The Heavy Photon Search Experiment at Jefferson Lab

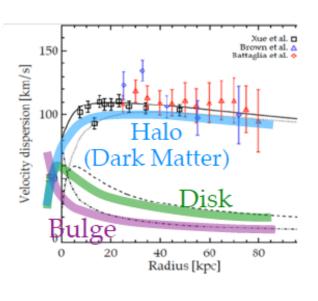
Takashi Maruyama, SLAC For the HPS Collaboration SLAC Summer Institute August 10-21, 2015





Ample evidence for Dark Matter





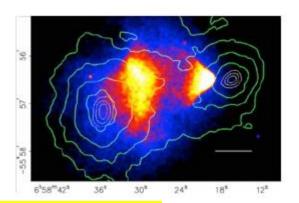
Rotation curves of spiral galaxies



Gravitational lensing DM Map: PRL 115,051301 (2015)



Cluster collisions



WIMP Dark Matter



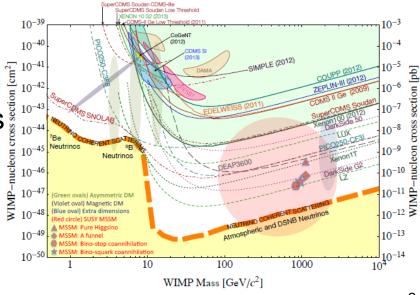
- Dark matter is Physics beyond the Standard Model.
 - Only gravitational interaction has been observed.
- WIMP miracle
 - SUSY WIMP with ~100 GeV mass was the most attractive candidate.
 - Dark matter relic density is consistent with

the weak interaction scale.

- experimental evidence

 Direct DM detector searches No experimental evidence

 - No SUSY at ATLAS/CMS
- Other possibilities?



Low Mass Dark Matter and Heavy Photon as a mediator

- Low mass dark matter, MeV ~ GeV.
 - Two PRL papers in the last few weeks: PRL 115, 021301 and 061301 (2015).
- Need low mass mediator
 - To be consistent with the DM relic density.

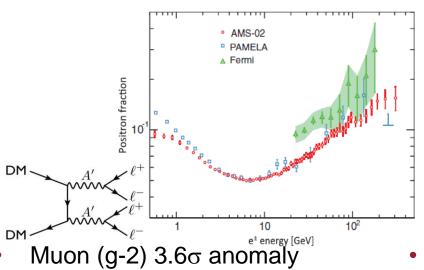
$$\Omega_{DM} pprox (\sigma v)^{-1}, \sigma v \propto \frac{m_\chi^2}{m_{med}^4}$$
 χ mediator f

- WIMPless miracle, PRL 101, 231301 (2008)
- HPS searches for the mediator.

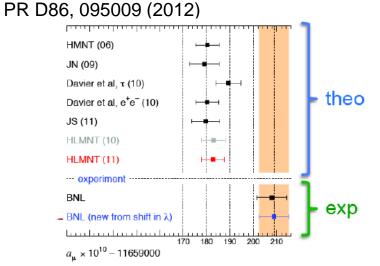
Anomalies got the heavy photon business going





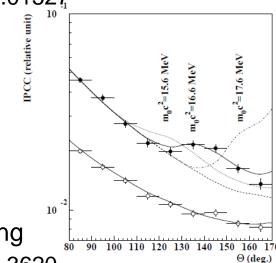


75 750 65766 (5546)

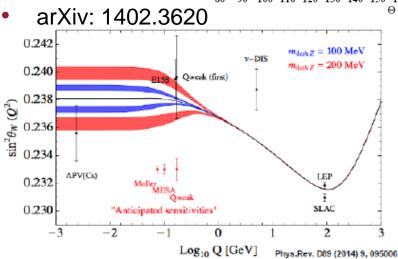


16.6 MeV anomaly





Moller scattering

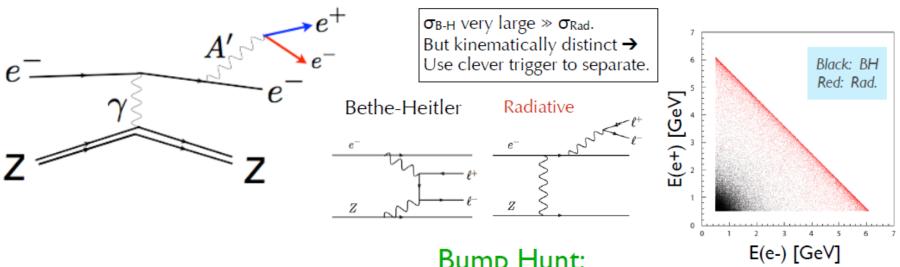


Ongoing heavy photon searches

SLAC

- Colliders
 - ATLAS and CMS at LHC
 - BaBar and Belle
 - KLOE at DAΦNE
 - PHENIX at RICH
 - WASA at COSY
- Fixed target
 - NA48 at SPS
 - APEX at JLab
 - HPS at JLab
 - DarkLight at JLab
 - A1 at MAMI, Mainz
 - MicroBooNE at FNAL
 - SeaQuest at FNAL (proposal)

Look for radiated A' decay to e^+e^- , $(\mu^+\mu^-)$



Very high luminosities: Intensity Frontier Physics.

P. Schuster, R. Essig et al, Intensity Frontier WS 'II summary paper.

Bump Hunt:

Look for signal over background.

Bump Hunt + Vertexing:

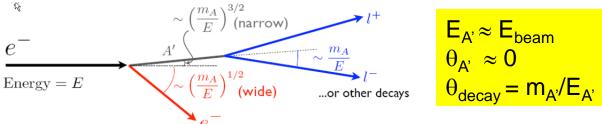
Look for signal over background, reduce background with vertexing.

BEST: Bjorken, Essig, Schuster, Toro, Phys.Rev. D80 (2009) 075018

HPS Design Choices



A' kinematics ⇒ very forward production



 Vertexing A' decays requires detectors close to the target. Invariant mass is an essential signature, so good momentum/mass resolution is also required. Vertexing and bump hunting need tracking and a magnet.

Want $\Delta m/m \sim 1\%$ for bump hunt Want $\Delta z \sim 1$ mm

• Trigger with a high rate, rad hard EM Calorimeter Placed downstream of the magnet, it can ID e⁺ and e⁻.

Beam's Eye View

e⁺ and e⁻

Entering ECal

HPS opts for large forward acceptance/moderate currents. This
requires placing sensors as close as possible to the beam.

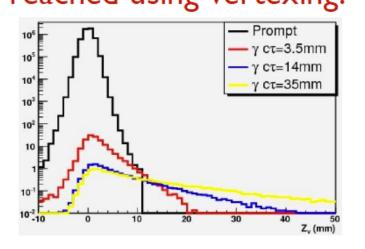
A' lifetime

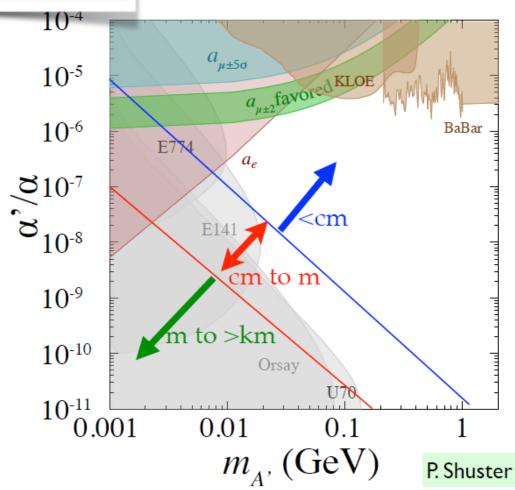
SLAC

$$\gamma c \tau \approx 1 \text{ mm} \left(\frac{\gamma}{10}\right) \left(10^{-8} \frac{\alpha}{\alpha'}\right) \left(\frac{100 \text{ MeV}}{m_{A'}}\right)$$

Lower α′, lower mass
→longer lifetime

Background is all prompt
 → Lower coupling can be reached using vertexing.





Controlling Beam Backgrounds

SLAC

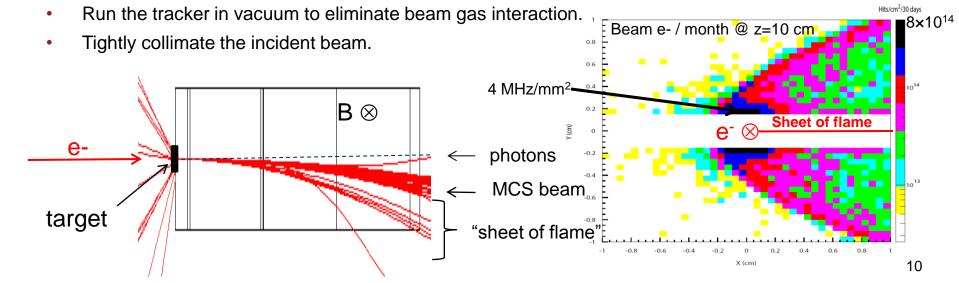
With sensors close to the beam (just ½ mm for the first Si sensor), background control, radiation damage, and beam stability become critical.

Constraints

- Avoid Multiple Coulomb Scattered (MCS) beam
- Avoid the "sheet of flame", the beam electrons which have radiated, lost energy, and been deflected in the horizontal plane by the magnet
- Avoid beam gas interactions.
- Avoid errant beam motions.

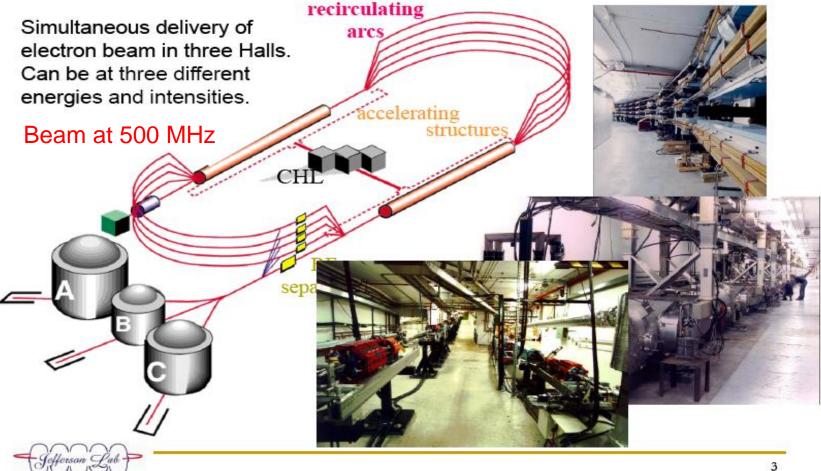
Design solutions

Split the detectors top-bottom to avoid the beam and the "sheet of flame.



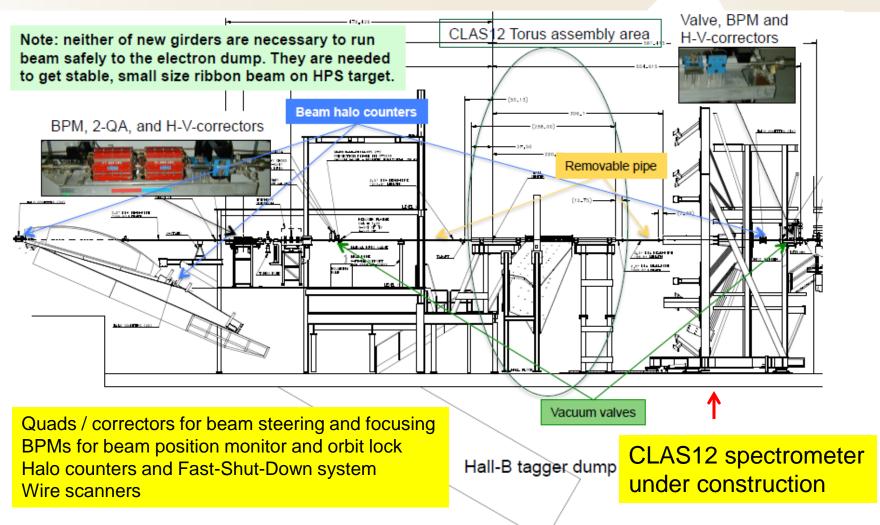
High Duty Cycle = CEBAF at JLab

CEBAF - Continuous Electron Beam Accelerator Facility



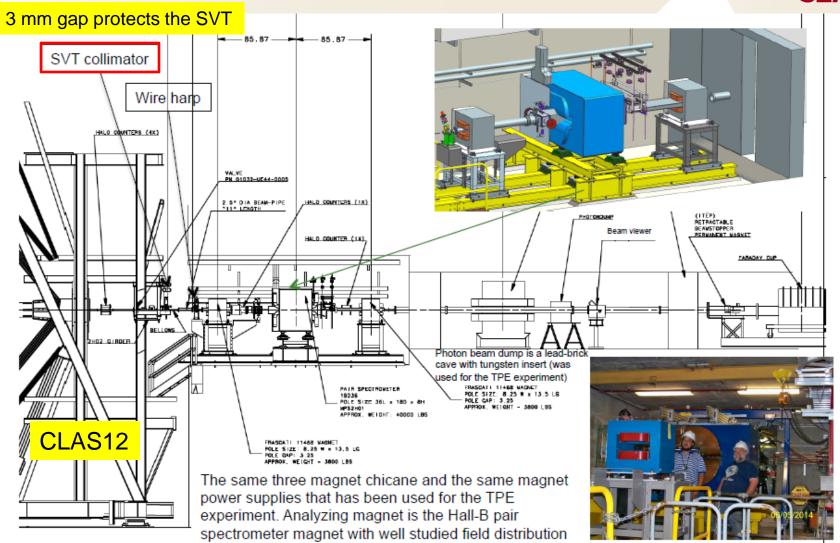
Beam line in Hall B



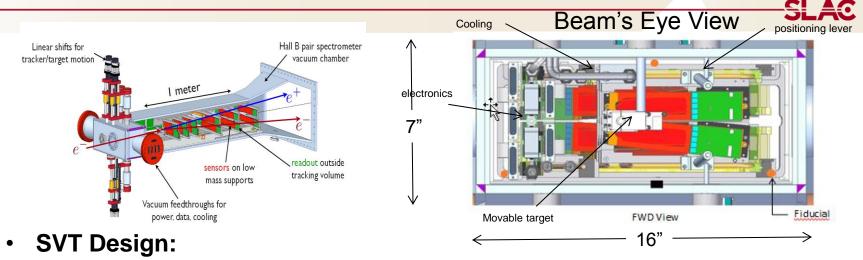


HPS in Hall B Alcove





HPS Apparatus: SVT



- Six layers of Si modules, split top-bottom, each with two sensors: axial and stereo
- 4x10 cm Hamamatsu microstrip sensor with 60 μm sense pitch.

Fast Readout:

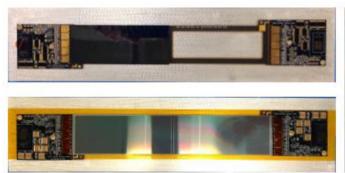
- CMS APV25 chip , 40 MHz continuous amplitude sampling with 3 μsec latency.
- Digitizing electronics and power distribution in vacuum.
- Power and control in/data out through vacuum feedthroughs.
- Electronics and sensors cooled < 0°C to remove heat and boost radiation hardness.

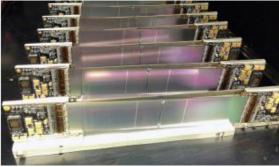
Precision Movers:

Position layers 1-3 close to the beam, do wire scans, and insert targets as needed.

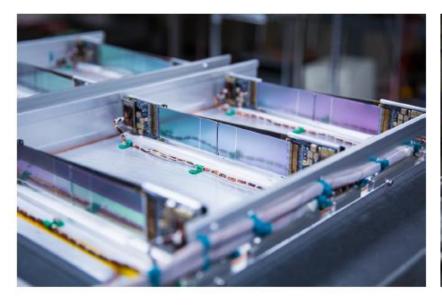
SVT Assembly at SLAC

SLAC











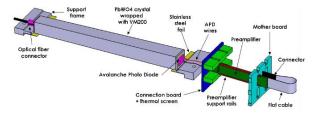
HPS Apparatus: ECal

SLAC

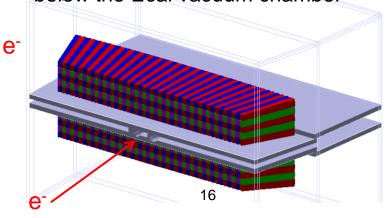
- Top and bottom modules
 - 5 layers each
 - 442 lead-tungstate (PbW0₄) crystals in all
- APD readout through custom preamplifiers
- Data recorded with 250 MHz 12 bit FADCs
- A thermal enclosure to keep crystal temperature constant to ~1°F to stabilize gains.

Ecal is downstream of SVT & magnet

PbWO₄ crystal with APD and preamp



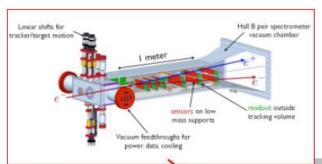
Crystals are arrayed above and below the Ecal vacuum chamber



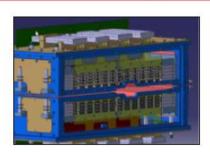
HPS Setup in Hall B Alcove

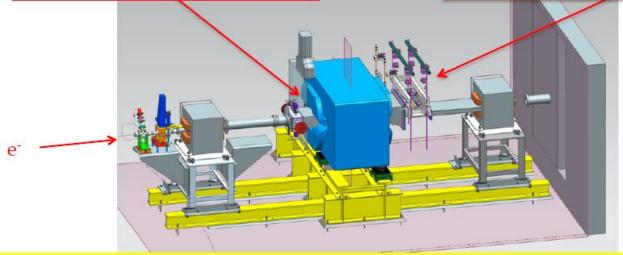


Si Vertex Tracker Installed Feb 23, 2015



PbWO₄ ECal Installed September, 2014



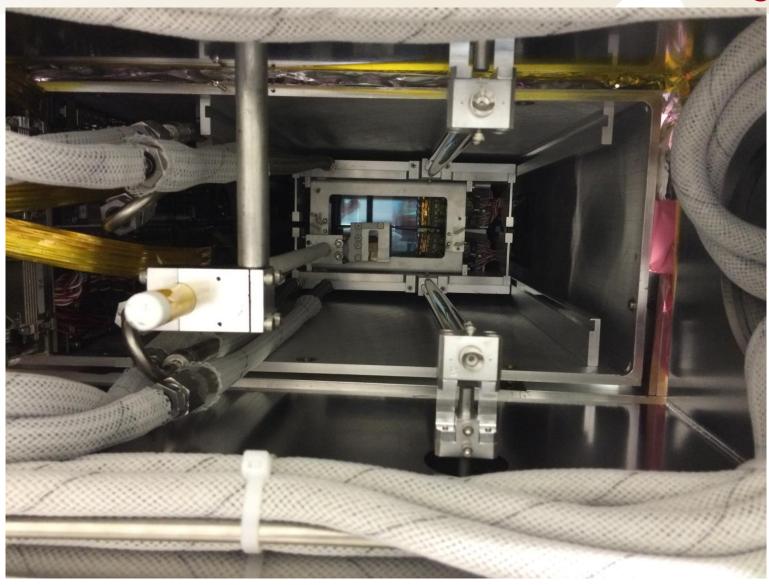


A magnet chicane directs the CEBAF 12 electron beam onto a W foil, producing heavy photons. They decay to e⁺e⁻ pairs, which are measured by the Si vertex tracker inside an analyzing magnet. A PbWO₄ ECal provides a fast trigger.

https://confluence.slac.stanford.edu/display/hpsg/Heavy+Photon+Search+Experiment

Beam's Eye View of the SVT

SLAC



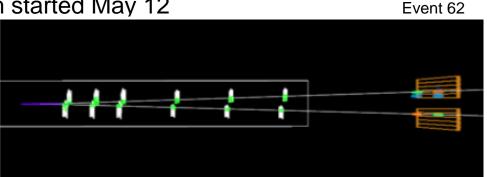
Spring 2015 Engineering Run

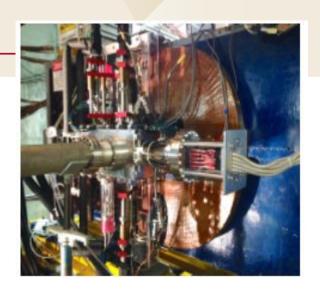
Opportunistic run: other Halls had a priority and the 12 GeV work was carried out during week-day day shift.

- Installed SVT end of February
- Commissioned Hall B beamline March-April
 - Calibrated bpms & established orbit locks
 - Set up SVT Protection Collimator
 - Checked beam position stability
- CEBAF down for two weeks after power outage
- Commissioned Trigger and Integrated SVT DAQ late April
- Explore SVT backgrounds as moved SVT closer to beam
- Production running at 1.5 mm started May 1
- Production running at 0.5 mm started May 12

Run ended May 18th.

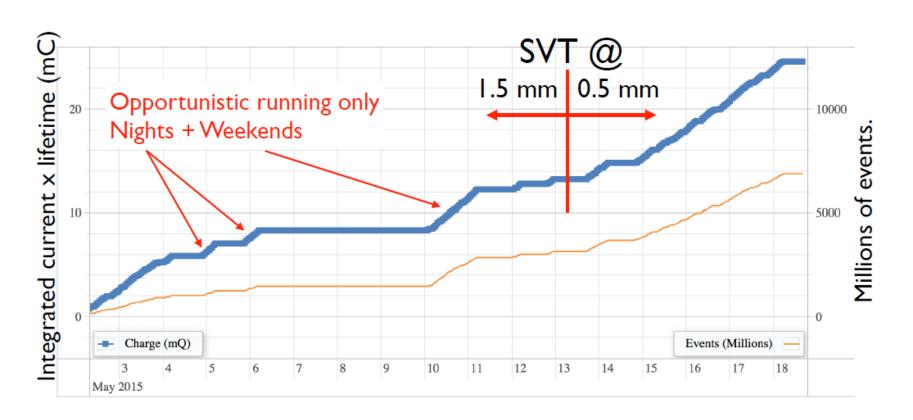
Layer 1 silicon sensors are just 0.5 mm above and below beam. Min. opening angle is $\theta y = 15$ mrad.





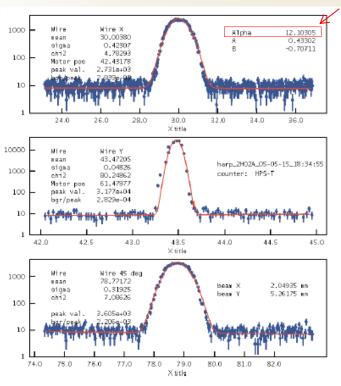
Run 5623

Proposal: 1 full week of 50 nA beam on target, 30mC Achieved: ~10 mC with SVT at 1.5mm, 10 mC at 0.5 mm



Beam Quality





X,Y and 45 degree beam profiles. May 5th, 2015

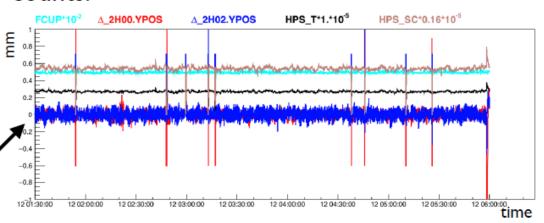
Very stable beam on May 12th.

Small skewness

HPS requires a very high quality beam, with very low halo.

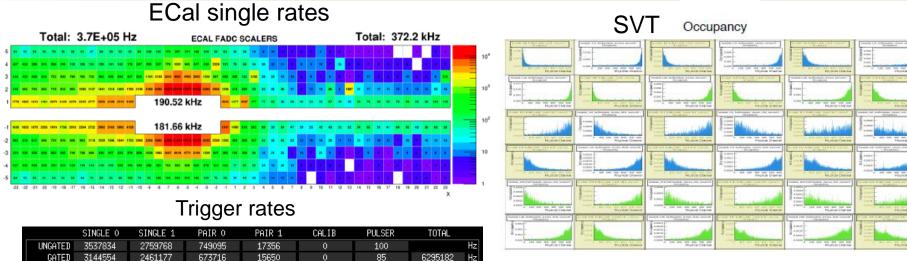
 $\sigma_X \sim$ 300 to 500 μm - To spread heat load. $\sigma_Y \sim$ 15 - 50 μm - To help vertexting & tracking.

The beam also needs to be very stable over time. A Fast Shut-Down stops the beam in <10 ms, if halo counters register above threshold counts.



Online data quality



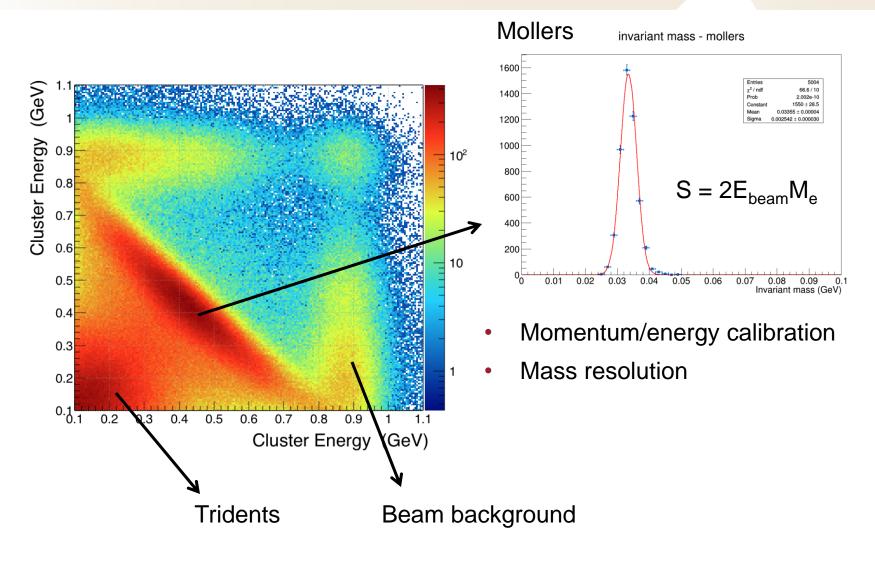


GATED 3144554 2461177 673716 15650 6295182 384 1202 658 15650 85 17979 PRESCALED 100 100.0 80.0 60.0 40.0 20.0 -20.0 -18.0 -16.0 -14.0 -12.0 -10.0 -8.0 time (min) SINGLE 0 SINGLE 1 PAIR 0 PAIR 1 PRESCALE FACTORS (2^N)

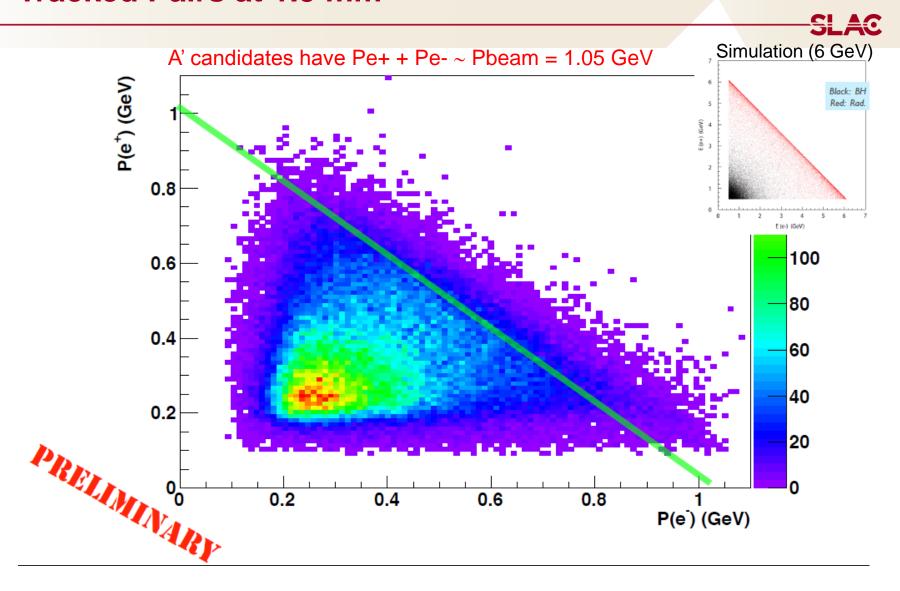
In good agreement with simulations.

Two Cluster Events





Tracked Pairs at 1.5 mm

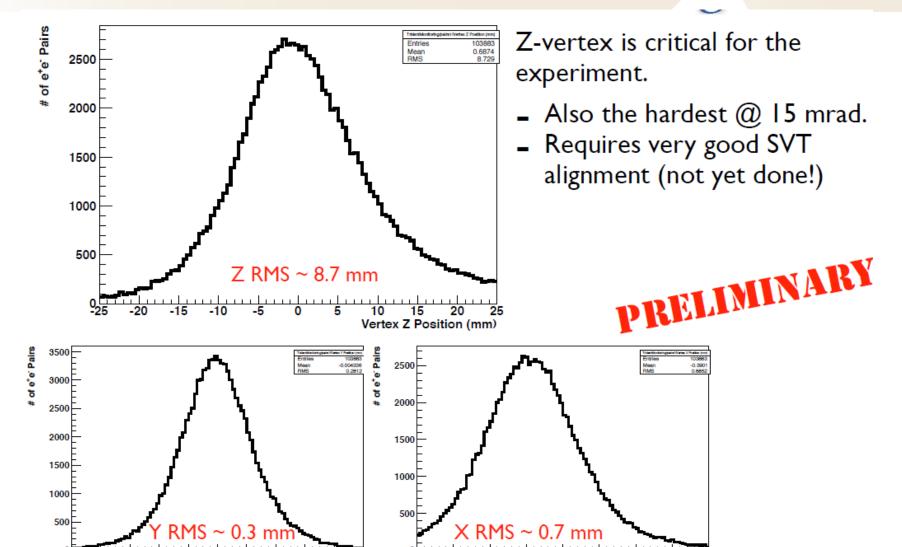


-0.4

-0.2

0.6

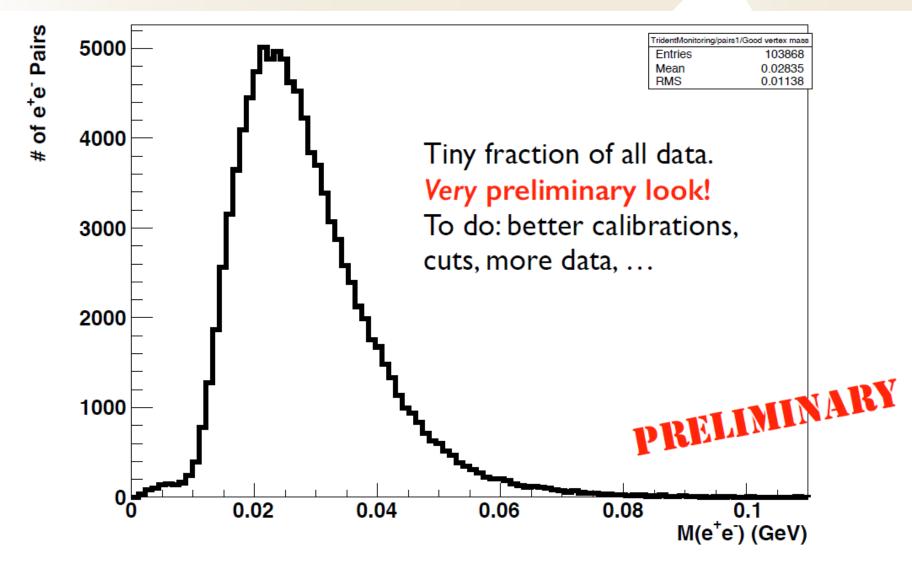
Vertex Y Position (mm)



Vertex X Position (mm)

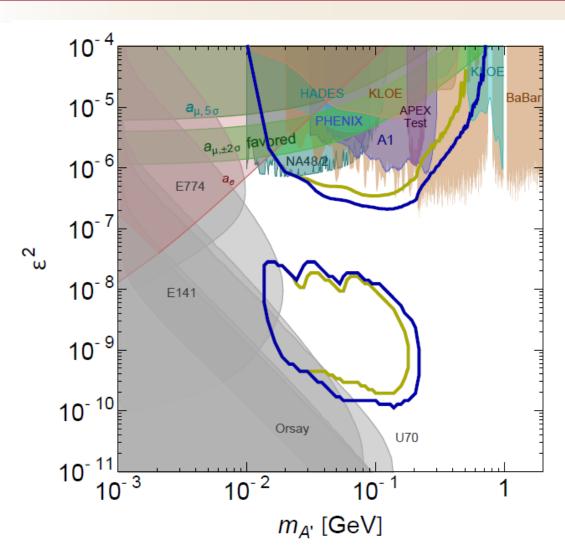
Pairs Mass Distribution





Full HPS Reach





Near term Running (Yellow)

1 week with 50nA @ 1.1 GeV 1 week with 200nA @ 2.2 GeV 2 weeks with 300nA @ 4.4 GeV

Additional Running (Blue):

2 weeks with 200nA @ 2.2 GeV 2 weeks with 300nA @ 4.4 GeV 3 weeks with 450nA @ 6.6 GeV

Times are "PAC" times = Calendar time/2

Opportunistic run Fall 2015 TBD Spring 2016

Summary



- We have roughly 1/3 PAC week with Si at 0.5 mm
 - 15 mrad acceptance
- Beamline, ECal, Trigger and SVT all worked well
 - Beam background and trigger rates are consistent with simulations.
- Lots of work to do ...
 - Check Trident Yield in the data
 - ECal energy calibration
 - SVT alignment
 - Understanding the vertex tails
- But a physics result may be in reach

HPS Collaboration JLAB + SLAC + FNAL + IPNO Orsay + INFN Genova + Universities (+ New Collaborators at Glasgow and INFN Catania, Torino, Sassari, Roma)

C! AG

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