

ASP Summary

- ASP is intended for **time-critical science** analysis tasks and so explicitly excludes tasks related to instrument performance monitoring and science analysis that is not time critical, such as blind pulsar searches.
- Inputs consist of Level 1 and possibly GBM data.
- Analysis is performed with ScienceTools, FTOOLs, and similar packages.
- Outputs comprise deliverables to the user community (e.g., GCN notices, source fluxes, source locations, etc.) and data products for further analysis by the collaboration.

ASP Tasks

- 1. Gamma-Ray Burst (GRB) Analyses. GRB time scales < O(10) s; "expected" number of LAT photons $O(10)-O(10^2)$.
 - (a) Refine parameters (position, time, duration) for GRBs detected on-board by the LAT.
 - (b) Refine parameters for GBM (or GCN) detected bursts (but not detected on-board).
 - (c) Perform blind searches for GRBs in L1 data (not detected on-board, by the GBM or GCN).
 - (d) Search for afterglows of detected bursts in LAT data on time scales of 10s of minutes to hours or longer (diffuse and instrument backgrounds must be considered).
- 2. Flaring Source Monitoring. Time scales of several hours to days or longer.
 - (a) Monitoring of a predefined list of sources. LAT Data Release Plan (DRP) presently calls for 11 (3EG) + 8 AGNs, 1 NS-Be binary system.
 - (b) Search for new transients and report if flux exceeds $2 \times 10^{-6} \text{cm}^{-2} \text{s}^{-1}$ s (DRP), presumably above 100 MeV.

Implementation

- Python scripts to drive ScienceTools, FTOOLS, etc. Two packages live outside of ST:
 - pyASP The Python scripts themselves (plus some SWIG-exposed classes from astro, etc.)
 http://glast.stanford.edu/cgi-bin/viewcvs/users/jchiang/pyASP/
 - **BayesianBlocks** Python module (implemented in C++) for performing temporal analysis. http://glast.stanford.edu/cgi-bin/viewcvs/users/jchiang/BayesianBlocks/
- Pipeline II to drive the scripts given various triggers/inputs.
- Examples:
 - Refining GRB parameters
 - GRB Blind Search



