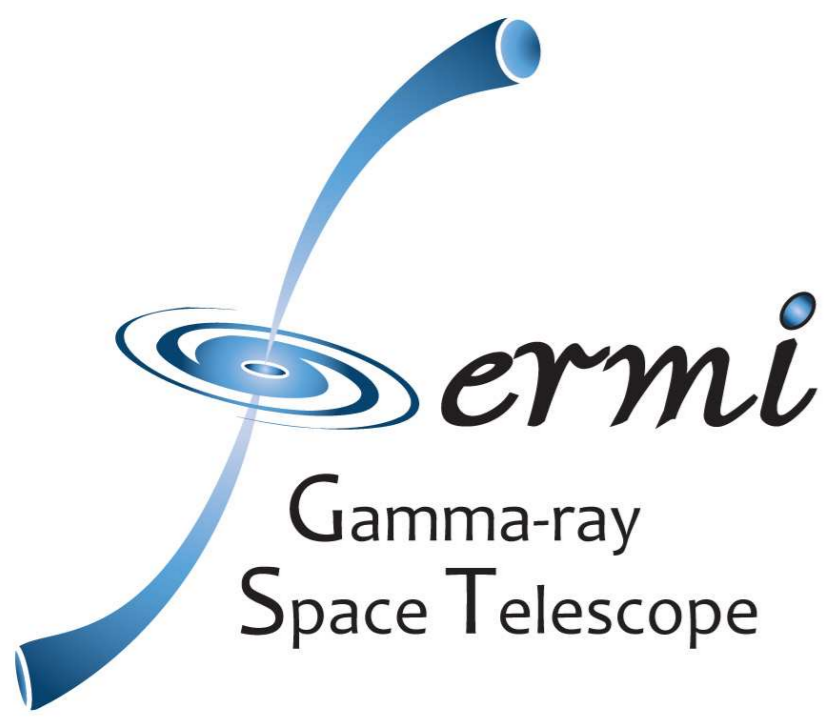


# Spectral features of *Fermi* Gamma-Ray Bursts revealed by the LAT Low-Energy technique

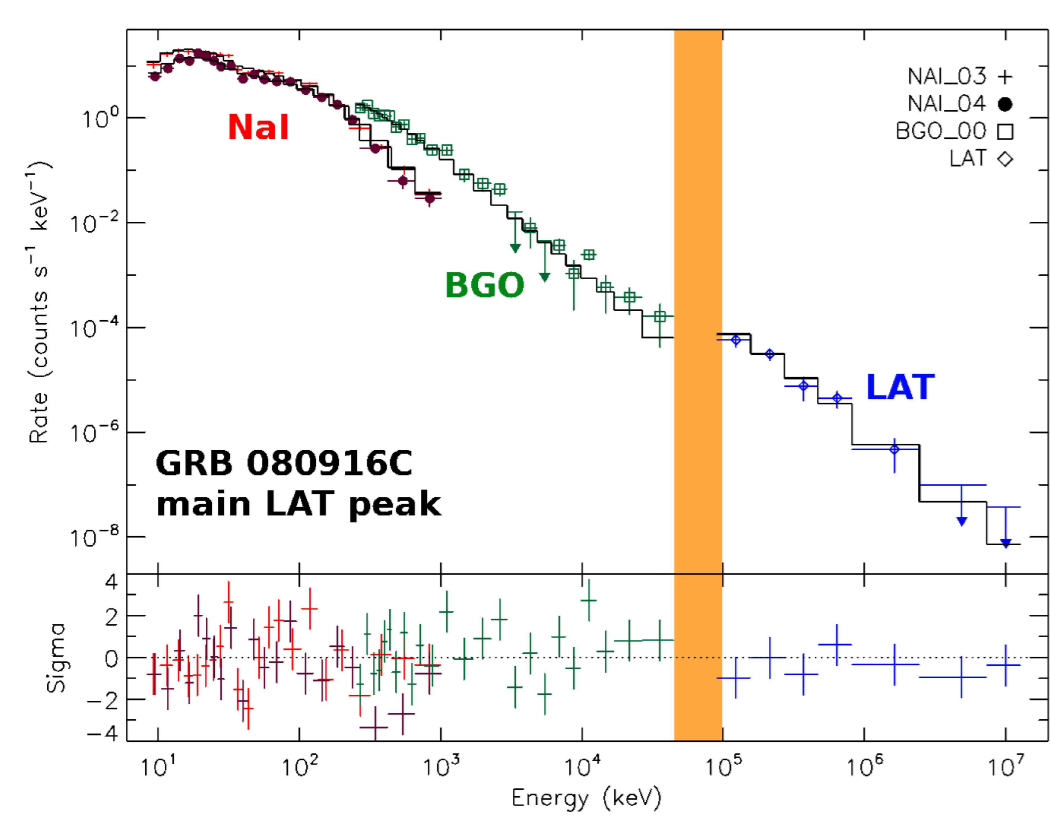
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on behalf of the *Fermi* LAT and GBM collaborations



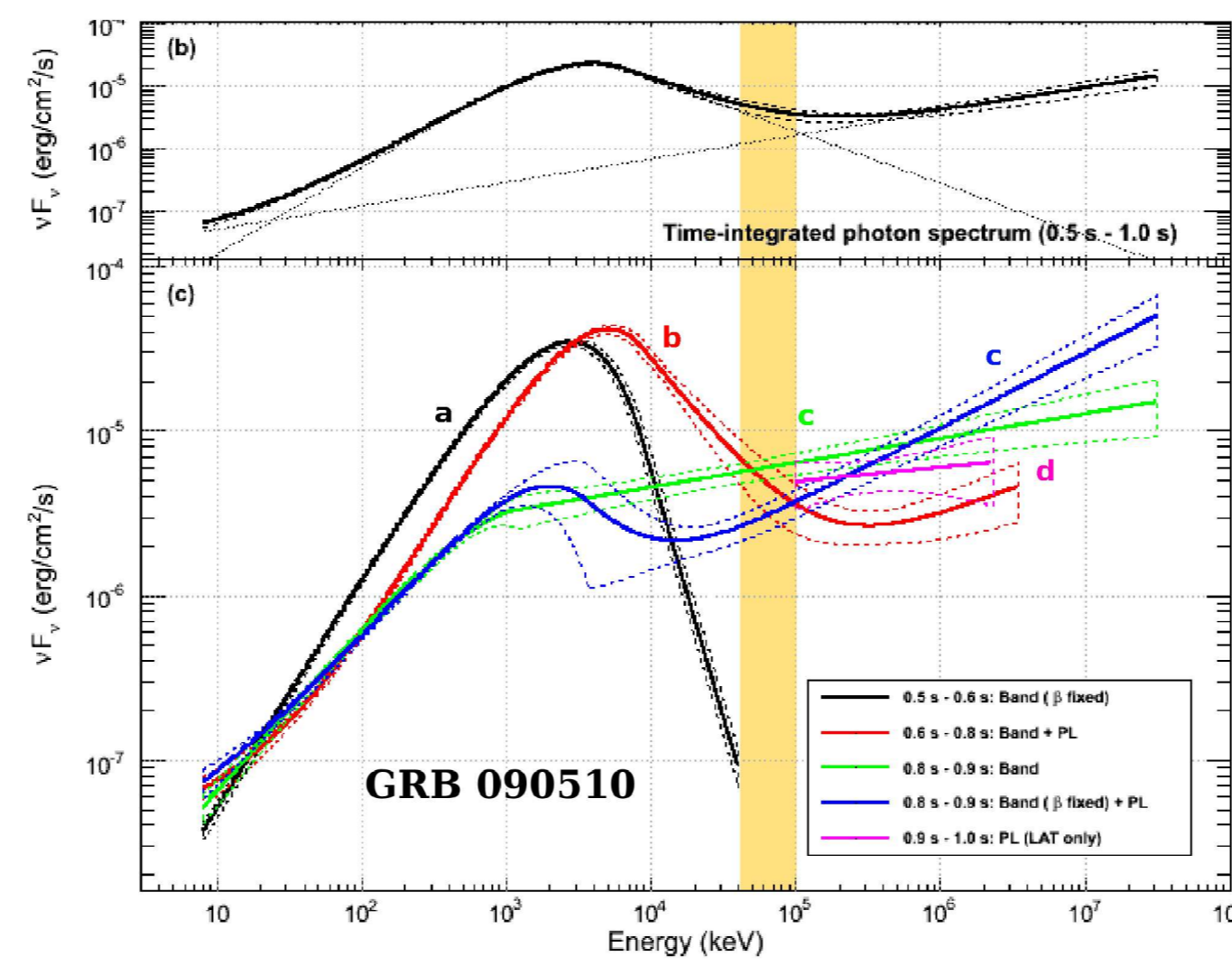
## Abstract

*Fermi* Large Area Telescope (LAT) data analyses based on event reconstruction and classification are so far restricted to events of measured energy larger than 100 MeV. We present a new technique to recover the signal from Gamma-Ray Bursts' (GRB) prompt emission between  $\sim 30$  MeV and 100 MeV, which differs from the standard LAT analysis. Filling the "gap" between Gamma-ray Burst Monitor and LAT observations allows to better constrain the high-energy spectra of GRB. The LAT Low-Energy (LLE) technique is described, first performance studies are presented, and preliminary spectral re-analyses of two *Fermi* GRBs are presented.

## The 40 MeV – 100 MeV gap



LAT standard event-by-event analysis starts at 100 MeV since detector response and background rejection are not fully understood yet below 100 MeV.



## Why looking there?

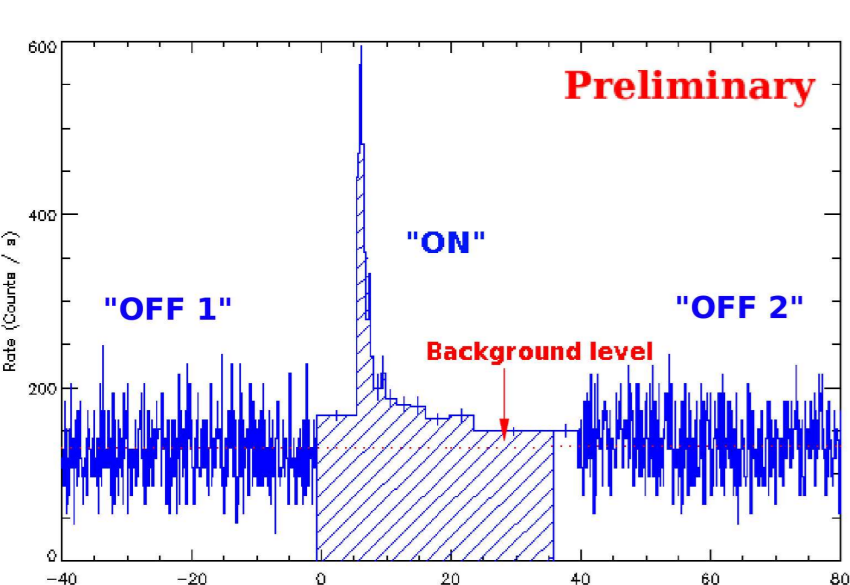
### Faint/soft spectra (e.g. GRB 090510 bin a : soft)

- no signal > 100 MeV, or very faint
- $\Rightarrow$  killed by the usual event selection
- $\Rightarrow$  few information on spectrum > 40 MeV

### Bright spectra (e.g. GRB 090510 bins b, c, d, e)

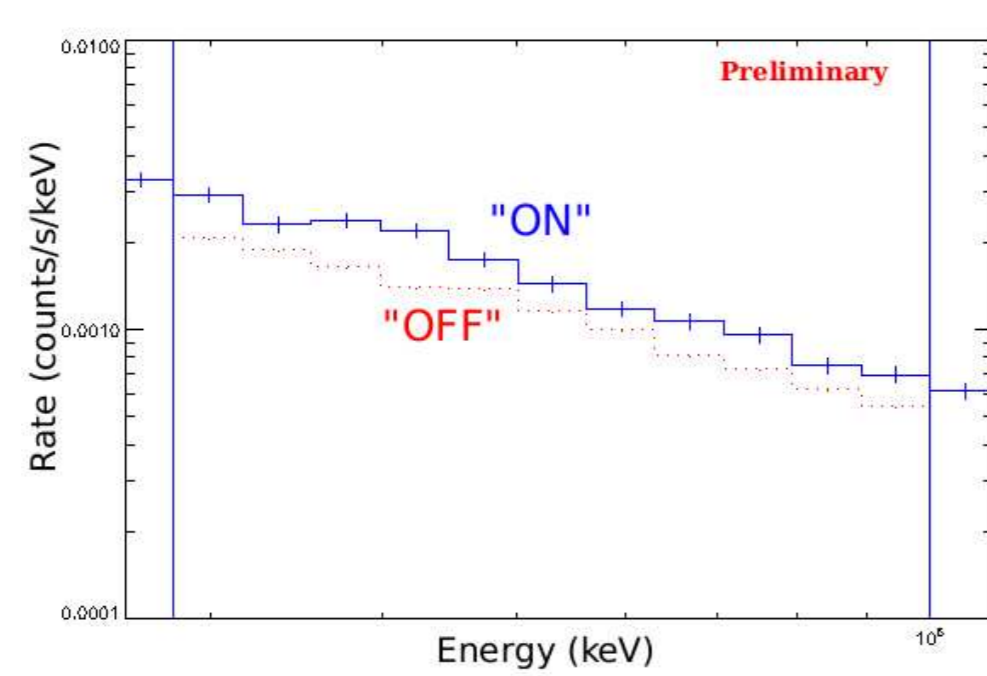
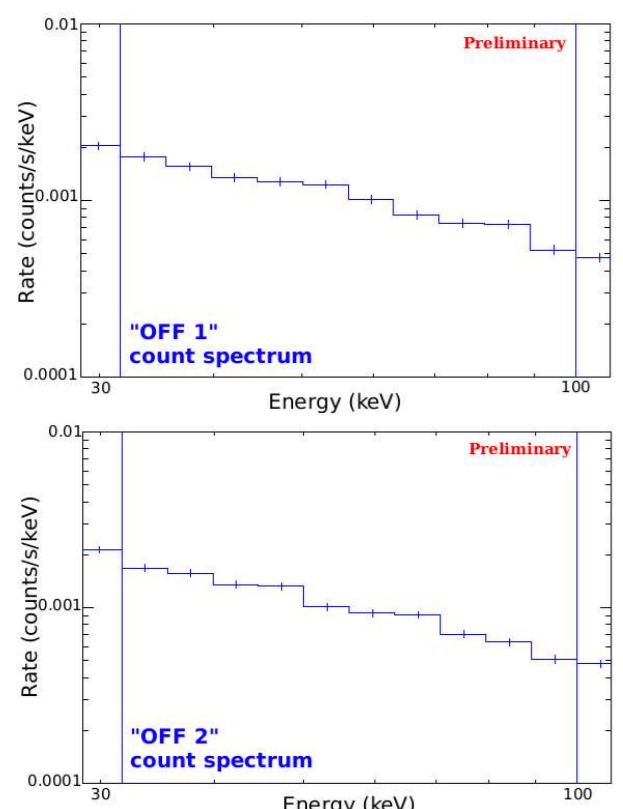
- better spectral constraints
- e.g. additional power-law component takes over the Band spectrum at  $\sim 100$  MeV

## LAT Low-Energy (LLE) analysis procedure

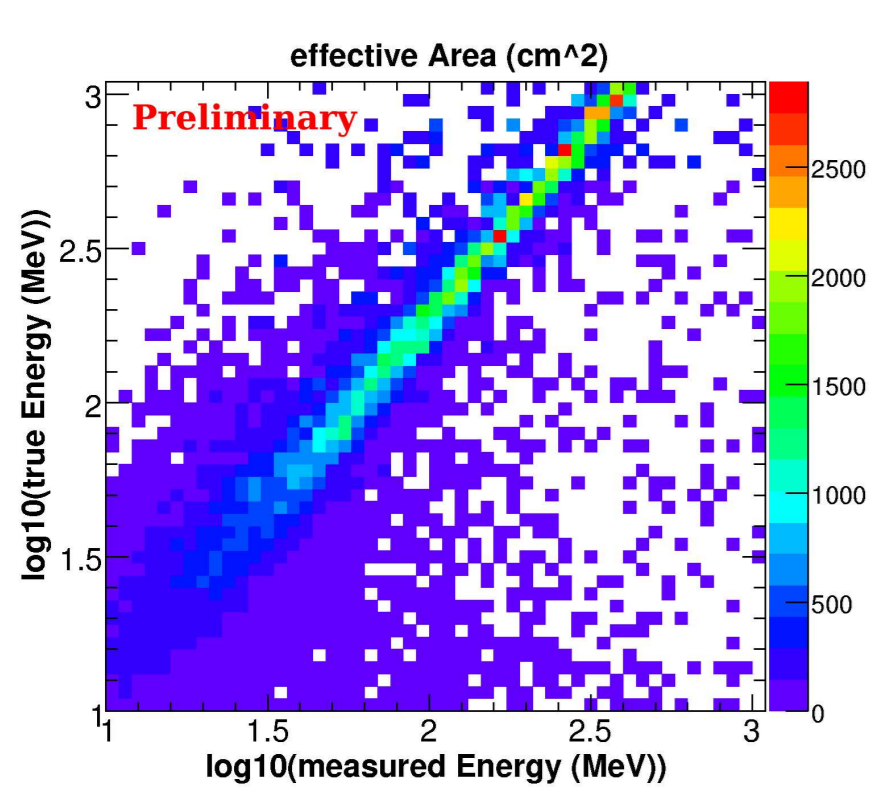


### "ON" - "OFF"

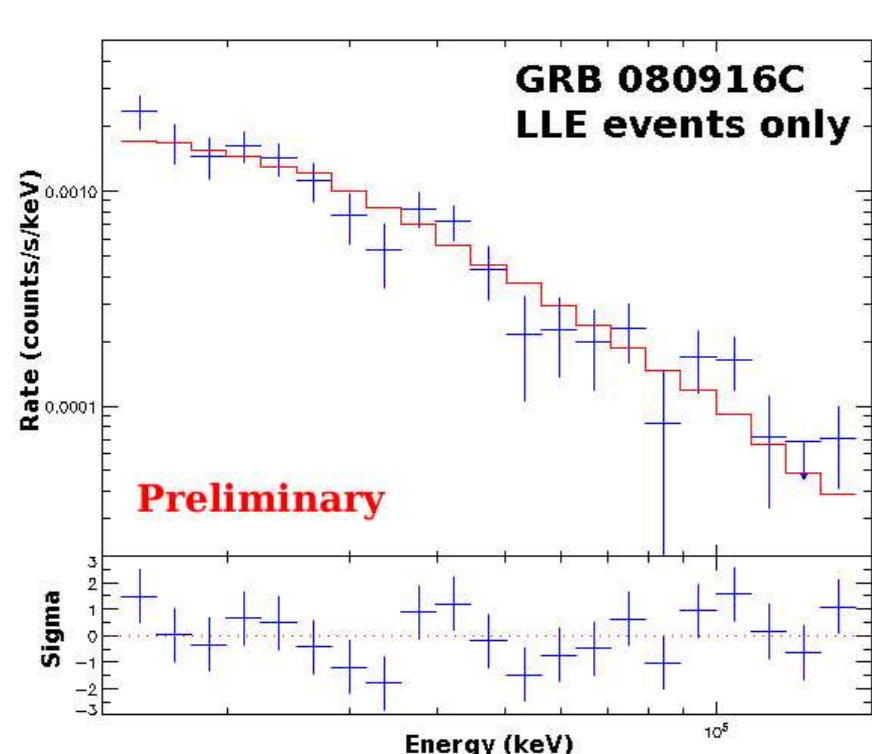
- not an event-by-event technique
- LLE data sample used
- no direction selection
- temporal selection of "ON" and "OFF" regions
- "ON" and "OFF" binned in measured energy
- background time profile polynomial fit in each energy bin
- "OFF" subtracted to "ON"



## Detector Response Matrix (DRM)



- extensive  $\gamma$  simulation with GLEAM (GEANT4)
- observation conditions of the GRB: same inclination angle and livetime over time
- LLE selection is applied
- distribution binned in true (Y-axis) and measured (X-axis) energy
- converted to area units

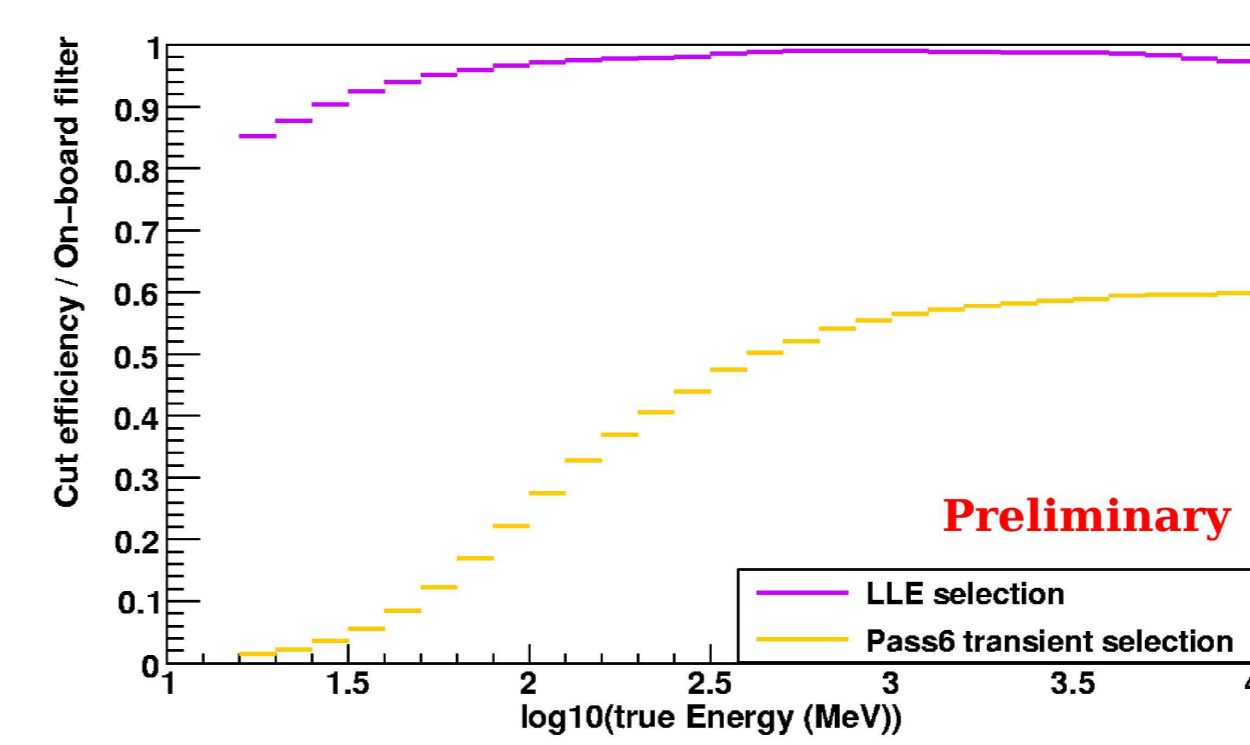


## Spectral analysis (RMfit)

The "ON"- "OFF" spectrum is fitted using the DRM (forward-folding).

## LLE selection

- on-board photon filter
- at least one associated track
- $\Rightarrow$  many more possible photons than the standard selection
- $\Rightarrow$  much more background too
- cuts efficiency compared to standard Pass6 transient class, from GLEAM photon simulations



LLE vs standard P6 transient class :

- $A_{eff} \times 50$  at 30 MeV
- $A_{eff} \times 4.5$  at 100 MeV
- $A_{eff} \times 2 > 1$  GeV
- $\Rightarrow$  photon statistics greatly improved at low energies, and may improve at high energies too

## Energy measurement

- based on tracker and calorimeter
- not the same variable as in standard analysis
- good correlation for events with at least one track (see DRM)
- resolution study from GLEAM photon simulations
- $\Rightarrow$  no very large bias, reasonable energy resolution (under improvement)

incoming photon inclination :  $\theta < 40^\circ$

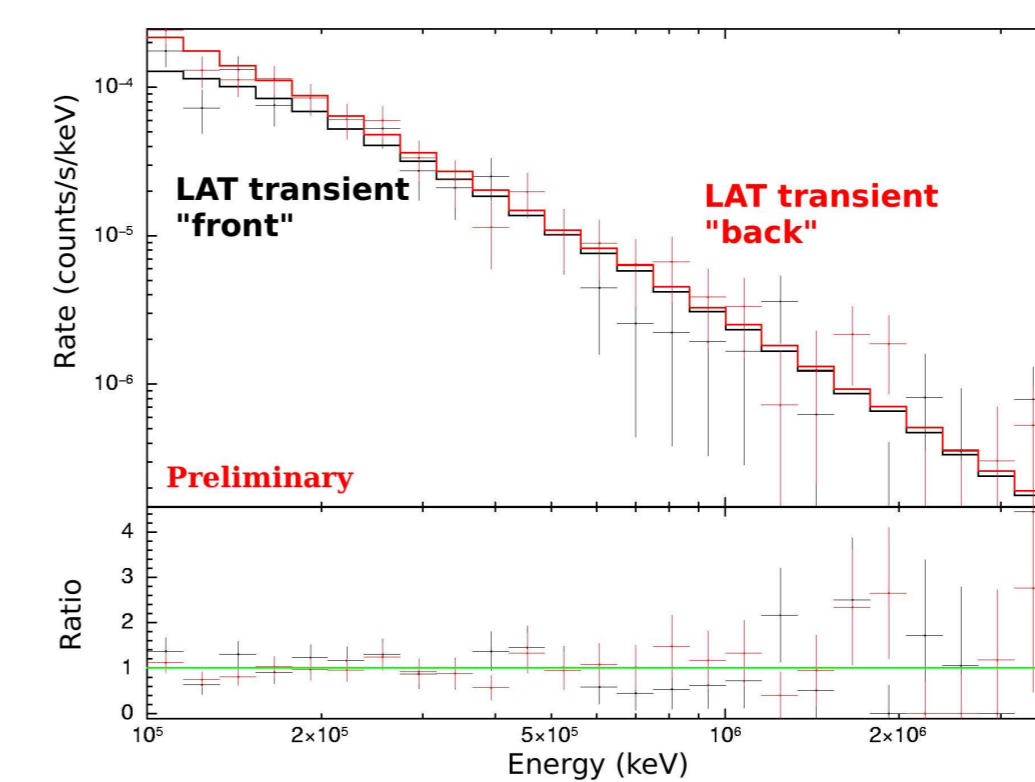
$E_{MC} (\pm 2 \text{ MeV})$	$\langle E_{mes} \rangle \pm \text{RMS (MeV)}$	$\text{RMS}/\langle E_{mes} \rangle$	$(\langle E_{mes} \rangle - E_{MC})/E_{MC}$
30	$27 \pm 10$	37%	-10%
50	$45 \pm 16$	36%	-10%
100	$90 \pm 27$	30%	-10%
500	$490 \pm 70$	14%	-2%

incoming photon inclination :  $40^\circ < \theta < 70^\circ$

$E_{MC} (\pm 2 \text{ MeV})$	$\langle E_{mes} \rangle \pm \text{RMS (MeV)}$	$\text{RMS}/\langle E_{mes} \rangle$	$(\langle E_{mes} \rangle - E_{MC})/E_{MC}$
30	$30 \pm 14$	47%	0%
50	$44 \pm 18$	41%	-12%
100	$85 \pm 34$	40%	-15%
500	$470 \pm 80$	17%	-6%

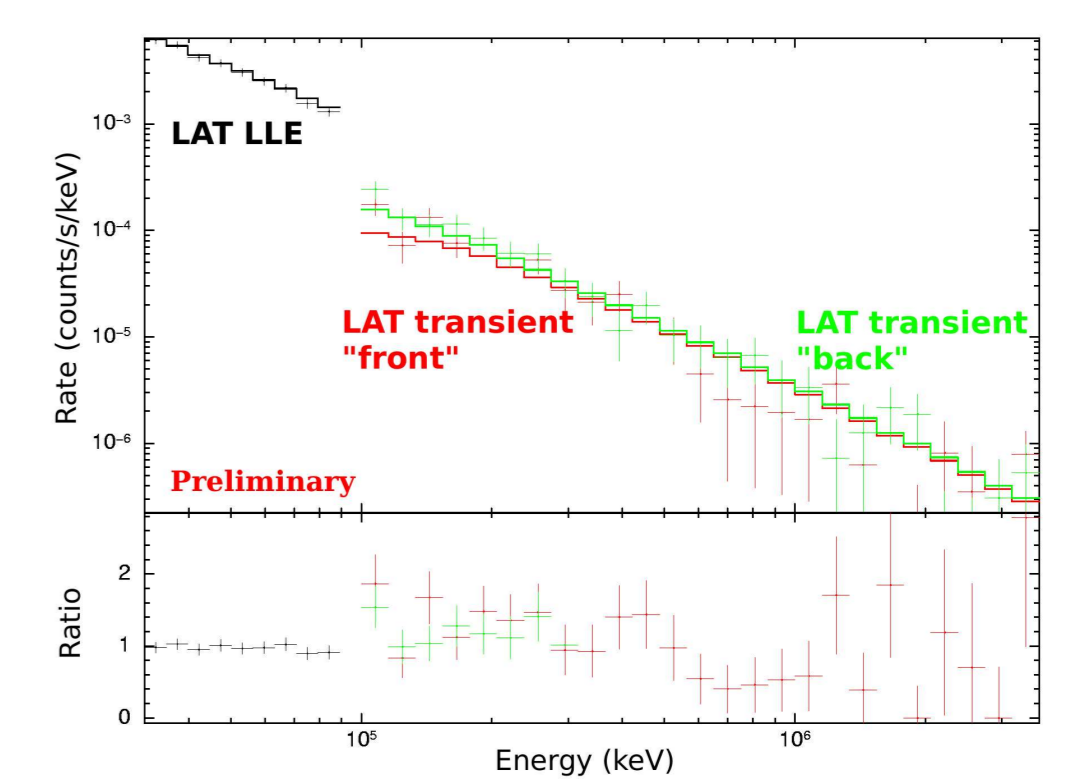
## Reconstruction capabilities

- simulated GRB
- spectrum :  $N(E) = N_0(E/E_0)^{-\beta}$
- $E_0 = 1 \text{ GeV}$
- $N_0 = 1.19 \times 10^{-9} \text{ ph.cm}^{-2} \cdot \text{s}^{-1} \cdot \text{keV}^{-1}$
- $\beta = 2.1$
- no background added
- LAT-only reconstruction : with and without LLE
- $\Rightarrow$  similar results are observed, with a smaller uncertainty on  $\beta$  when using LLE events, yet with a systematic error which will be investigated after the technique has been fully calibrated.



$$N_0 = 1.36 \pm 0.13 \times 10^{-9}$$

$$\beta = 2.21 \pm 0.06$$



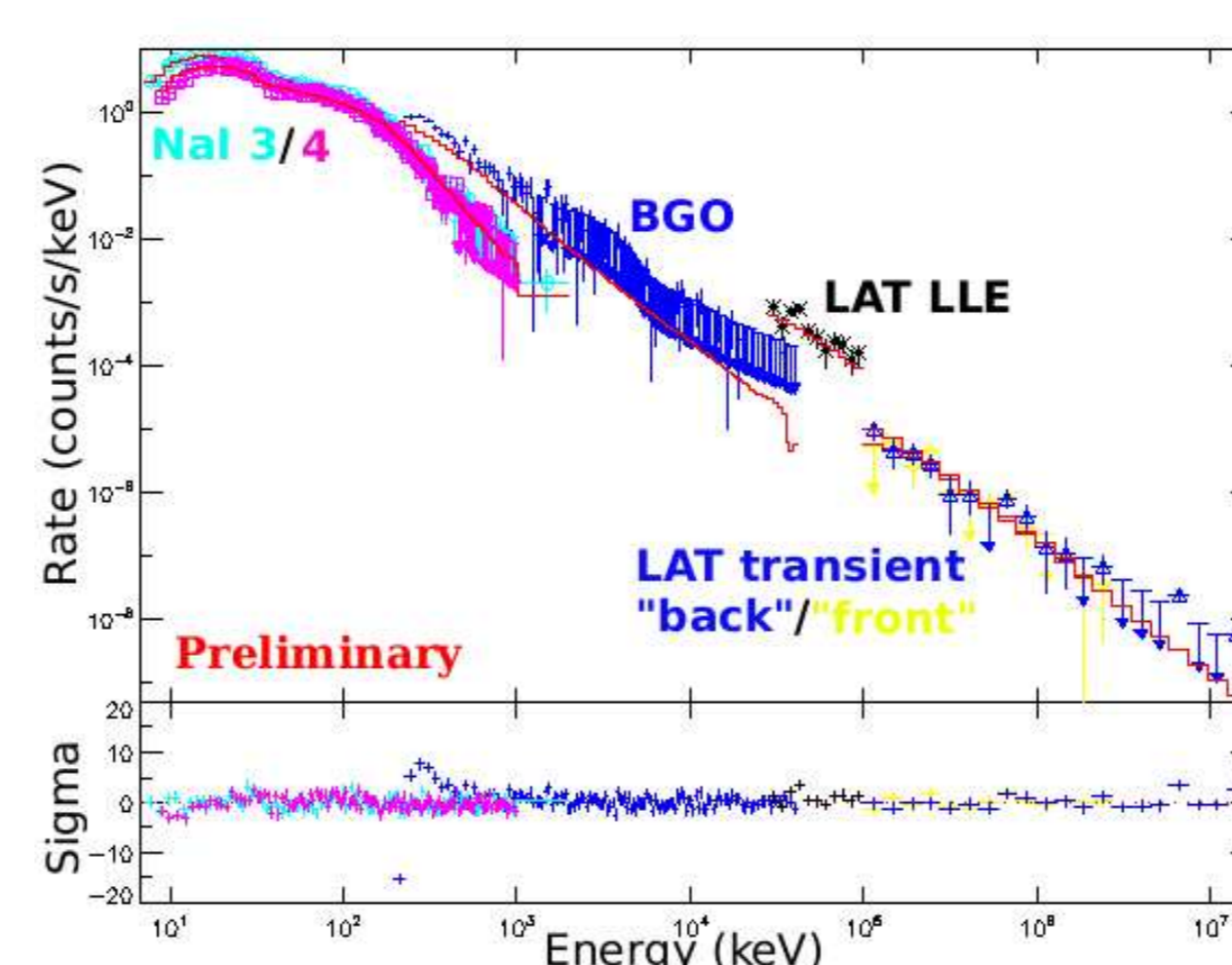
$$N_0 = 1.60 \pm 0.12 \times 10^{-9}$$

$$\beta = 2.02 \pm 0.02$$

(statistical errors)

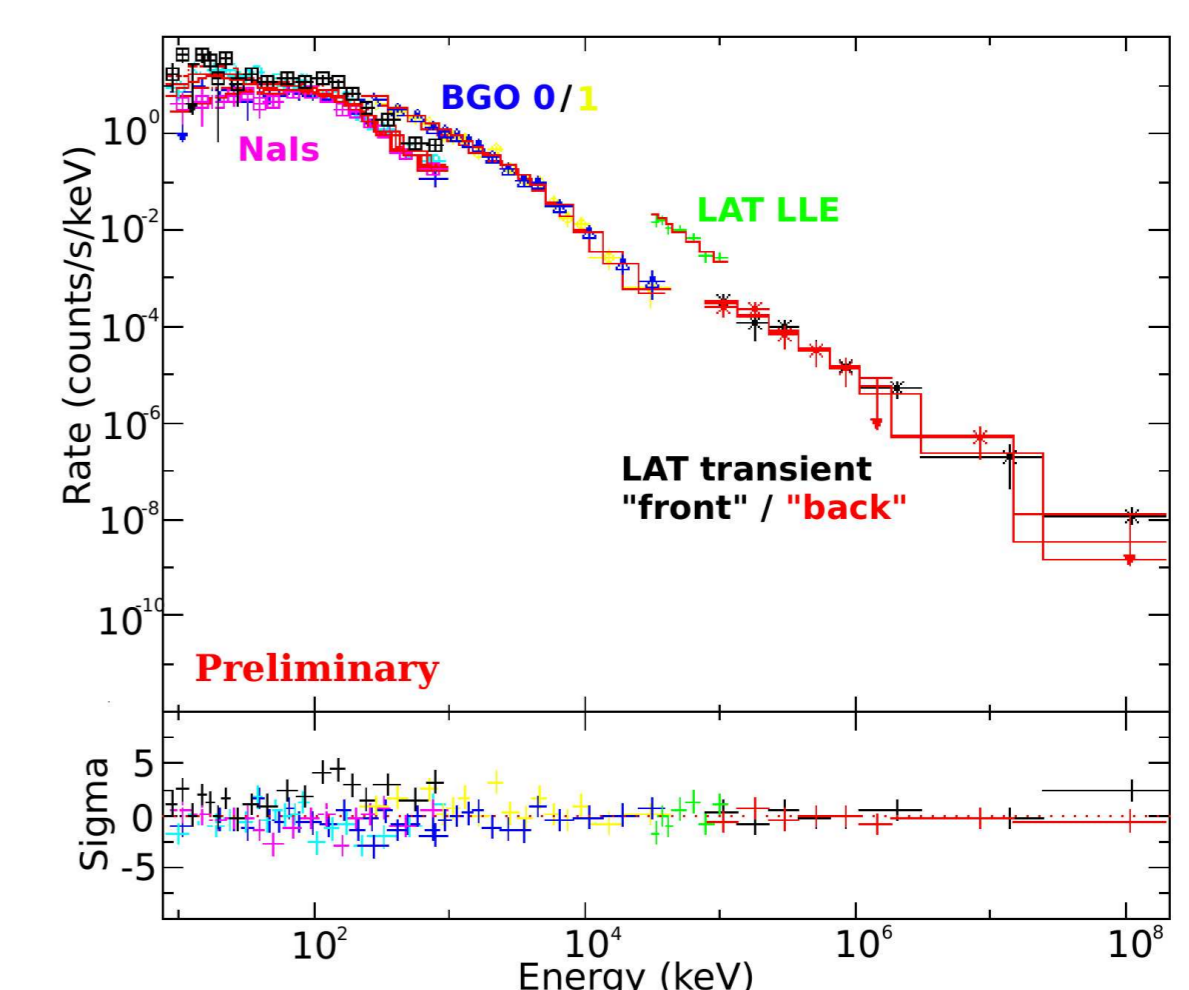
## Preliminary re-analyses of real *Fermi* GRBs

### bright long GRB 080916C



- time integrated spectrum, Band function fit
- LLE data superimposed to other datasets but not fitted
- $\Rightarrow$  good residuals
- $\Rightarrow$  consistent with the standard analysis

### bright short GRB 090510



- time-integrated spectrum, Band + power-law fit
- LLE and other data fitted together
- $\Rightarrow$  high-energy additional component even more significant :  $N_\sigma = 8.9$  (5.6 without LLE)
- $\Rightarrow$  spectral evolution may better show up if using the LLE data in the time-resolved spectroscopy

## Conclusions

- very promising technique for GRB studies
- can be used also for other sources (see poster P2-96 by J.M. Burgess)
- this is a non-standard analysis, which uses non-public information
- validation and improvement still ongoing (acceptance, energy calibration, systematics studies)

## References

- W.B. Atwood et al., "The Large Area Telescope on the *Fermi* Gamma-ray Space Telescope", *ApJ*, 697, 1071 (2009) (arXiv : 0902.1089)
- A.A. Abdo et al., "*Fermi* observations of high-energy gamma-ray emission from GRB 080916C" *Science*, 323, 1688 (2009)
- A.A. Abdo, et al. "*Fermi* Observations of GRB 090510 : A bright, short burst with a hard power-law component above 100 MeV", *ApJL*, in prep.