

Results and Future Prospects for HPS

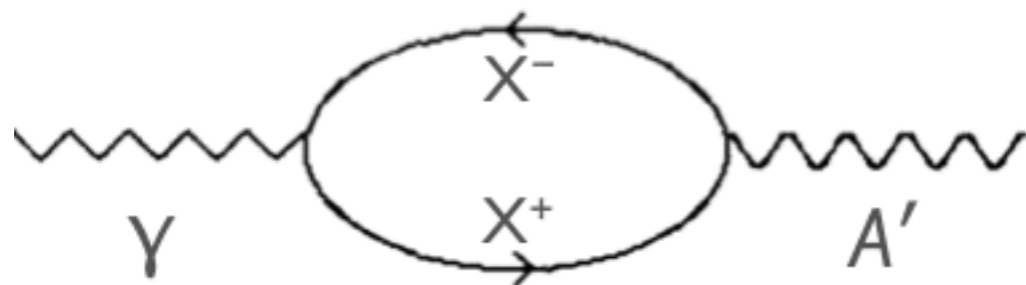
Matt Graham, SLAC
on behalf of the HPS Collaboration
APS April Meeting, 2018

The Vector Portal and Kinetic Mixing

an old idea: if there is an additional $U(1)$ symmetry in nature, there will be mixing between the photon and the new gauge boson

Holdom, Phys. Lett B 166, 1986

$$\mathcal{L}_{U(1)'} = -\frac{1}{4} V_{\mu\nu}^2 - \boxed{\frac{\epsilon}{2} V_{\mu\nu} F^{\mu\nu}} + |D_\mu \phi|^2 - V(\phi)$$



Kinetic Mixing term

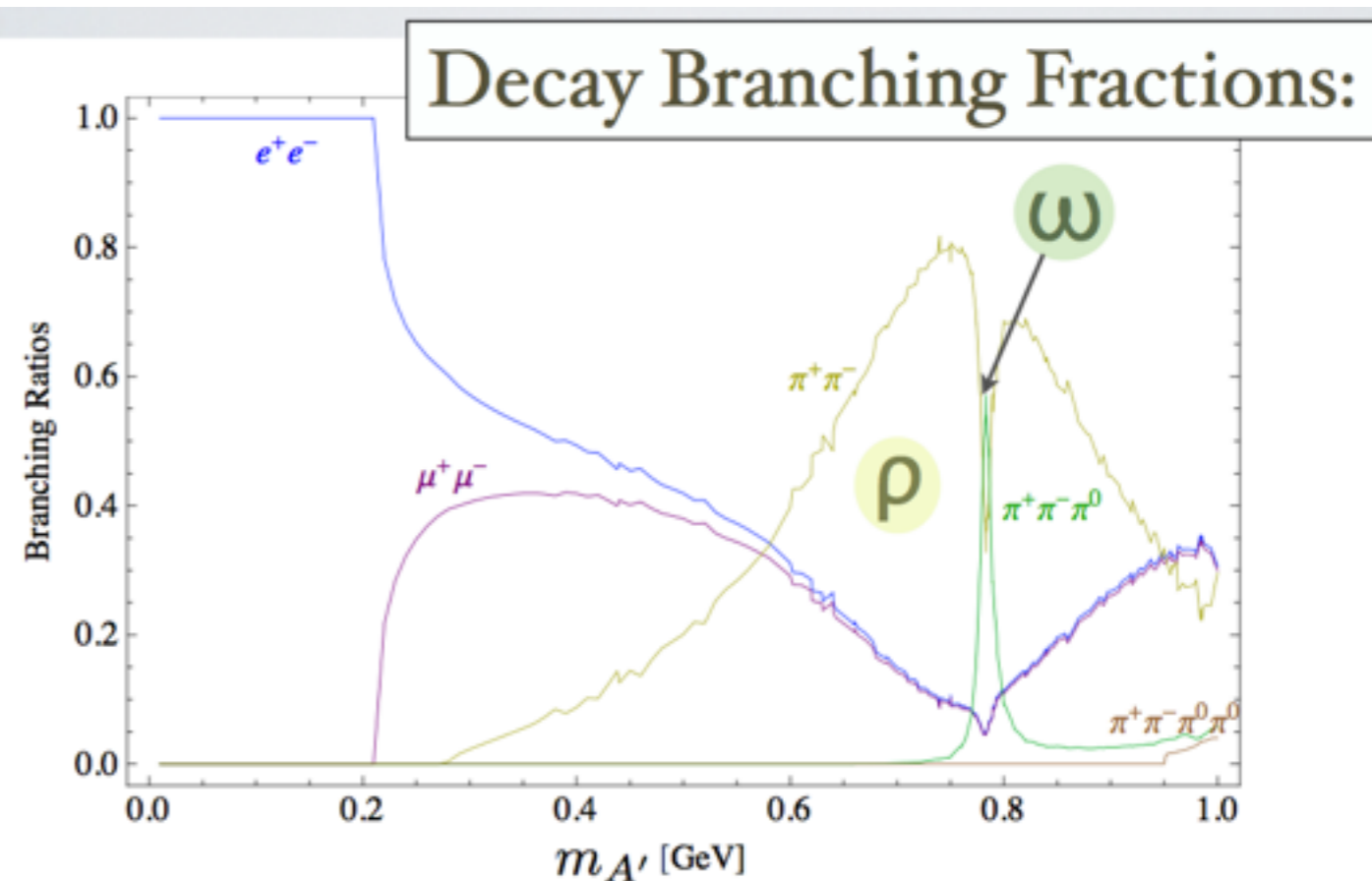
- extremely general conclusion...even arises from broken symmetries
- gives coupling of normal charged matter to the new “heavy photon” $q=\epsilon e$
- in everything I show, I’ll assume a “simple” dark sector, where the A' is part of an Abelian gauge group...other cases are interesting too!

Decays of the A': Two Scenarios

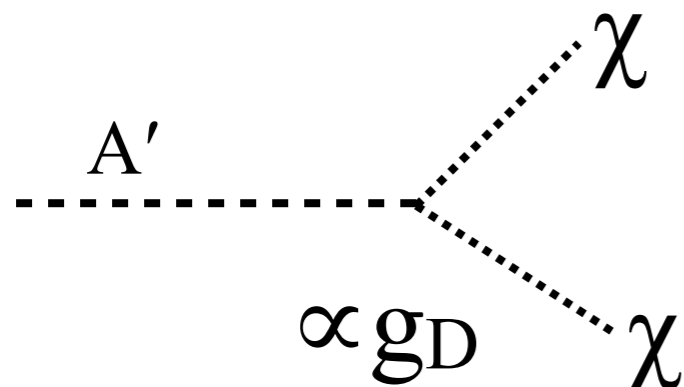
If $m(A') < 2m(\chi)$

..the A' will decay to SM particles just like a virtual γ^* "decays", i.e. the ratio R!

Which means, for $m(A') < \text{GeV}$ away from QCD resonances, a lot of lepton pairs



If $m(A') > 2m(\chi)$



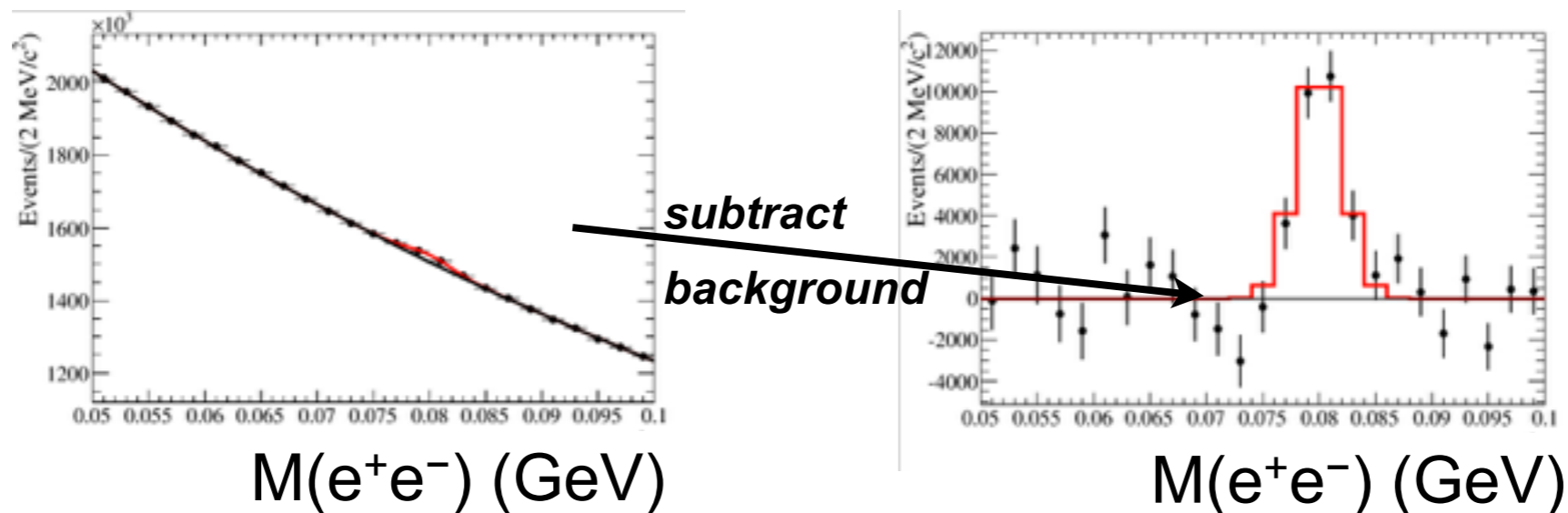
...for relatively lighter χ , A' decays into the dark sector will dominate...

the experimental picture of vector-mixed dark sectors is not complete without this scenario!

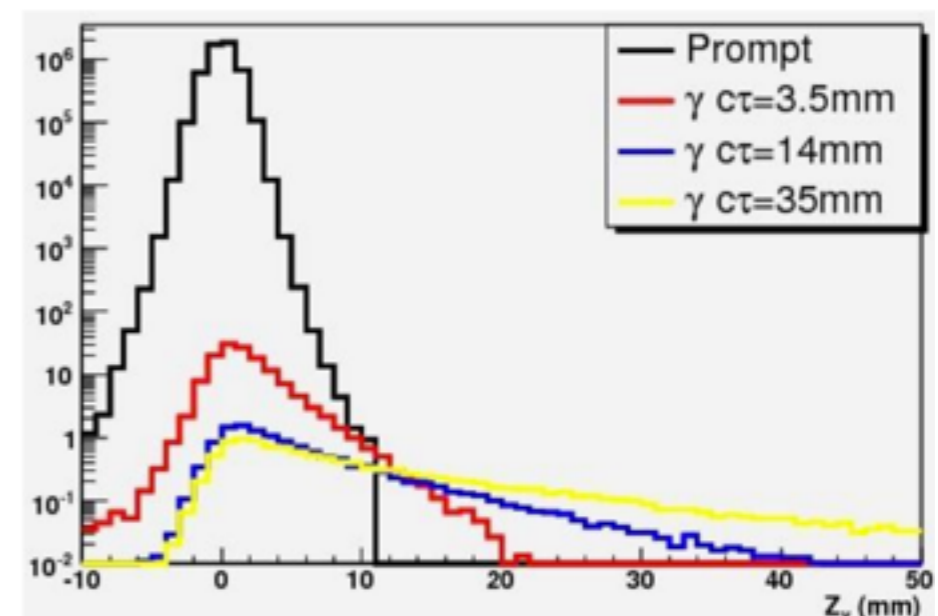
BUT, for this talk I'll focus on visible decays

Two ways to look for a visibly decaying A'

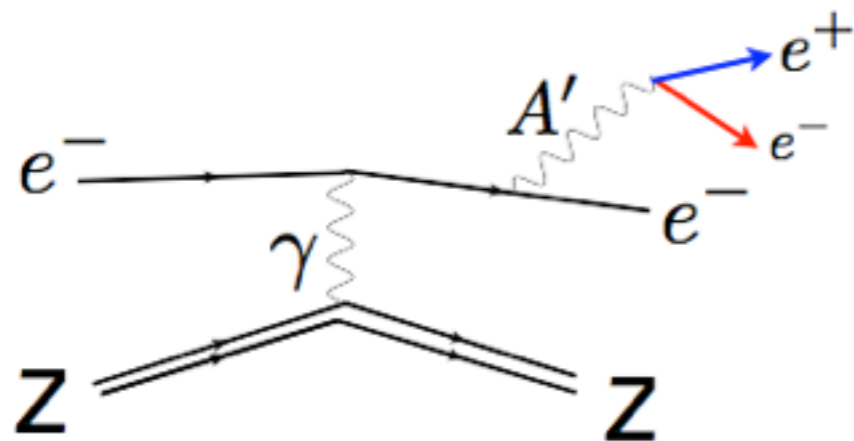
- Massive
 - when we say “dark” photon we typically also mean “heavy”
 - look for peak in the invariant mass spectrum



- Non-zero lifetime
 - some regions of parameter space will have decays that happen far from production target
 - backgrounds typically decay promptly



A' production & decay at a fixed target experiment



Production is analogous to bremsstrahlung:

$$\frac{d\sigma}{dx} \approx \frac{8Z^2\alpha^3\epsilon^2 x}{m_{A'}^2} \left(1 + \frac{x^2}{3(1-x)}\right) \mathcal{L}og$$

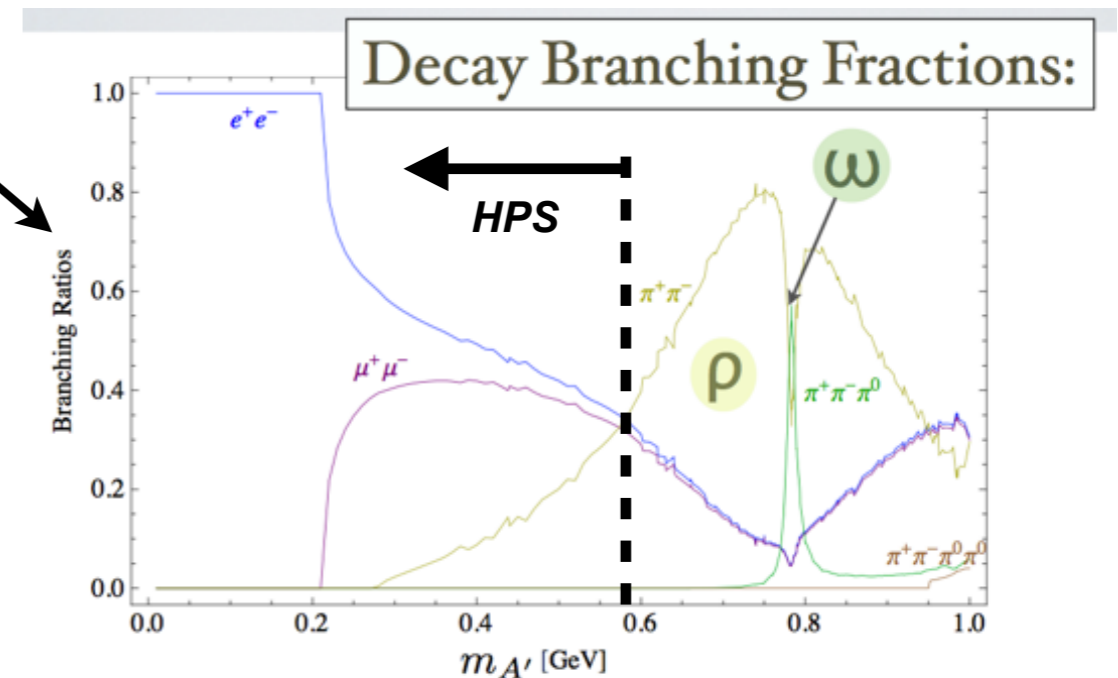
- prefers $x \sim 1$ (i.e. $E_{A'} = E_{beam}$)
- small angle emission dominates

A' **decays** back to charged SM fermions with BFs taken from $R(e^+e^- \rightarrow \text{hadrons}/e^+e^- \rightarrow \mu^+\mu^-)$

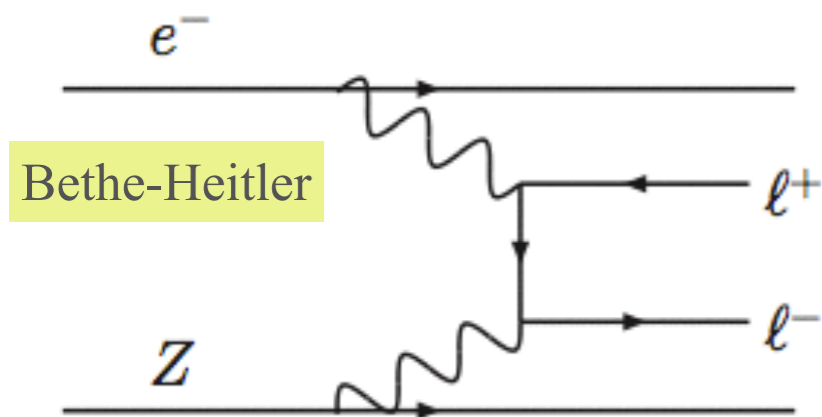
The decay length depends on $m_{A'}$ and ϵ :

$$\begin{aligned} \ell_0 &\equiv \gamma c\tau \simeq \frac{3E_1}{N_{eff} m_{A'}^2 \alpha \epsilon^2} \\ &\simeq \frac{0.8\text{cm}}{N_{eff}} \left(\frac{E_0}{10\text{GeV}}\right) \left(\frac{10^{-4}}{\epsilon}\right)^2 \left(\frac{100\text{MeV}}{m_{A'}}\right)^2 \end{aligned}$$

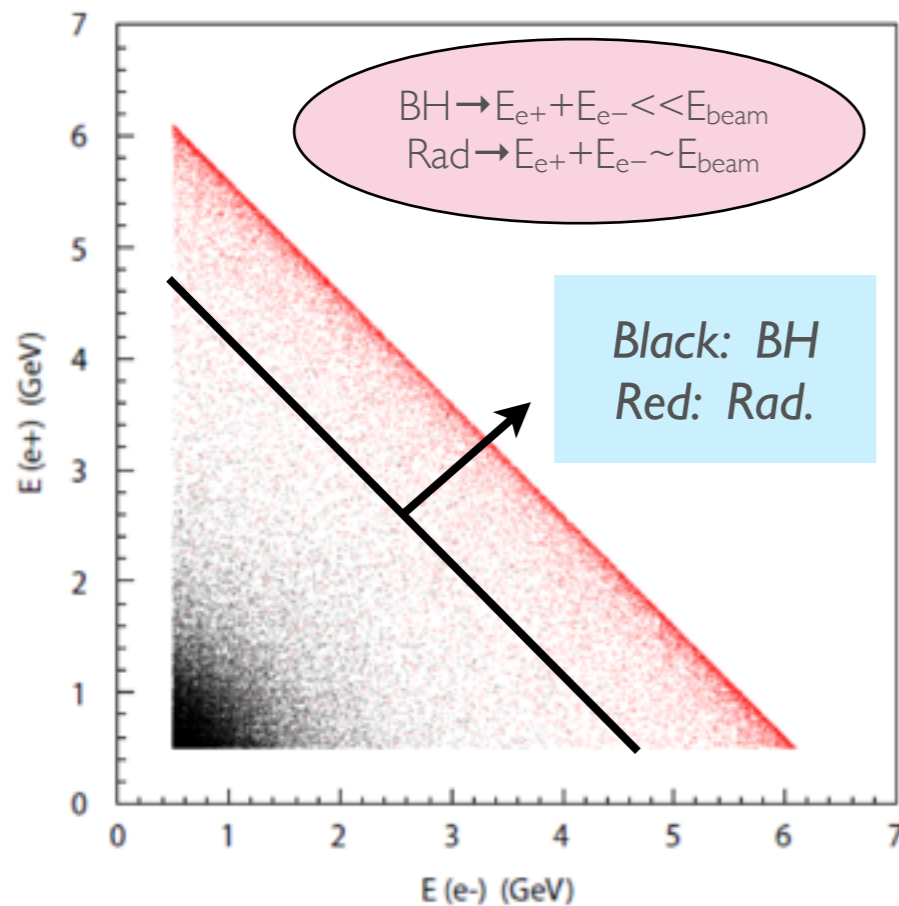
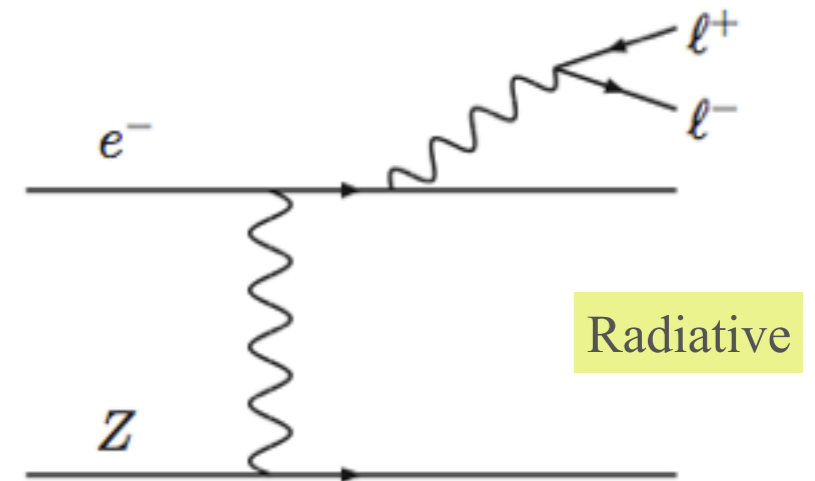
HPS is sensitive to A's with decays ~5-100mm and to e^+e^- final states.



A' backgrounds at a fixed target experiment



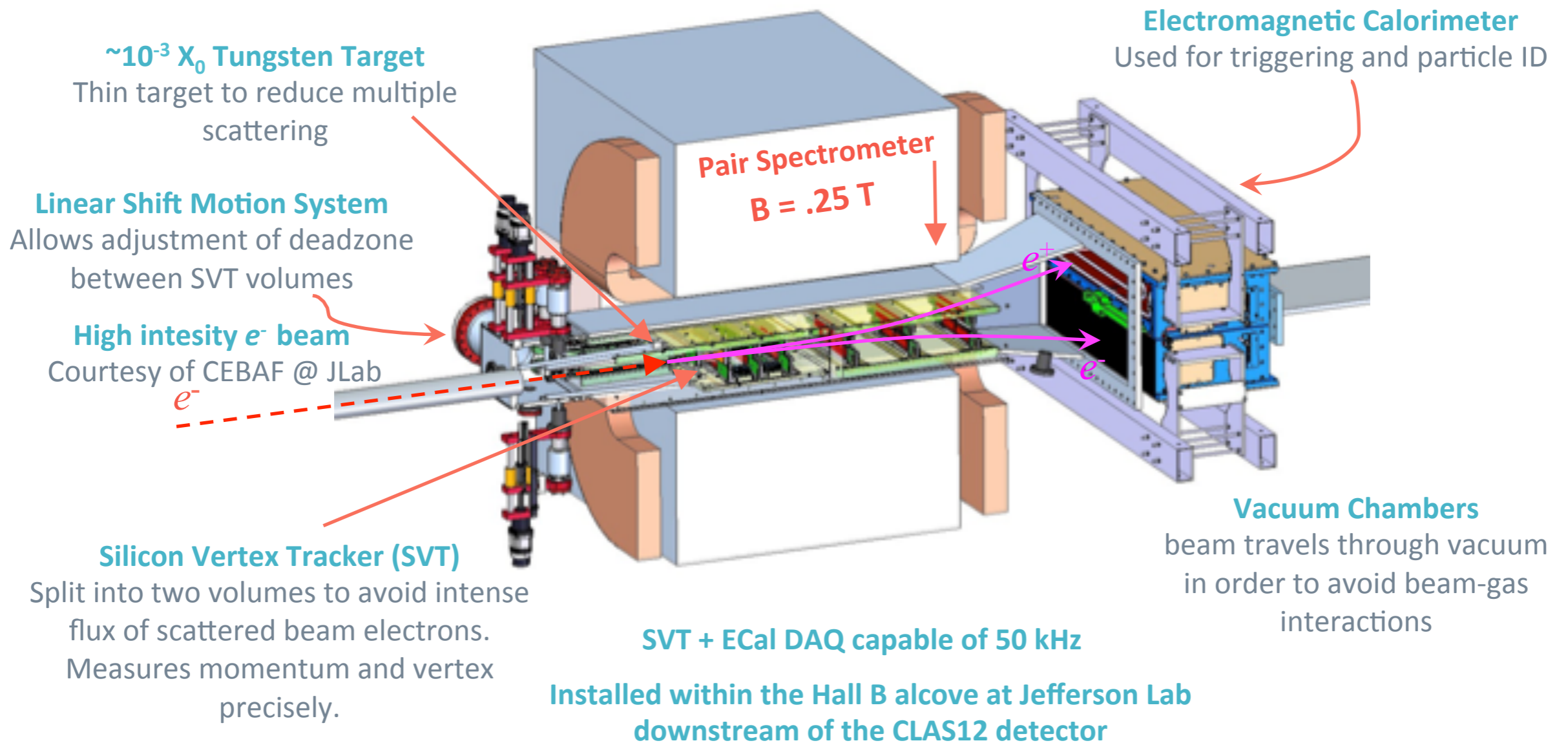
Two physics backgrounds, collectively known as "tridents"



- BH and Radiative cross-sections calculated by MadGraph at NNLO
- BH cross section is huge, but dominated by $E(e^+) + E(e^-) \ll E_{\text{beam}}$
 - this background is reducible, but still large ($\sim 10x$ radiative) after $E(e^+) + E(e^-) > 0.8 E_{\text{beam}}$
- Radiative tridents have the same kinematics as A' decays...only invariant mass & decay vertex can resolve these two
- All trident events decay promptly!

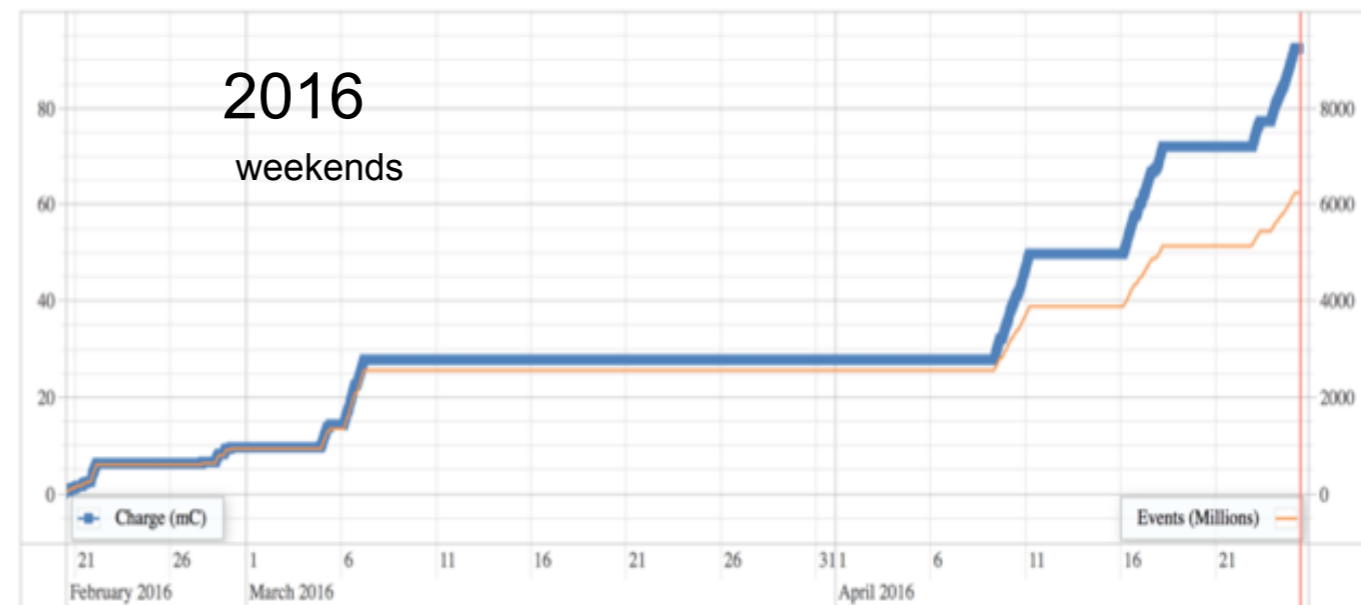
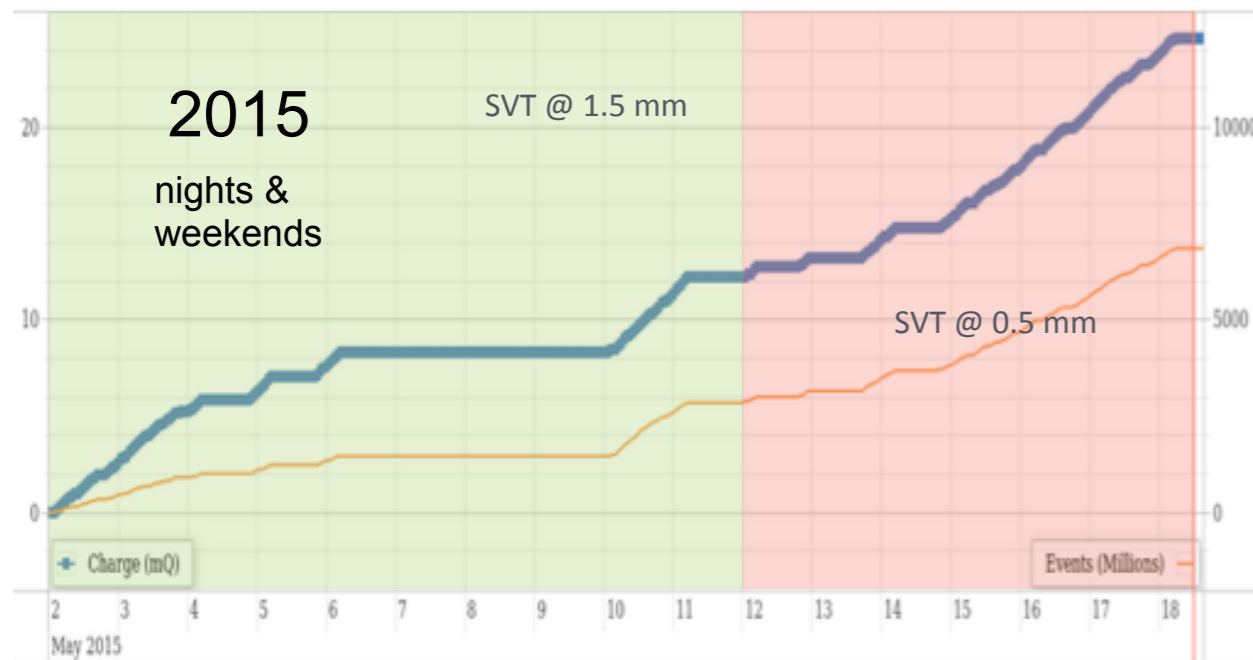
...there is one other background, which I will discuss shortly!

HPS: Heavy Photon Search Experiment



HPS Engineering Runs

- HPS has taken two runs, technically defined engineering runs but which we got some physics quality data
- In May 2015, took **1.7 PAC days** of data each with SVT at 0.5mm (nominal) and 1.5mm **@ 1.05 GeV**
- In spring 2016, ran **@ 2.3 GeV** and took **5.4 PAC days** all with SVT at 0.5mm
- **Analysis presented here covers 1.7 PAC days @ 1.05 GeV with SVT at 0.5 mm**
- HPS has been officially approved (no more “engineering runs”) for 180 PAC days
- We hope to have our next run in 201X (??? fingers crossed ???)



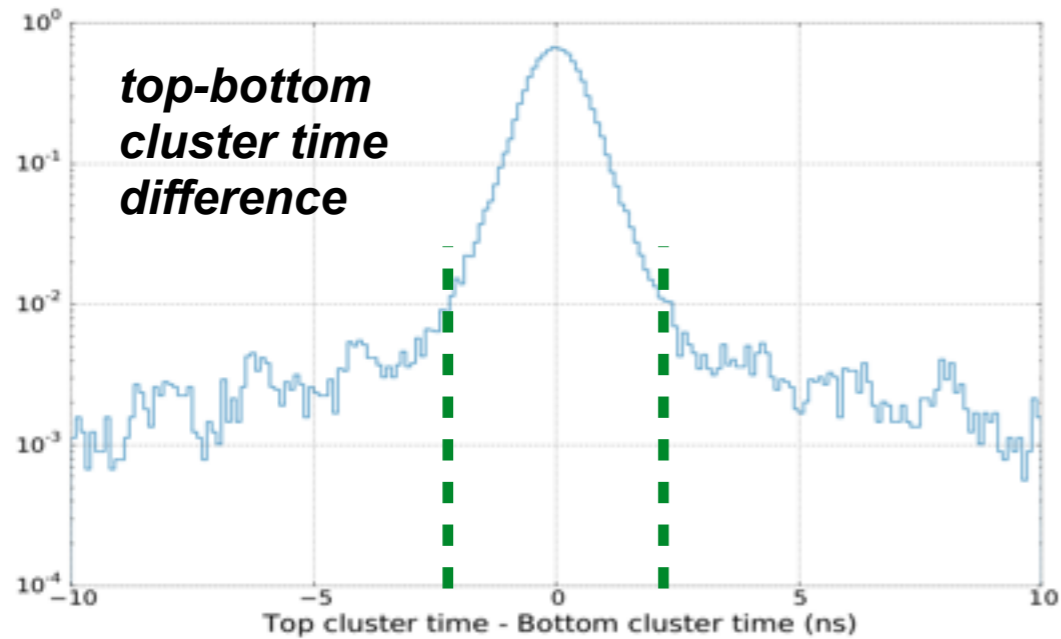
The HPS Bump-hunt Analysis

Ratio of A' to radiative rate in $\delta m \sim A'$ width

$$\frac{d\sigma (e^- Z \rightarrow e^- Z (A' \rightarrow l^+ l^-))}{d\sigma (e^- Z \rightarrow e^- Z (\gamma^* \rightarrow l^+ l^-))} = \frac{3\pi\epsilon^2}{2N_{eff}\alpha} \frac{m_{A'}}{\delta m}$$

- As mentioned, A' production related to “radiative trident” production... we use this relation to relate our extracted signal yields to $m(A')$ vs ϵ space
- Ingredients going into this:
 - the fraction of events in our final selected sample that are radiative events
 - use MG5 to simulate full-diagram and radiative-only tridents and then simulate detector simulation, use nominal reconstruction and cuts
 - experimental mass resolution
 - simulated with MG5-generated A' events at range of masses
- Absolute rates are not used directly, which cancels a lot of potential systematic effects!

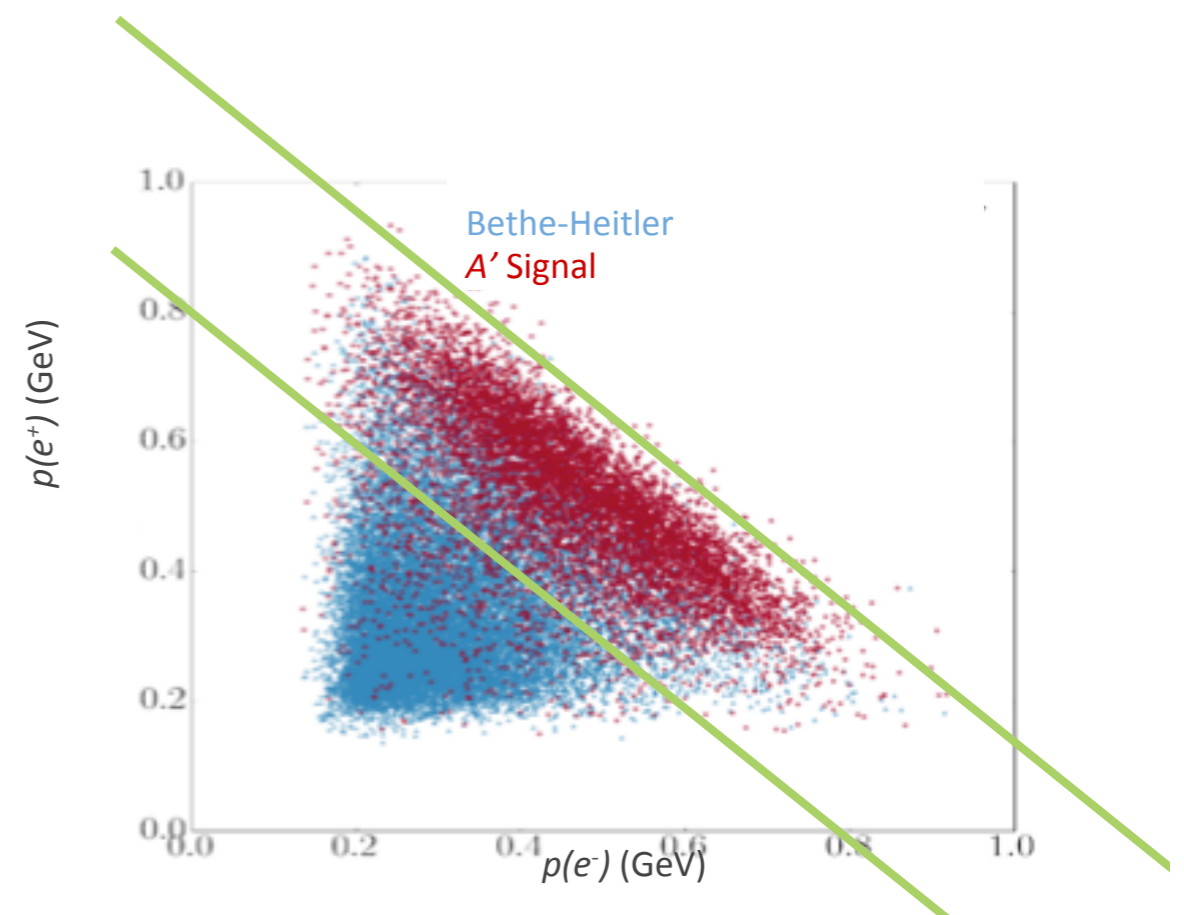
Bump-Hunt event selection



- We generally to to keep the KISS approach, particularly for this first analysis:
- loose track reconstruction and track-ECal matching cuts
 - ~ 2 sigma cut on relative ECal cluster times
 - $< 1\%$ accidental e^+e^- pairs

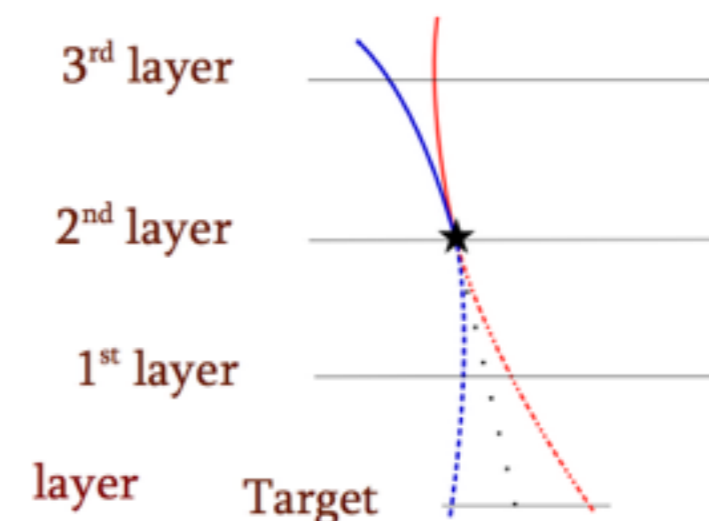
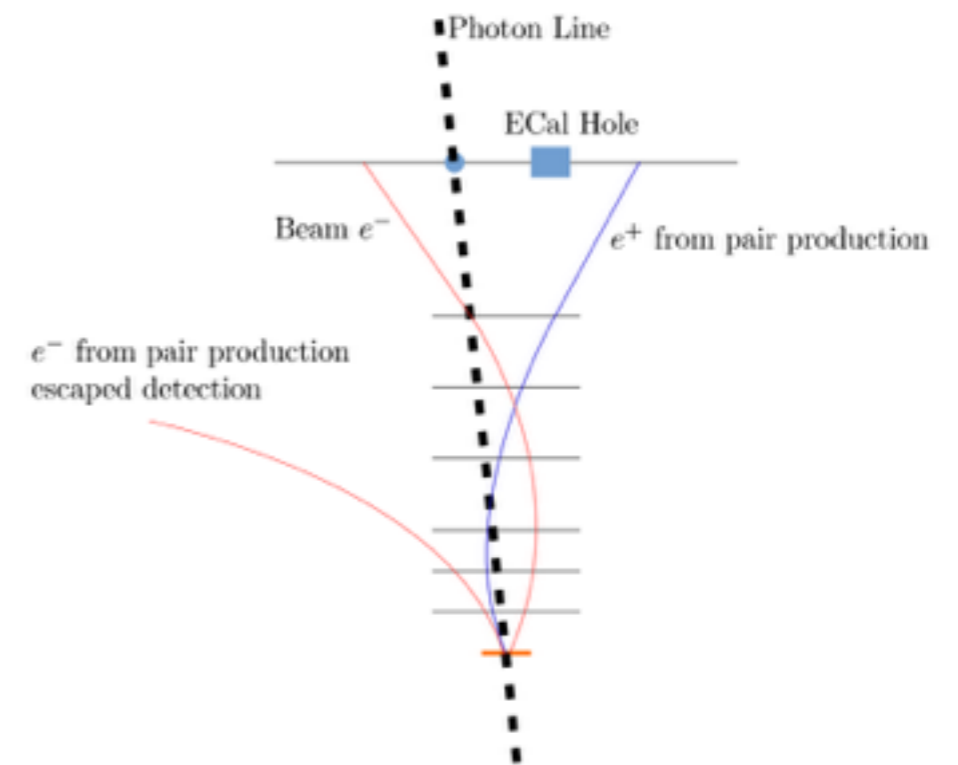
... also require that the e^+e^- momentum sum be greater than $0.8 \times E_{\text{beam}}$

This is the single-best discriminant against BH-like trident background

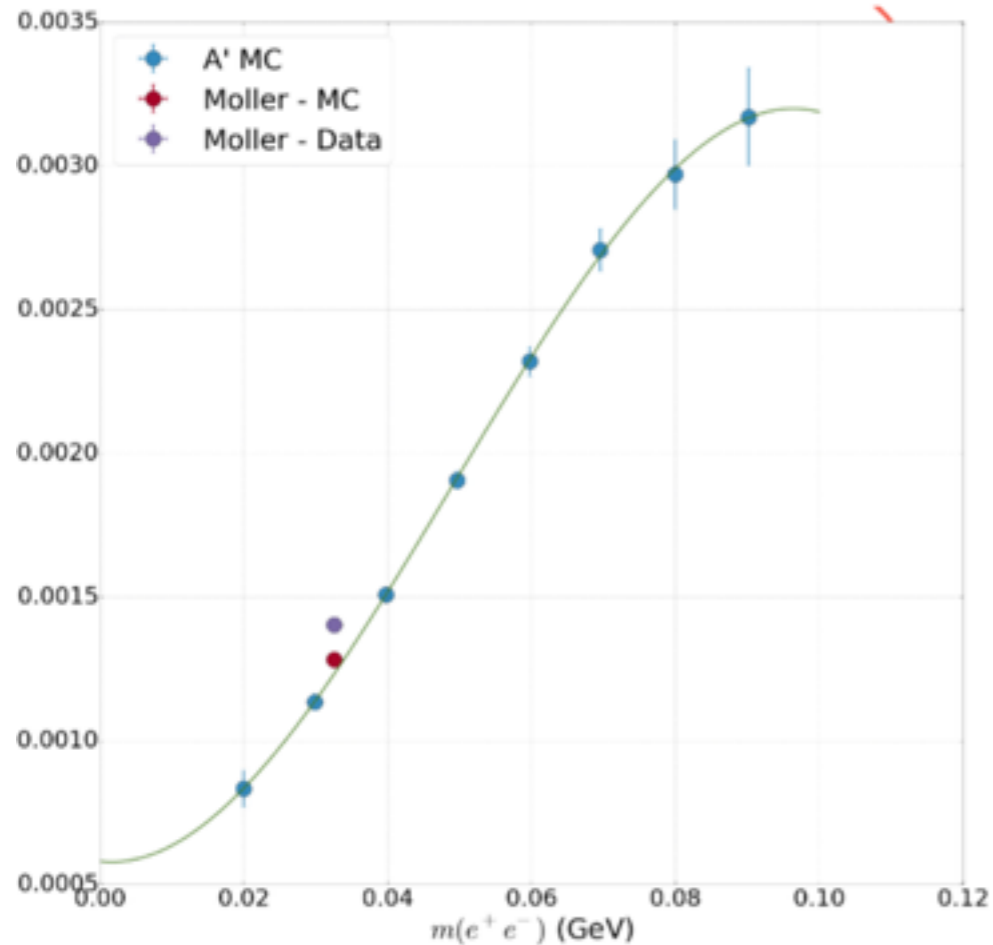


Reducing WAB gamma converted events

- One thing we missed in proposal: wide-angle Brems (WABs)!
- How could we miss this? None of the usual event generators simulate it correctly!
 - because it is expensive and the rate is “very small” (3-body)
- WAB: $e^- Z \rightarrow e^-_{\text{large } \theta} \gamma_{\text{large } \theta} Z$
 - then, the γ converts in L1 or L2 SVT with the positron forward
 - $E(e^-e^+) \sim E_{\text{beam}}$ just like radiative tridents
- We can reduce this background quite a bit
 - require positron hits in L1
 - requirements on positron DOCA and $p_T(e^-e^+)$



Mass resolution

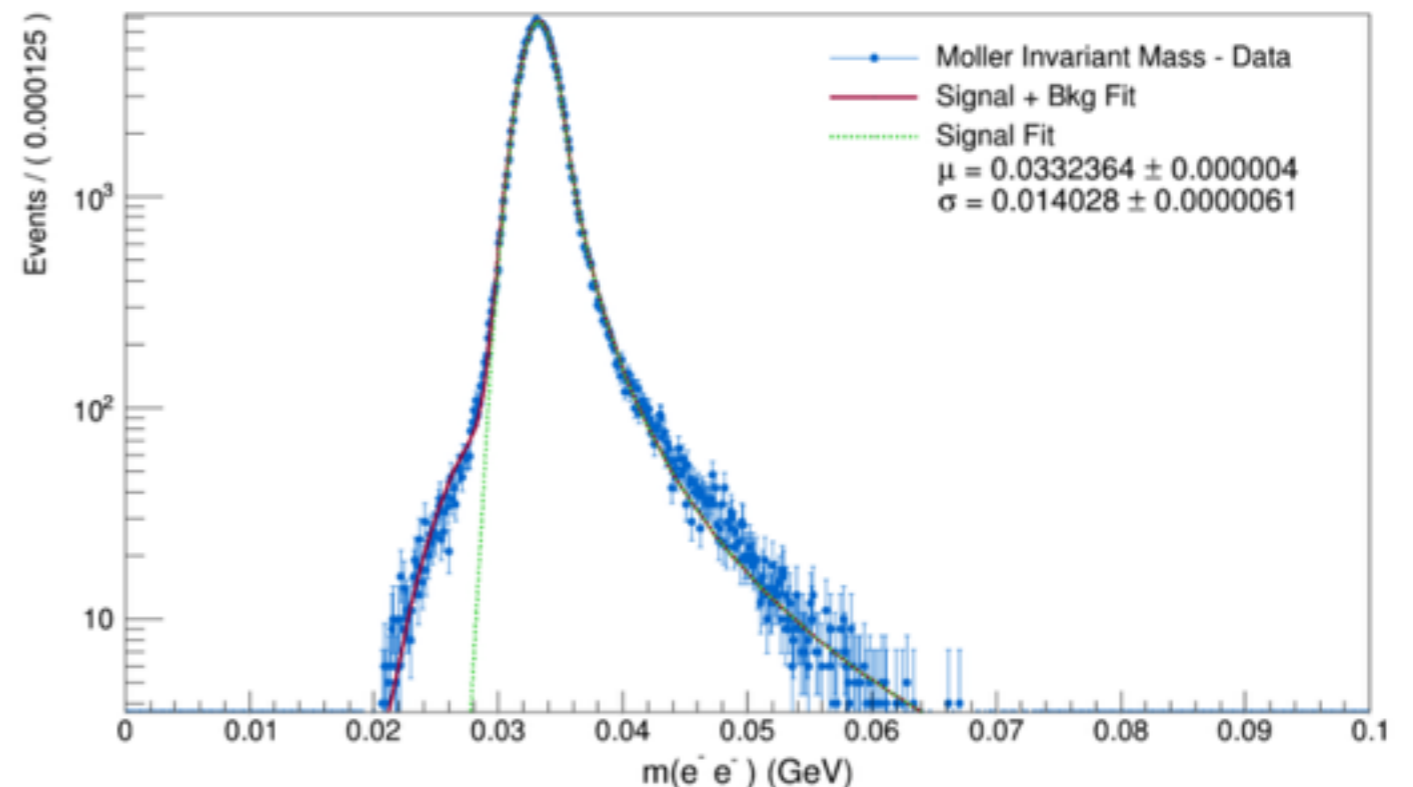


Use mass resolution from MC at various A' mass hypotheses ...how do we calibrate?

Luckily, at 1.05 GeV, we have a large number of Moller-scattered events in our detector ($e^-e^- \rightarrow e^-e^-$) ...these events have defined invariant mass (and well determined p vs angle)

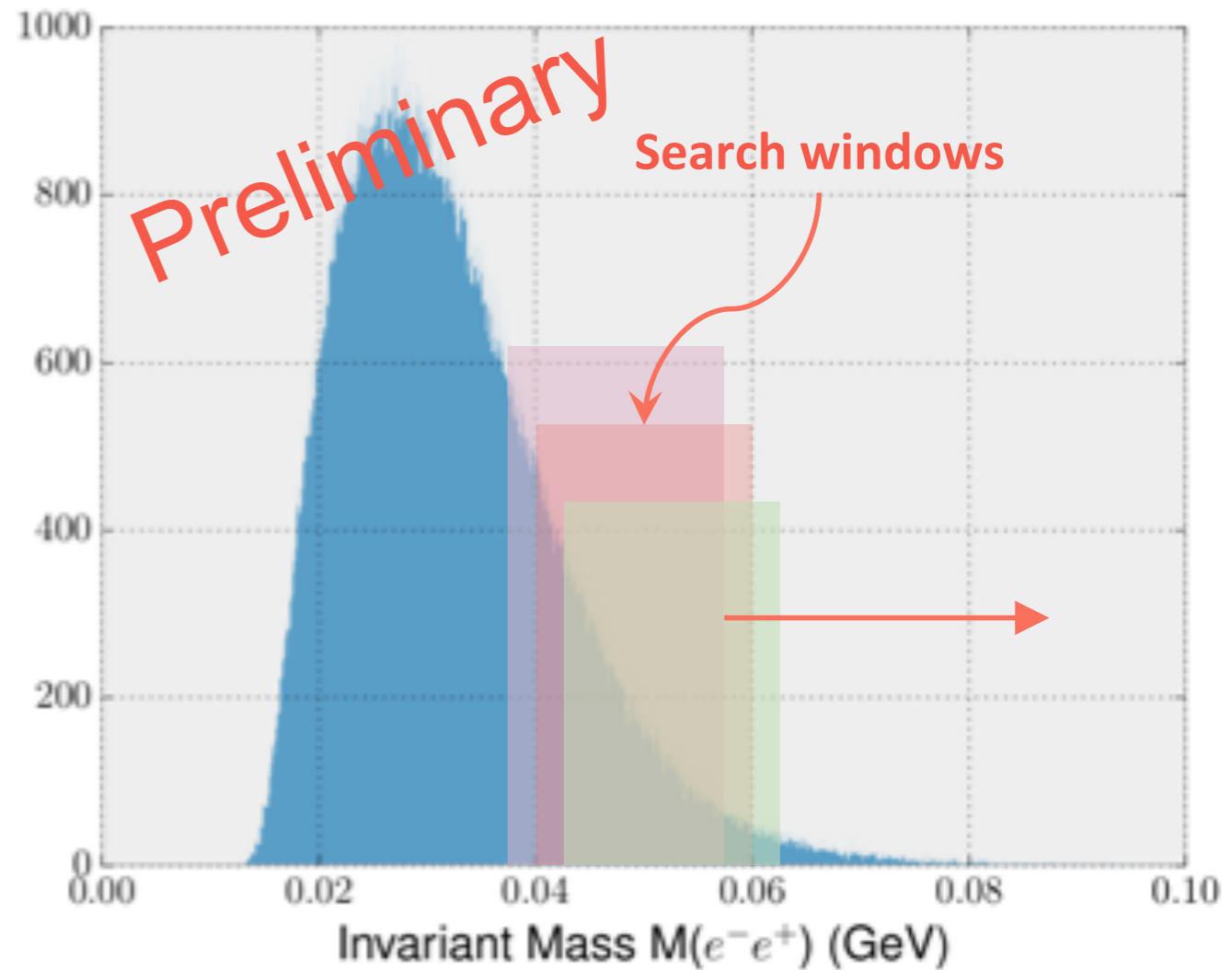
We see ~10% difference between Moller data and MC

In the bump-hunt fits, actually use the MC resolution scaled up by this 10%

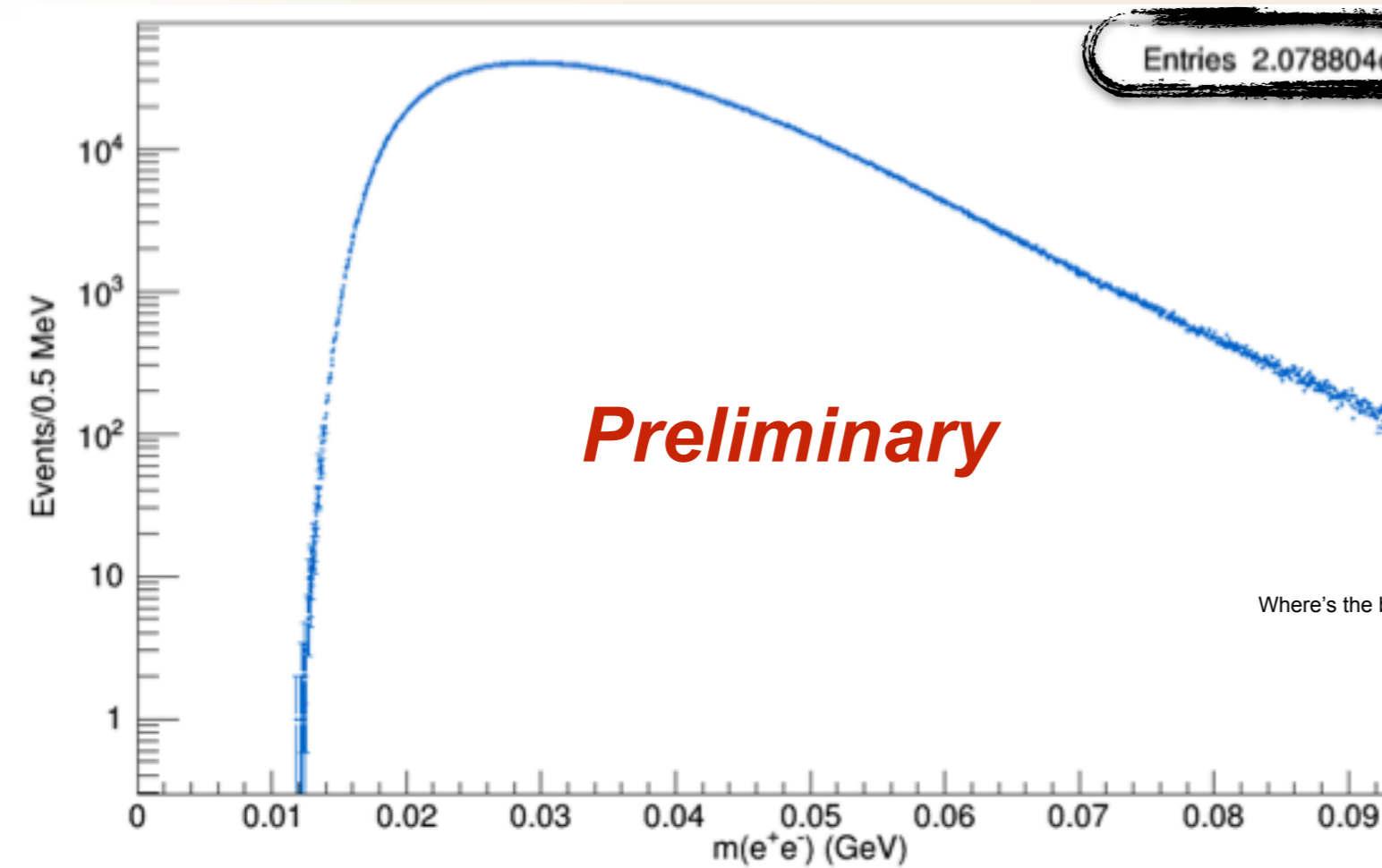


Mass spectrum scan

- Scan through the mass spectrum and search for a bump!
- Background: P7, parameters floating
- Signal: Gaussian, mass fixed (scanned), sigma fixed to resolution
- Search window is “many times” the resolution
- Step size: 1 MeV
- “local” signal significance found using likelihood ratio of signal+bkg vs bkg only

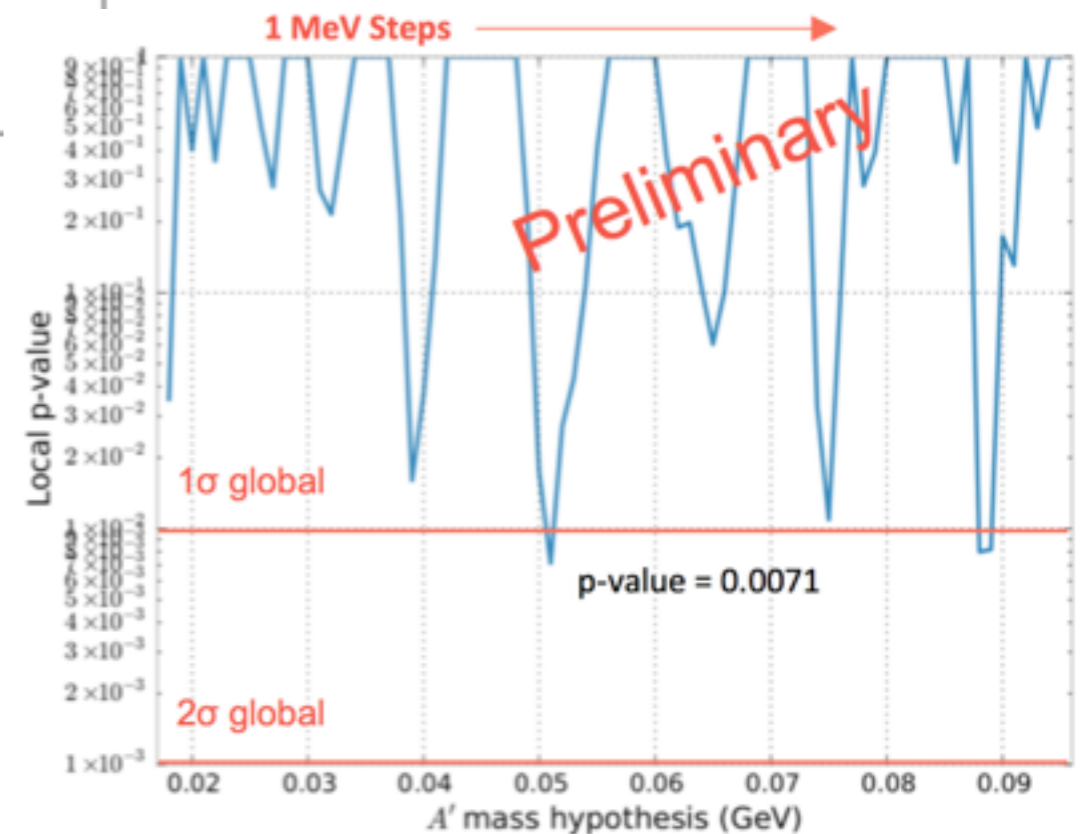


HPS 2015 Engineering Run invariant mass spectrum



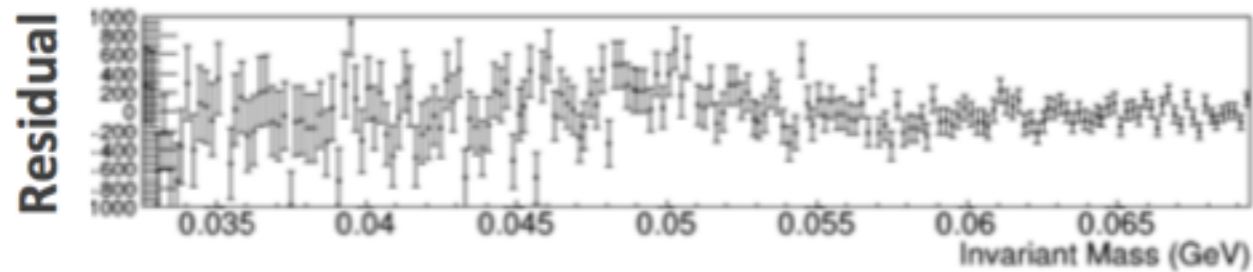
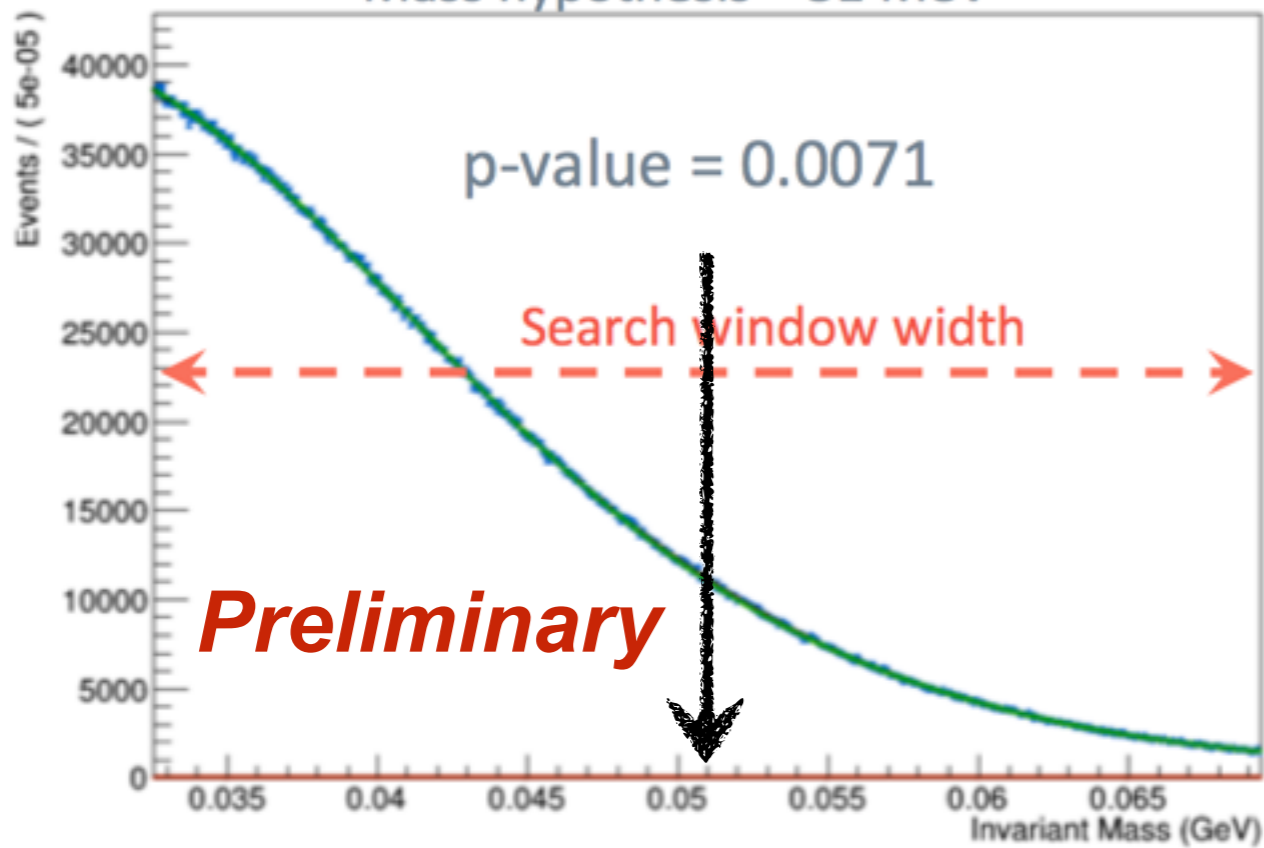
in 1.7 days!

Where's the bump?

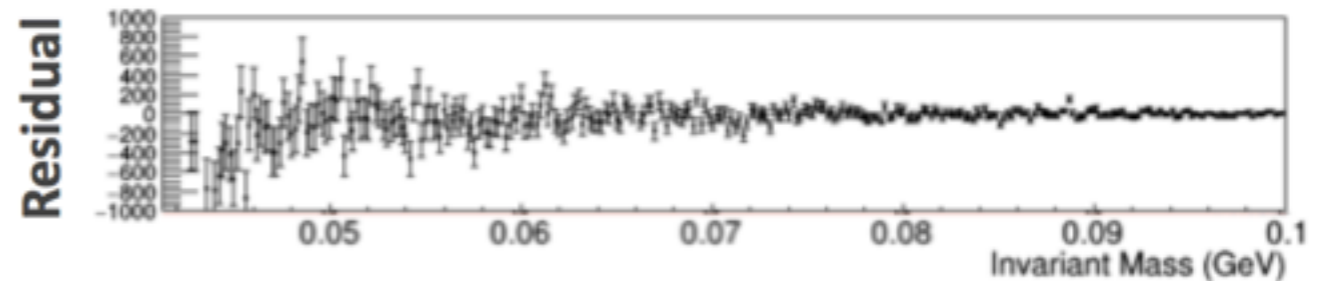
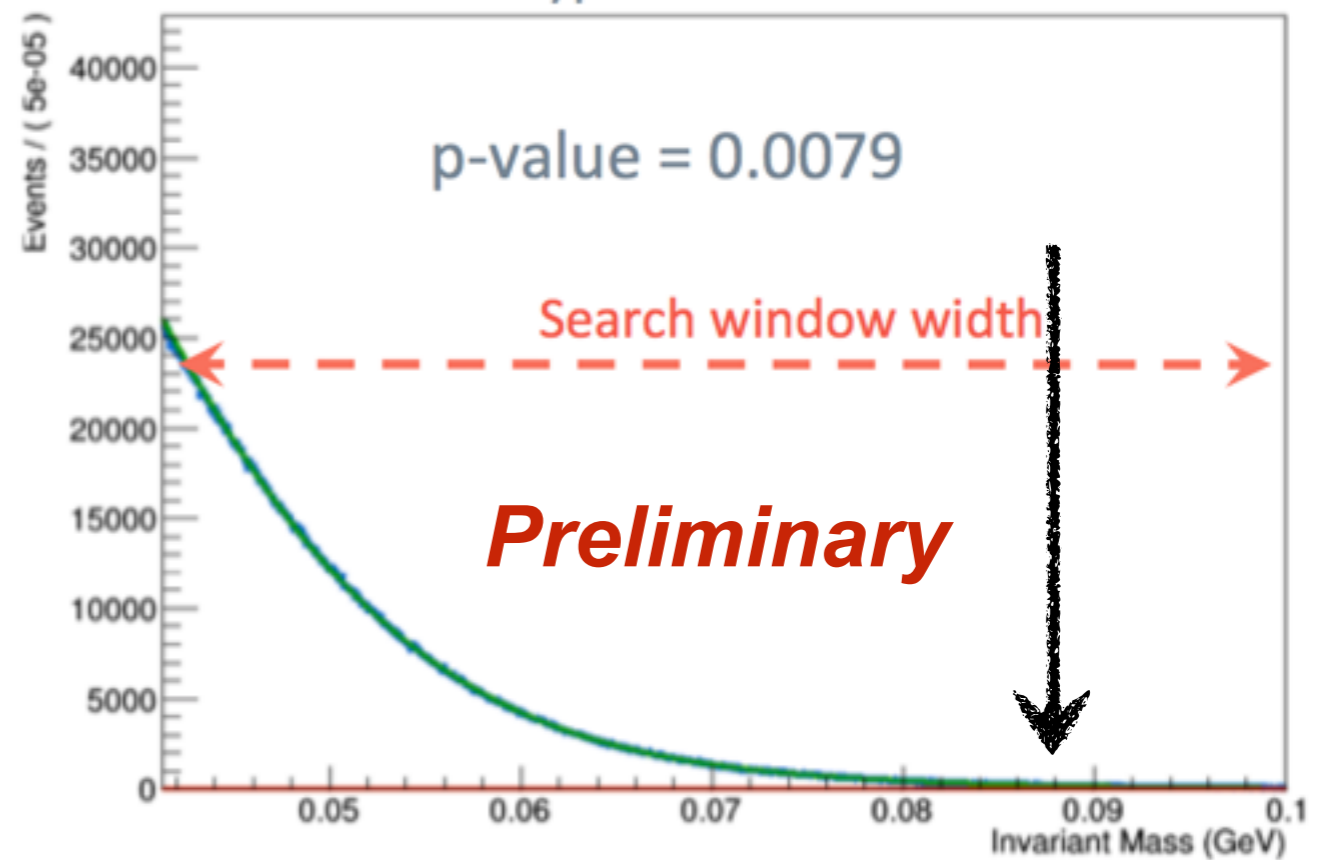


HUGE BUMPS!

Mass hypothesis = 51 MeV

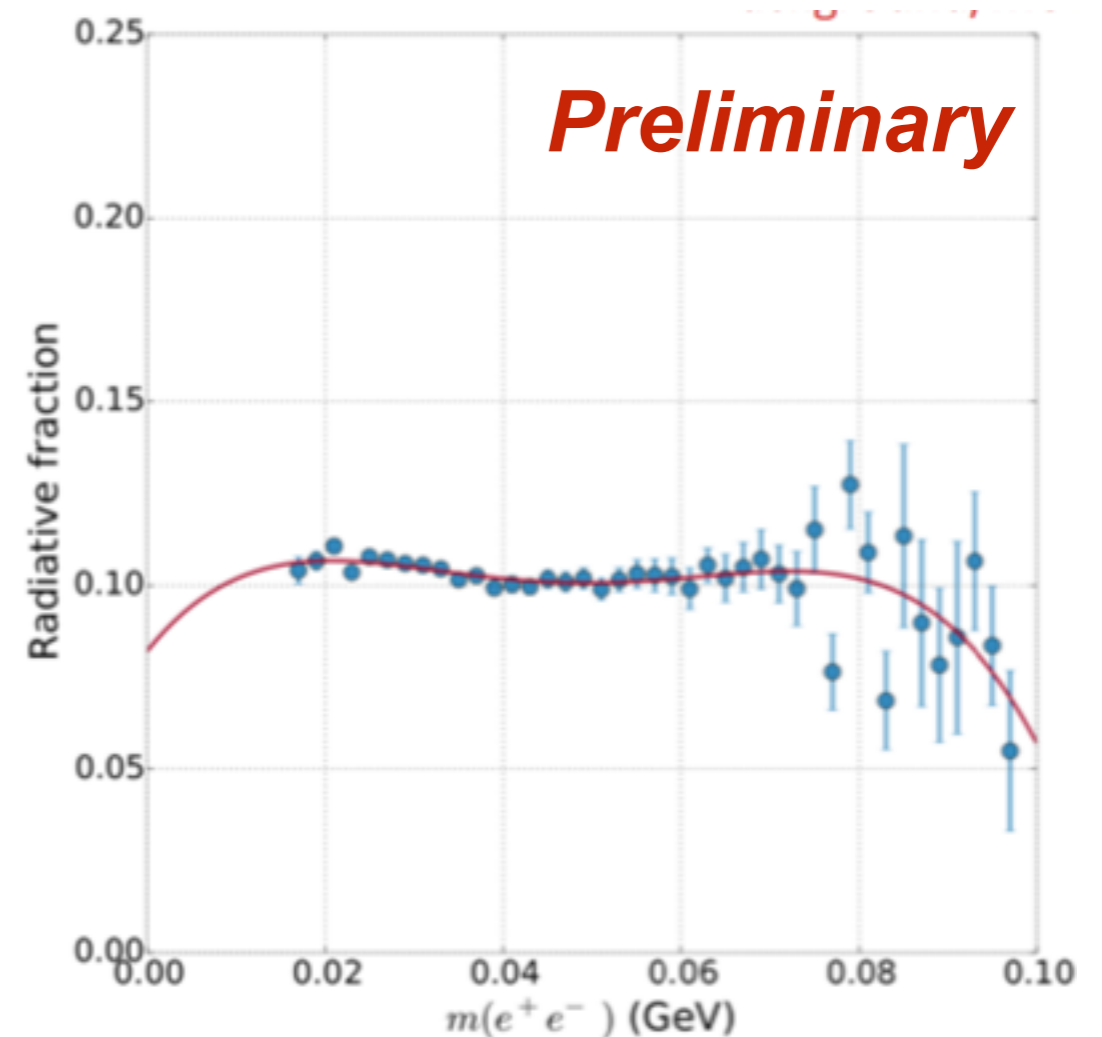
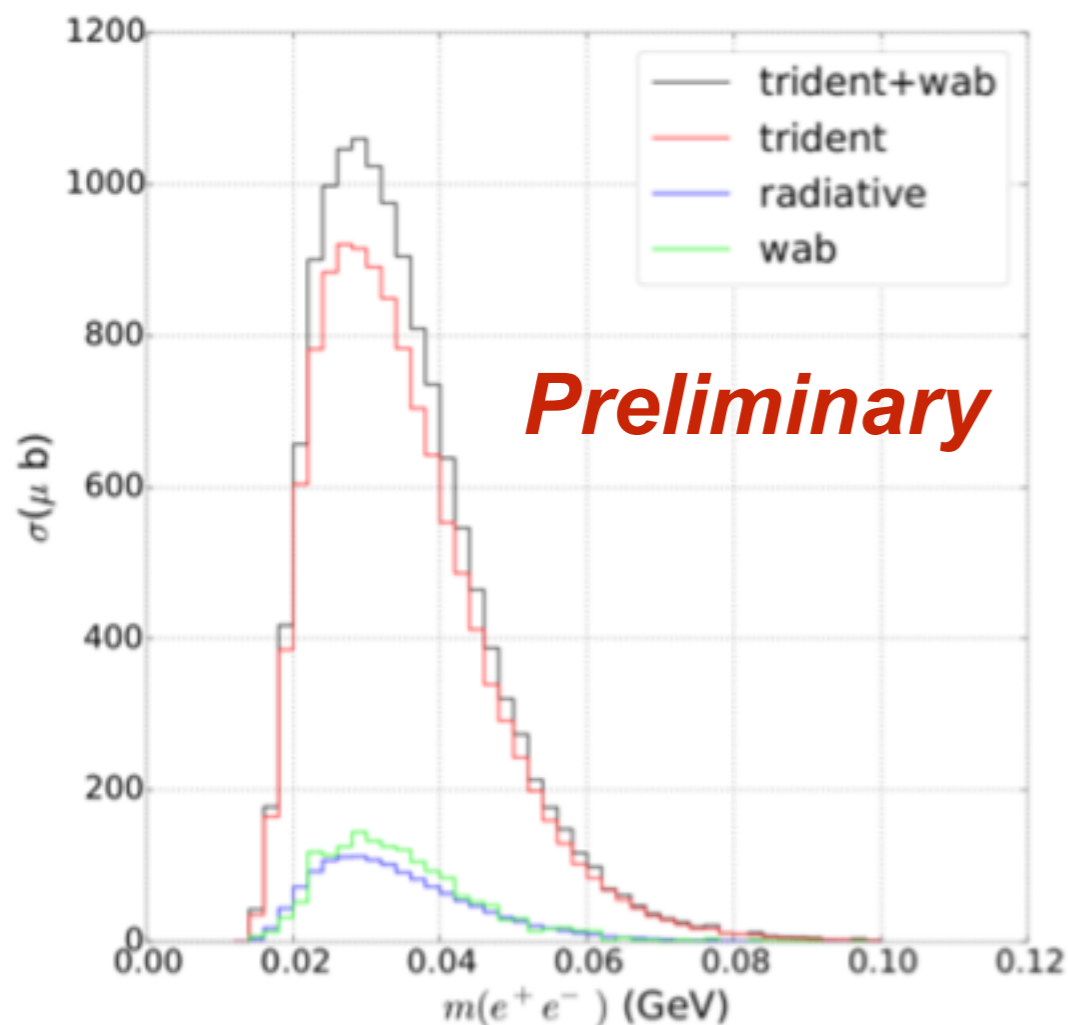


Mass hypothesis = 88 MeV

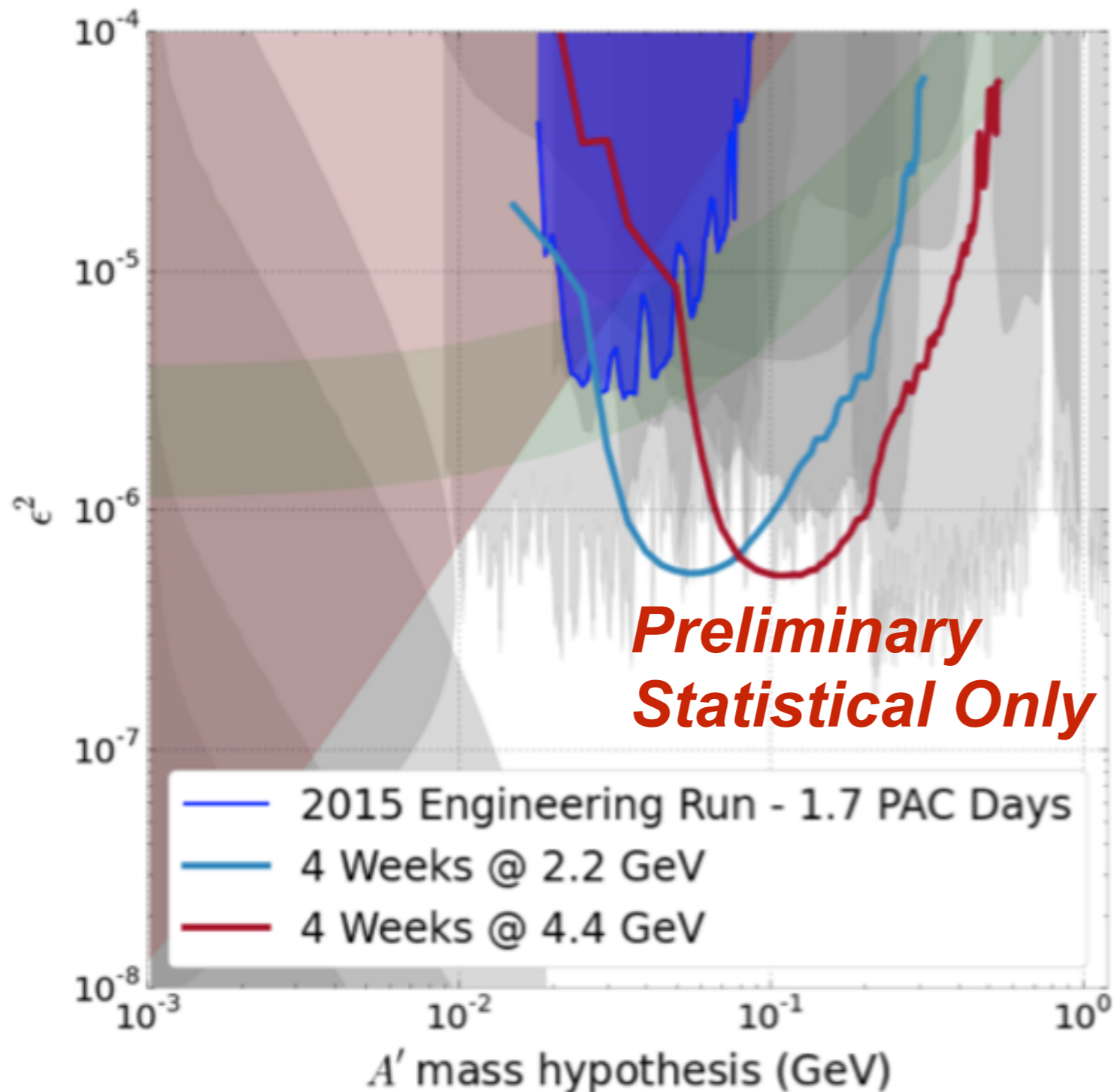


Determining the Radiative Fraction

- As mentioned, fraction of radiative events determined from MG5 + full detector simulation
- We include WAB events in our definition of “radiative fraction”



Beyond the HPS 1.05 GeV Bump-Hunt Upper Limit

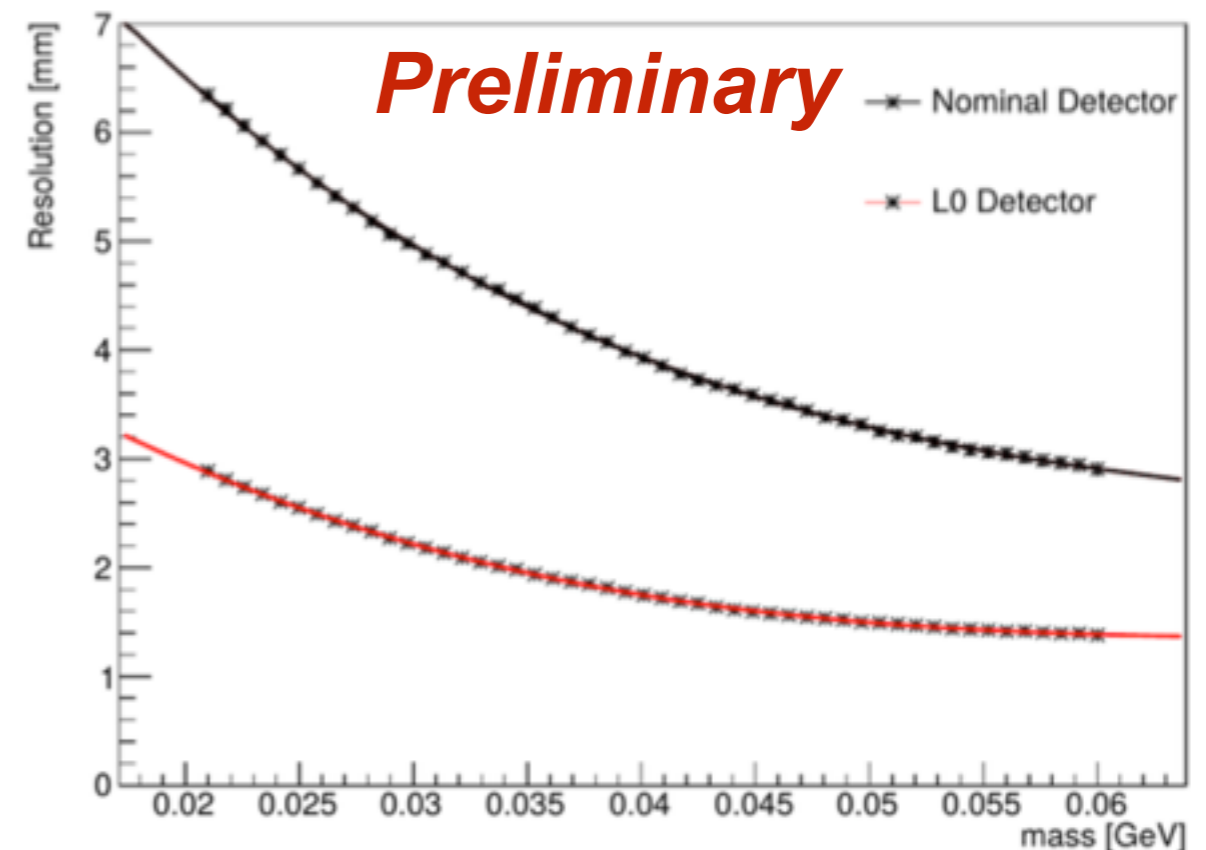
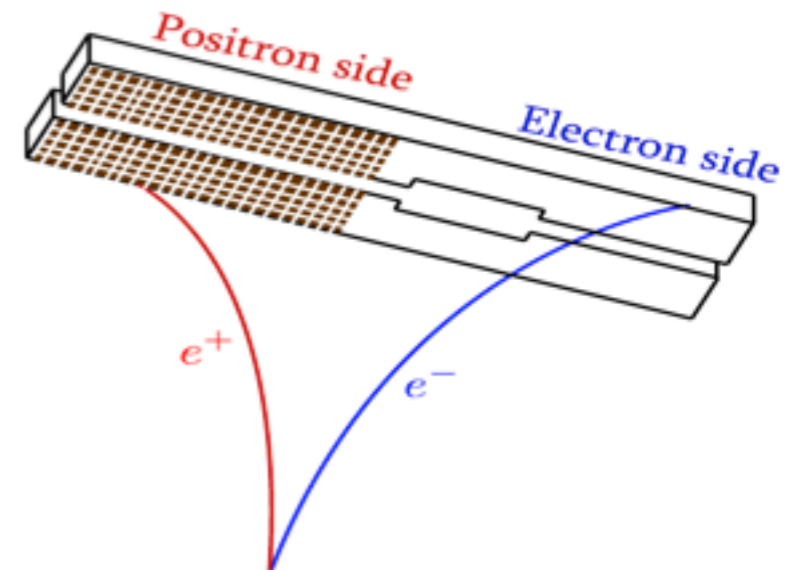


The limits for 2.2 & 4.4 GeV are simply scaled from the 1.7 days @ 1.05 GeV

Wait...why isn't this as good as your proposal?

What's next? Possible upgrades!

- We are finding out proposal reach was too optimistic...a few reasons:
 - somebody (me) got the acceptance the ECal wrong...didn't account for the electron hole
 - this cuts down low-mass acceptance by a lot!
 - didn't account for WABs (smallish effect)
 - for displaced vertices (doesn't effect bump-hunt results), assumed flat efficiency vs decay distance (up to layer 1)
- To address some of these, two proposed upgrades:
 - add a hodoscope in front of positron-side of the ECal...trigger on just positrons
 - adds a lot of low-mass acceptance, including in the ECal hole
 - add an SVT layer at 5cm (layer 0)
 - reduces vertex position resolution by $\sim x2$



Final Thoughts/Take Home Message

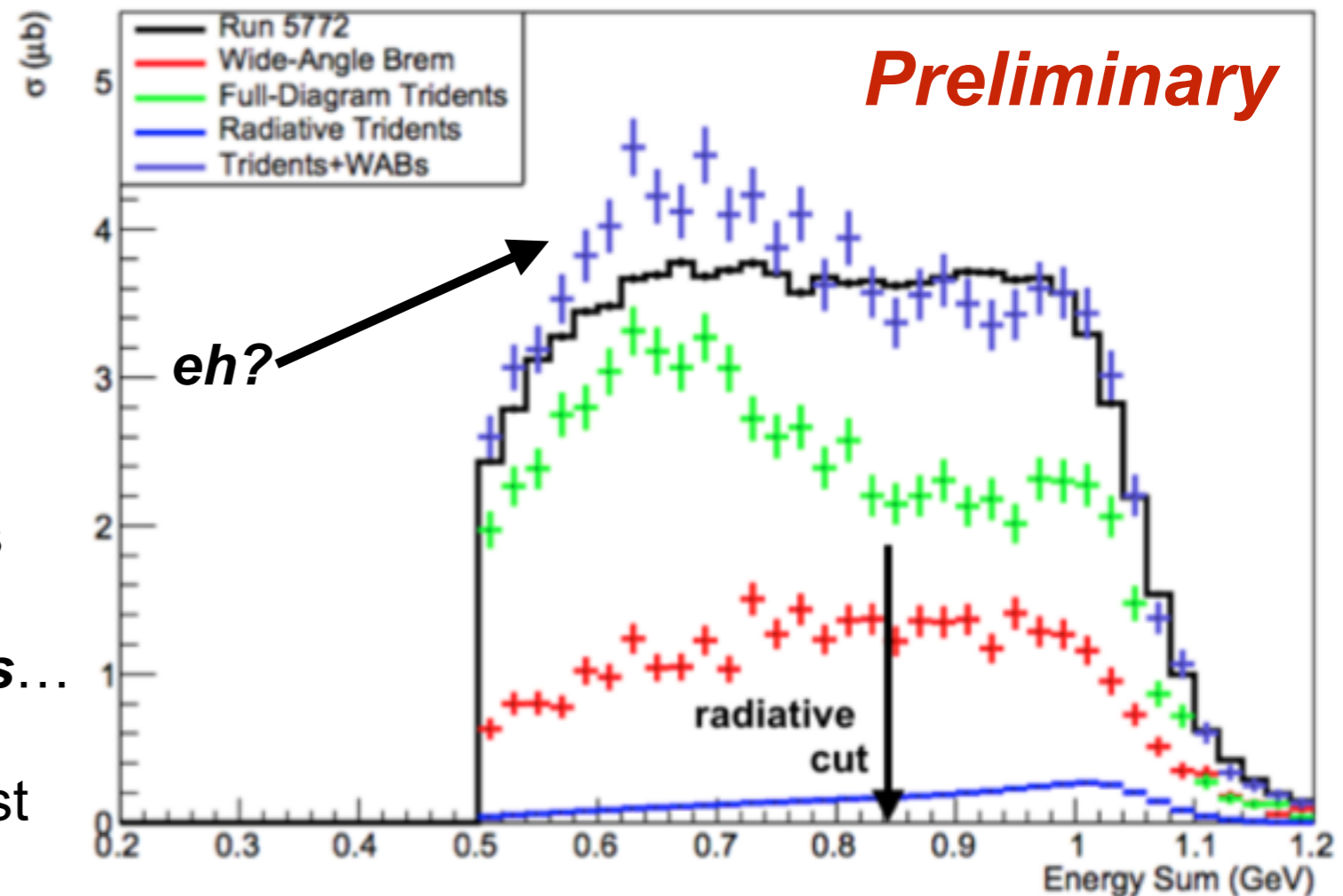
- Nobody knows what dark matter is...really, nobody
- Light, thermal dark matter is really a possibility and it has not been probed in detail
 - LDM *requires* a light force carrier
- Dark/Hidden/Secluded sector physics (where LDM could live) is pretty much mainstream now
 - 10 years ago it was considered a wacky, far out idea...now we have a huge community, almost every experiment has folks working on it, and DOE is hosting workshops about what the field should do and **build** in the next 10 years
- HPS **finally** has a public result, though no new territory yet
 - Submittal of Eng. Run bump-hunt results *very soon*
 - we live & die, as an experiment, through our displaced vertex reach and we hope to have an upgraded expected reach, based on what we observe in the Eng. Run datasets, in the next few weeks/months
 - sneak peak, 1 week of beam time will not be enough ... good thing we have 180 PAC-days

QED Tridents @ 1.1 GeV

signal yield limit \rightarrow (m, ϵ) limit requires the fraction of radiative events in the sample...**so you better believe the events you're seeing are tridents!**

we did detailed studies of the e^+e^- rates and distributions measured in our detector...**and found lots of mysteries...**

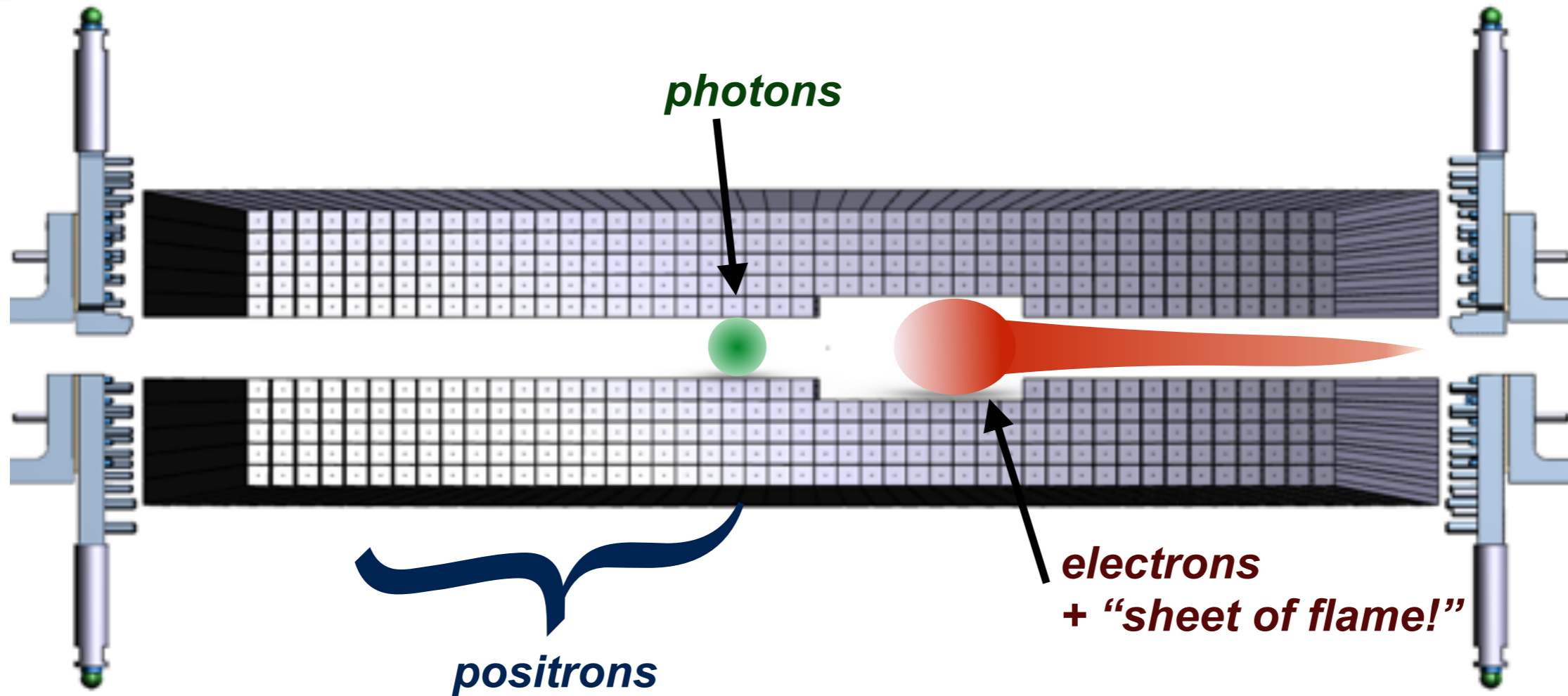
- event generators (MG4 vs MG5 vs first principles
 - alpha at 1GeV \neq alpha (80GeV)
- WABs
- efficiency and acceptance effects...
- **Don't trust anything, ever**



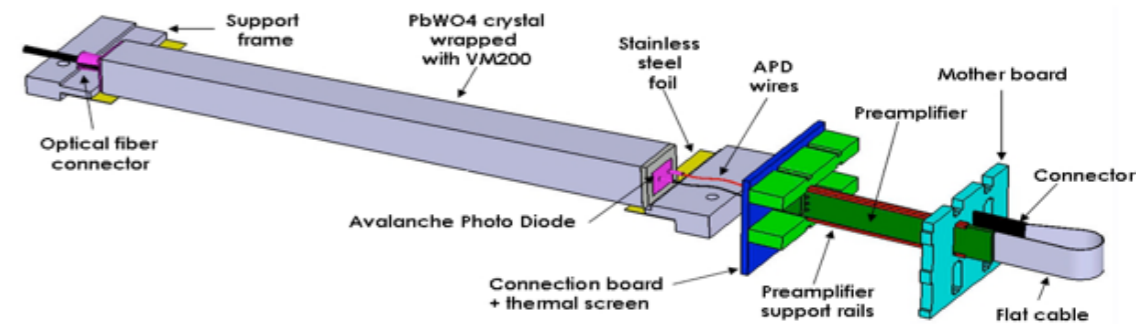
Lots of improvements made and now we believe our trident sample

Understanding this is a big part of what made it take so long to get this result out...

The HPS ECal



- ECal is 442 PbWO_4 crystals with APD readout
- Digitizer is JLAB-designed 250 MHz flashADC
- Used to trigger on e^+e^- pairs...cluster in top +bottom/left+right quadrants with $\sim 8\text{ns}$ resolution



HPS Silicon Vertex Tracker (SVT)

SVT is the key ingredient! Needs to measure momenta & vertex pairs with extreme purity in a busy environment. **Require low material & very fast.**

Si μ strip sensors

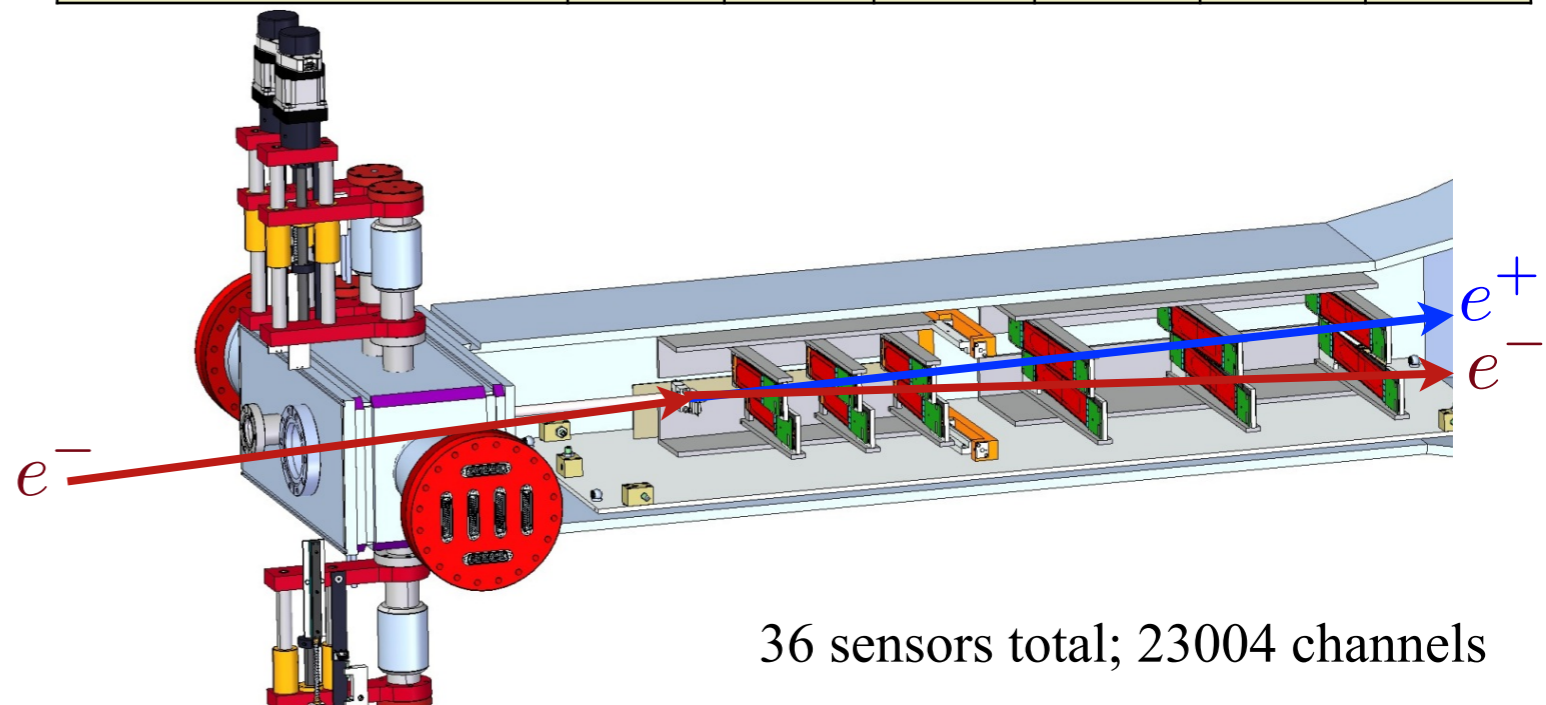
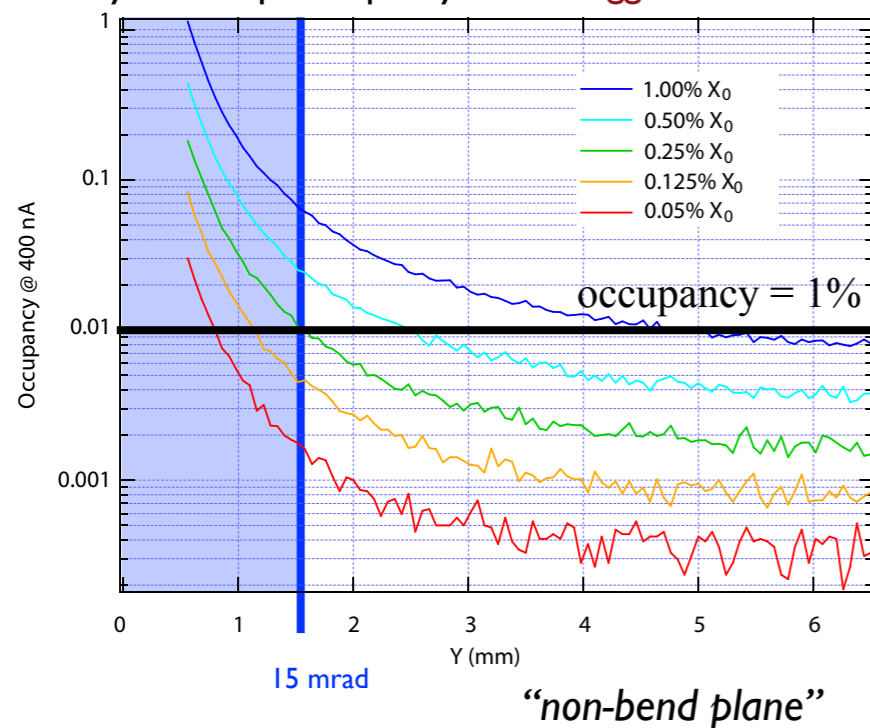
- Rad hard, thin (320 μ), 60 μ /30 μ readout/sense pitch & \$\$\$=Free (from RunIIb)

APV25 readout chip

- S/N>25 & ~2ns timing resolution

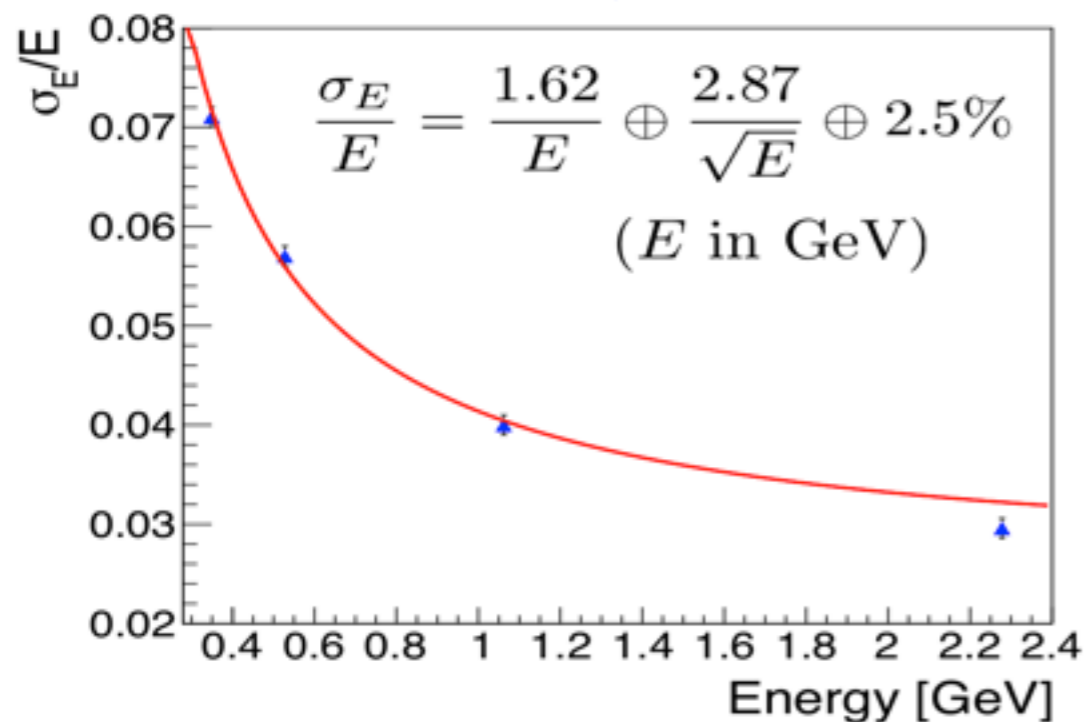
	Layer 1	Layer 2	Layer 3	Layer 4	Layer 5	Layer 6
z position, from target (cm)	10	20	30	50	70	90
Stereo Angle (mrad)	100	100	100	50	50	50
Bend Plane Resolution (μm)	≈ 60	≈ 60	≈ 60	≈ 120	≈ 120	≈ 120
Non-bend Resolution (μm)	≈ 6	≈ 6	≈ 6	≈ 6	≈ 6	≈ 6
# Bend Plane Sensors	2	2	2	4	4	4
# Stereo Sensors	2	2	2	4	4	4
Dead Zone (mm)	± 1.5	± 3.0	± 4.5	± 7.5	± 10.5	± 13.5

Layer I strip occupancy / 8 ns trigger window

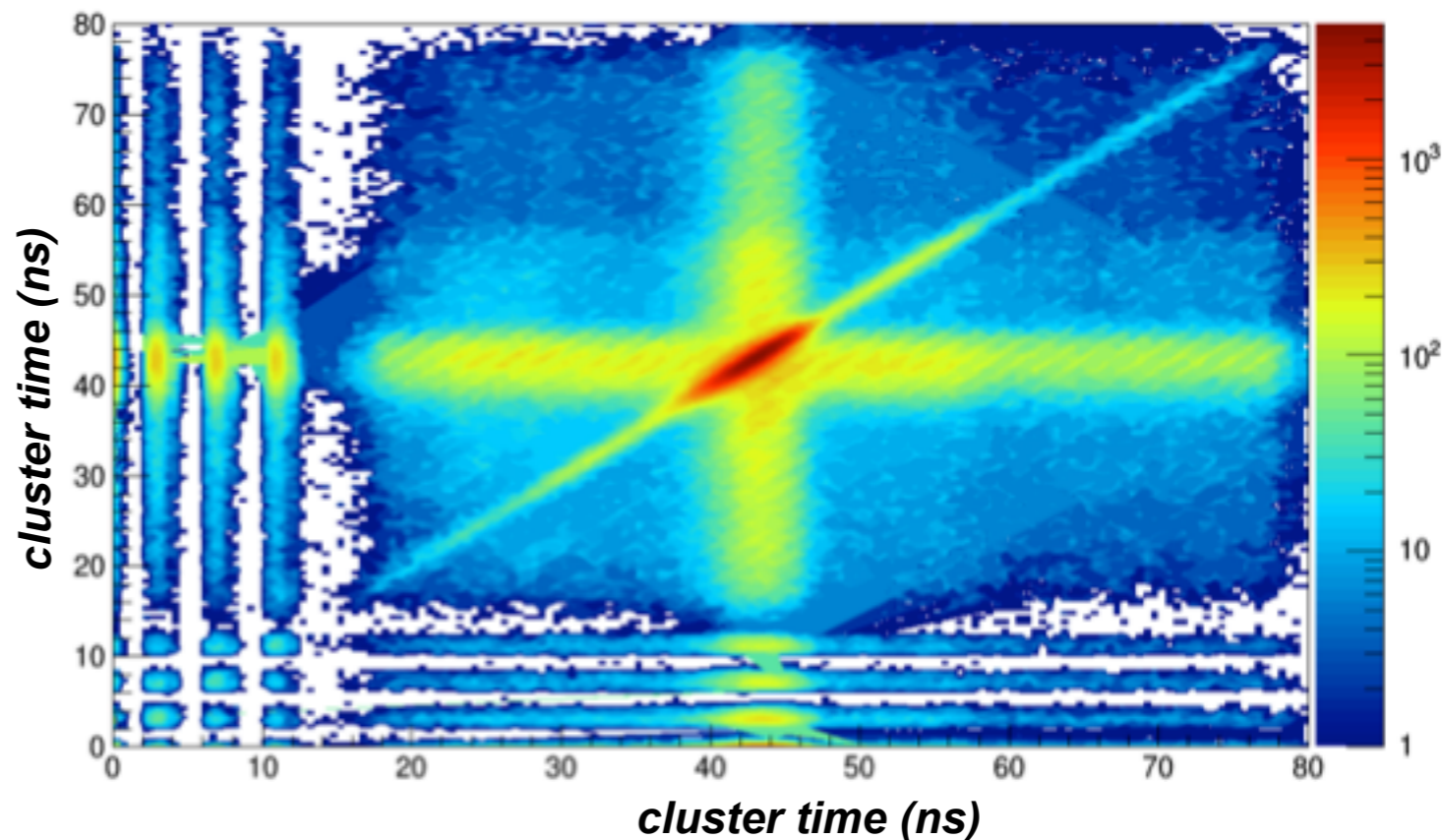
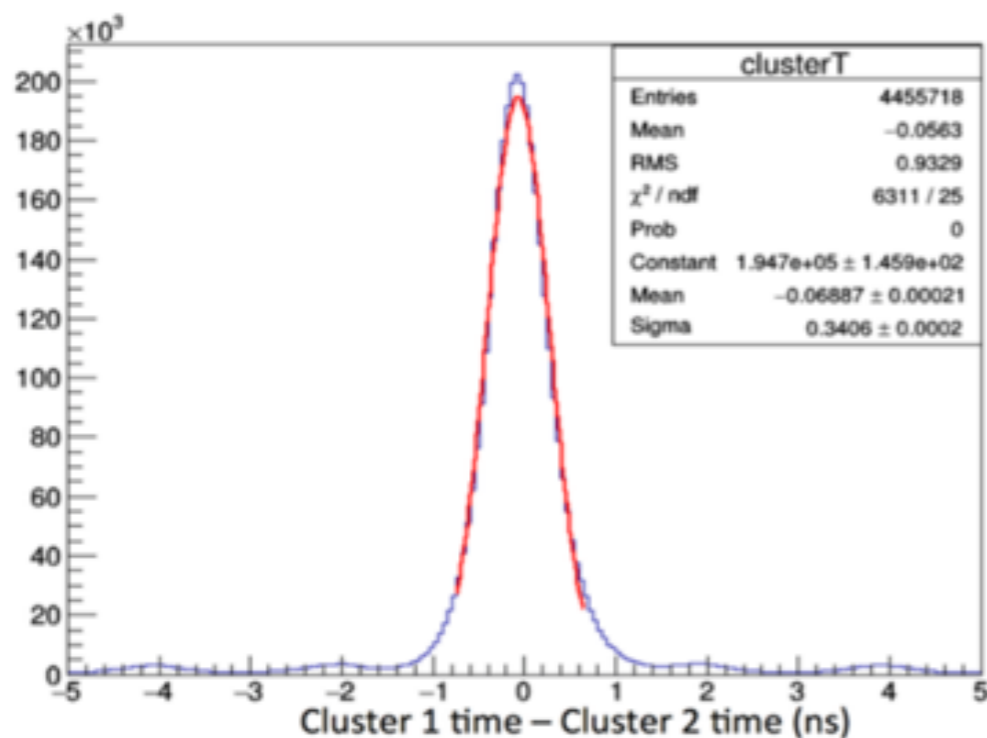


36 sensors total; 23004 channels

ECal & Trigger performance from 1.05 GeV run

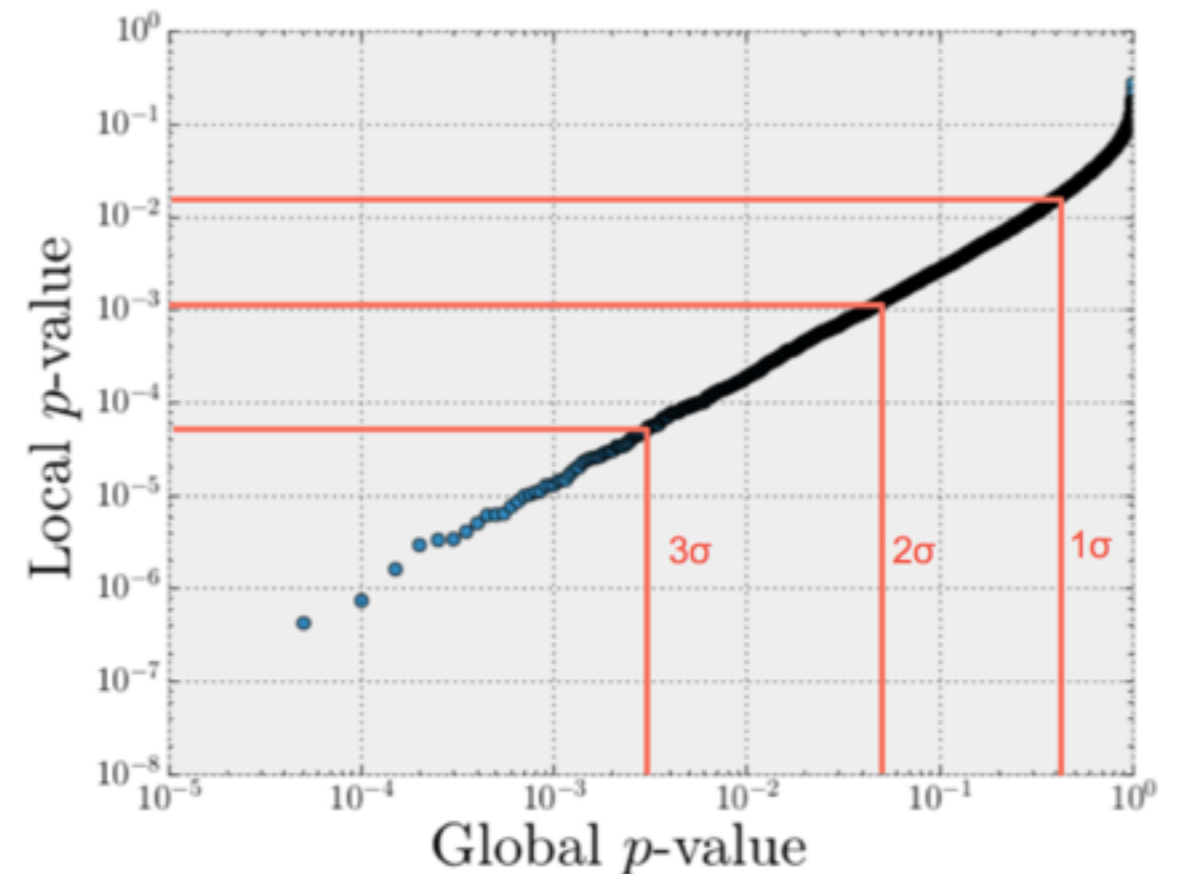


- ECal performed as expected!
 - <2ns relative cluster time resolution
 - reasonable energy resolution (as expected given the design)
- trigger/DAQ capable at >100kHz running, though SVT occupancy considerations max us out at ~20kHz

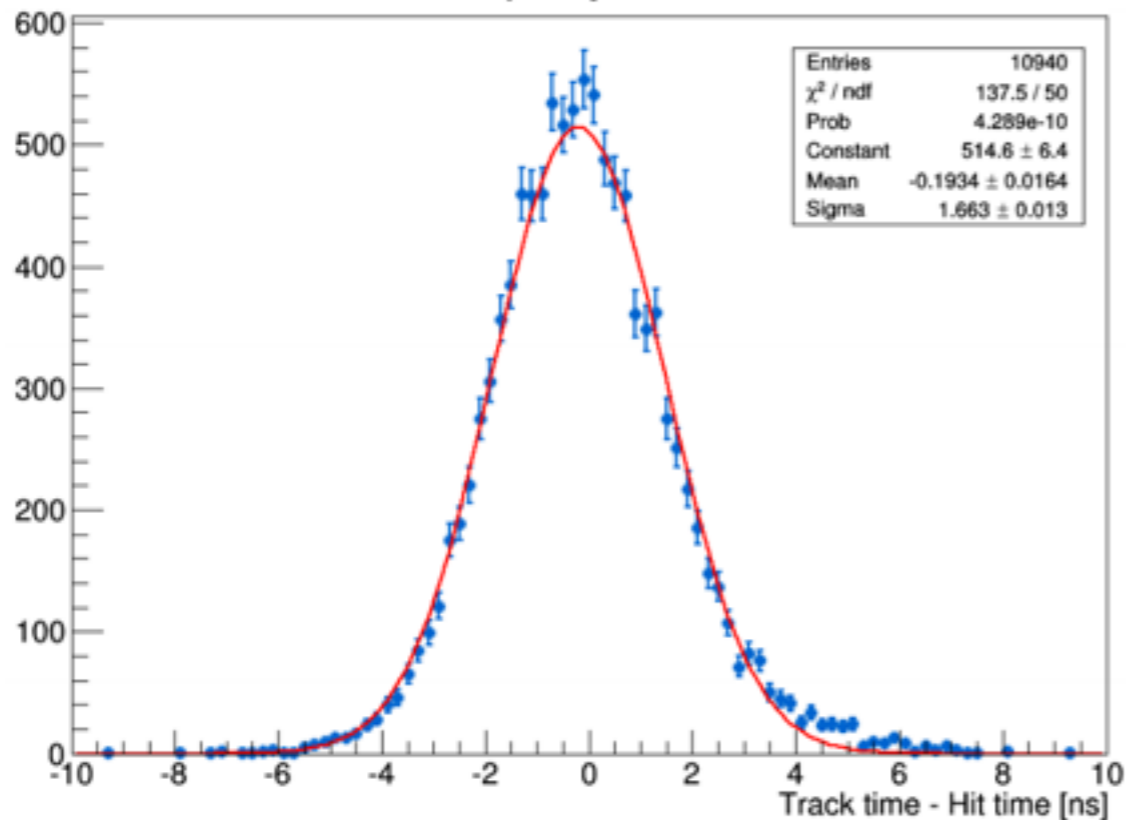


Accounting for look-elsewhere effect

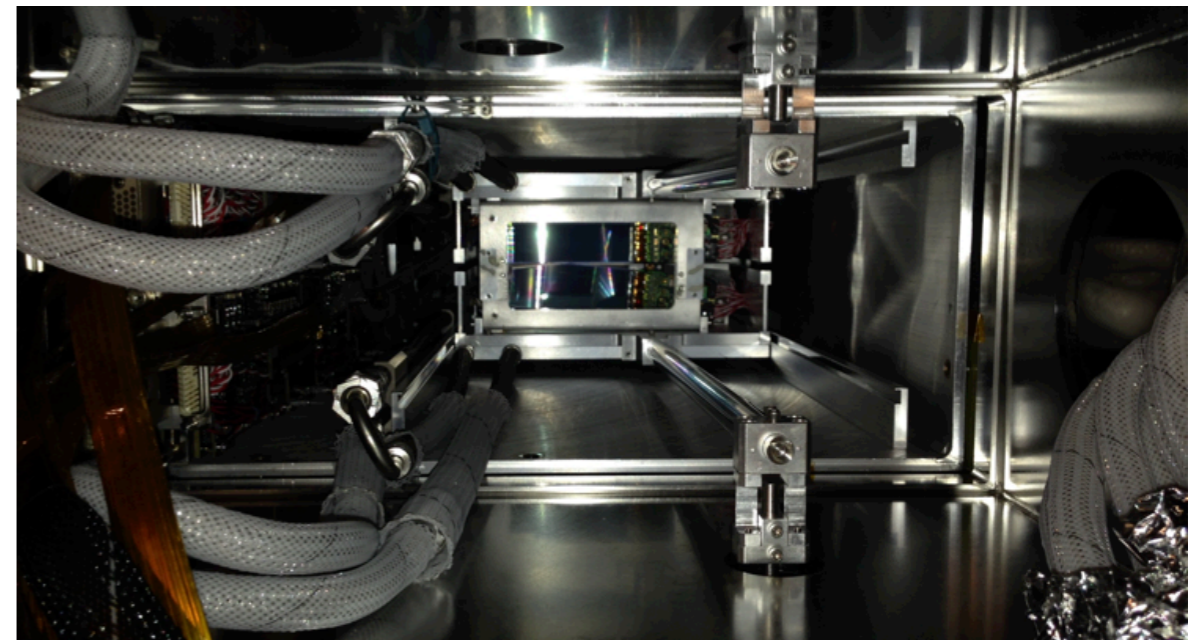
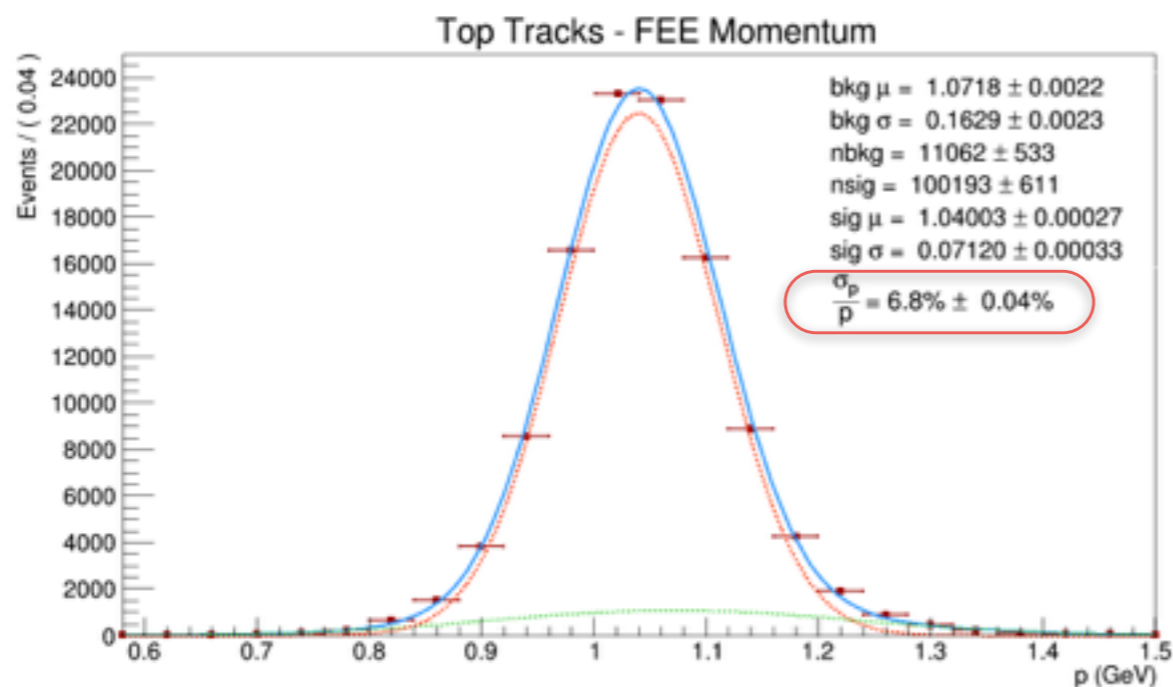
- When you're doing a scan like this, must ask the right questions
- “What is the probability I see a signal significance fluctuation of at least X% given the scan I'm doing”
- We account for this by doing a bunch of toy experiments, scanning, and plotting the p-value of the largest fluctuation
- Correct the “X-sigma” criteria for this effect



HPS SVT Performance from 1.05 GeV run



- SVT also performed well!
 - <<1% dead or noise channels
 - <2ns relative cluster time resolution
 - momentum and angular resolution as expected (dominated by MS)
 - occupancies roughly as expected from MC

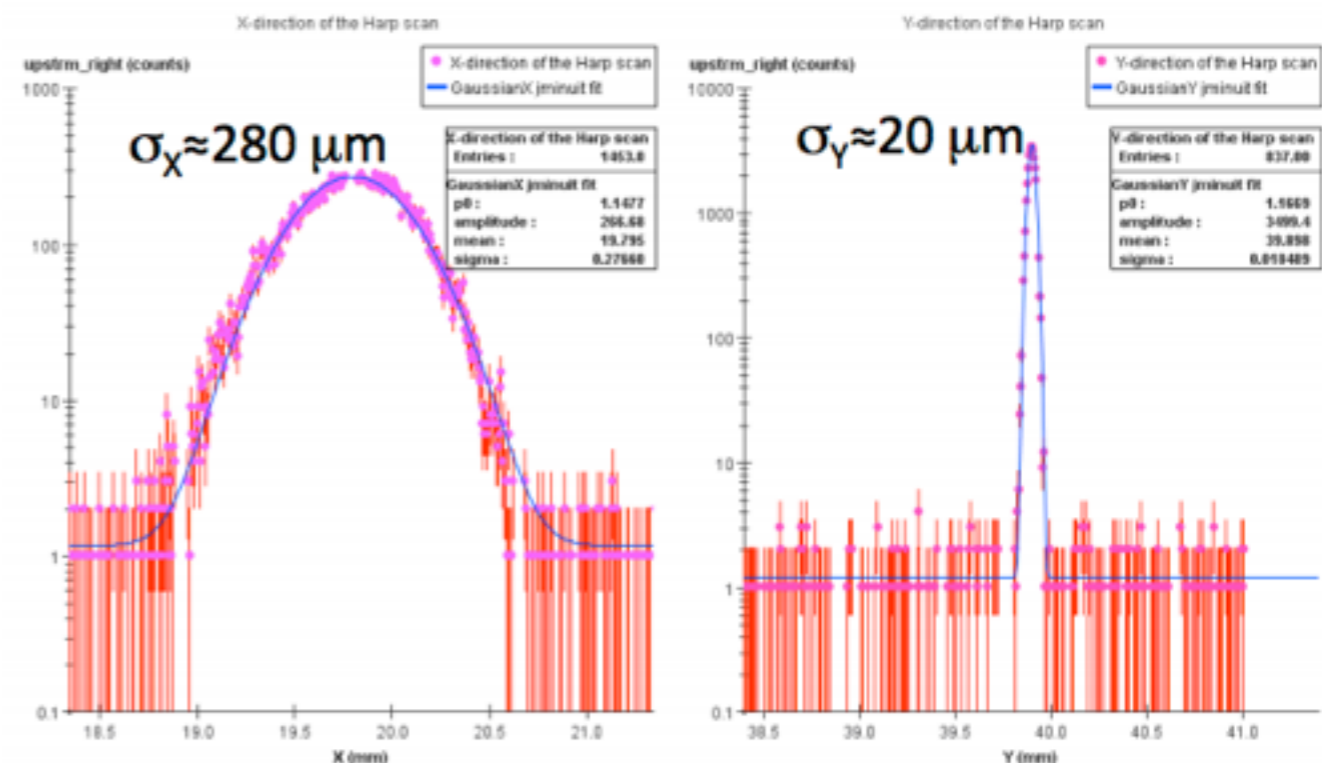
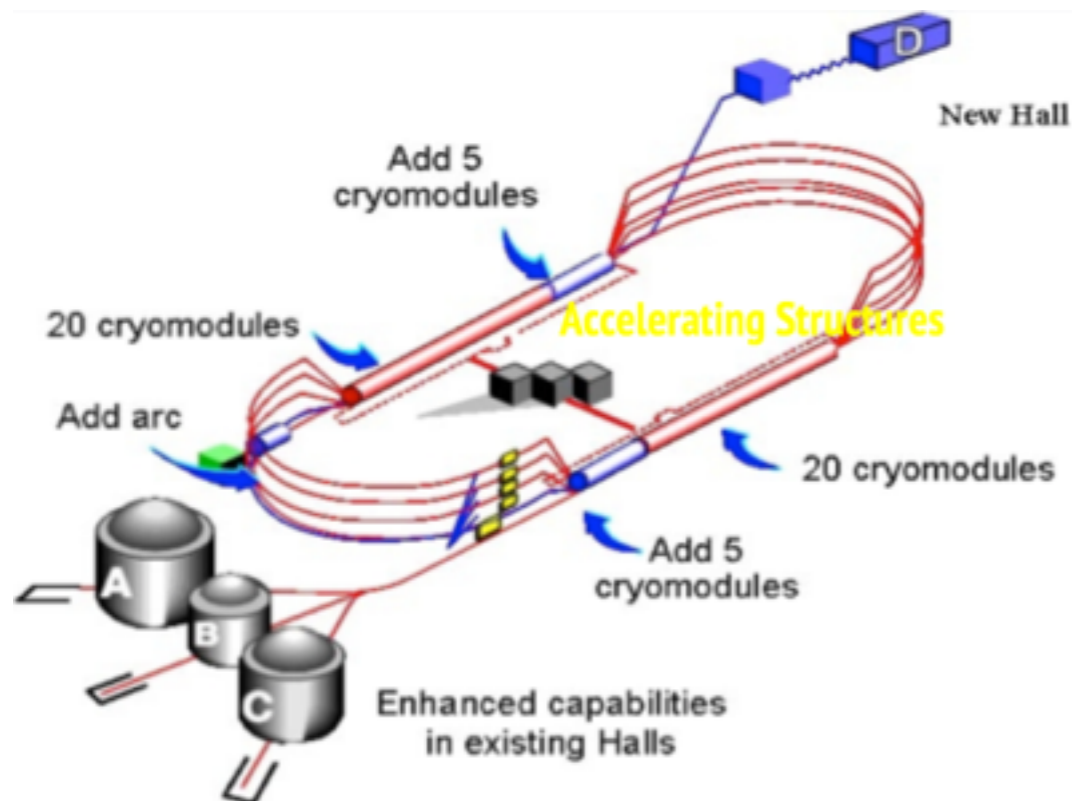


*beam's-eye view of the SVT
(looking downstream)*

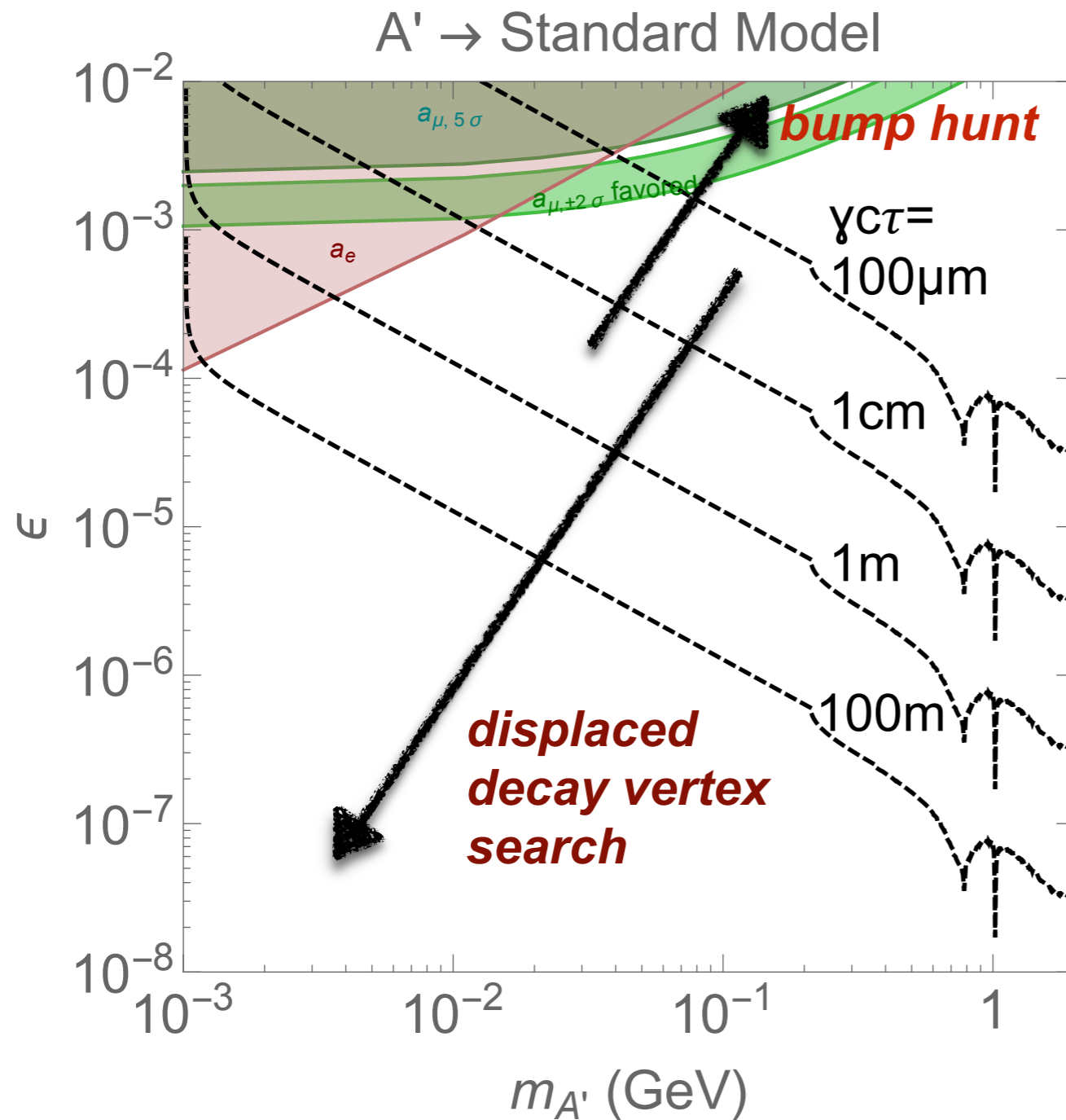
CEBAF and Hall-B @ JLAB

HPS sits in an alcove in *JLAB Hall-B*...behind CLASS, in front of beam dump

- energy: *1.1-11.1 GeV*
- current: up to *~700nA*
- roughly CW...*2ns* bunch spacing
- can focus beamspot *~300μ × 20μ* at target with small beam halo
- small beamspot lets us use the IP as a constraint for long-lived decay search



Two (roughly) complimentary types of searches



thinking about direct searches...

if ($\gamma c\tau < \sim 100 \mu\text{m}$)
bump-hunt;

if ($\gamma c\tau > 1 \text{ m}$)
displaced decay/"light shining through wall";

...some mushy middle where both handles are useful;