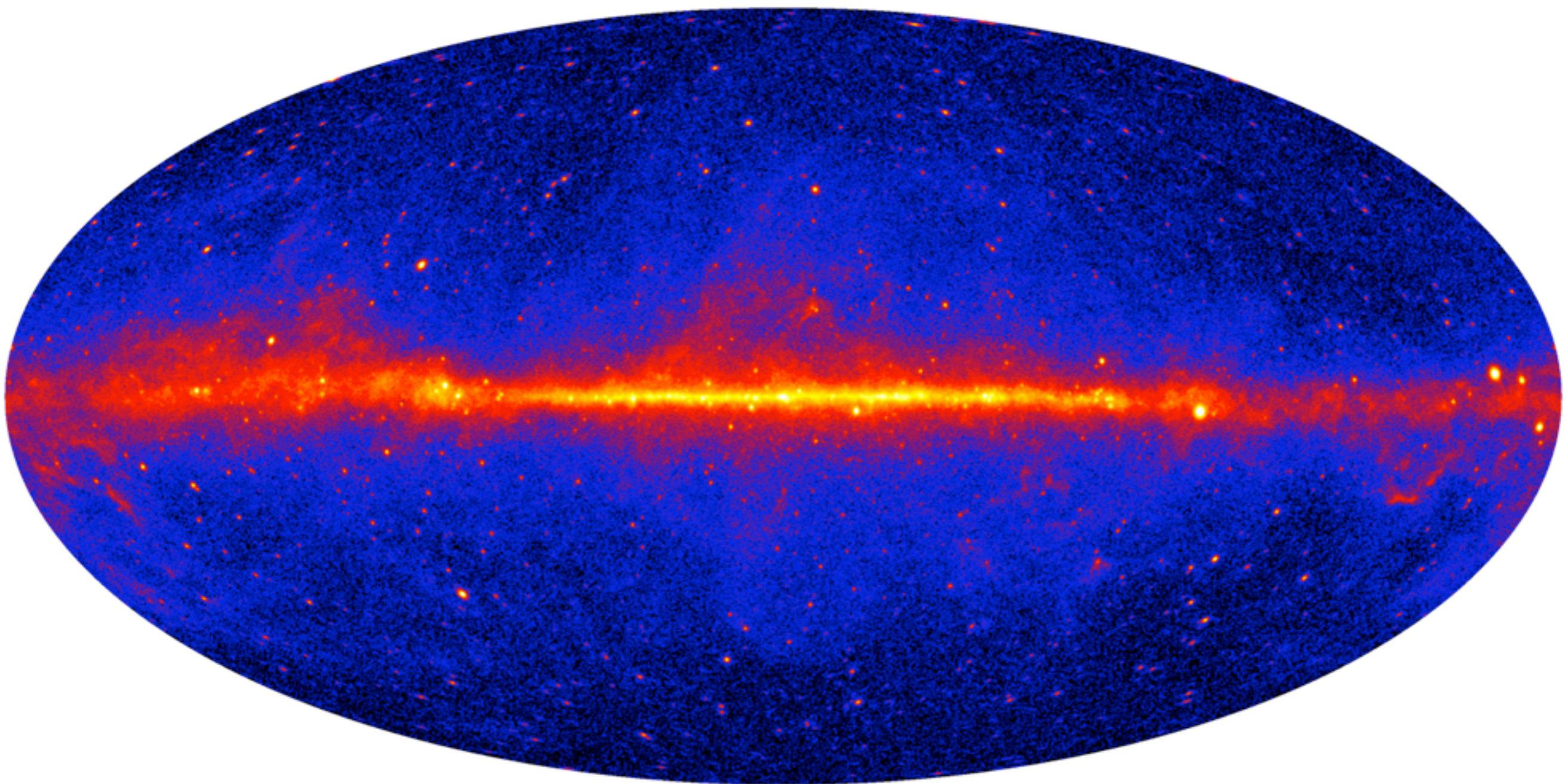
- $\gamma$  rays as a cosmic-ray tracer
- **Cosmic-ray acceleration in supernova remnants**
- Cosmic rays in massive star-forming regions
- Large-scale propagation of cosmic rays in the Milky Way
- Cosmic rays in external galaxies
- Challenges and summary

# Supernova remnants as CR sources

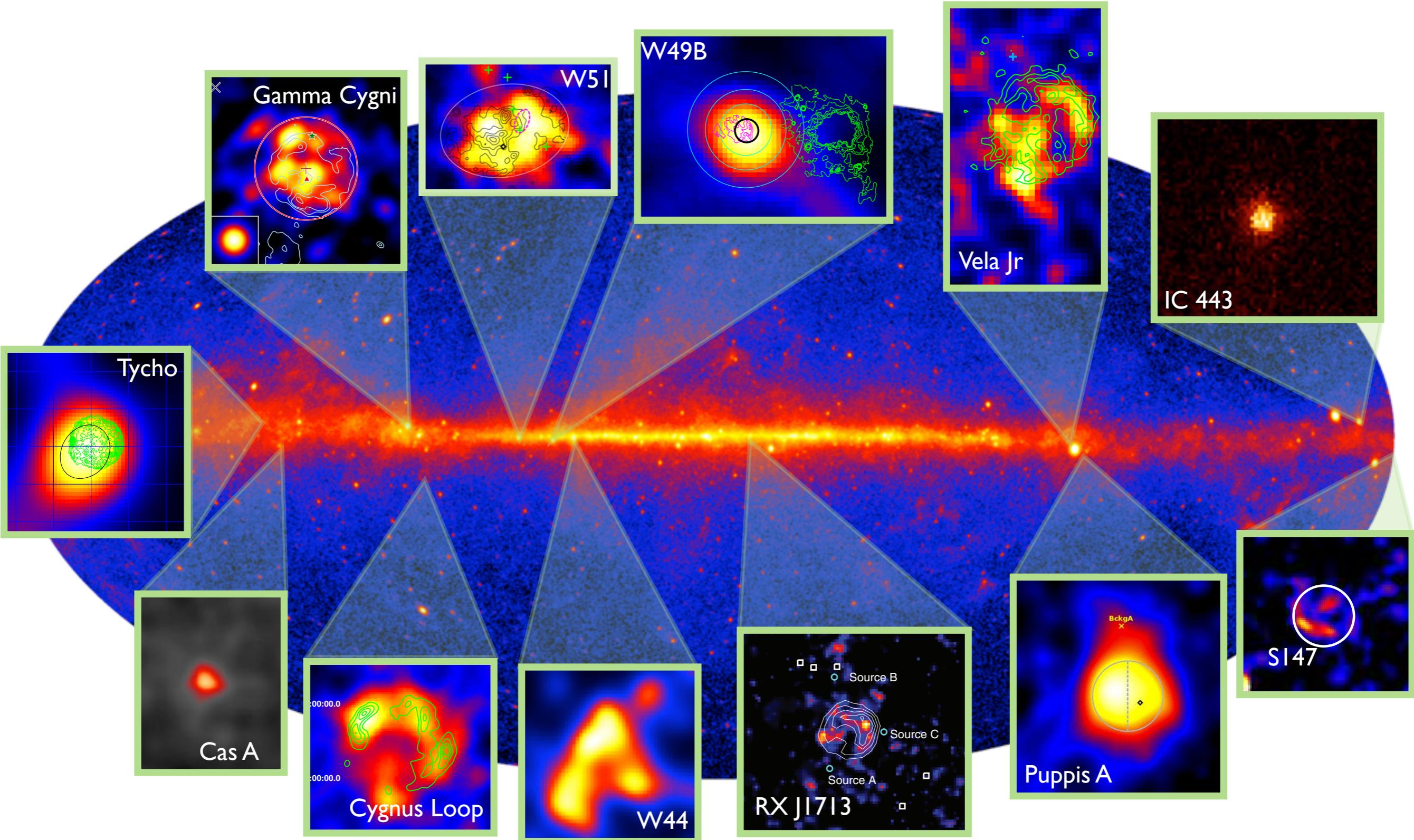
- energetic and numerous enough
- non-linear diffusive shock acceleration
- SNRs accelerate
  - electrons
  - nuclei? up to the knee?



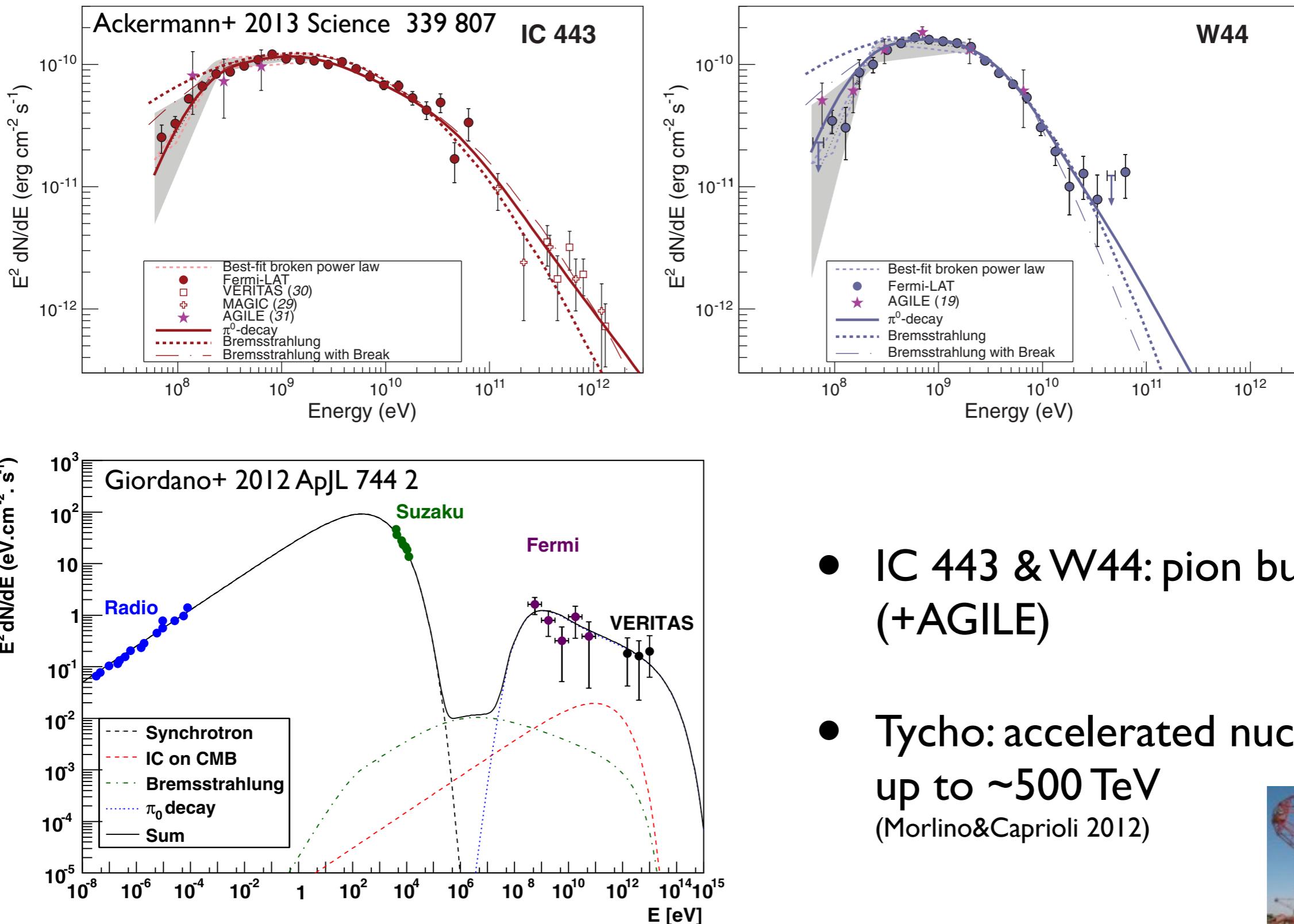
# Supernova remnants in the $\gamma$ -ray sky



# Supernova remnants in the $\gamma$ -ray sky



# Accelerated nuclei!

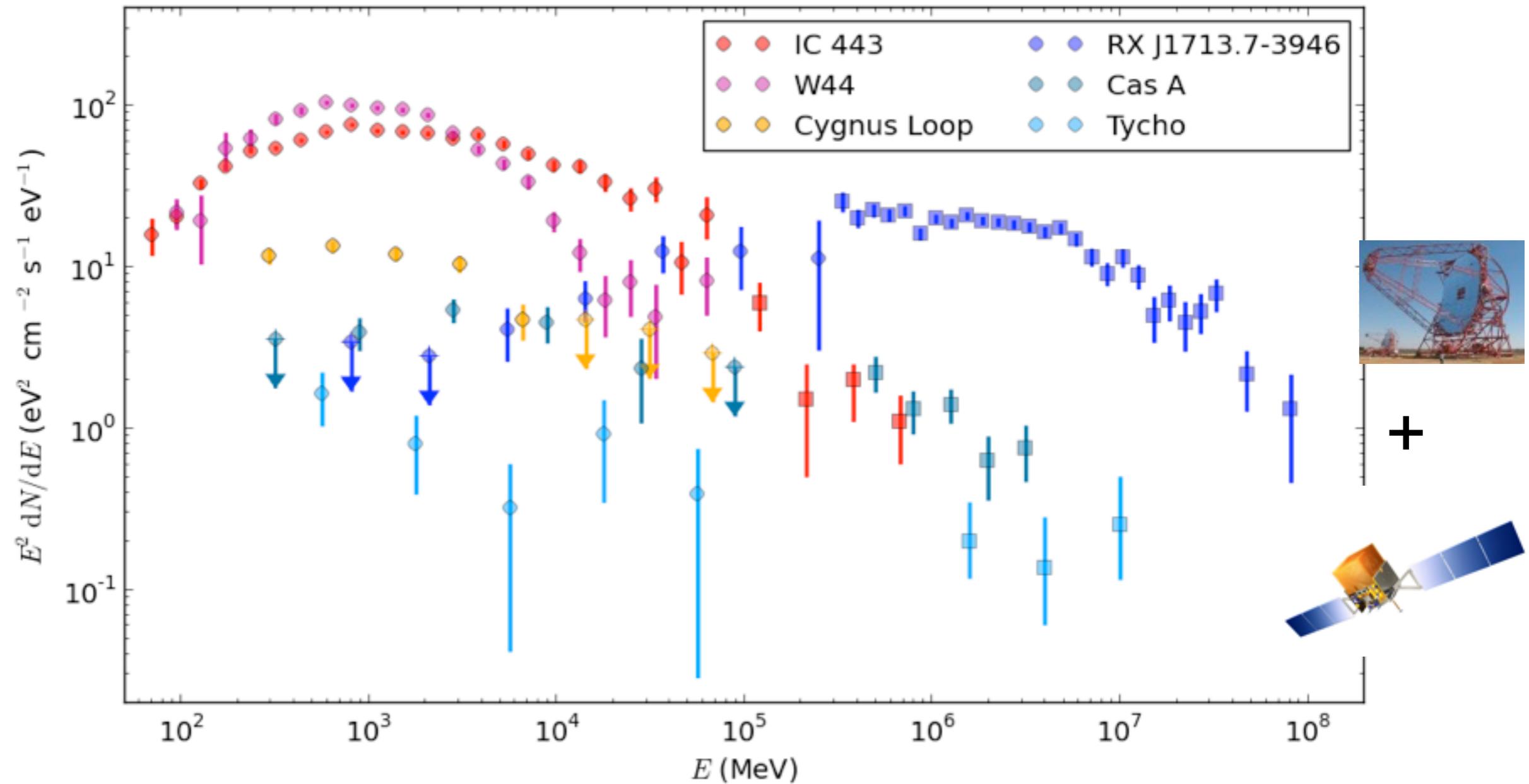


- IC 443 & W44: pion bump (+AGILE)

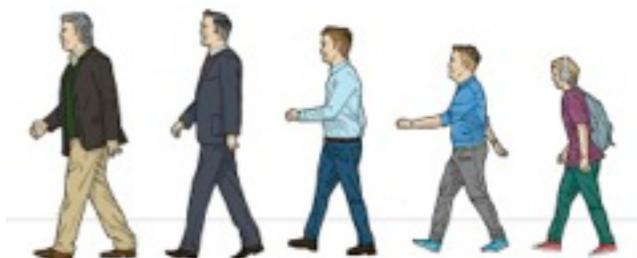
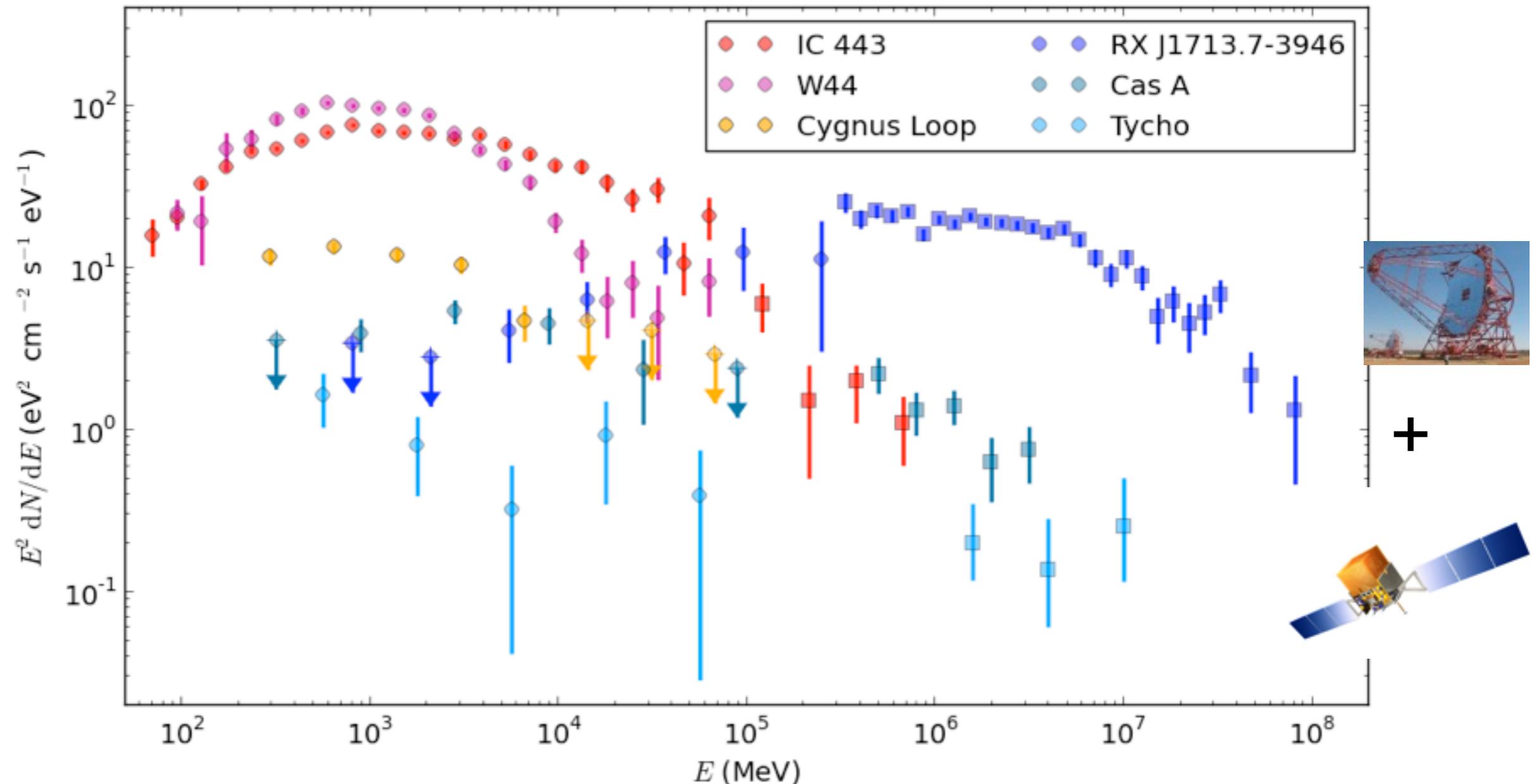
- Tycho: accelerated nuclei up to  $\sim 500$  TeV  
(Morlino&Caprioli 2012) ???



# The ages of supernova remnants

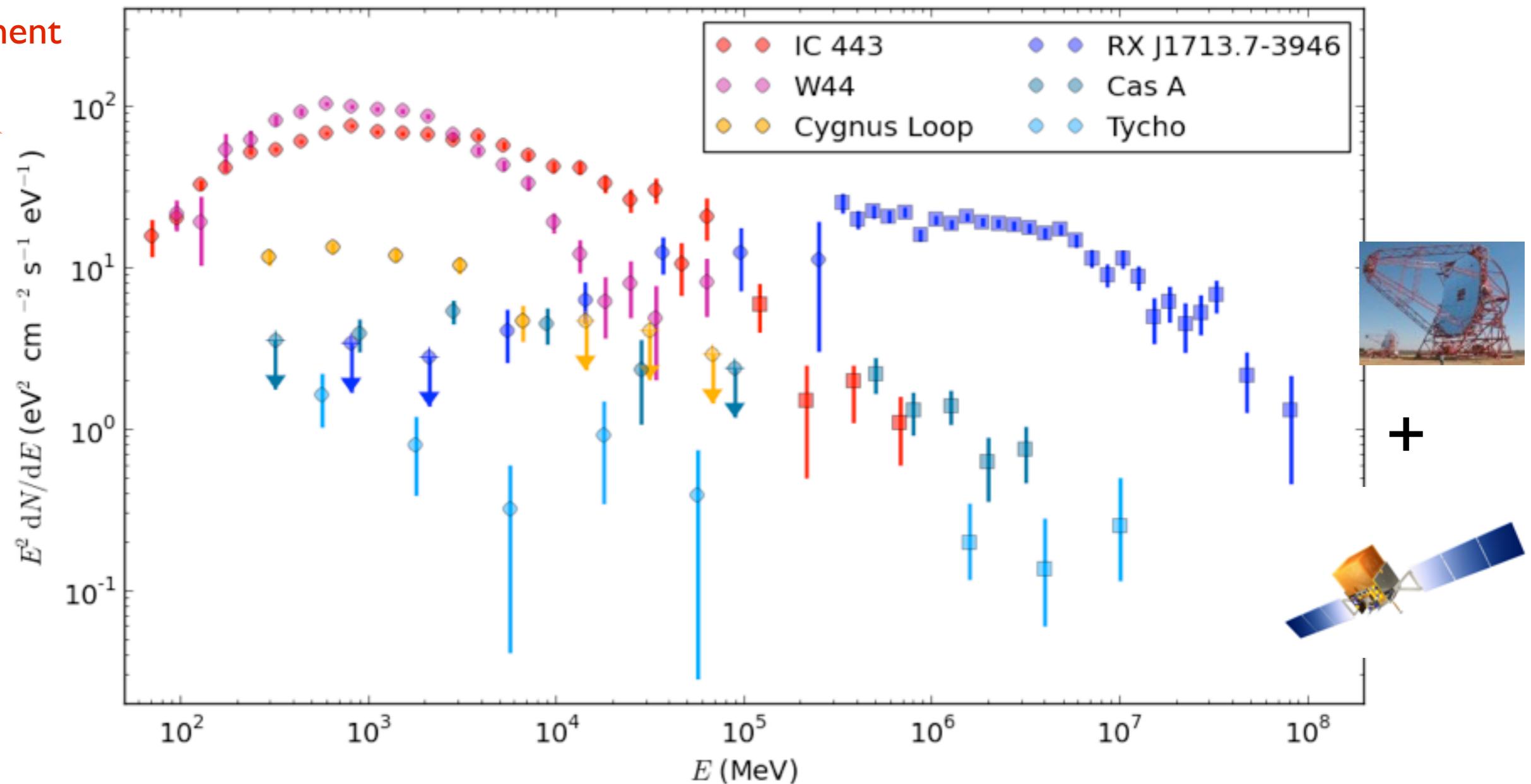


# The ages of supernova remnants



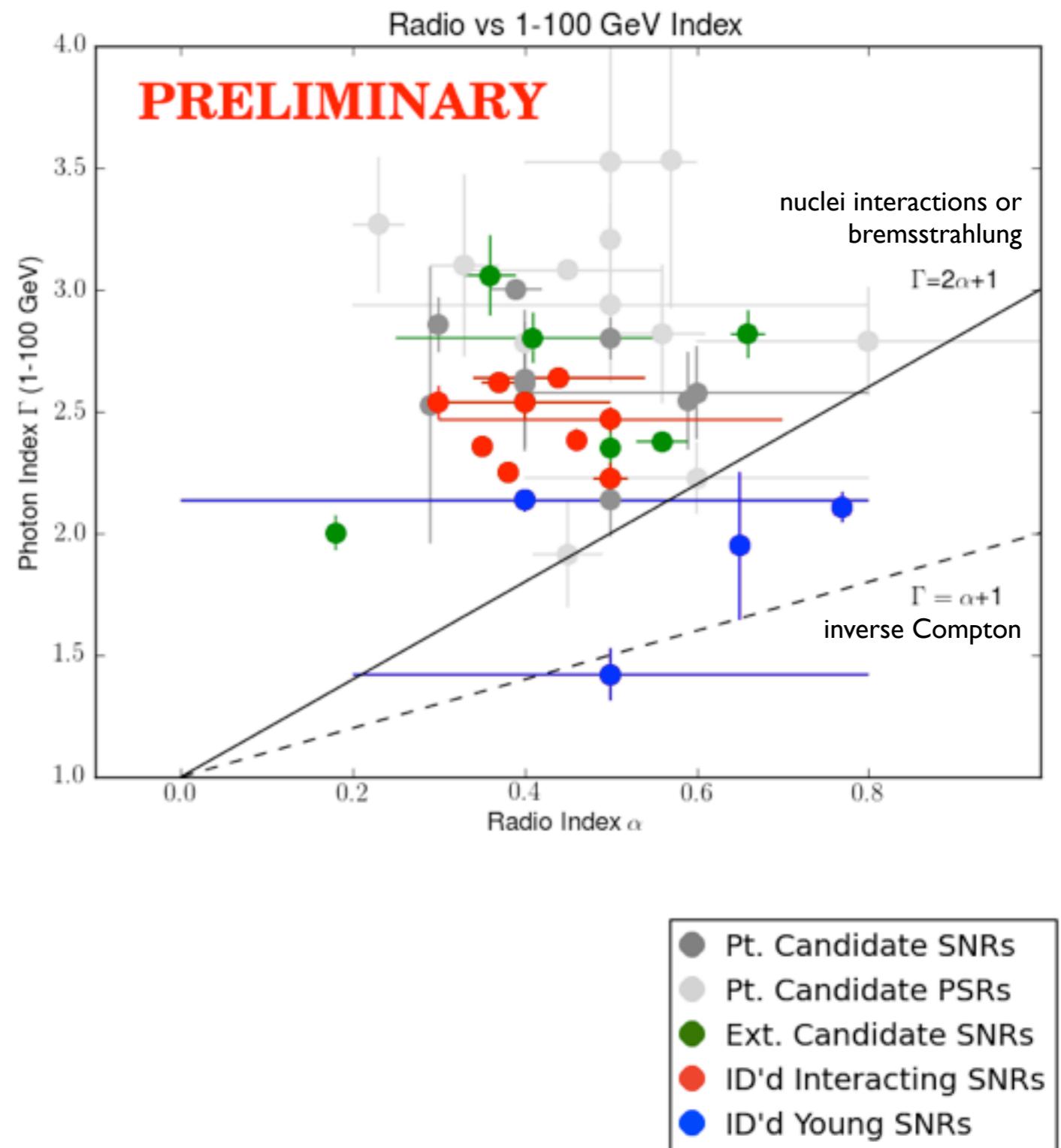
# The ages of supernova remnants

denser  
environment



# The first LAT SNR Catalog

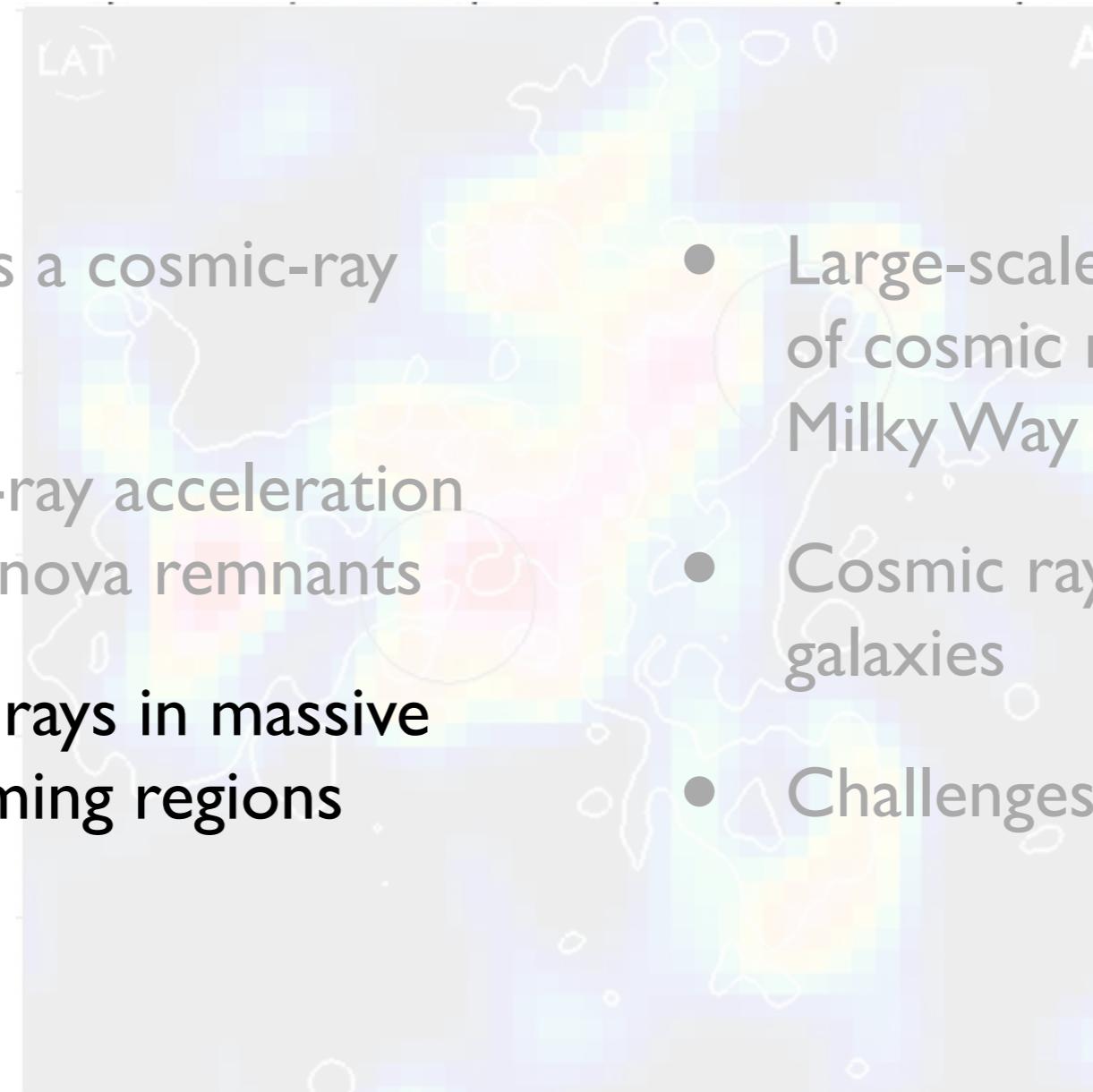
- systematic/uniform characterization of radio SNRs
- SNRs as a population of CR sources



talks by  
Brandt, Hewitt

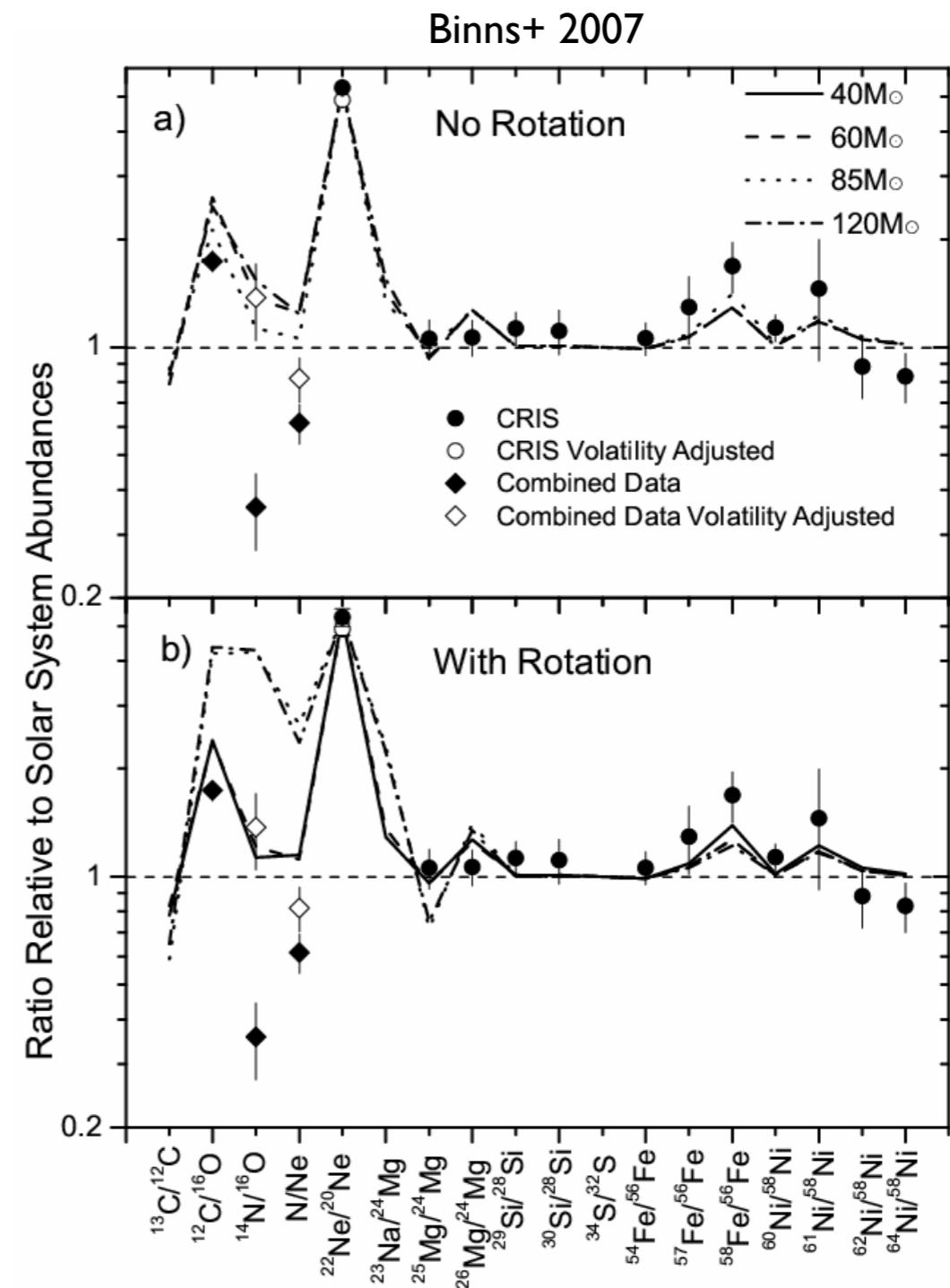
# Outline

- $\gamma$  rays as a cosmic-ray tracer
- Cosmic-ray acceleration in supernova remnants
- **Cosmic rays in massive star-forming regions**
- Large-scale propagation of cosmic rays in the Milky Way
- Cosmic rays in external galaxies
- Challenges and summary

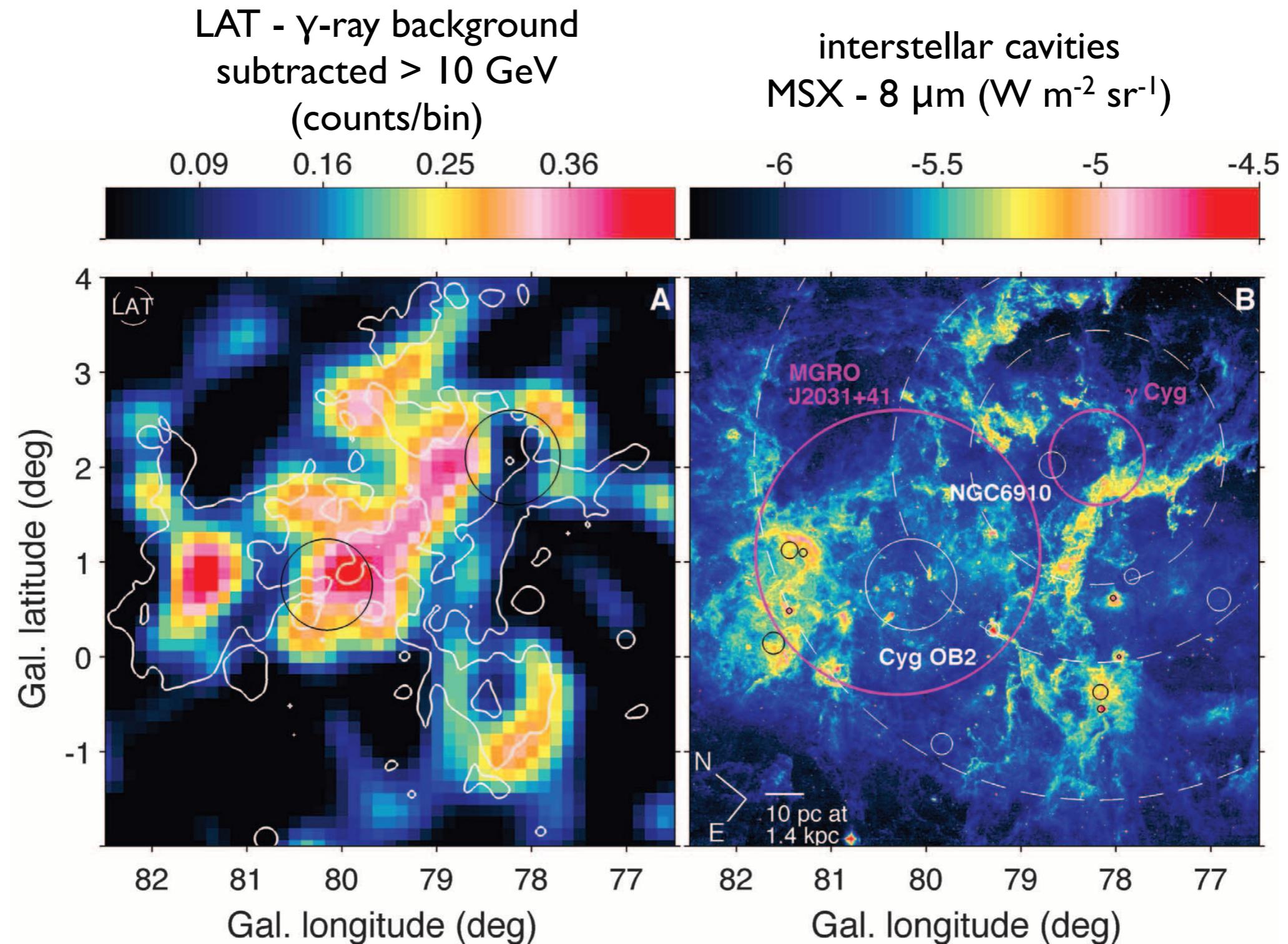


# A link with massive-star forming regions?

- isotopic abundances of WR stars ( $^{22}\text{Ne}$ , > Fe)
- ~80% of supernovae in massive-star clusters
- superbubbles?
- impact of massive-star environment on young CRs?



# The *Fermi* LAT view of Cygnus X



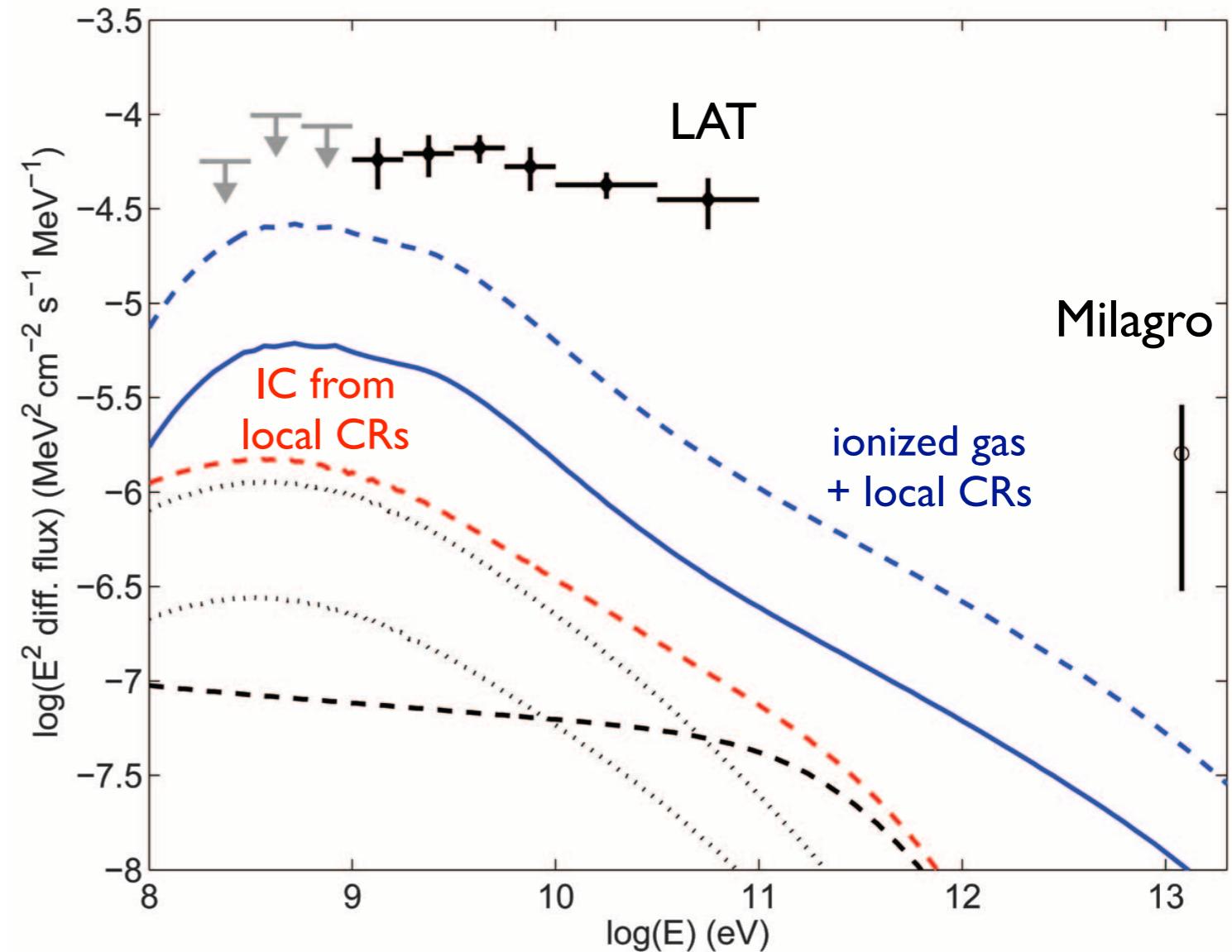
Ackermann+ 2011 Science 334 1103

# A cocoon of young cosmic rays

- requires **freshly-accelerated CRs**
  - hadronic → too soft
  - leptonic → too soft and faint

$$\frac{dN}{dE} \times (1.5 - 2) \left( \frac{E}{10 \text{ GeV}} \right)^{0.3}$$

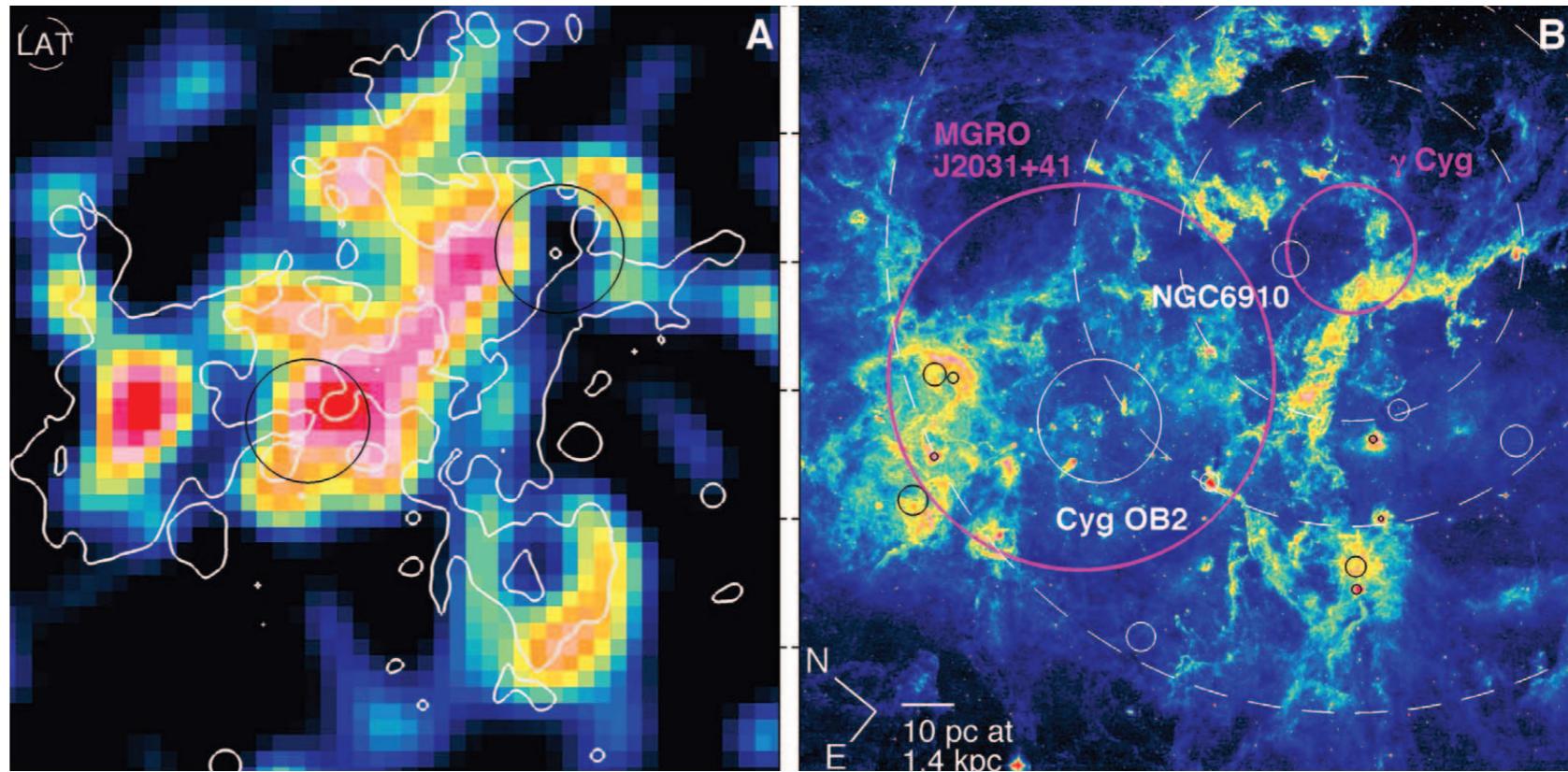
$$\frac{dN}{dE} \times 60 \left( \frac{E}{10 \text{ GeV}} \right)^{0.5}$$



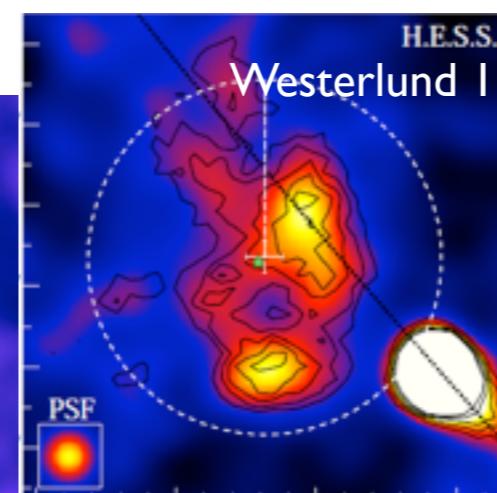
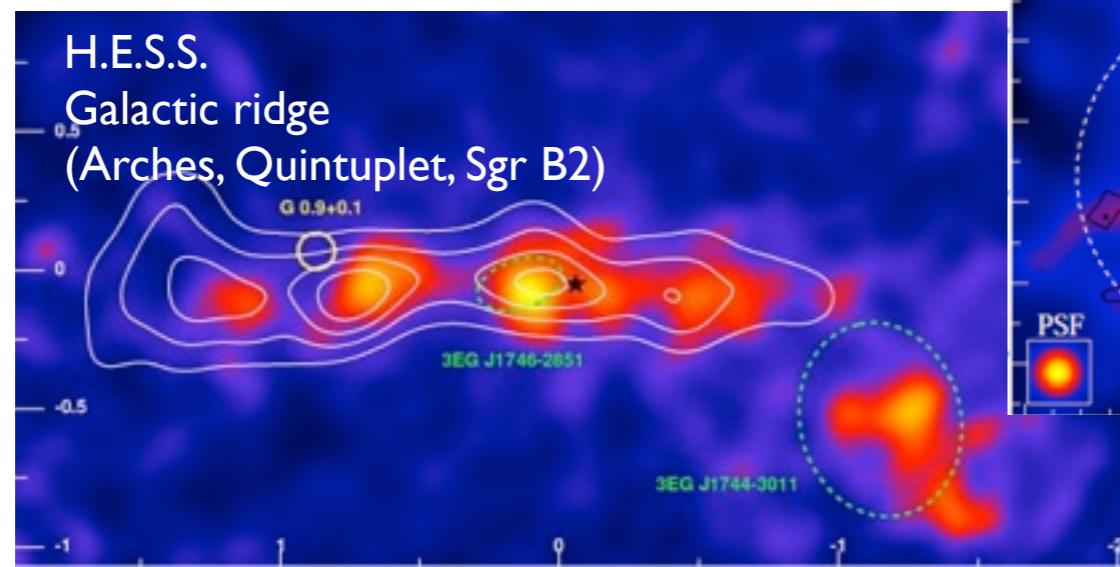
Ackermann+ 2011 Science 334 1103

# Origin and propagation

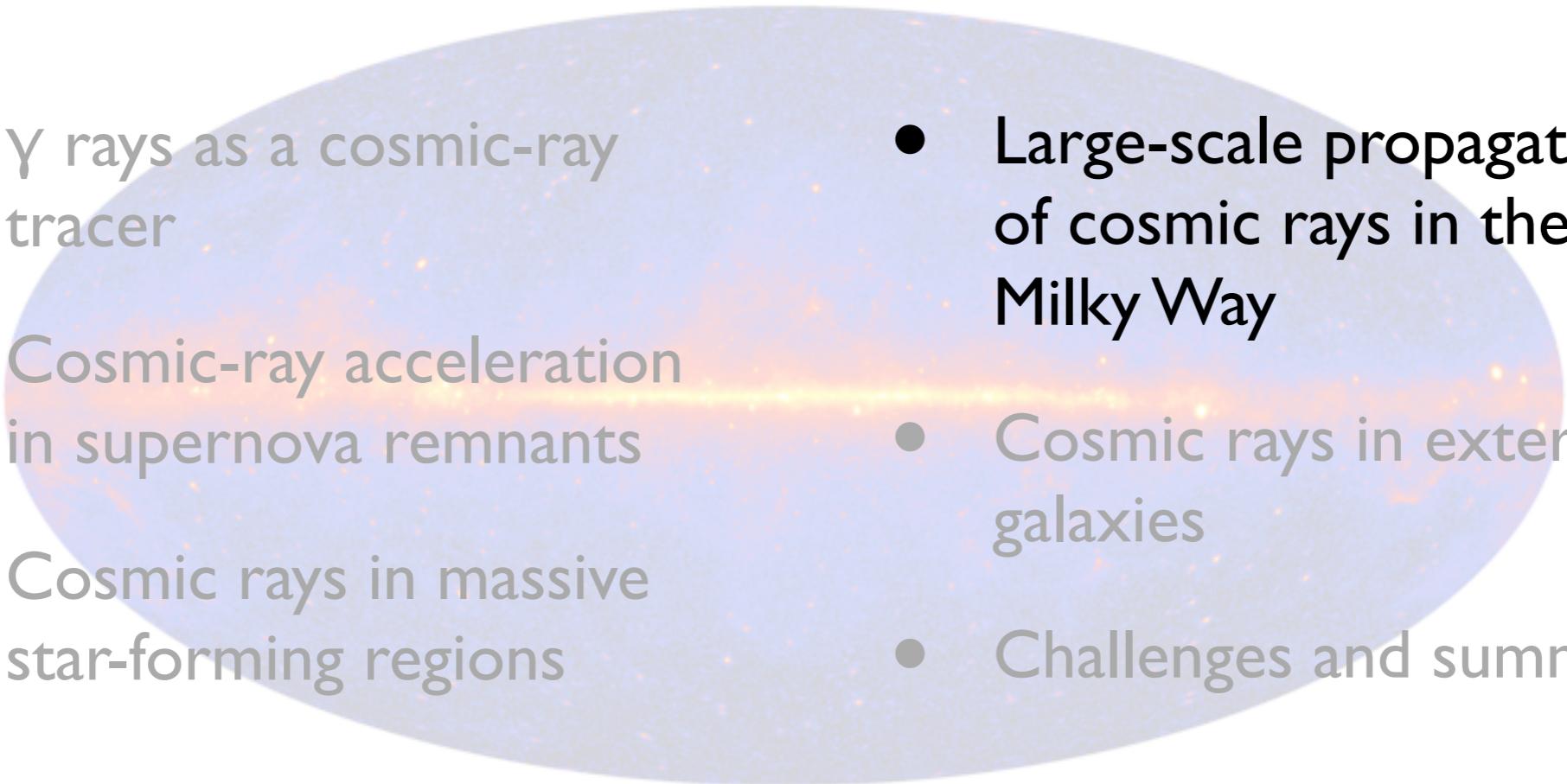
- Gamma Cygni supernova remnant?
- stellar-wind superbubble?
- active airlock?



Ackermann+ 2011 Science 334 1103

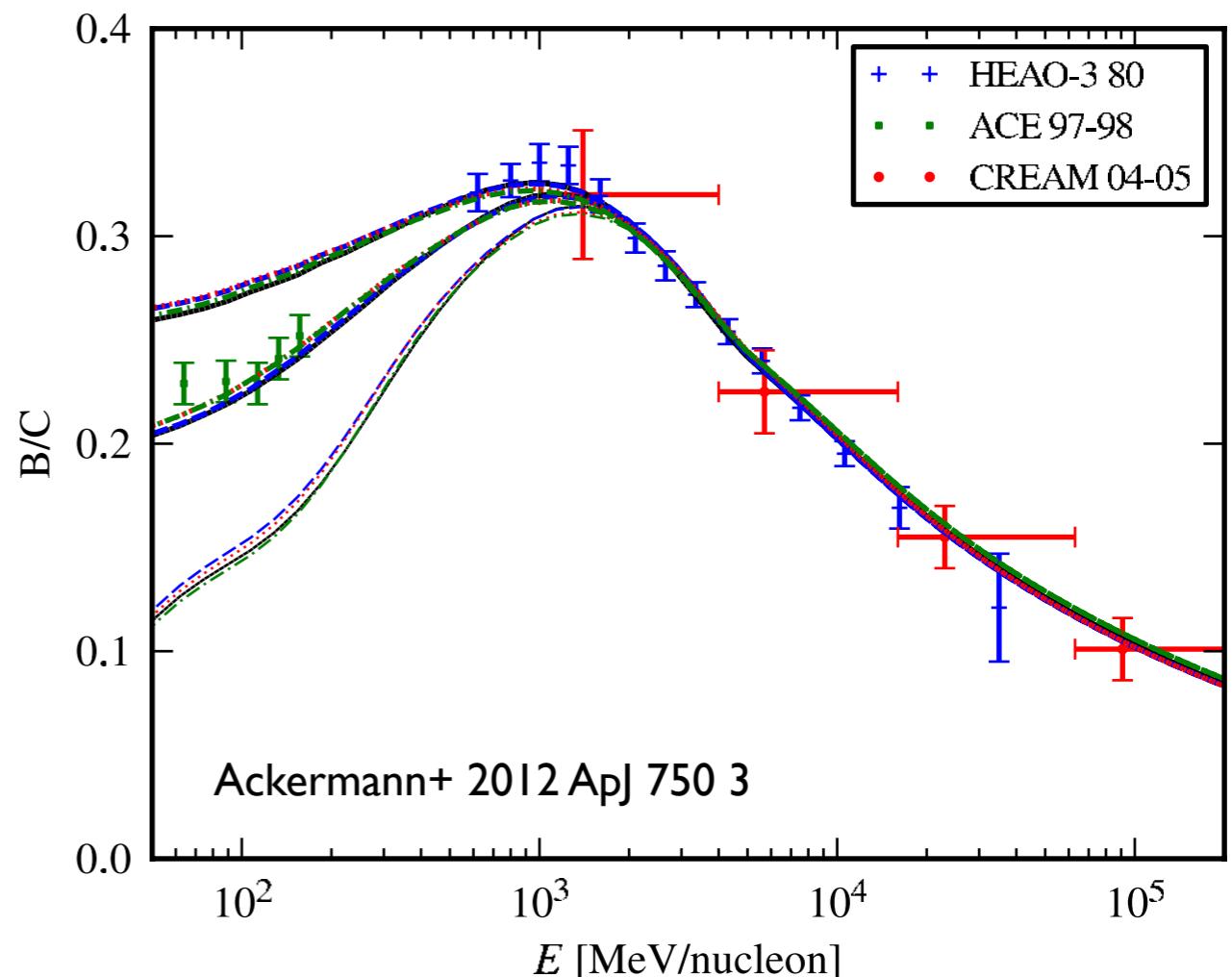


# Outline

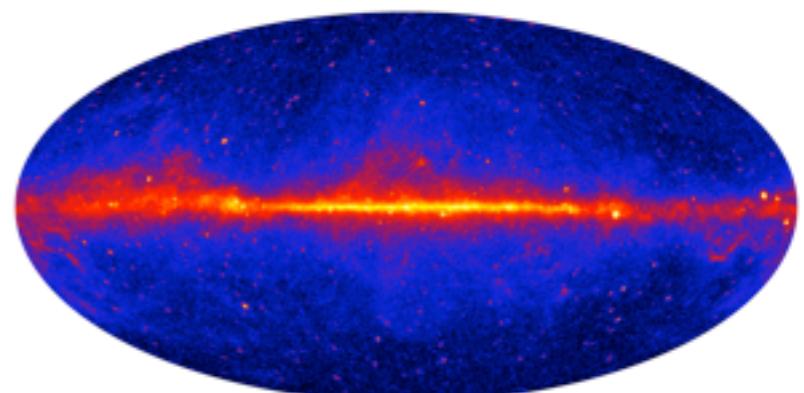
- 
- $\gamma$  rays as a cosmic-ray tracer
  - Cosmic-ray acceleration in supernova remnants
  - Cosmic rays in massive star-forming regions
  - Large-scale propagation of cosmic rays in the Milky Way
  - Cosmic rays in external galaxies
  - Challenges and summary

# A travel through the Galaxy

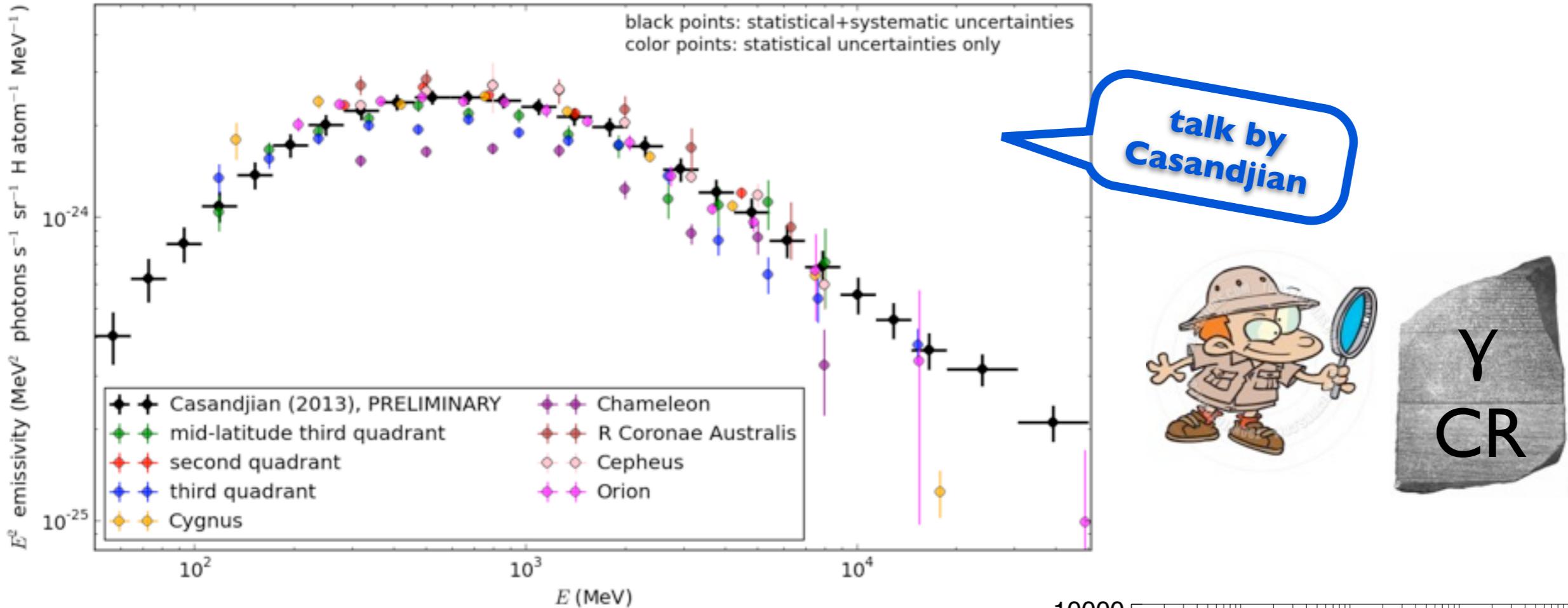
- diffusion on magnetic fields
$$D = D_0 \left( \frac{R}{R_0} \right)^\delta$$
  - $\delta = 1/3, 1/2, 0.7?$
  - breaks in D and/or CR spectra?
  - size of the propagation halo?
- convection? reacceleration?



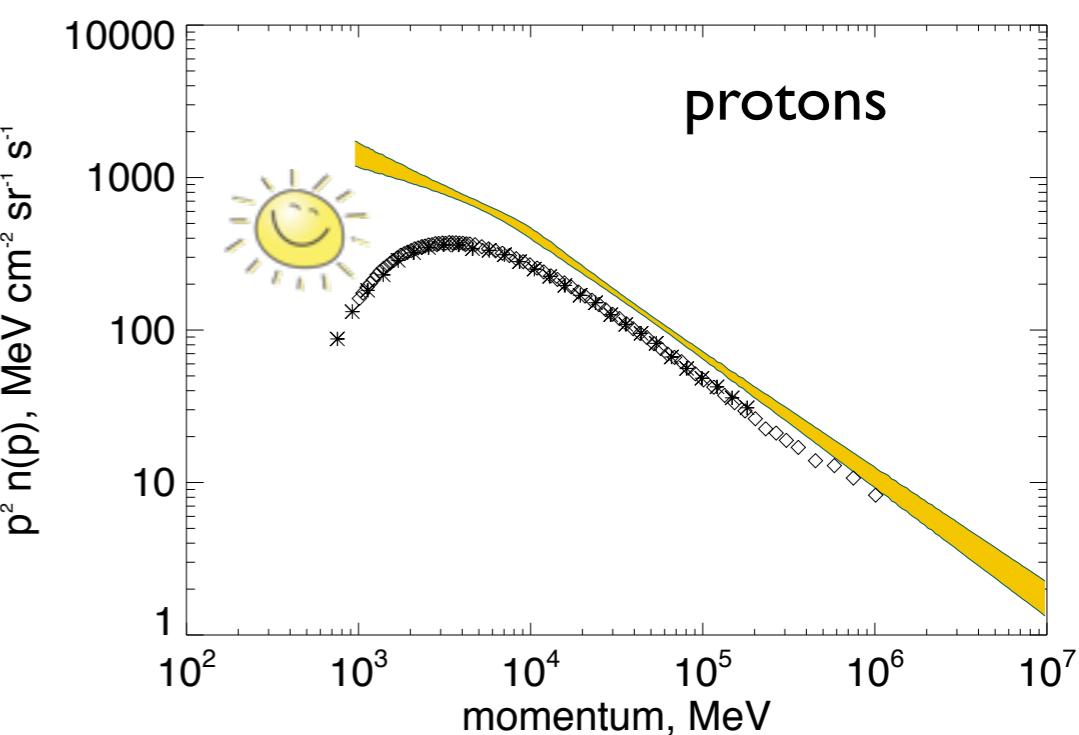
cosmic-ray interactions produce  
~60% of these  $\gamma$  rays!



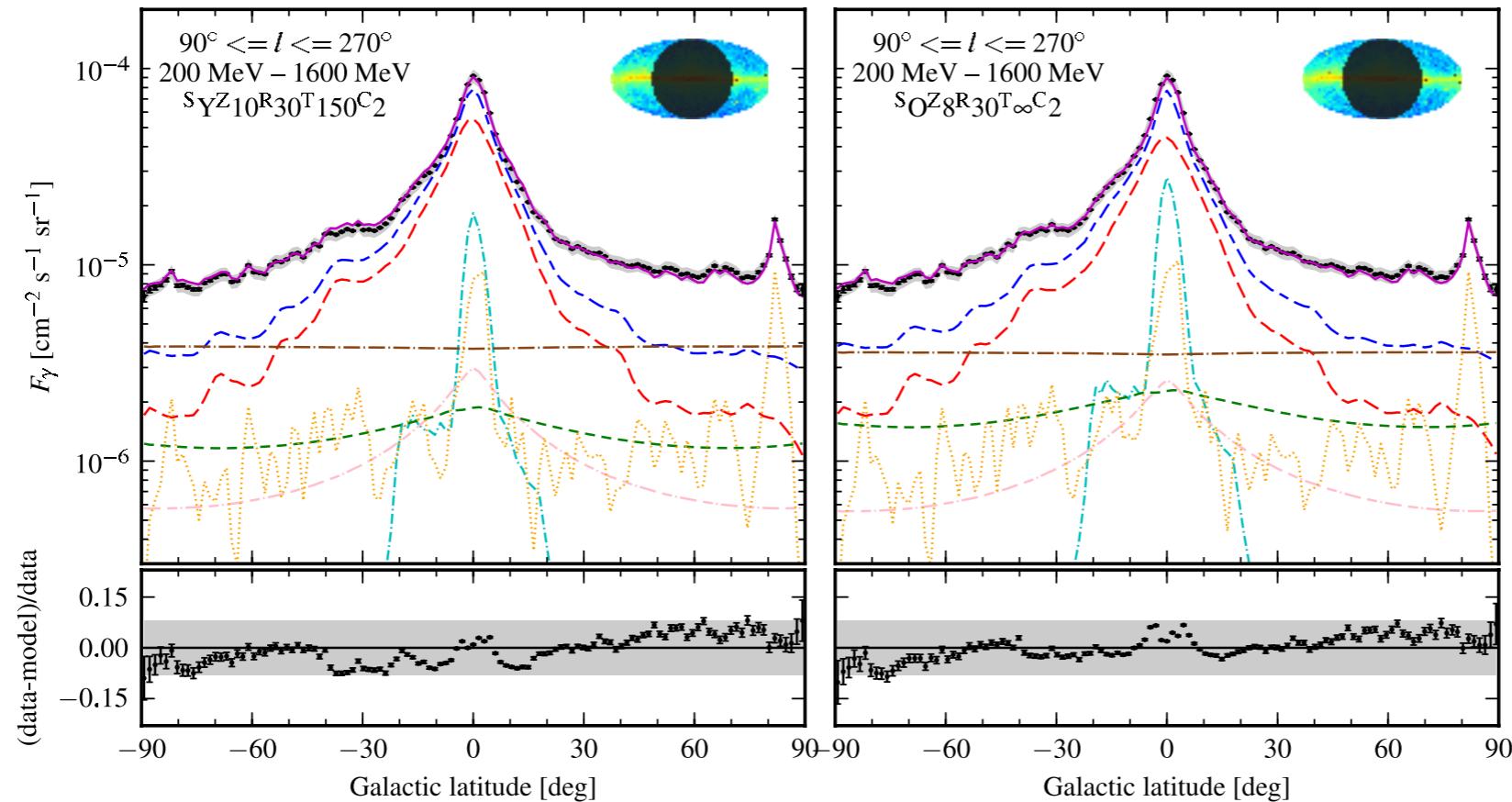
# The Rosetta stone of diffuse $\gamma$ rays



- $\gamma$ -ray emission rate per H atom in the local interstellar medium
  - propagation
  - solar modulation



# The modeling of large-scale propagation



Ackermann+ 2012 ApJ 750 3

talk by  
Moskalenko

- large-scale structures reproduced at ~15%
- degeneracies between sources and propagation
- unmodeled features (e.g. bubbles)

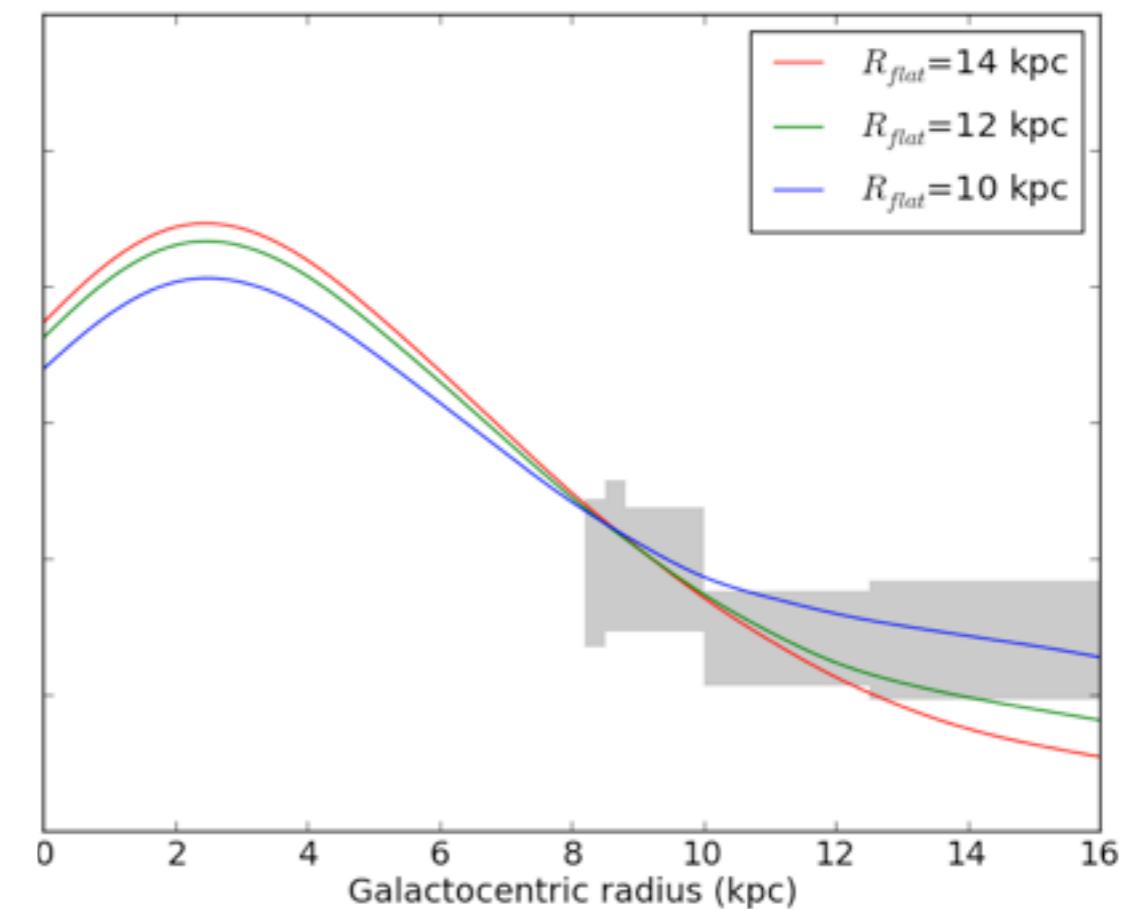
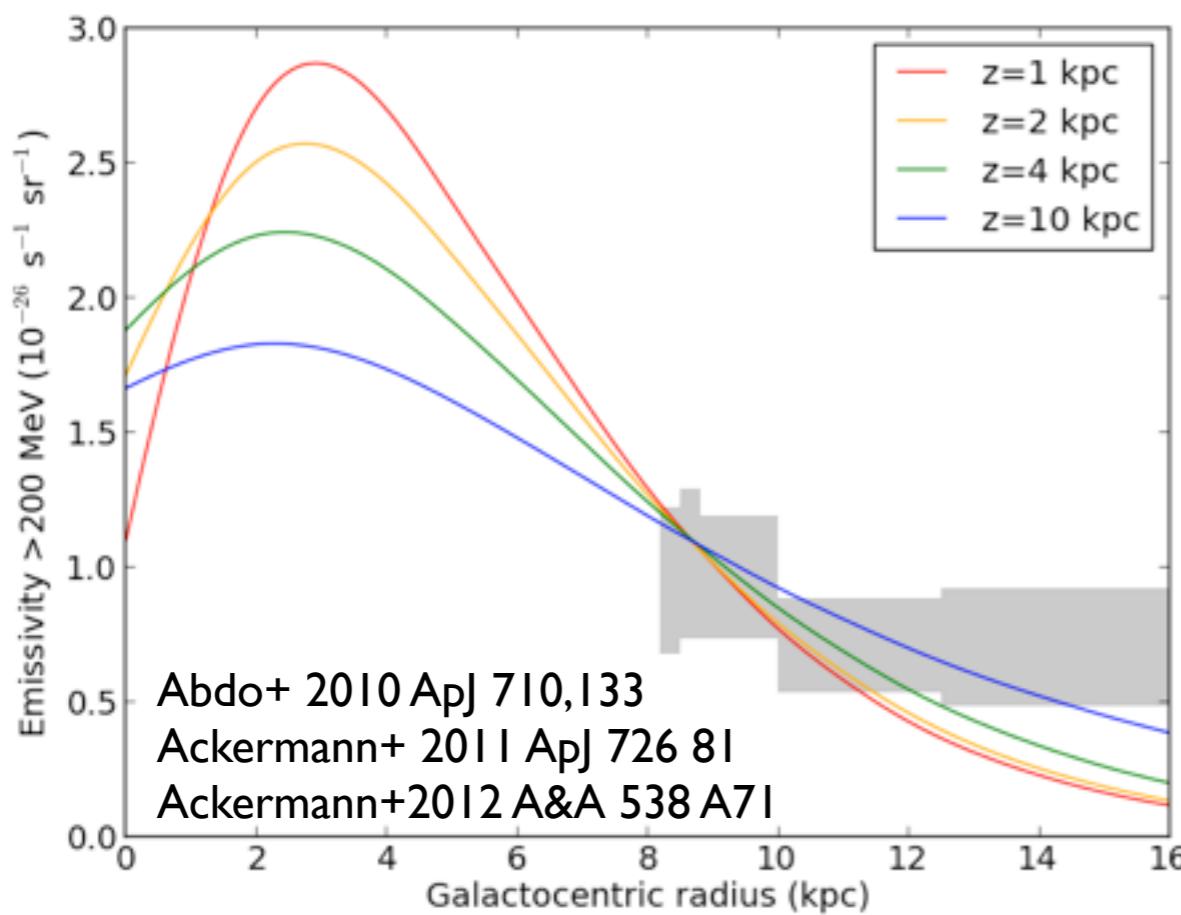
talk by  
Franckowiak

# The gradient problem



CR densities larger than expected in outer Galaxy

- large propagation halo
- more sources
- missing gas
- varying diffusion coefficient  
(e.g. Evoli+ 2012)



# Outline

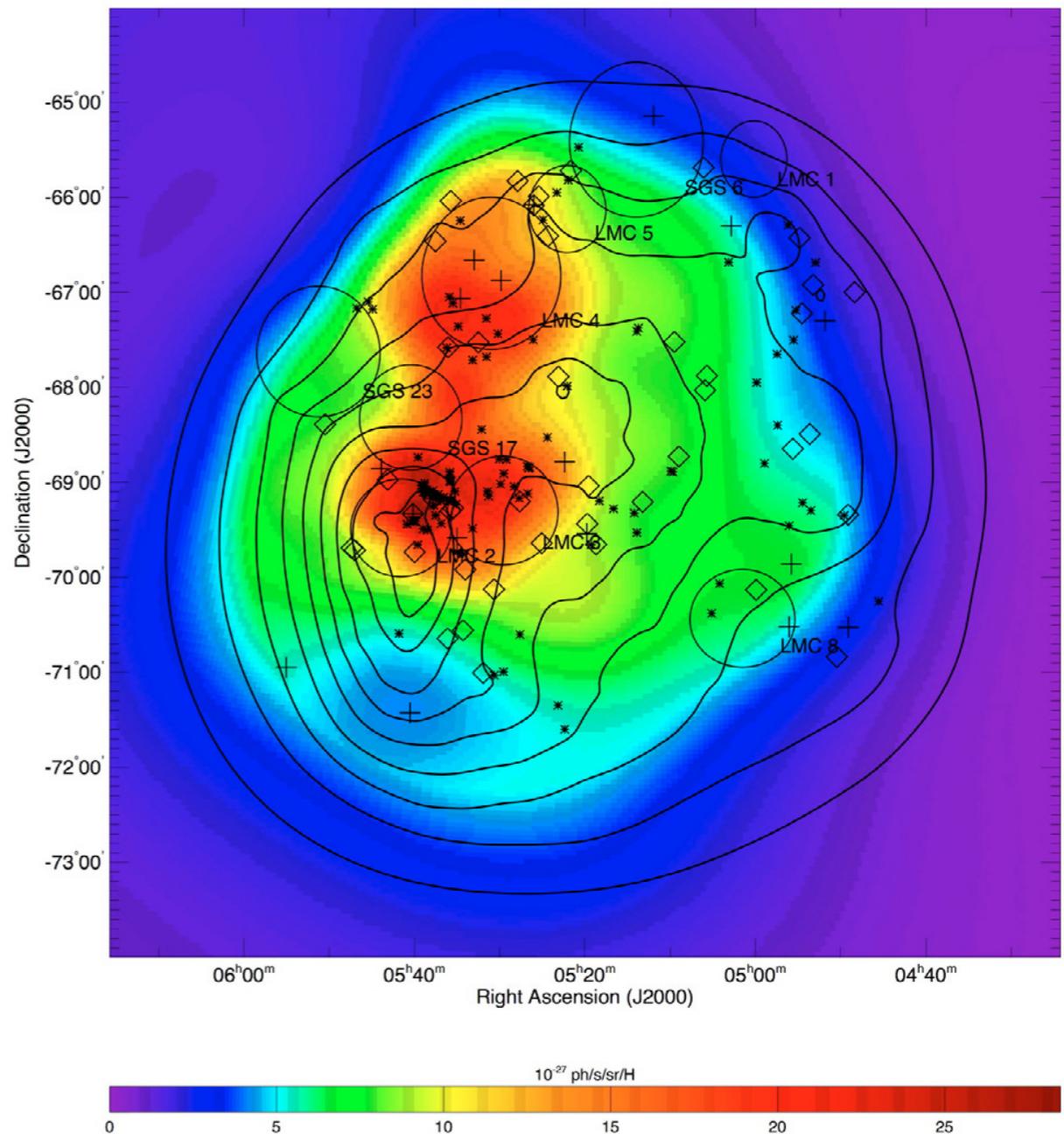
- $\gamma$  rays as a cosmic-ray tracer
- Cosmic-ray acceleration in supernova remnants
- Cosmic rays in massive star-forming regions
- Large-scale propagation of cosmic rays in the Milky Way
- **Cosmic rays in external galaxies**
- Challenges and summary



# Probing cosmic rays in external galaxies

- EGRET: CRs  $< 10^{15}$  eV are Galactic in origin
- *Fermi* images CR propagation in nearby galaxies

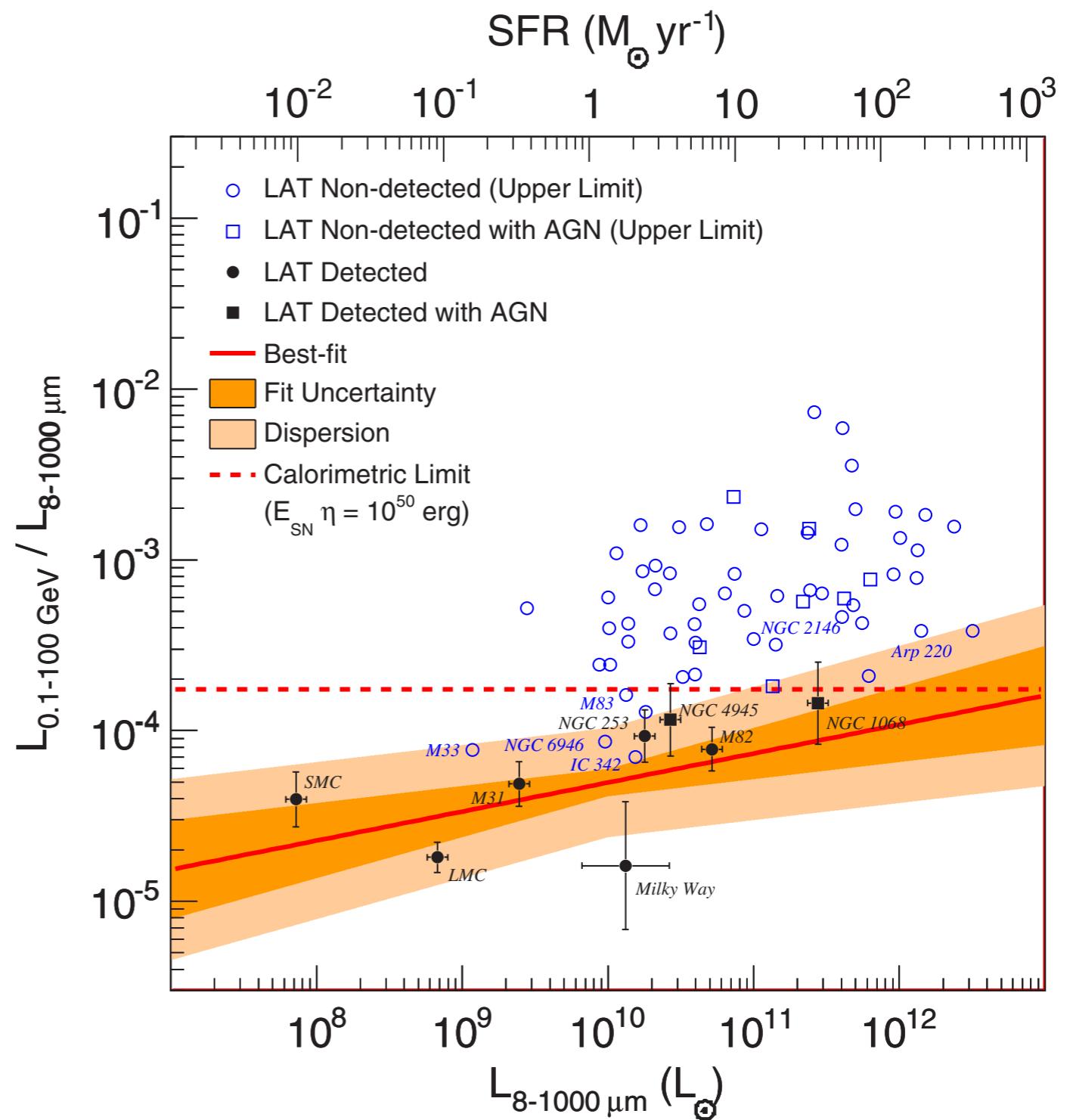
Large Magellanic Cloud:  
 $\gamma$ -ray emissivity map  
(Abdo+ 2010 A&A 512 A7  
Murphy+ 2012 ApJ 750 126)



# The star formation rate- $\gamma$ correlation

Ackermann+ 2012 ApJ 755 164

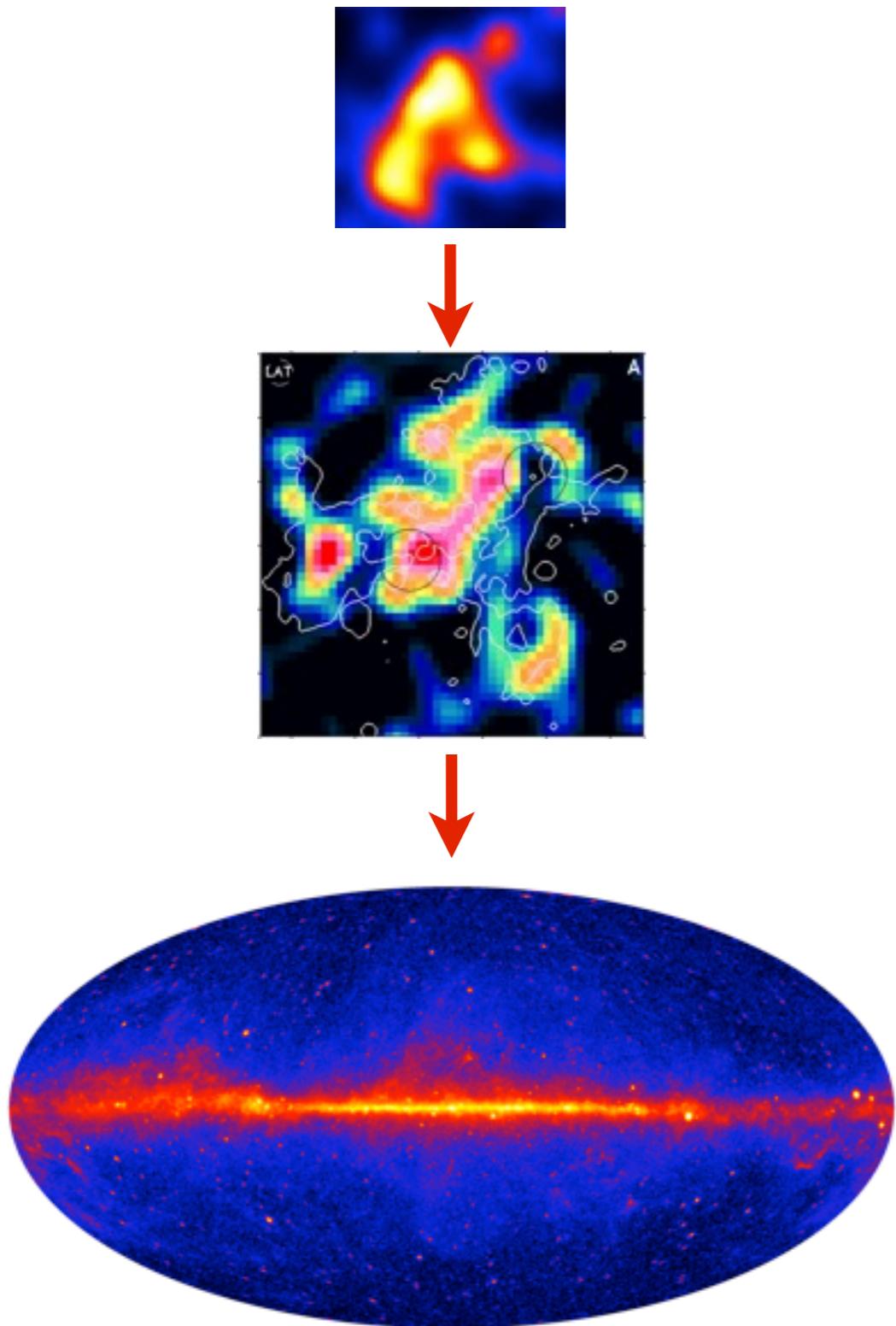
- quasi-linear scaling  $\gamma$  luminosity with radio/IR
- large fraction of energy in CRs escapes
- starburst galaxies: E-independent CR cooling?



# Outline

- $\gamma$  rays as a cosmic-ray tracer
- Cosmic-ray acceleration in supernova remnants
- Cosmic rays in massive star-forming regions
- Large-scale propagation of cosmic rays in the Milky Way
- Cosmic rays in external galaxies
- **Challenges and summary**

# Summary



- supernova remnants
- massive-star forming regions
- galaxies