

The Large Synoptic Survey Telescope (LSST) Big Data and Processing

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August 21, 2012



LSST: What

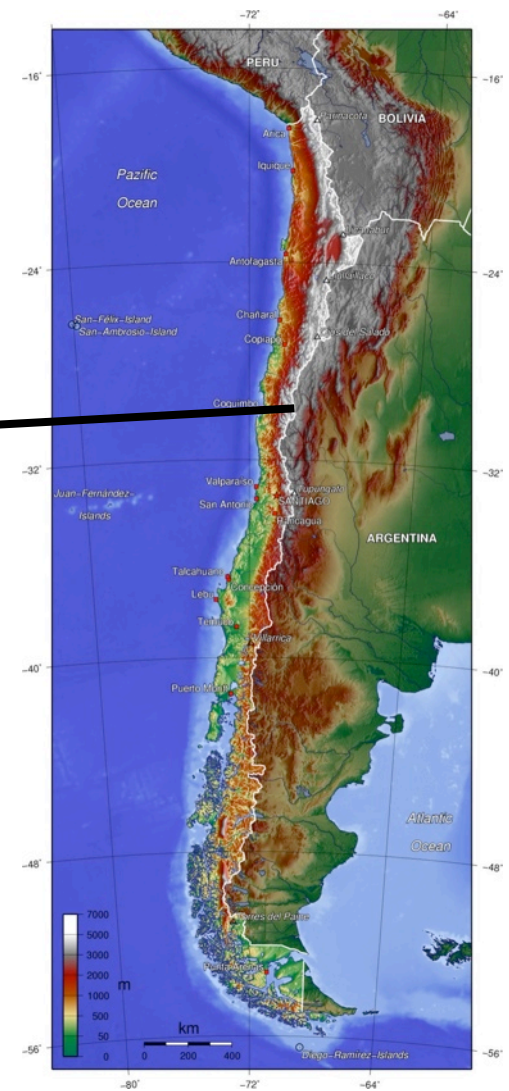
LSST: Why

LSST: Data

Data-Intensive Science



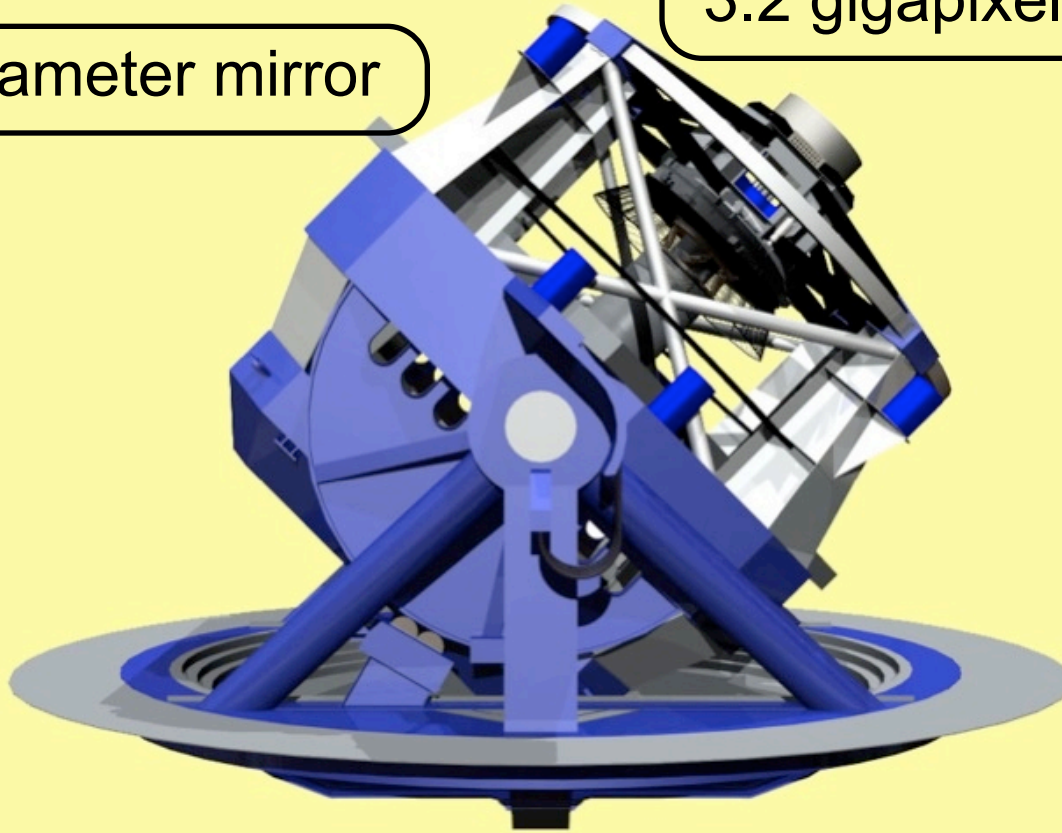
Proposed telescope to be built in Chile



More than proposed: 26 tons of glass made into mirror, soon to be polished

8.4 meter diameter mirror

3.2 gigapixel camera



Wide

Deep

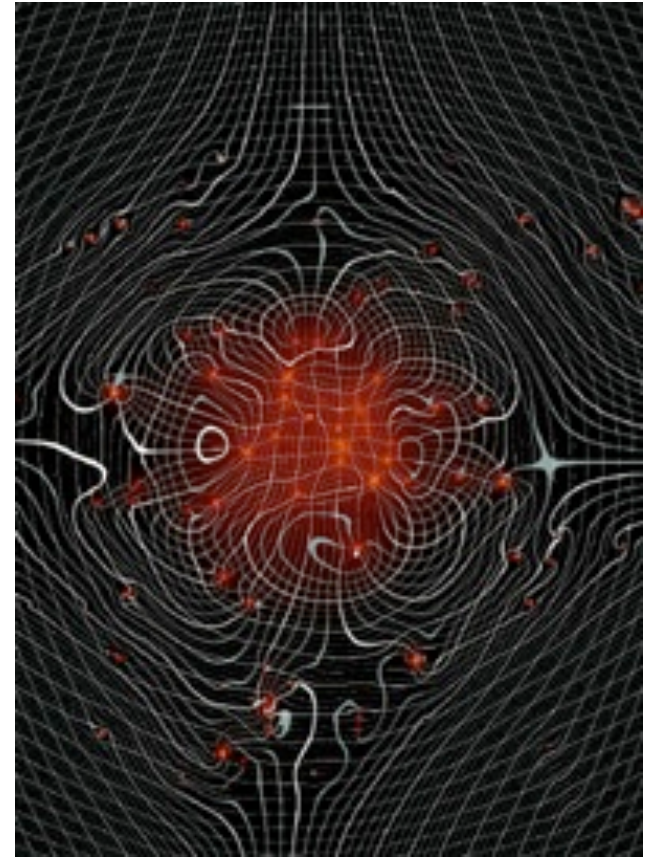
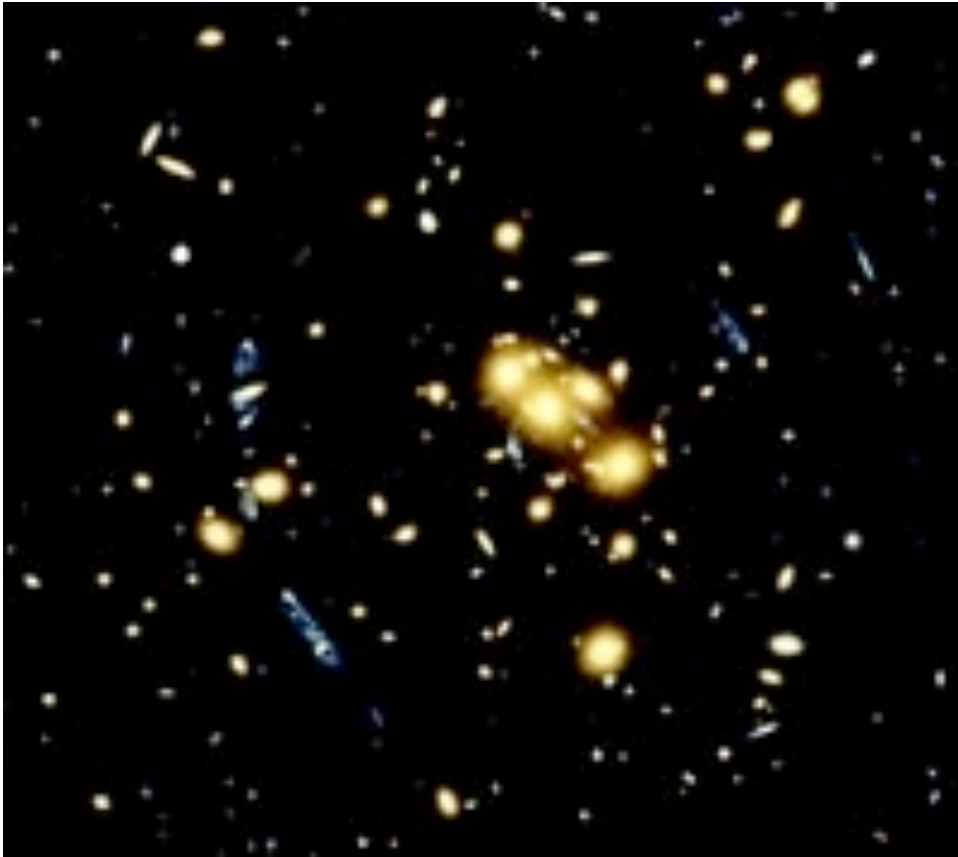
Fast

Wide: surveys entire sky, not just small area

Deep: gathers faint and distant objects

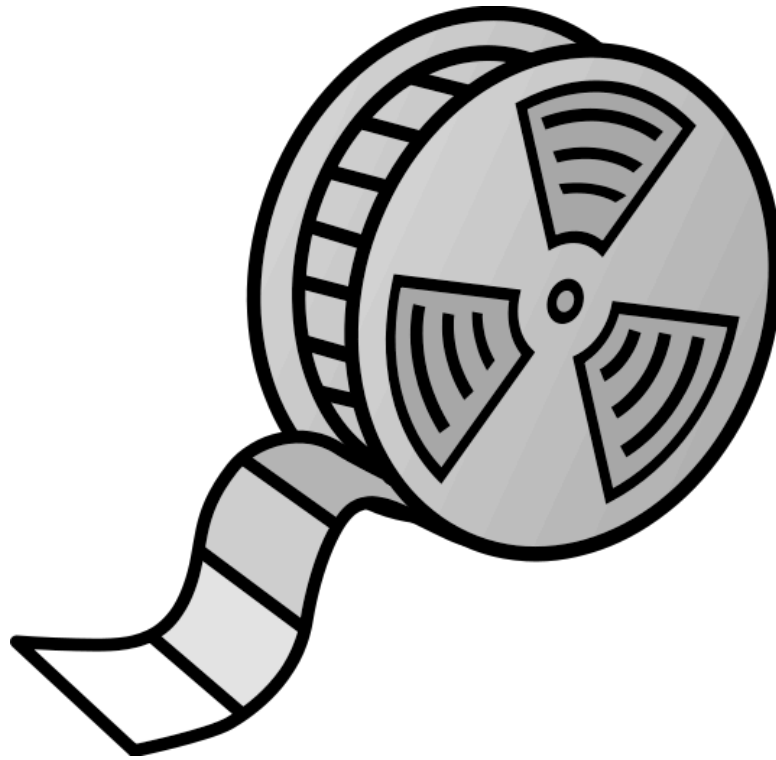
Fast: image every 15 seconds; entire sky twice a week

Brings in time domain, unlike almost all other surveys

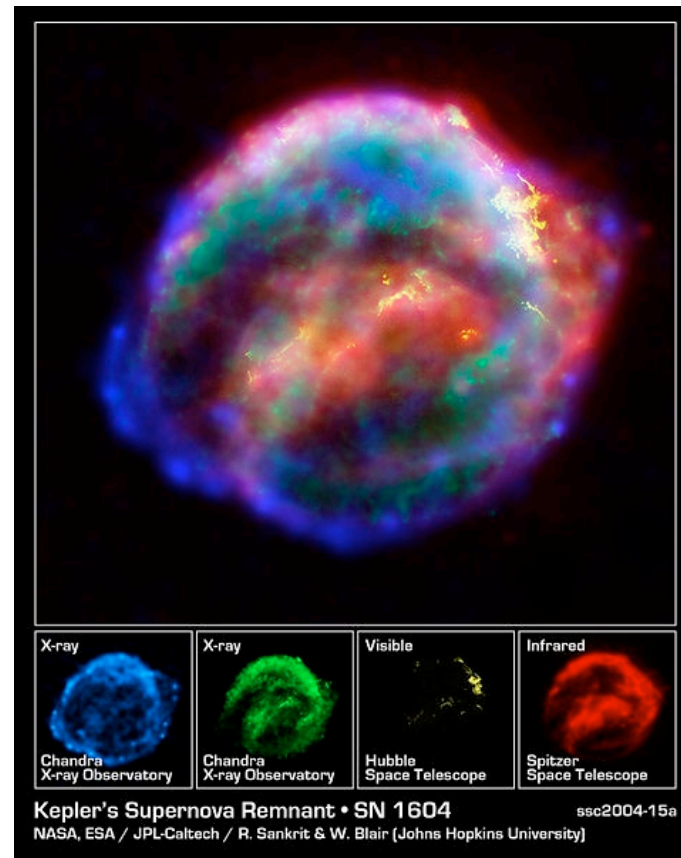


Credit: J. A. Tyson, W. Colley, E. L. Turner, and NASA

Looking at distortions of millions or billions of galaxies
Gives information to map dark matter and dark energy



Get a thousand-frame movie of every visible object showing variations over time



Supernovae appear unexpectedly
We will detect them as new objects that appear (or disappear)



Credit: D. Roddy, Lunar and Planetary Institute

Very interested in objects much closer — in the solar system — that move and might collide with us

Images

Database

Two main types of data

Images - pixel data - raw and processed

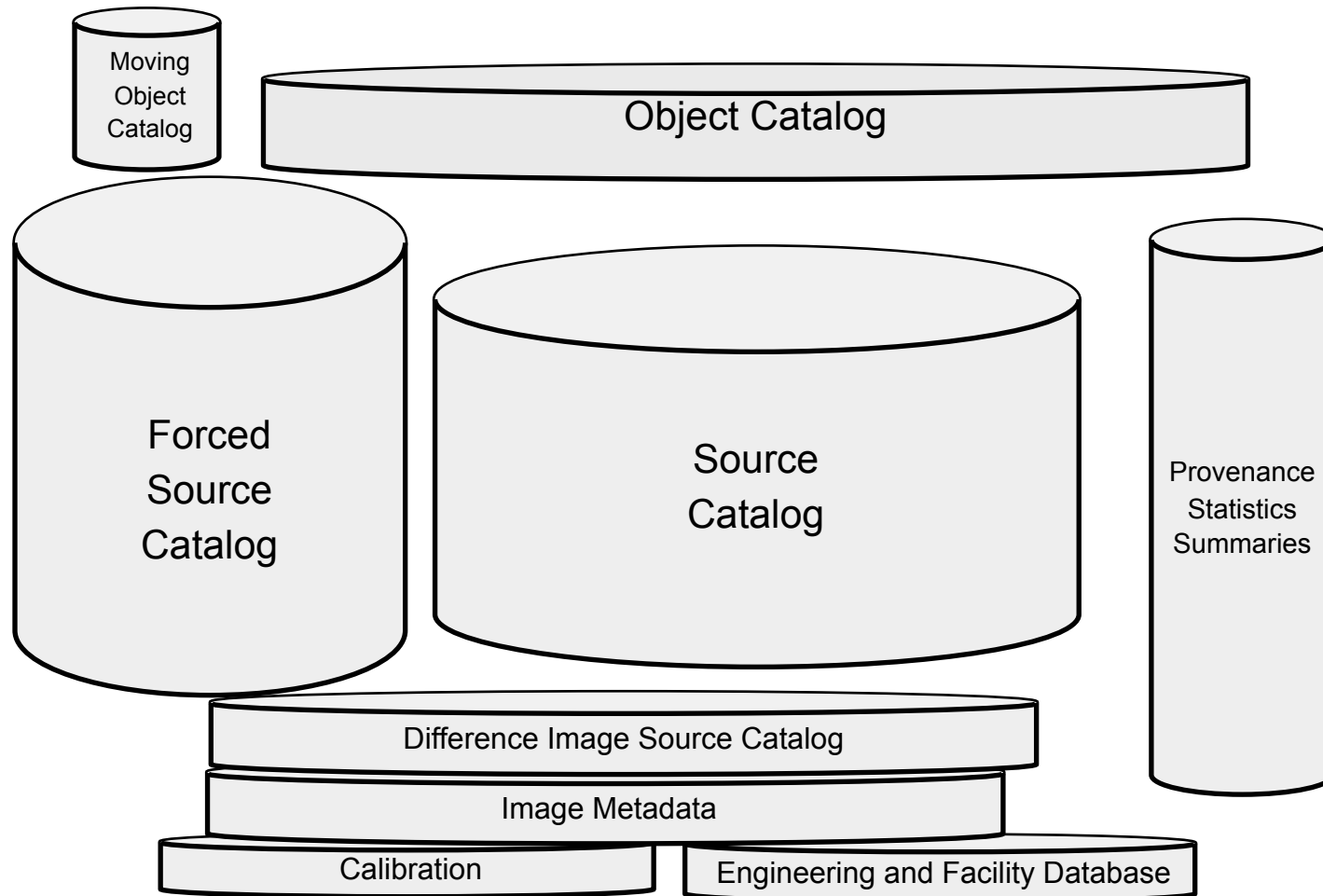
Database - everything else, including:

metadata and provenance

astronomical objects

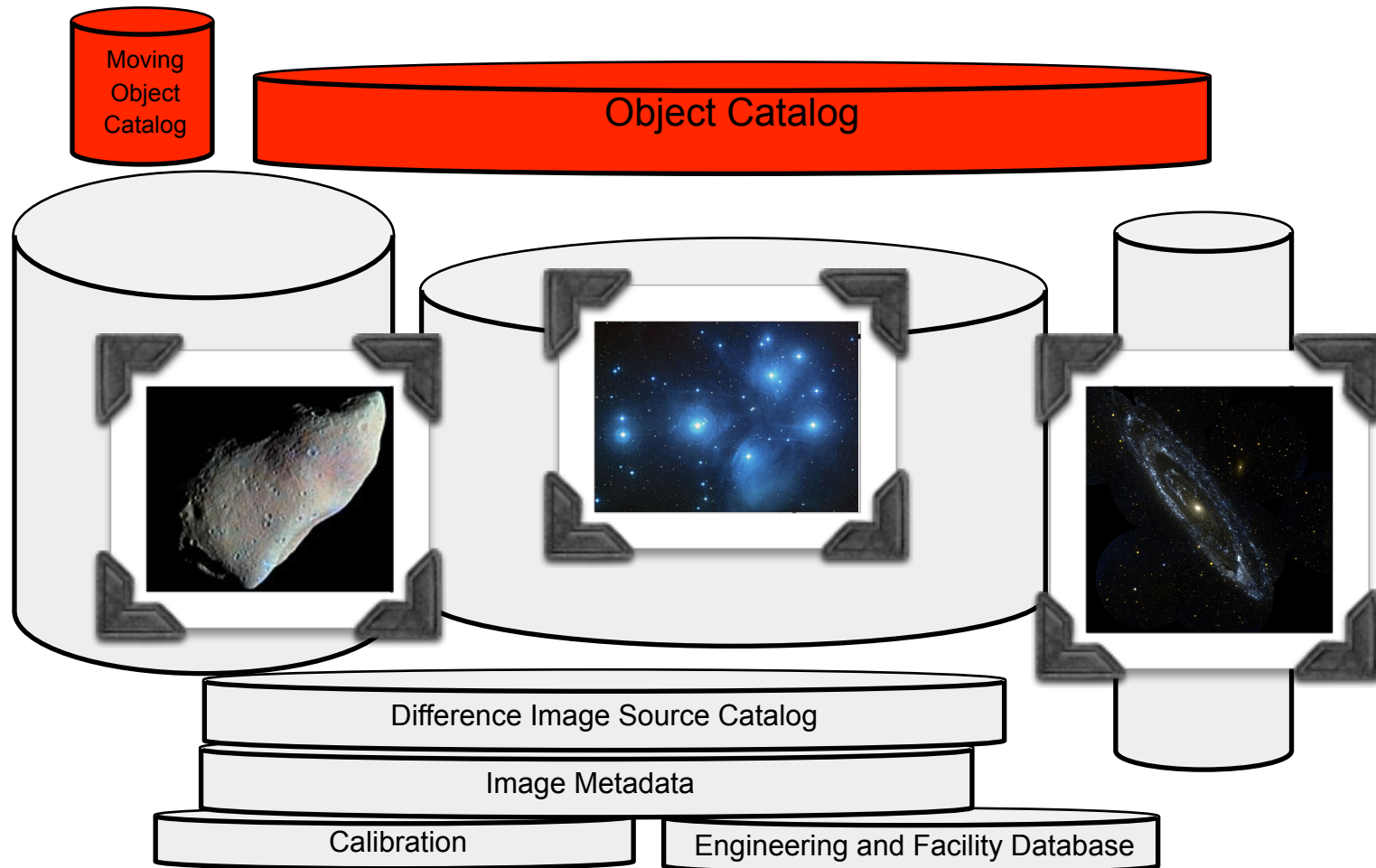
sources (or detections of objects)

Database: Components



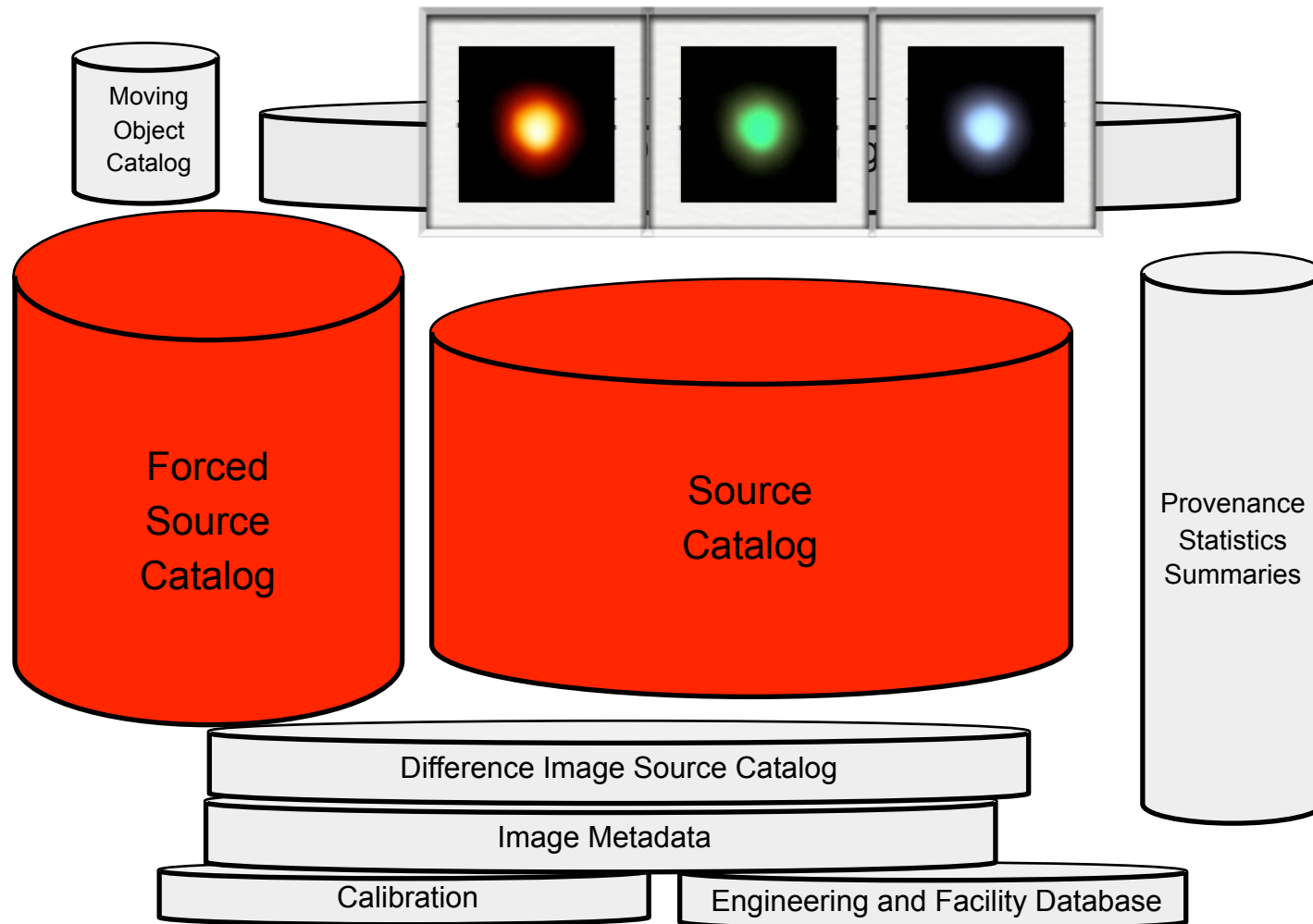
Database is composed of many catalogs, each of which has several tables

Astronomical Objects

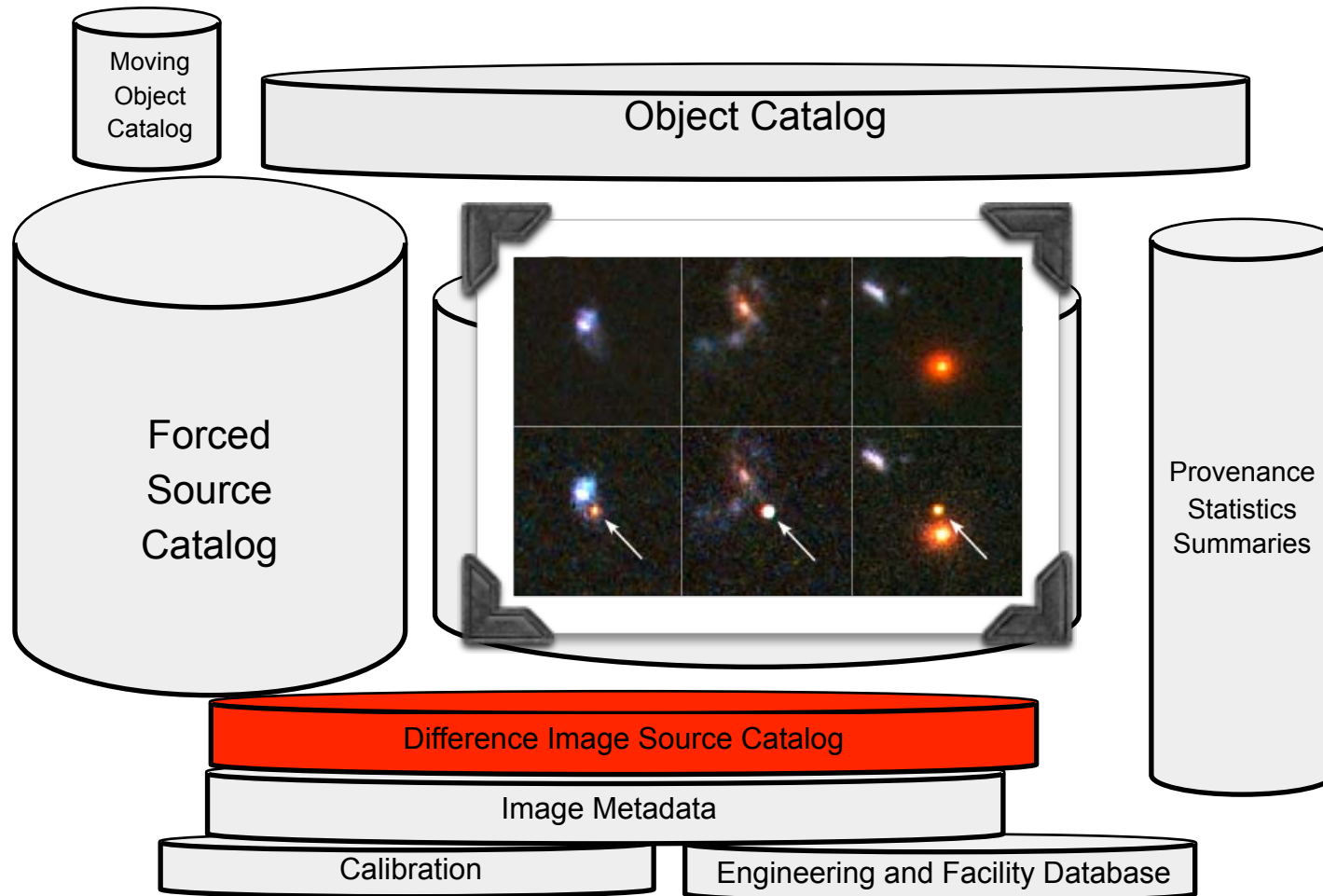


Asteroids, stars, and galaxies

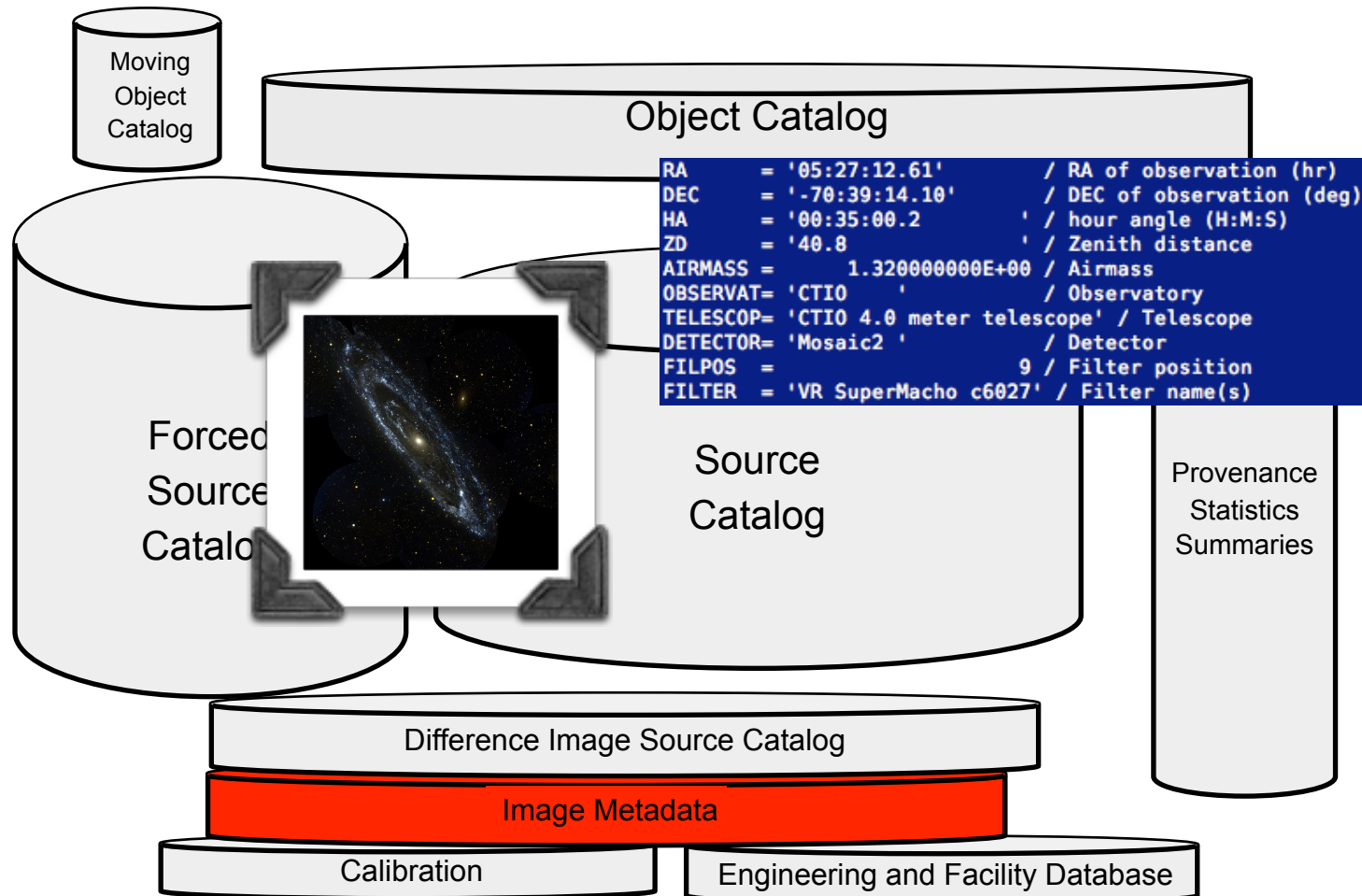
Each has many attributes, some fixed, many averaged over time



Sources are observations or measurements of objects
At different times, through different filters
Forced Sources are low-signal-to-noise
Looks rather like commercial data warehouse



Difference images are how we detect changes
Subtract an average known template from the current image
Anything that has changed stands out
Sources from these difference images are kept separately



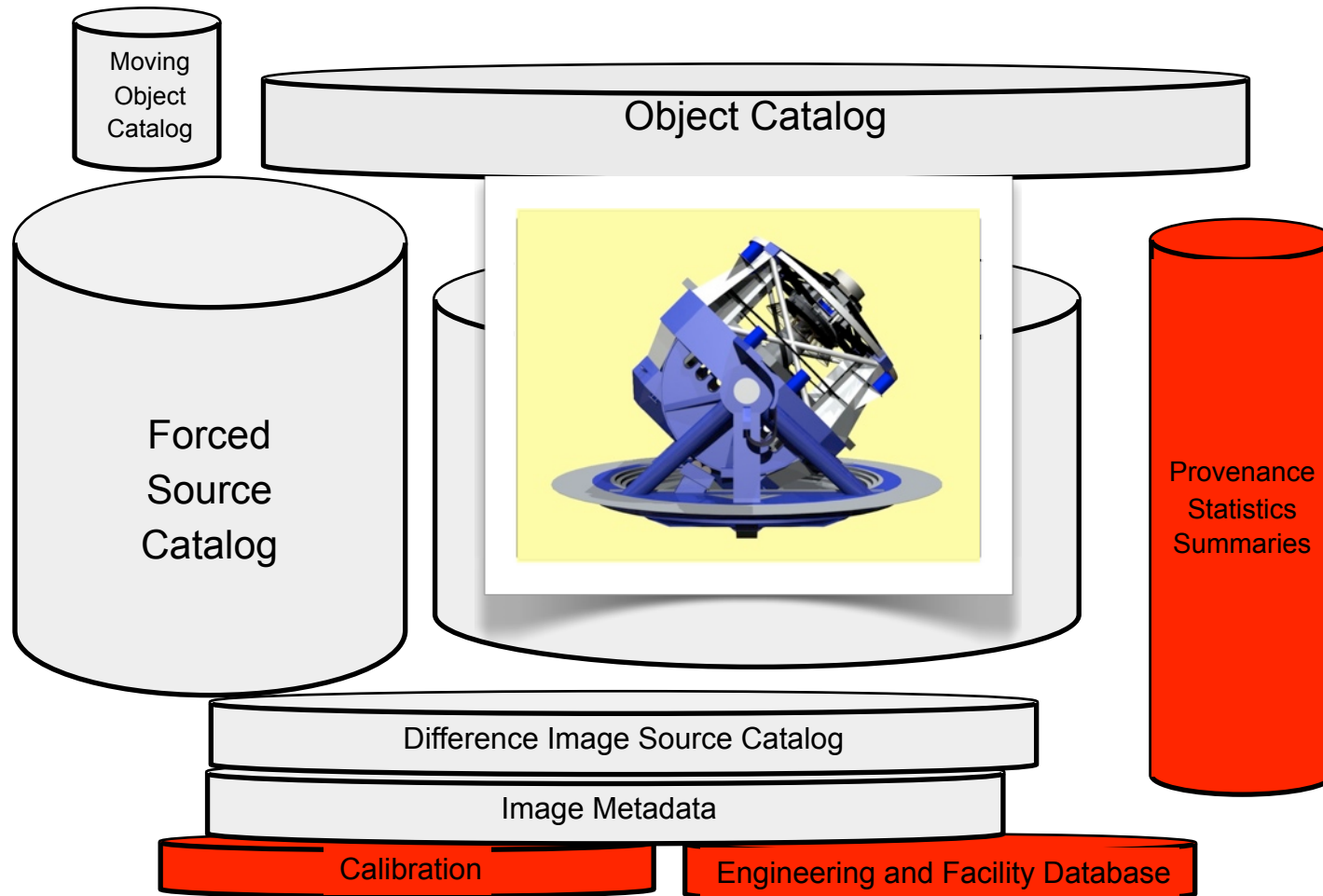
Each image has lots of metadata

Observation conditions: pointing, time, filter, atmospheric conditions

Image analysis for artifacts: cosmic rays, satellites and airplanes

Quality control flags

Location on disk or tape



Calibration information and information about the telescope itself must be retained

Also need provenance of all elements in database, various aggregate statistics

38 *billion* objects

5 *trillion* sources

32 *trillion* forced sources

8 *quadrillion* pixels

Mountain

Base Site

Archive Site

Data Access Center

EPO Center

Where is the data?

Five locations

Mountain is where images are generated

Base Site is where images are sent; real-time processing for alert generation may occur here

Archive Site is permanent repository, science processing

Data Access Center is where end users (US/Chile/Intls) query the DB and access images

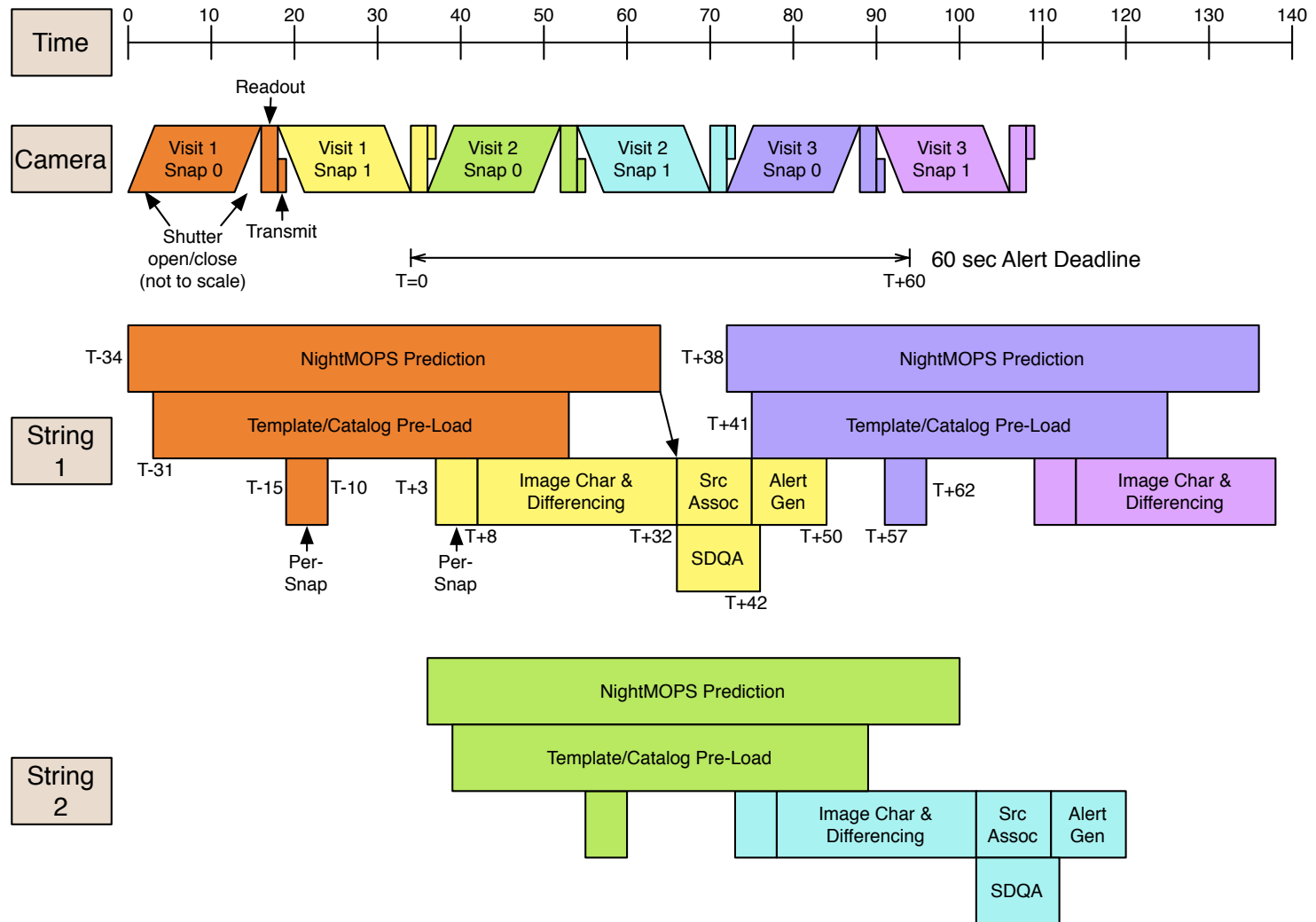
Education and Public Outreach is specialized for the general public (Google Sky/WWT)

Data Size

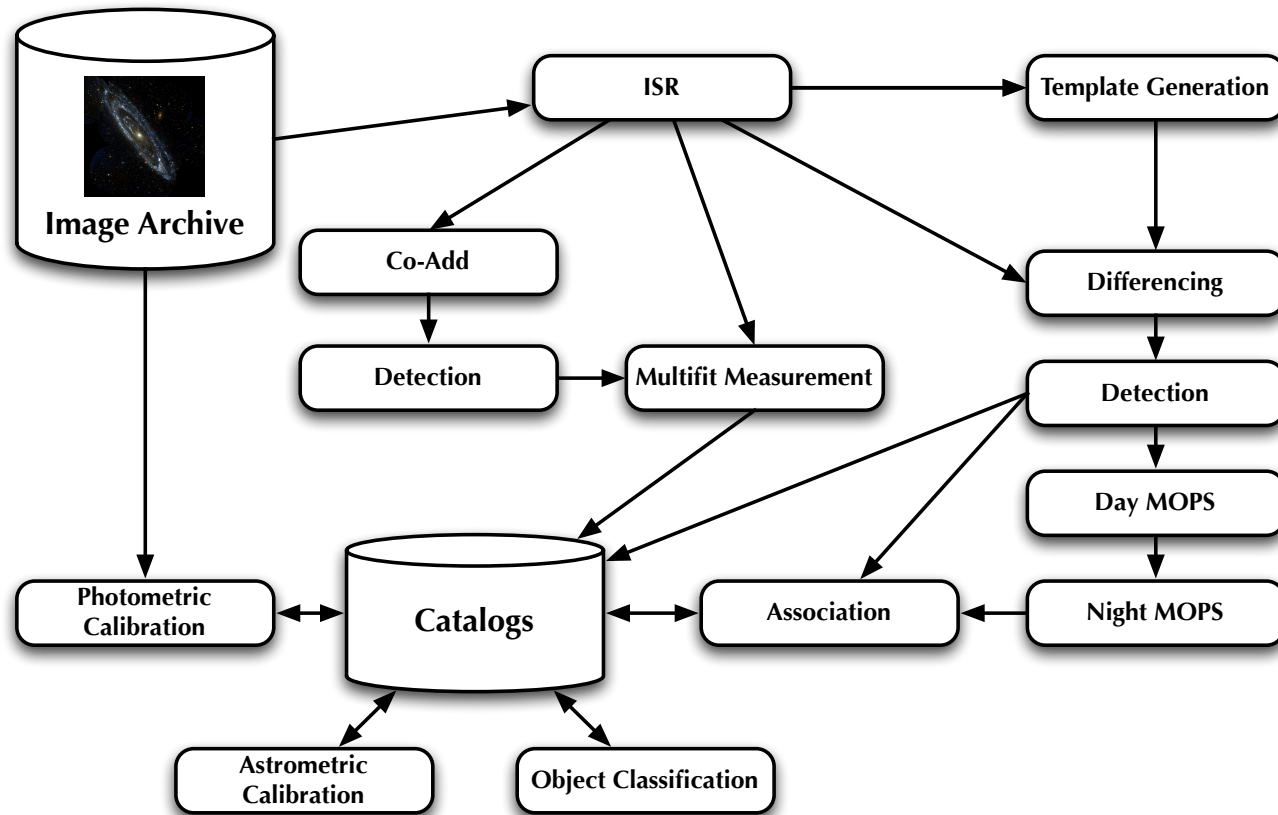
		Archive Site	Base Site
Compute	Teraflops (sustained)	175 → 1000 (required) 270 → 1600 (on-the-floor)	30 → 45 (required) 50 → 80 (on-the-floor)
	Nodes	1700 → 1400 (on-the-floor)	300 → 60 (on-the-floor)
	Cores	70K → 280K (on-the-floor)	12K → 13K (on-the-floor)
	Memory Bandwidth	25 → 130 TB/s (required)	4 → 6 TB/s (required)
	File I/O Bandwidth	120 → 425 GB/s (required)	25 → 30 GB/s (required)
Storage	Capacity	13 → 25 PB (required) 19 → 100 PB (on-the-floor)	8 → 8 PB (required) 12 → 23 PB (on-the-floor)
	Drives	1500 → 1100 (on-the-floor)	950 → 275 (on-the-floor)
	Disk Bandwidth	120 → 425 GB/s (required)	27 → 31 GB/s (required)
Mass Storage	Capacity	7 → 75 PB (required) 8 → 83 PB (on-the-floor)	7 → 75 PB (required) 8 → 83 PB (on-the-floor)
	Tapes	1000 → 3800 (near line) 1000 → 3800 (offsite)	1000 → 3800 (near line)
	Tape Bandwidth	6 → 24 GB/s (required)	6 → 24 GB/s (required)
Database	Teraflops (sustained)	16 → 193 PB (required) 16 → 199 PB (on-the-floor)	12 → 126 PB (required) 12 → 133 PB (on-the-floor)
	Storage Capacity	4 → 50 PB (required) 10 → 128 PB (on-the-floor)	3 → 40 PB (required) 7 → 95 PB (on-the-floor)
	Disk Drives	1400 → 2600 (on-the-floor)	1000 → 2000 (on-the-floor)
	Disk Bandwidth (sequential)	125 → 625 GB/s (required)	95 → 425 GB/s (required)
	Database Nodes	100 → 190 (on-the-floor)	80 → 130 (on-the-floor)
Facilities	Floorspace	950 → 850 sq ft (1100 high water mark)	460 → 350 sq ft (470 high water mark)
	Power	720 → 640 kW (880 high water mark)	180 → 110 kW (210 high water mark)
	Cooling	2.4 → 2.1 mmbtu (3.0 high water mark)	0.6 → 0.4 mmbtu (0.7 high water mark)

First number is survey year 1; second number is survey year 10
 Requirements go up because data size goes up and processing rate goes up
 In last year, need to process 10 years of data in just one year

Alert Production



Alerts issued within 60 seconds of taking picture
 Near-realtime processing involving complex image processing



Scalable

Fast

Fault-tolerant

Cost-effective

Open Source

Reuse existing components where possible:

xrootd for distributed processing and communication

Fermi/GLAST workflow software contending for workflow management

Computing, but not super-computing



Credit: NCSA/University of Illinois

Mining large databases



Credit: National Institute for Occupational Safety and Health

Types of Queries

- All about an object
- All objects meeting criteria
- All objects near objects meeting criteria
- All objects with interesting time series
- All pairs of objects with similar time series

Bandwidth

Bandwidth

Bandwidth

For input
For (re-)processing
For queries

LSST is an example of a Big Data experiment
requiring data-intensive scientific computing