# The Heavy Photon Search Experiment

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SLAC National Accelerator Laboratory
June 27, 2017





# Introduction

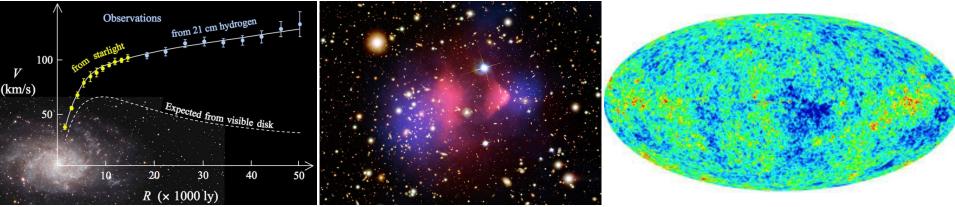
#### SLAC

- A heavy photon (or dark photon, or A') is a hypothetical vector boson that couples indirectly to electromagnetic particles
- The Heavy Photon Search (HPS) is a fixed target experiment at Jefferson Lab dedicated to searching for this hypothetical vector boson, an A'
- HPS uses two distinct methods to search for A's a resonance search and a displaced vertex search
- Motivation for A's and light dark matter, the experimental setup, initial results, and the future of HPS are presented

# The Existence of Dark Matter

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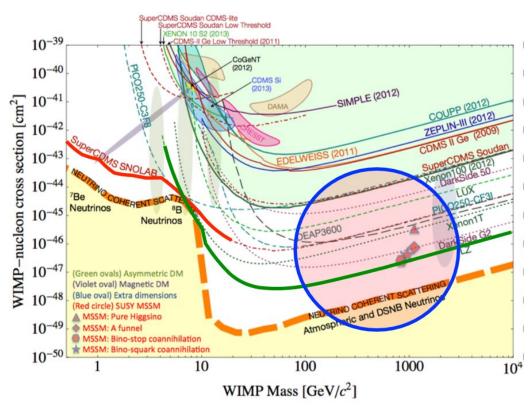
- There is clear evidence for the **existence of dark matter** (DM)
  - Galactic rotation curves, gravitational lensing, and CMB
- The fundamental nature of DM is a **central puzzle in particle physics**
- What are some ideas for what DM could be?



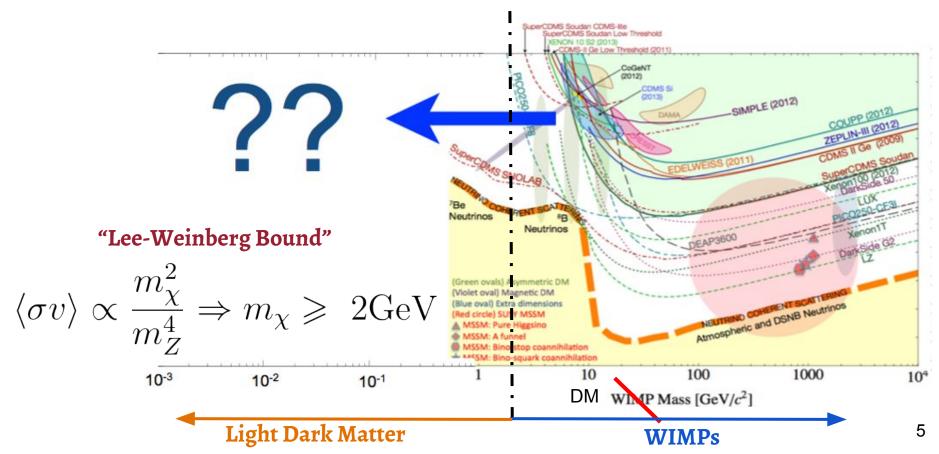
# **Weakly Interacting Massive Particles**

- Assume DM is thermal (i.e. was in equilibrium with SM matter in the early universe)
- WIMPs are most popular model for thermal DM, but have yielded null results so far
- WIMPs will either be found or most of the accessible parameter space will be ruled out within the next few years
- Are there other places to look?

# arXiv:1310.8327



Lighter dark matter requires a new, light force carrier

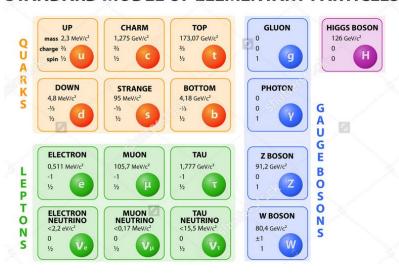


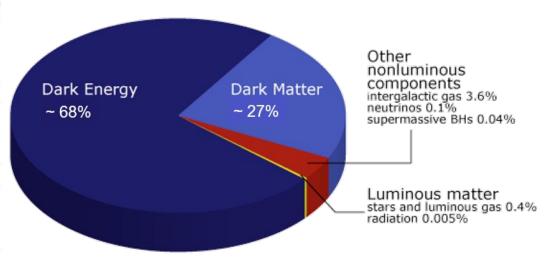
## The Standard Model

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- The SM is the best description of nature at the fundamental level, but is only 4% of the universe
- SM is also complicated with several forces and many types of particles. Why should DM be any different/simpler?

#### STANDARD MODEL OF ELEMENTARY PARTICLES





# **Heavy Photon Primer**

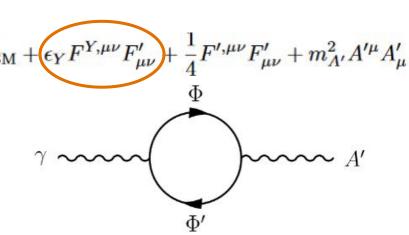
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Suppose nature contains an additional

Abelian gauge symmetry U'(1) (analogous

$$\mathcal{L} = \mathcal{L}_{\rm SM} + \epsilon_Y F^{Y,\mu\nu} F'_{\mu\nu} + \frac{1}{4} F'^{,\mu\nu} F'_{\mu\nu} + m_{A'}^2 A'^{\mu} A'_{\mu}$$

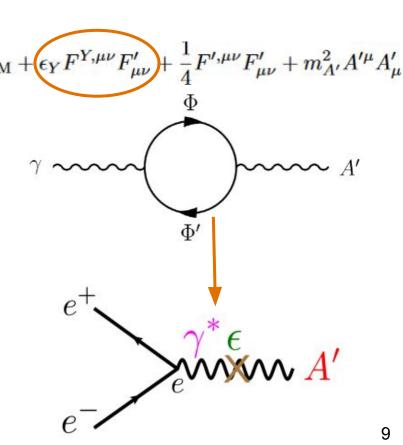
- Suppose nature contains an additional
   Abelian gauge symmetry U'(1) (analogous to EM)
- This gives rise to a kinetic mixing term
  where the SM photon mixes with a new
  gauge boson (an A') through interactions
  of massive fields (i.e. a "vector portal")



- Suppose nature contains an additional **Abelian gauge symmetry** U'(1) (analogous to EM)
- This gives rise to a kinetic mixing term where the SM photon mixes with a new gauge boson (an A') through interactions of massive fields (i.e. a "vector portal")
- Induces a weak effective coupling of Ee to **SM** fermions

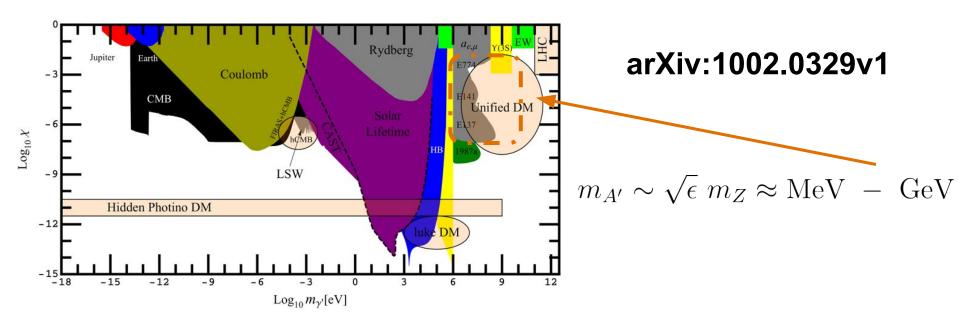
$$\epsilon \sim \frac{g_Y g_D}{16\pi^2} \ln\left(\frac{m_\Phi}{m_{\Phi'}}\right) \sim 10^{-3} - 10^{-1}$$

GUT theories motivate  $\log(\epsilon) \sim -5$  to -3

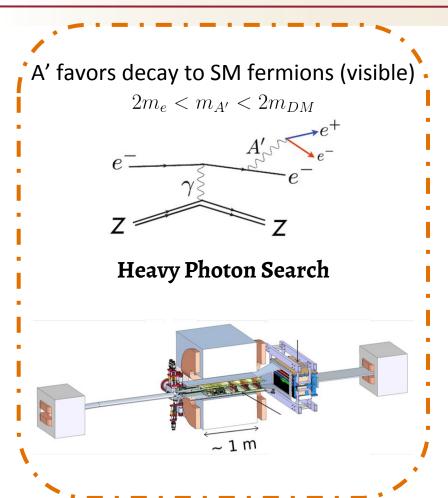


Possible origin for mass related to mass of Z by small parameter

(SUSY + kinetic mixing → scalar coupling to SM Higgs)

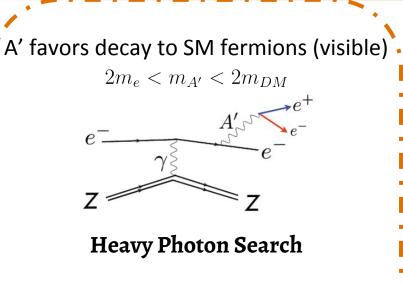


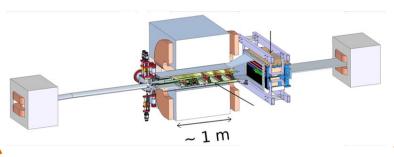
This is the SM-like mass range. Must consider both visible and invisible decays



# **Heavy Photon Signatures**



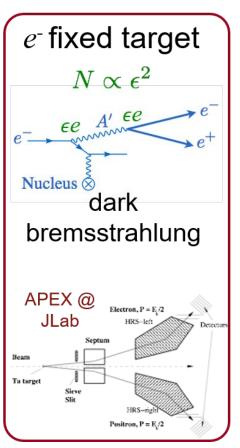


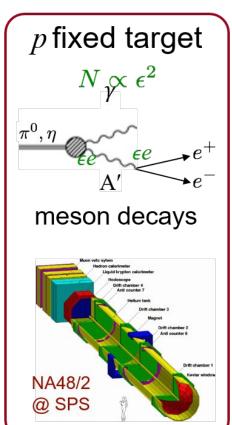


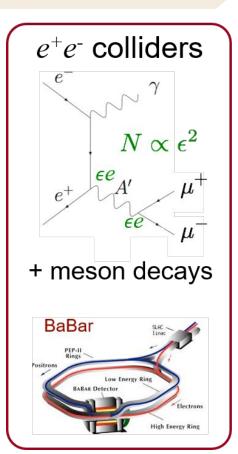


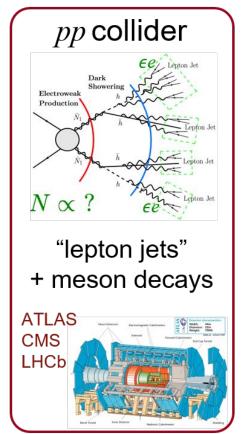
# **Experiments Searching for Heavy Photons**

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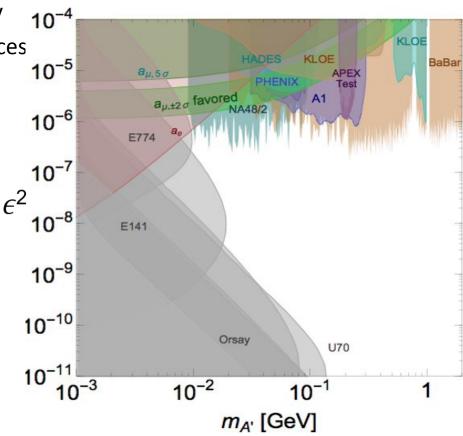




# **Existing Heavy Photon Constraints**

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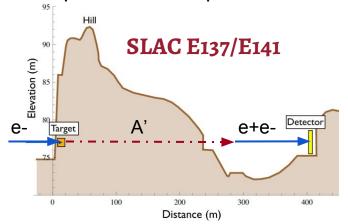
• Large coupling searches are generally "bump hunts" for  $m(l^+l^-)$  resonances

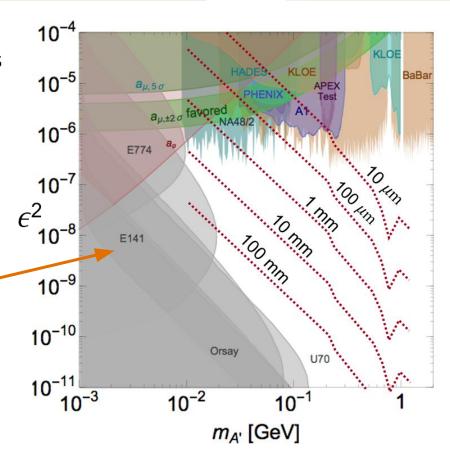


- Large coupling searches are generally "bump hunts" for  $m(l^+l^-)$  resonances
- A's with small coupling are long-lived

$$\gamma c\tau \propto \frac{1}{\epsilon^2 m_{A'}^2}$$

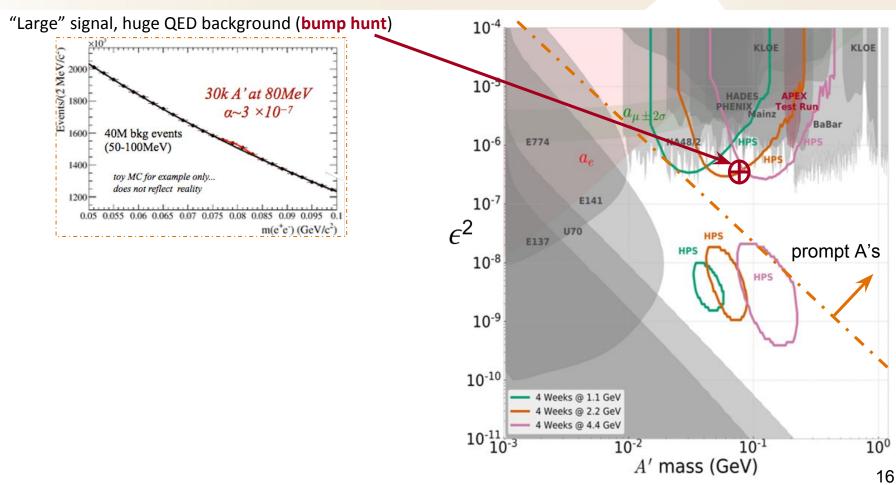
 Constraints from "beam dump" experiments are possible





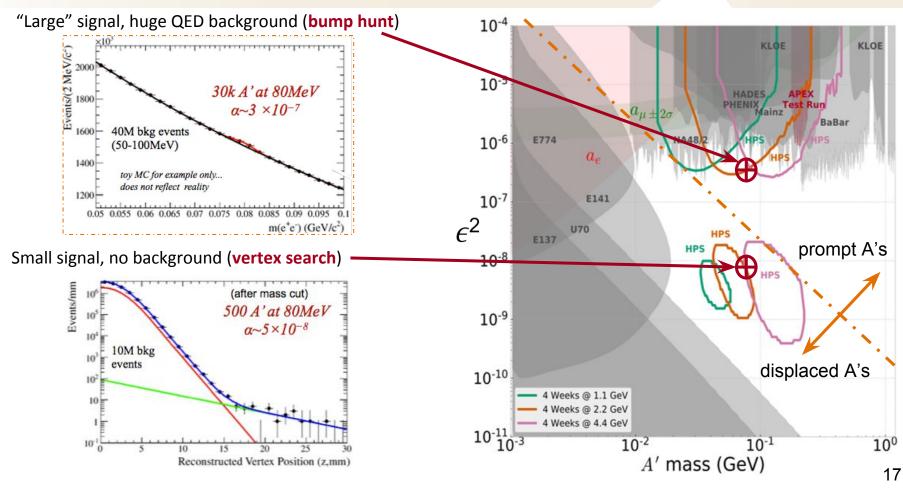
# **Heavy Photon Signatures in HPS**



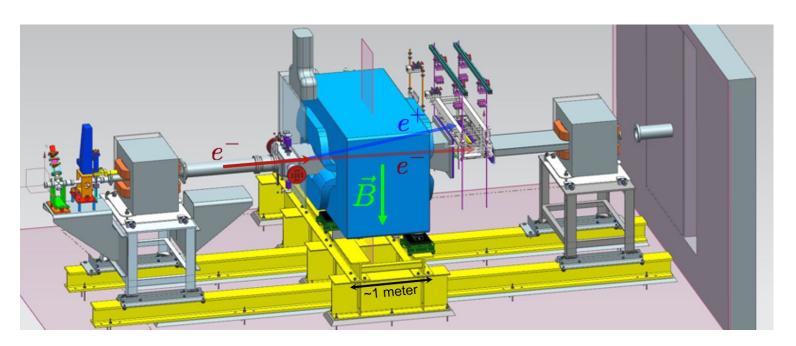


# **Heavy Photon Signatures in HPS**



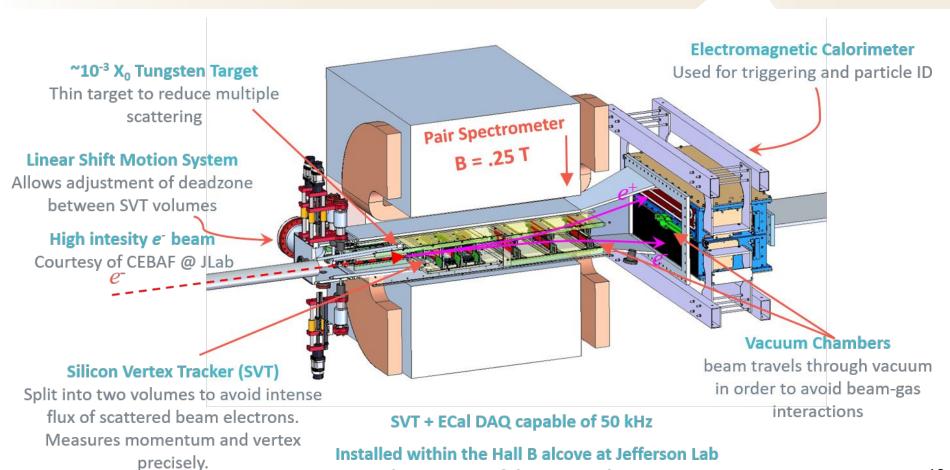


- HPS is a fixed-target search for visibly decaying dark photons using the CEBAF electron beam (1-11 GeV) in Hall B at Jefferson Lab
- A's can be produced in a process analogous to Bremsstrahlung in a thin W foil
- Large dipole magnet spreads e+e- pairs and provides momentum measurement



## **HPS Detector**

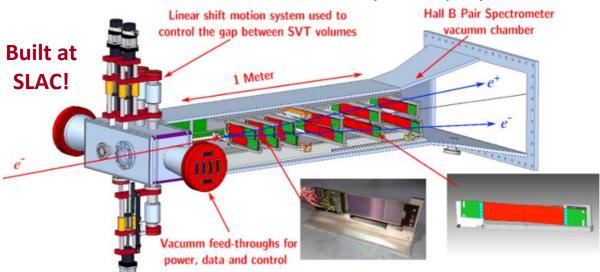
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downstream of the CLAS12 detector

19

- SVT measures trajectories of e+e- and reconstructs mass and vertex position
- 6 layers of silicon microstrips (~0.7% radiation length per layer)
- Each layer has axial/stereo strips (100 mrad) for 3D hit position
- 50 kHz max trigger rate
- L1-L3 vertically retractable from beam
- L4-L6 are double wide for acceptance purposes





# Jefferson Laboratory and CEBAF

JLab (Newport News, VA) has the Continuous Electron Beam Accelerator Facility (CEBAF) that can simultaneously deliver intense electron beams of different energies to 4 experiment halls

- 2.2 GeV per pass up to 12 GeV and 2 ns bunch pulse
- Provides small beam spot with small tails

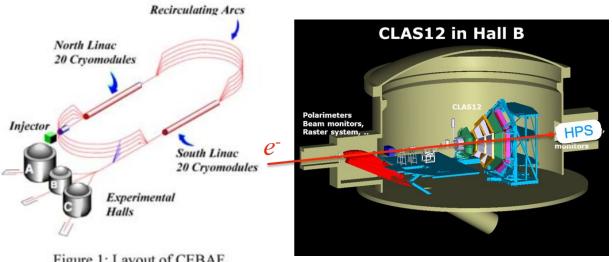
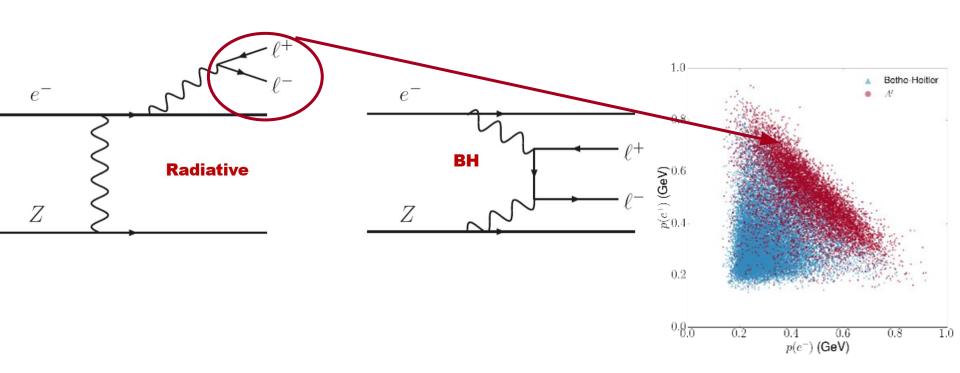


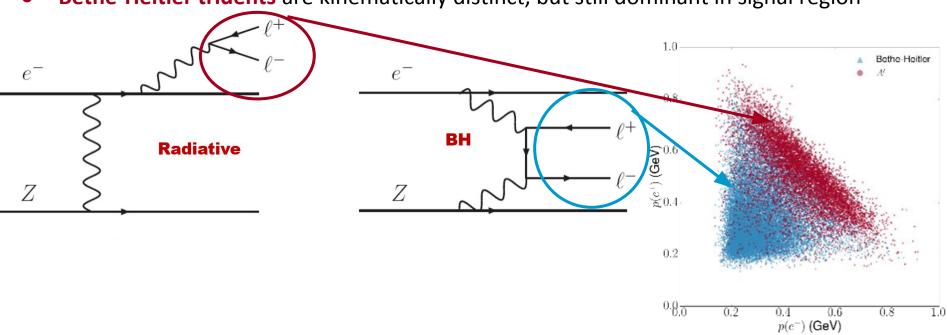


Figure 1: Layout of CEBAF.

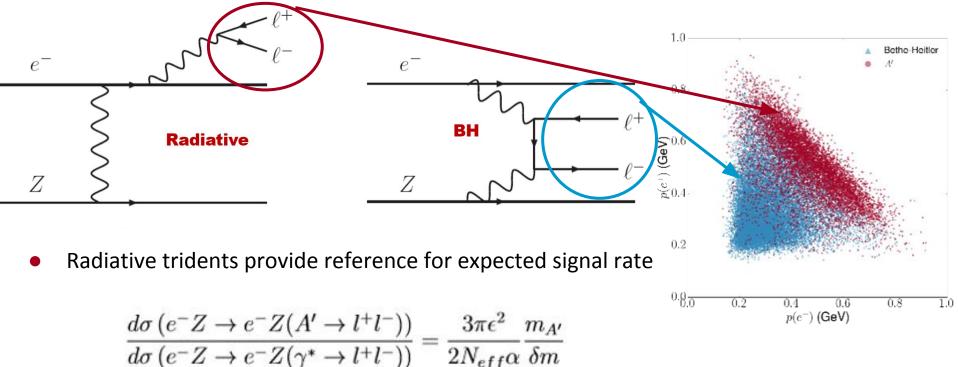
Radiative tridents have identical kinematics for e+e- pair (irreducible)



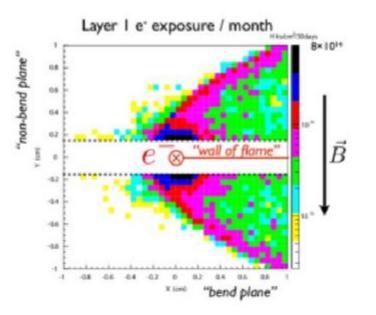
- Radiative tridents have identical kinematics for e+e- pair (irreducible)
- Bethe-Heitler tridents are kinematically distinct, but still dominant in signal region

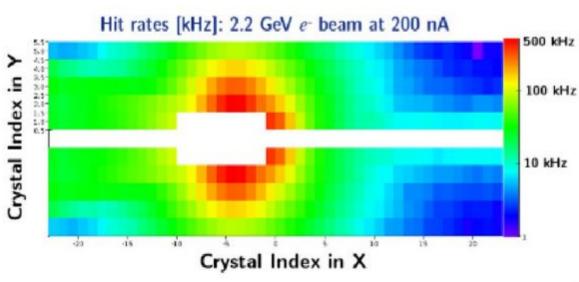


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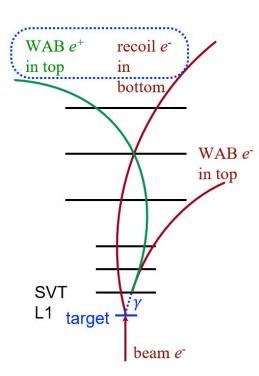


- Background is dominated by electrons scattering in the target
- Detector (vertical) acceptance down to +/- 15 mrad (which means L1 of SVT is
   0.5 mm from beam axis!)
- This provides challenges for occupancies, data rates, and radiation tolerances





- Converted photons in tracker or target are common, but pairs are in the same hemisphere. Recoils are generally soft
- Recoil electron and a conversion positron in opposite hemispheres can trigger: rate comparable to tridents



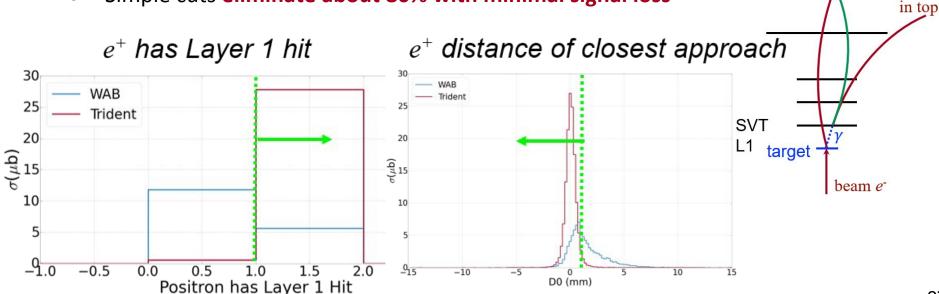
recoil e

botton

WAB  $e^+$ 

in top

- Converted photons in tracker or target are common, but pairs are in the same hemisphere. Recoils are generally soft
- Recoil electron and a conversion positron in opposite hemispheres can trigger: rate comparable to tridents
- Simple cuts eliminate about 80% with minimal signal loss



WAB e



#### **2015** Engineering Run

50 nA at 1.06 GeV 1.7 days (10 mC) of physics data

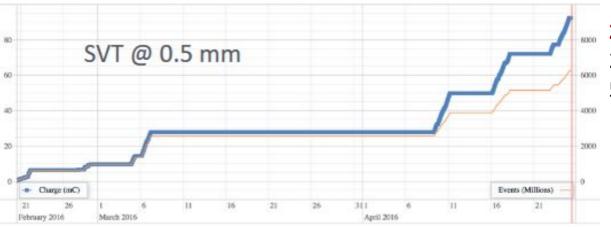
# 2015 & 2016 Engineering Runs





#### **2015** Engineering Run

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#### **2016 Engineering Run**

200 nA at 2.3 GeV 5.4 days (92.5 mC) of physics data

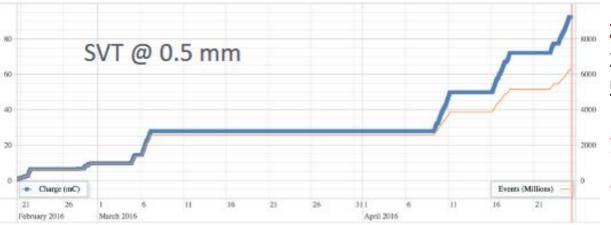
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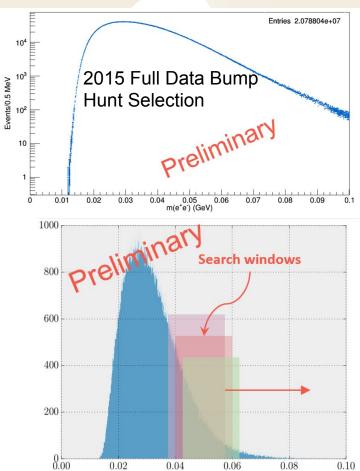
# 180 days of data taking approved by JLab PAC!

- 2015 Bump Hunt analysis is complete
  - Presented at JLab seminar in May of 2017 (Omar Moreno)
- 2015 Vertexing analysis is currently in progress
  - We might have interesting physics sensitivity (SIMPs), but not for the minimal A' model...
- Brief overview and status of these analyses will be presented in the next few slides
- 2016 Bump Hunt and Vertexing analysis are waiting (we are learning A LOT from the 2015 data)

# **Resonance Search**



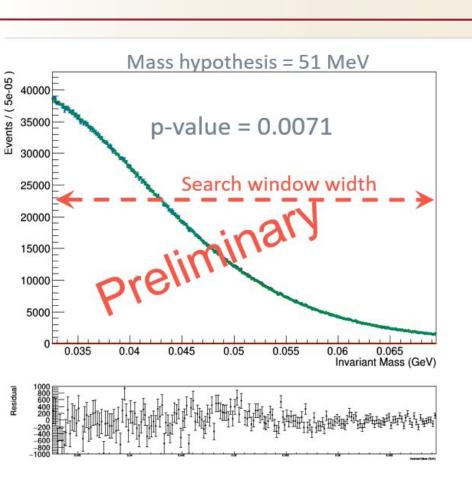
- Scan mass spectrum between 17-80 MeV using search windows
- Search for a bump by fitting a polynomial background plus a Gaussian signal
- Use a likelihood ratio to quantify the significance of any bump
- Bound upper limit with expected limit to avoid exaggerated exclusion from downward fluctuations (power constrained limit)
- Plot the 2σ exclusion in A' parameter space

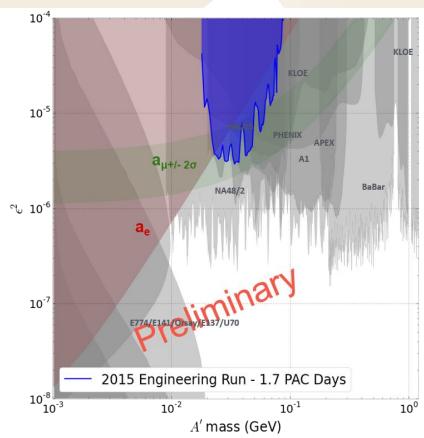


Invariant Mass  $M(e^-e^+)$  (GeV)

# 2015 Resonance Search Results



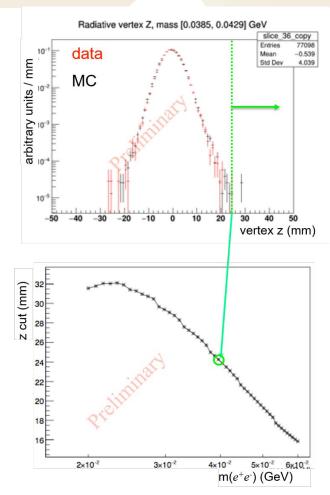




# **Vertex Search**



- First analysis requires e+e- hits in SVT L1 (L1L1)
  - This is only part of the acceptance, require L1L2 and L2L2 exclusive categories in the future for full acceptance
- Plot vertex z in mass slices. Use MC to fit background as gaussian and exponential tail in each slice
- Define "zcut" as 0.5 background events expected for z > zcut
- Any event with a z vertex great than zcut is a candidate for signal



# **Vertex Search**

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 Vertex search is in progress, but will not have sensitivity for minimal A' model with 2015 data at 1.05 GeV

But...

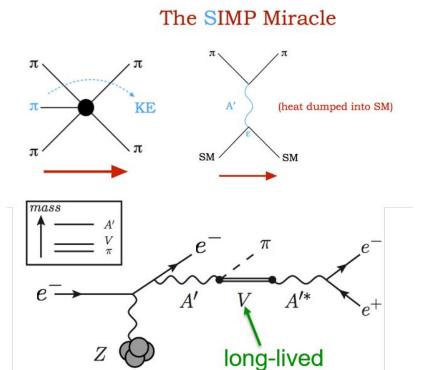
## **Vertex Search**

SLAC

 Vertex search is in progress, but will not have sensitivity for minimal A' model with 2015 data at 1.06 GeV

But...

- Possibly sensitive to Strongly Interacting
   Massive Particles (SIMPs) in both 2015
   and 2016 datasets
- Small upgrade projects planned for next run (additional SVT layer and positron trigger)
- 95% of data is still to come!

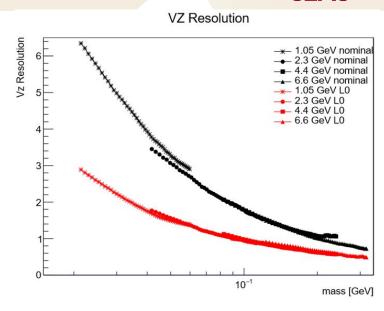


arXiv:1402.5143

## **HPS Upgrades**

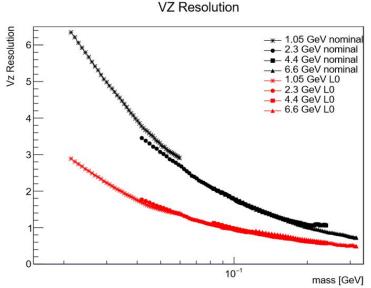
#### SLAC

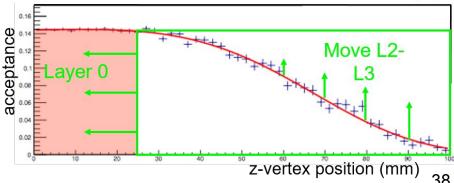
- Add a tracking layer (Layer 0) between target and current first layer
  - Dramatically improves vertex resolution, hence the vertex reach



## **HPS Upgrades**

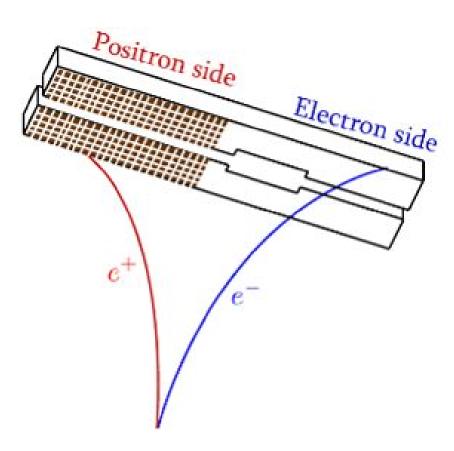
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- Move L2-L3 slightly towards beam
  - Improves acceptance for longer-lived A's



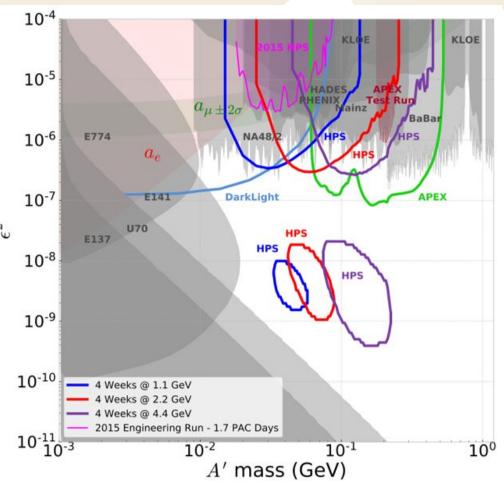


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  - Dramatically improves vertex resolution, hence the vertex reach
- Move L2-L3 slightly towards beam
  - Improves acceptance for longer-lived A's
- Add positron hodoscope inside vacuum chamber
  - Reduces acceptance losses in the "Ecal hole"
- Relatively simple. Currently awaiting formal approval



- Latest HPS reach estimates including minor upgrades (LO and positron trigger) for 4 weeks at 3 different beam energies
- Without upgrades, there is no vertexing reach at 4 weeks



- Heavy photons are still motivated as a proposed model for more "complex" dark matter and could be linked to light dark matter
- HPS has successfully run at two different beam energies (1.06 GeV in 2015 and 2.3 GeV in 2016)
- First bump hunt results recently released!
- Many more upcoming analyses to come out including vertexing, 2016 bump hunt, and possibly SIMPs
- HPS upgrades are small projects but provide dramatic improvements
- HPS is getting ready for a lot of data taking in 2018 with potential upgrades!

## **Questions?**

#### -SLAC

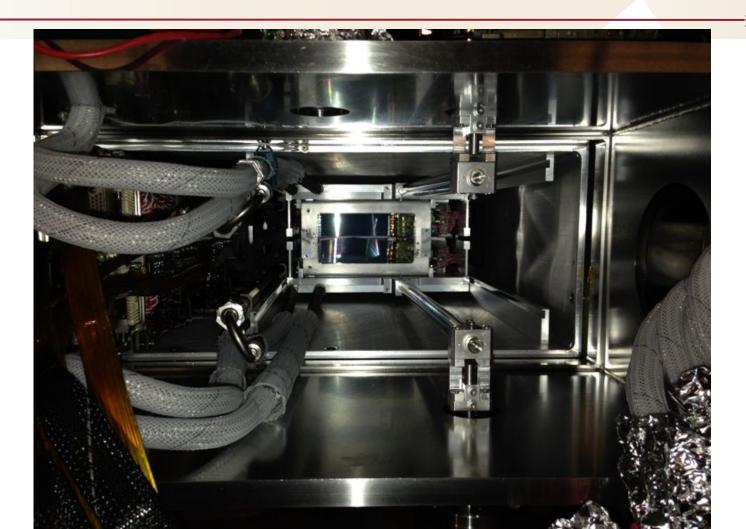


#### **HPS Collaboration**

May 3 - 5, 2017 Jefferson Lab • Newport News, VA

# **HPS SVT**

## -SLAC



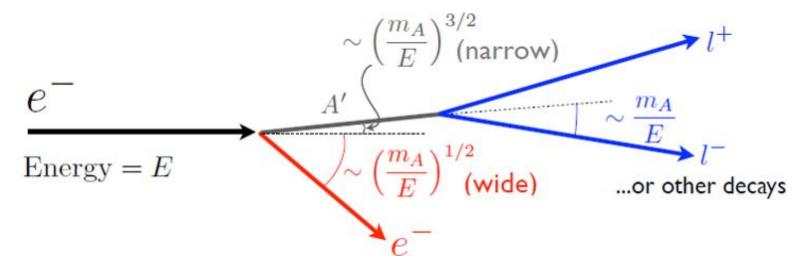
# **HPS SVT**



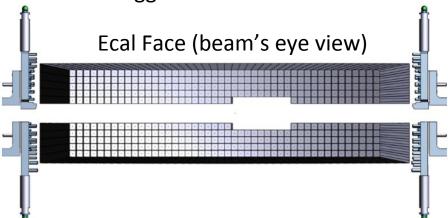
## **Heavy Photon Kinematics and Design Considerations**

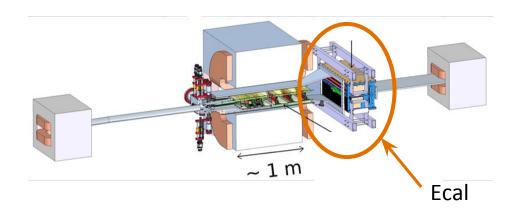
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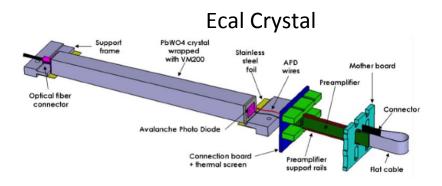
- A' decay products have very small opening angle (take most of beam energy)
- Detector acceptance must be very forward (very close to beam plane)
- Bump Hunt: need good mass resolution
- Vertexing: requires excellent vertex resolution, thin target, tracker under vacuum, and minimal tracker material
- Small couplings -> small cross-section (rates). Need high intensity beam



- Ecal made out of 442 lead tungstate (PbWO4) and built by JLab/Orsay/INFN
- Provides e+e- trigger with precision timing
- >100 kHz max trigger rate with 8 ns trigger window



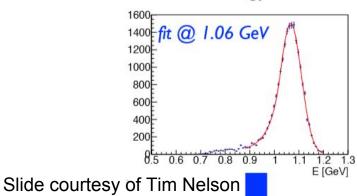


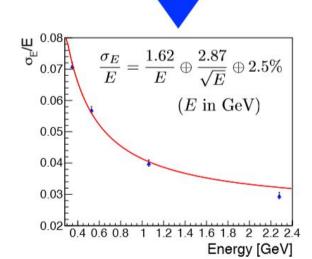


#### **Ecal Performance**

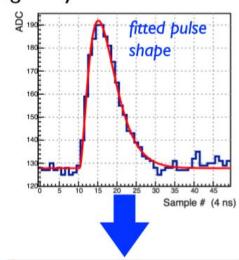


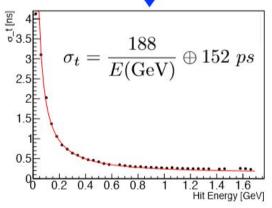
#### cluster energy resolution

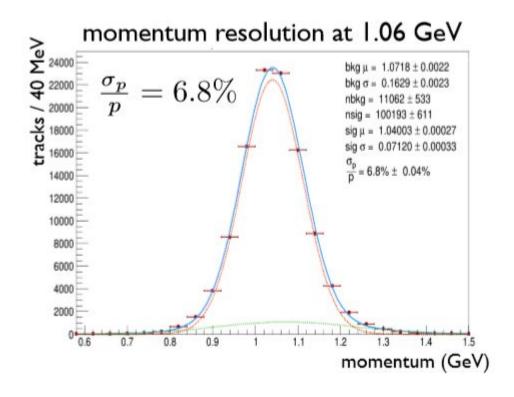




#### single-crystal time resolution

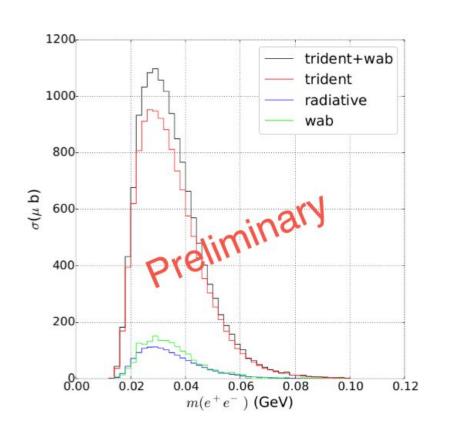


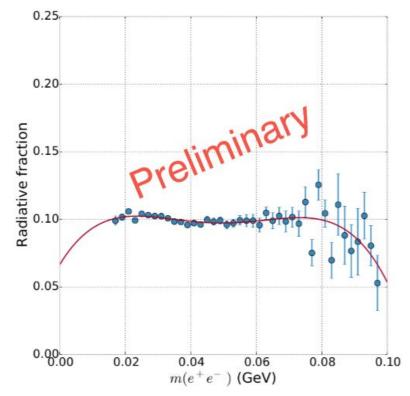




### ~10% of total yield are radiative tridents

Slide courtesy of Tim Nelson

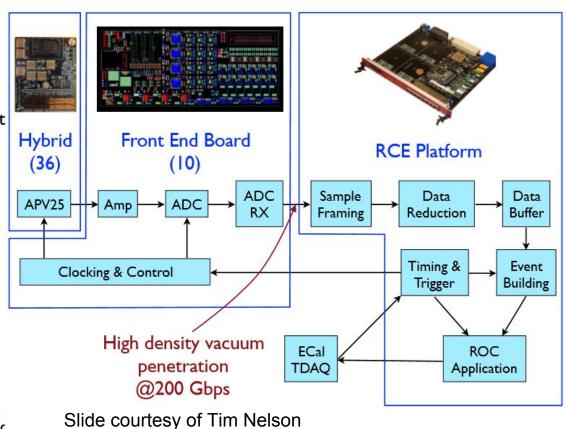




Based upon SLAC RCE platform (ATLAS upgrade, DUNE, LSST...)

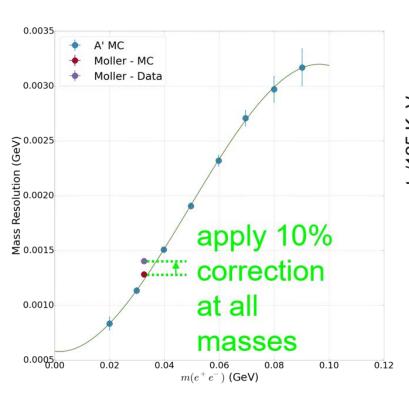
#### Some unique challenges too...

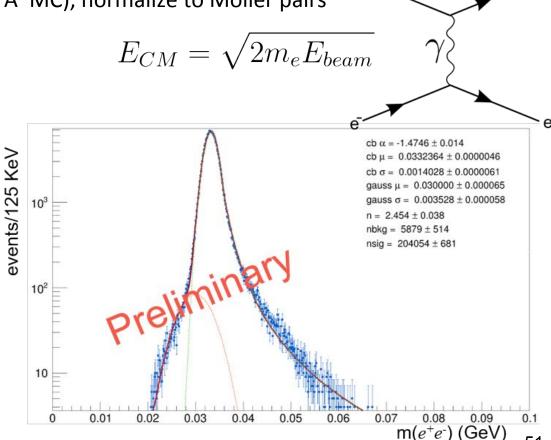
- CMS APV25 multi-peak readout for 2 ns time resolution
- In-vacuum ADC, voltage generation and power distribution/control on very dense Front End Boards
- Vacuum penetration for digital signals via high-density PCB through flange w/ external optical conversion.
- Supports trigger rates up to 50 kHz, raw data rates in excess of 100 gbit/sec.



#### **HPS Mass Resolution**

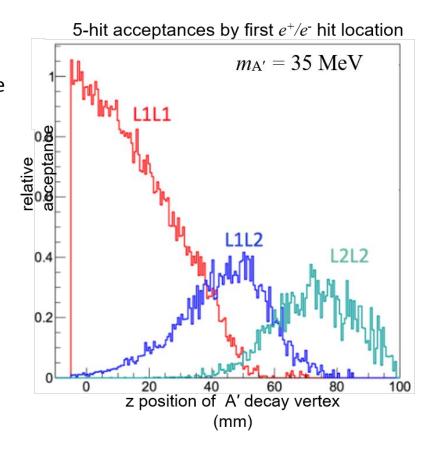
Mass resolution is linear (from A' MC), normalize to Moller pairs





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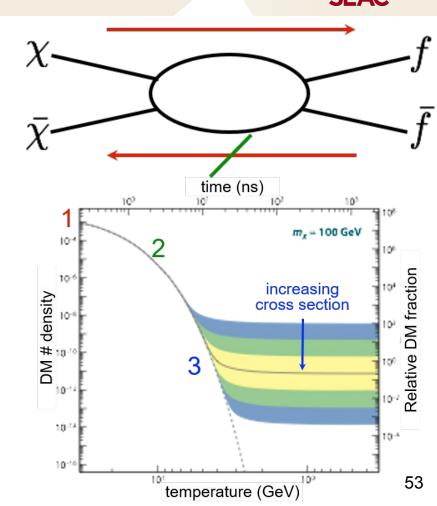


# **Thermally Produced Dark Matter**

- Very general/simple mechanism for DM production in the early universe
- Assume DM is in thermal equilibrium with SM particles
- Universe cools so SM cannot produce DM pairs
- 3. Universe expands so DM stops annihilating "freeze out"

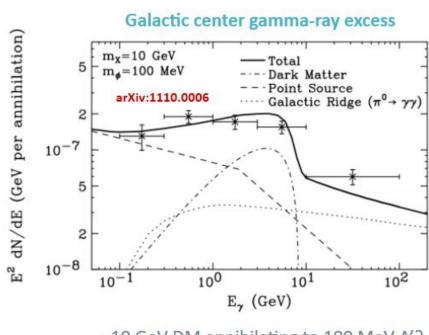
$$\Omega_\chi \propto rac{1}{\langle \sigma v 
angle}$$
 "WIMP Miracle"

$$\langle \sigma v \rangle \propto \frac{m_{\chi}^{2}}{m_{\chi}^{4}} \Rightarrow m_{\chi} \approx 100 \text{ GeV}$$



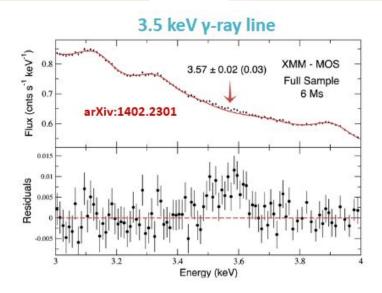
### Other anomalies with DM/heavy photon interpretation

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~10 GeV DM annihilating to 100 MeV A'?

Slide courtesy of Omar Moreno



#### **eXciting Dark Matter?**

