

GLAST - TeV Meeting: Collaborative efforts:

- MW campaigns:
 - Summary of planned campaigns, ToO thresholds, ...
 - Blazar monitoring and alerts
- Proposals to be submitted to major observatories (are we willing to submit joint proposals for greater chances of success?)



Some GeV - TeV science questions

- What are the dominant AGN emission mechanisms, and where is the AGN gamma-ray emission originating?
 - Temporal behavior TeV:GeV:X: Optical:radio
- What is the gamma-ray horizon vs energy (EBL studies)?
 - Disentangling EBL effects from blazar evolution effects
- what kinds of Galactic sources await us?
 - PWN dynamics and populations, binaries & microquasars, TeV UNIDs (!),
 - What is the origin of the cosmic rays, finally?
- Are there signals of new physics?
 - CDM and hunting for clumping
 - Lorentz invariance violation effects
 - Challenge is to exclude all astrophysics effects first!
- What are the energy budget and afterglow profiles of Gamma-Ray Bursts, and what are they telling us?
- MOST IMPORTANT: what new questions will the data present?

...... These could find the answers working together

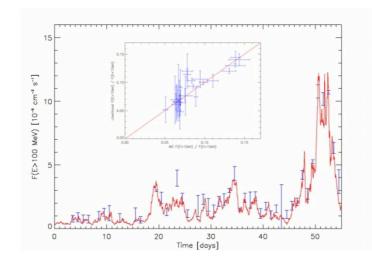


GLAST-LAT & Blazars

- Coverage of about 20% of the sky at any instant with good sensitivity
- The entire sky will be observed every 3 hours
- Uniform exposure in survey mode
- Broad energy range (20 MeV 300 GeV)



- □Daily sampled LC can be easily obtained for most of the bright blazars → Variability on timescales >= 1 day can be well investigated.
- □Intra-day (hours) variations can be detected for the brightest gamma-ray blazars.
- ☐ Detailed spectral variation analysis and intrabands delays studies may be performed
- Multiepoch SEDs can be obtained.



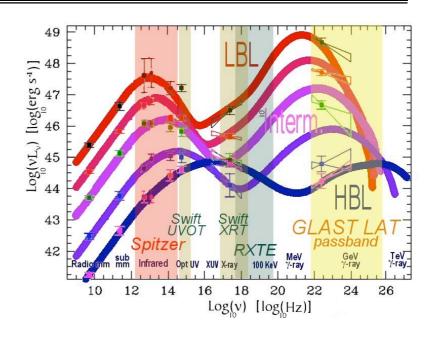
Simulation of a daily light curve as will be measured by the LAT for 3C279. The inset displays the true F(E>1 GeV)/F(E<1 GeV) hardness ratios versus the measured ones.

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The AGN Science Group MW Plan for the first year

- Target of Opportunity (ToO)
 - When a source will be in a bright state in gamma ray
- Planned Intensive Campaign (PIC, months)
 - On a few selected sources
- Planned Long-Term Campaign (PLTC , >=1 year)
 - On a sample of selected sources
- Ad-hoc Intensive Campaign: (AIC)
 - e.g. the ongoing 3C 454.3 MW campaign



Activities based on the AGN SG Science Goals Document in coordination with the LAT MW Group



ToO Campaigns

http://glast.gsfc.nasa.gov/ssc/data/policy/LAT_Monitored_Sources.html

0208-512	3EGJ0210-5055
0235+164	3EGJ0237+1635
PKS 0528+134	3EGJ0530+1323
PKS 0716+714	3EGJ0721+7120
0827+243	3EGJ0829+2413
ОЈ 287	3EGJ0853+1941
Mrk 421	3EGJ1104+3809
W Com	3EGJ1222+2841
3C 273	3EGJ1229+0210
3C 279	3EGJ1255-0549
1406-076	3EGJ1409-0745
11.4.426 ; 420	
H 1426+428	NA
1510-089	NA 3EGJ1512-0849
	1001
1510-089	3EGJ1512-0849
1510-089 PKS 1622-297	3EGJ1512-0849 3EGJ1625-2955
1510-089 PKS 1622-297 1633+383	3EGJ1512-0849 3EGJ1625-2955 3EGJ1635+3813
1510-089 PKS 1622-297 1633+383 Mrk 501	3EGJ1512-0849 3EGJ1625-2955 3EGJ1635+3813 NA
1510-089 PKS 1622-297 1633+383 Mrk 501 NRAO 530	3EGJ1512-0849 3EGJ1625-2955 3EGJ1635+3813 NA 3EGJ1733-1313
1510-089 PKS 1622-297 1633+383 Mrk 501 NRAO 530 1ES 1959+650	3EGJ1512-0849 3EGJ1625-2955 3EGJ1635+3813 NA 3EGJ1733-1313
1510-089 PKS 1622-297 1633+383 Mrk 501 NRAO 530 1ES 1959+650 PKS 2155-304	3EGJ1512-0849 3EGJ1625-2955 3EGJ1635+3813 NA 3EGJ1733-1313 NA 3EGZ158-3023

- The Automatic Science Processing (ASP) tool will analize the LAT data on daily and weekly time scales.
- Data will be made public for about 20 of these monitored sources
- For all other sources an alert will issue when a flare over 2x10⁻⁶ ph cm⁻² s⁻¹ will be observed
- Fast communication of a flaring event (Web, Atel, e-mail)
- Alerts on sources below the DRP theshold will be given on a best effort basis
- In 1 yr of LAT, we expect ~24 sources flaring over $2x10^{-6}$ in the LAT field of view
- A LAT- AGN Flare Advocate will be available to coordinate LAT activities on a flaring source

GLAST-TeV Evo meeting May 14, 2008



MW – Planned Intensive Campaings (1-3 months)

Source Name	Epoch (mm,yyyy)	Campaign Manager
PKS 0528+134	02/11,2008	B. Lott
3C 279	01,2009	G. Madejski
3C 273	ТоО	J. Chiang
Mrk 501	Ongoing	
1ES 1959+650	05 40 0000	. .
1 120 1303 : 000	05-10,2008	D. Paneque
Mrk 421	03-05,2008	D. Paneque
	,	D. Paneque B. Giebels

In coll. with HESS

Other Sources of interest for MW-PIC

PKS 0735+178, PKS 0537-441, AO 0235+164, OJ 287, PKS 1510-08 S5 0716+714, W Com (ON 231), 3C 66A, 3C 454.3



Long-Term MW Campaign: Bright sources

GLAST always observes all blazars in the sky:waiting for MW all Sky Monitors

A public web-page with a table of 206 "VIP" AGN/blazar targets for GLAST and MW analysis: http://glastweb.pg.infn.it/blazar/

- □ Bright AGN/blazars assembled trying to collect basic data with direct links to existing databases (avoiding duplication/obsolescence).
- ☐ This web-table can be useful to plan long term MW campaign, to select input sources for science Proposals to MW facilities (other satellites, ground-based obs.), to perform simultaneous MW analysis joint with the GLAST data, spectral-temporal analysis/comparison based on long-term historical datasets, etc.
- ☐ It is also the starting point of dedicated MW blazars database for GLAST and for ASP monitoring



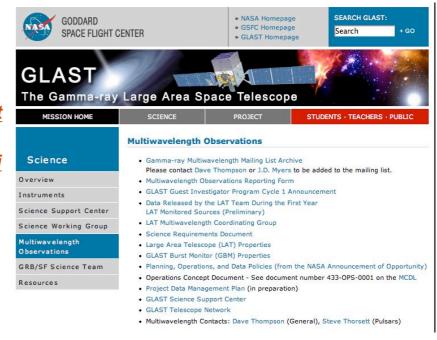
of course it will be revised when GLAST will be is in orbit

Please check the list and send us your comment, suggestions, sources



GLAST MW Info and Coordination

- Multiwavelength observations are key to many science topics for GLAST.
- GLAST welcomes collaborative efforts from observers at all wavelengths
 - For campaigners' information and coordination, see
 - http://glast.gsfc.nasa.gov/science/mult
 i
 - https://confluence.slac.stanford.edu/di splay/GLAMCOG
 - To be added to the Gamma Ray Multiwavelength Information mailing list, contact Dave Thompson:
 - David.J.Thompson@nasa.gov
 - For Information for Multiwavelength Observers about Working with the LAT Team see:
 - https://confluence.slac.stanford.edu/di splay/GLAMCOG/GLAST+LAT+Multiw avelength+Coordinating+Group



https://confluence.slac.stanford.edu/download/attachments/3169/Guidelines_Outside_Observers5.pdf



Some Personal Suggestions (from Chicago 2007 meeting)

 An International Coordinated Network (ICN) for MW Observations should be established:

Possible Actions:

- Definition of the strategies for MW campaigns on a list of sources as part of some Key Projects or ToO
- Definition of Common policies and guidelines for participation to MW campaigns
- Sharing of the schedules of the observations
- Definition of a Common format for high level data
- A central Web page (with restricted access) where people can have information about the ongoing MW campaigns (http://castor.adlerplanetarium.org/MWL/index.php/Main_Page)
- Promote the Fast reduction of the Data, using standard procedures, during a Campaign
- Coordination of proposals to use other facilities for some very important Key Projects
- Promote the MW approach among young people
- etc