

Breaking the Enigma of the X-ray Quasar jets with GLAST

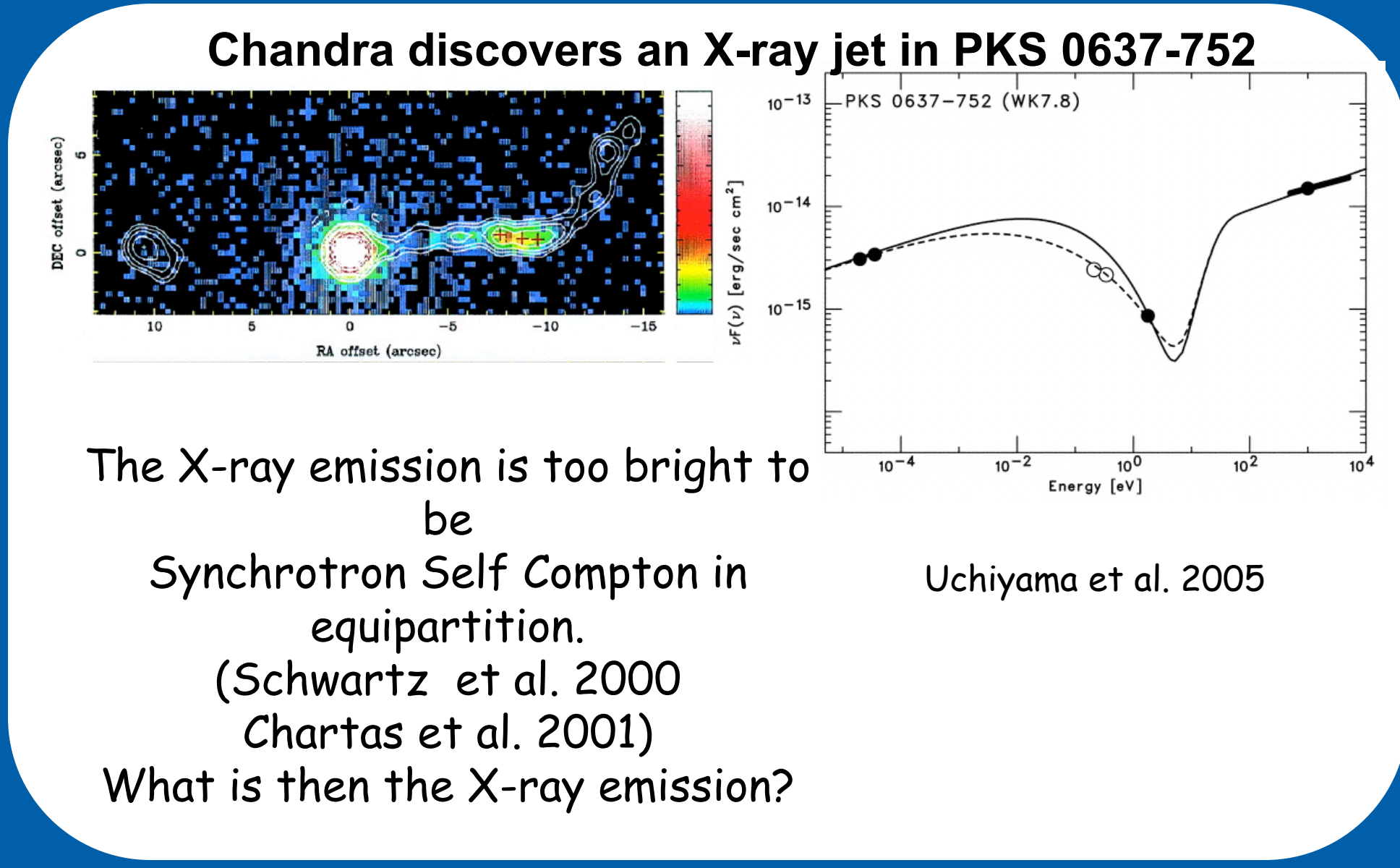


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What is the X-ray emission mechanism?

The X-ray predicted GLAST & TeV scopes

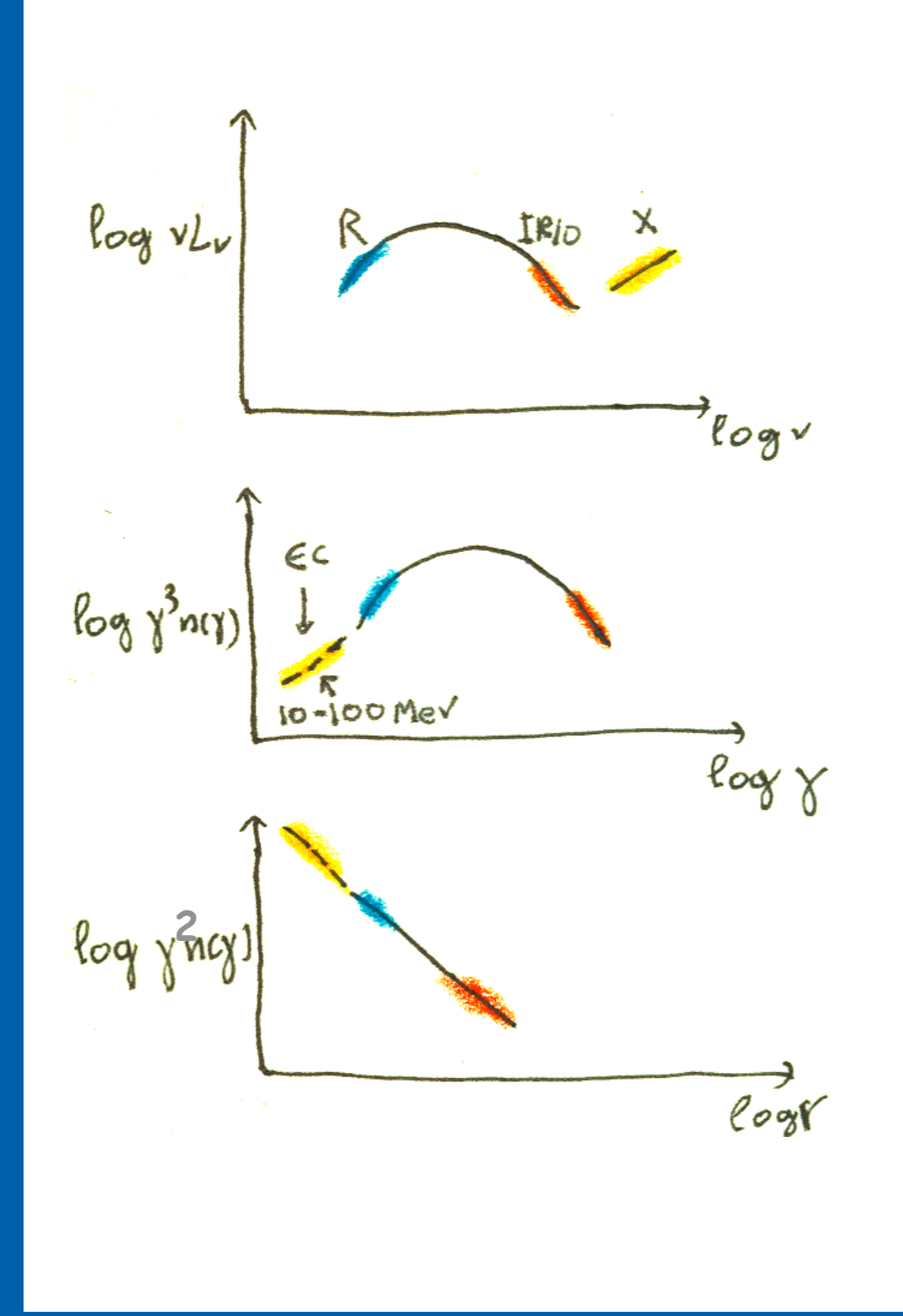


A. Inverse Compton scattering off the CMB (EC/CMB) (Tavecchio et al. 2000, Celotti et al. 2001)

Extends the electron energy distribution (EED) down to 10-100 MeV energies

Requires relativistic large scale jets ($\delta \sim 10$)

Increased jet power requirements, radiatively inefficient

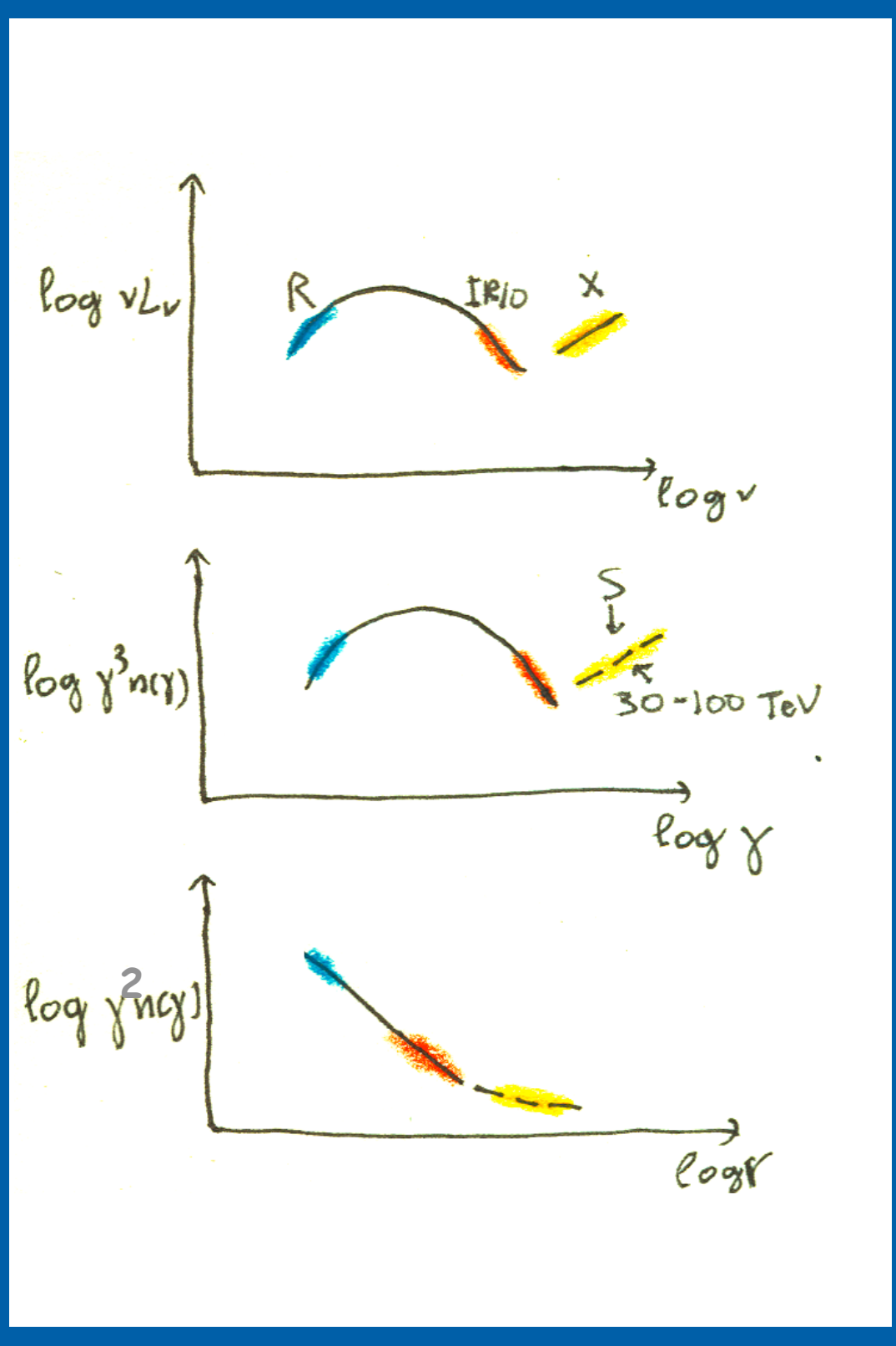


B. Synchrotron (e.g. Harris et al. 2004, Hardcastle 2006)

Additional EED component at ~1-100 TeV energies

No need for highly relativistic large scale jet

More economical in jet power, radiatively efficient



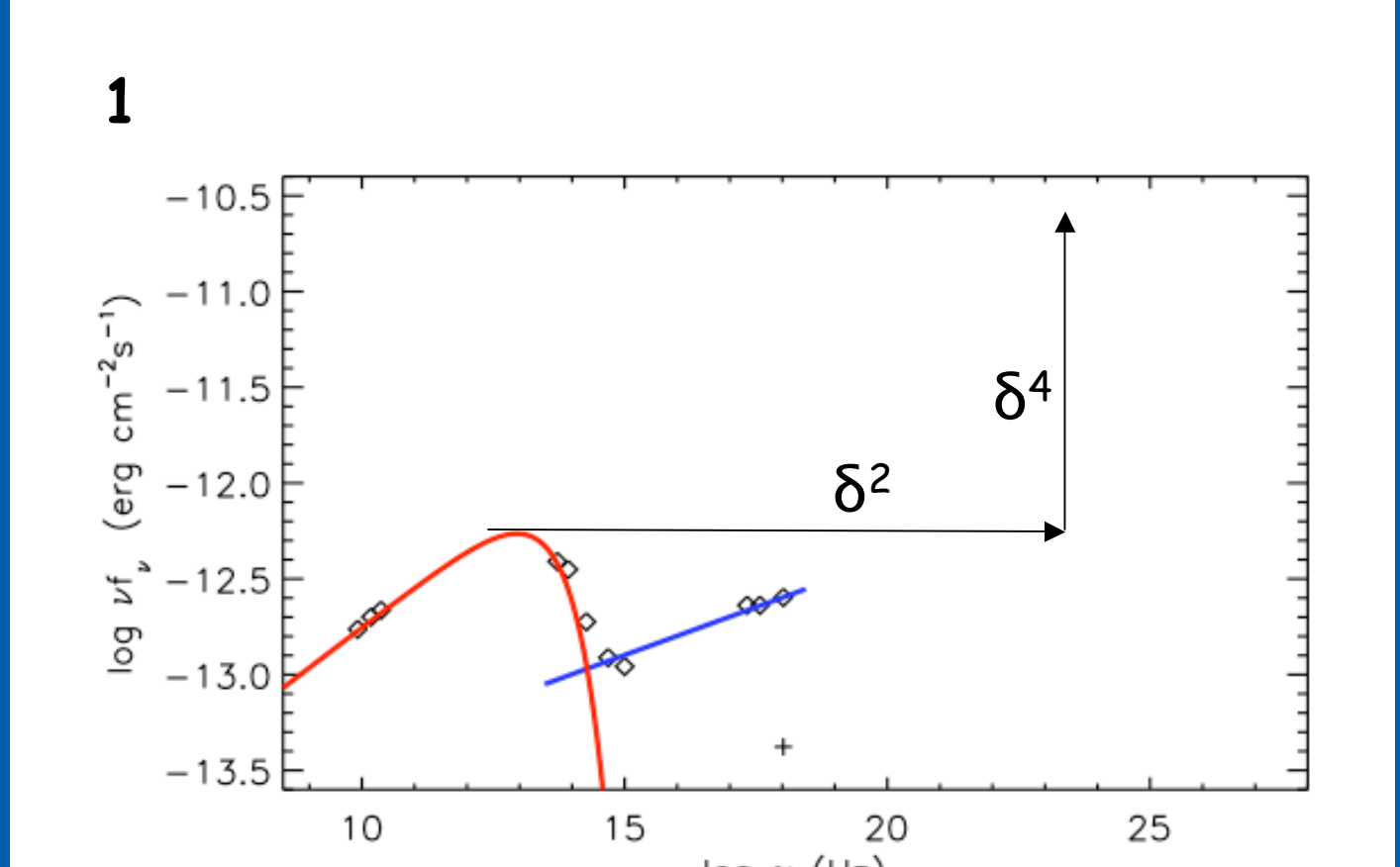
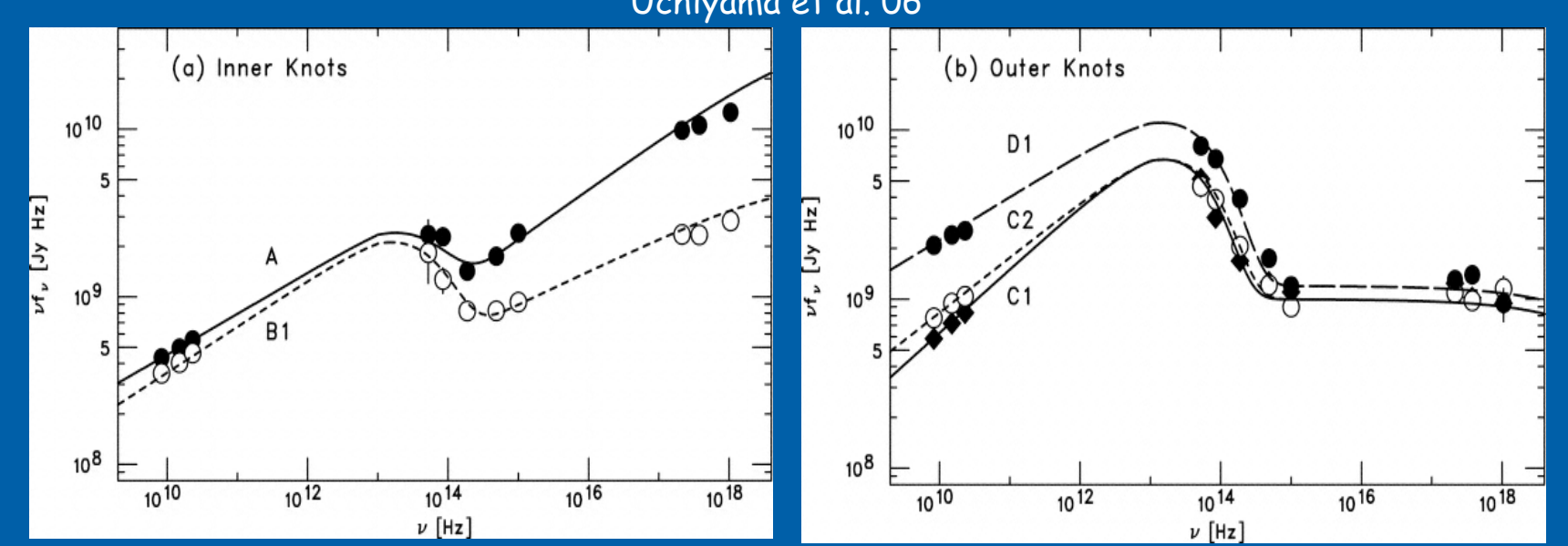
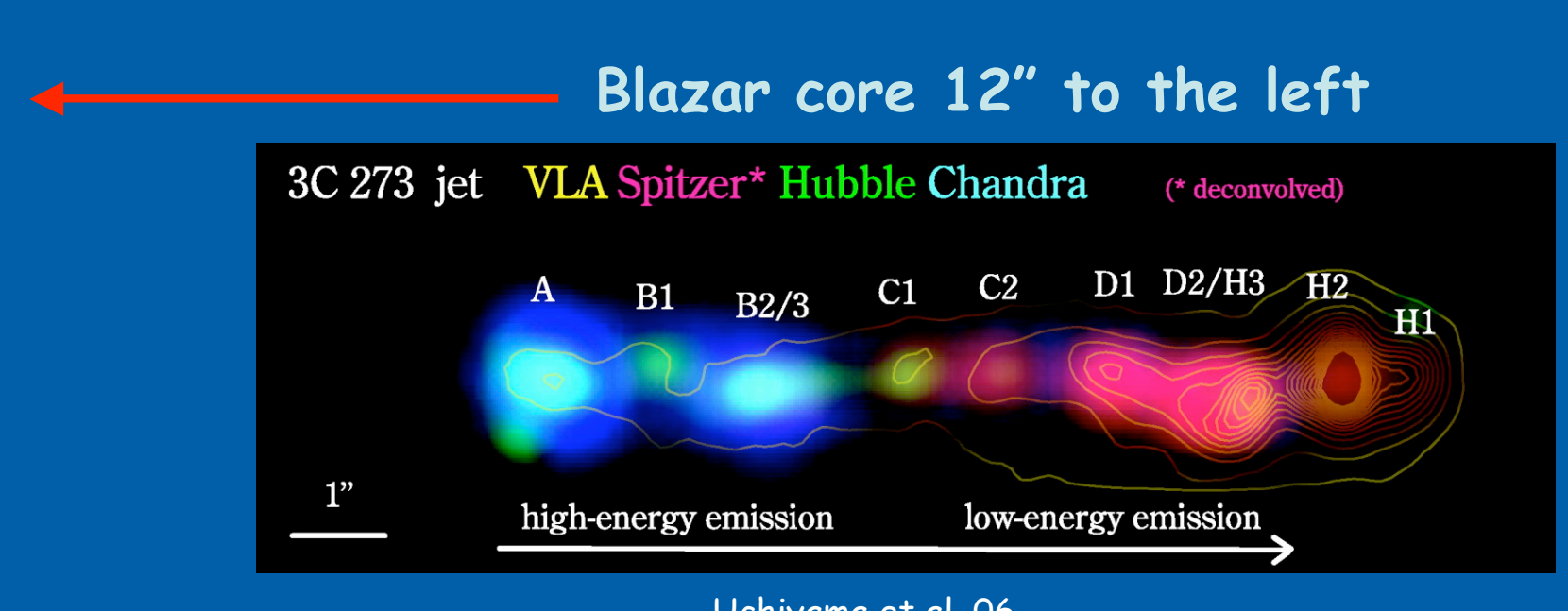
Two competing models, two γ -ray bands, and a helpful source

- The EED in both models have sub-TeV electrons: Sub-TeV electrons + CMB \rightarrow GeV emission from both models
- The EED of the synchrotron model has multi-TeV electrons: Multi-TeV electrons + CMB \rightarrow TeV emission only from the synchrotron model

Let's apply these to a bright, nearby source: 3C 273 at $z=0.158$

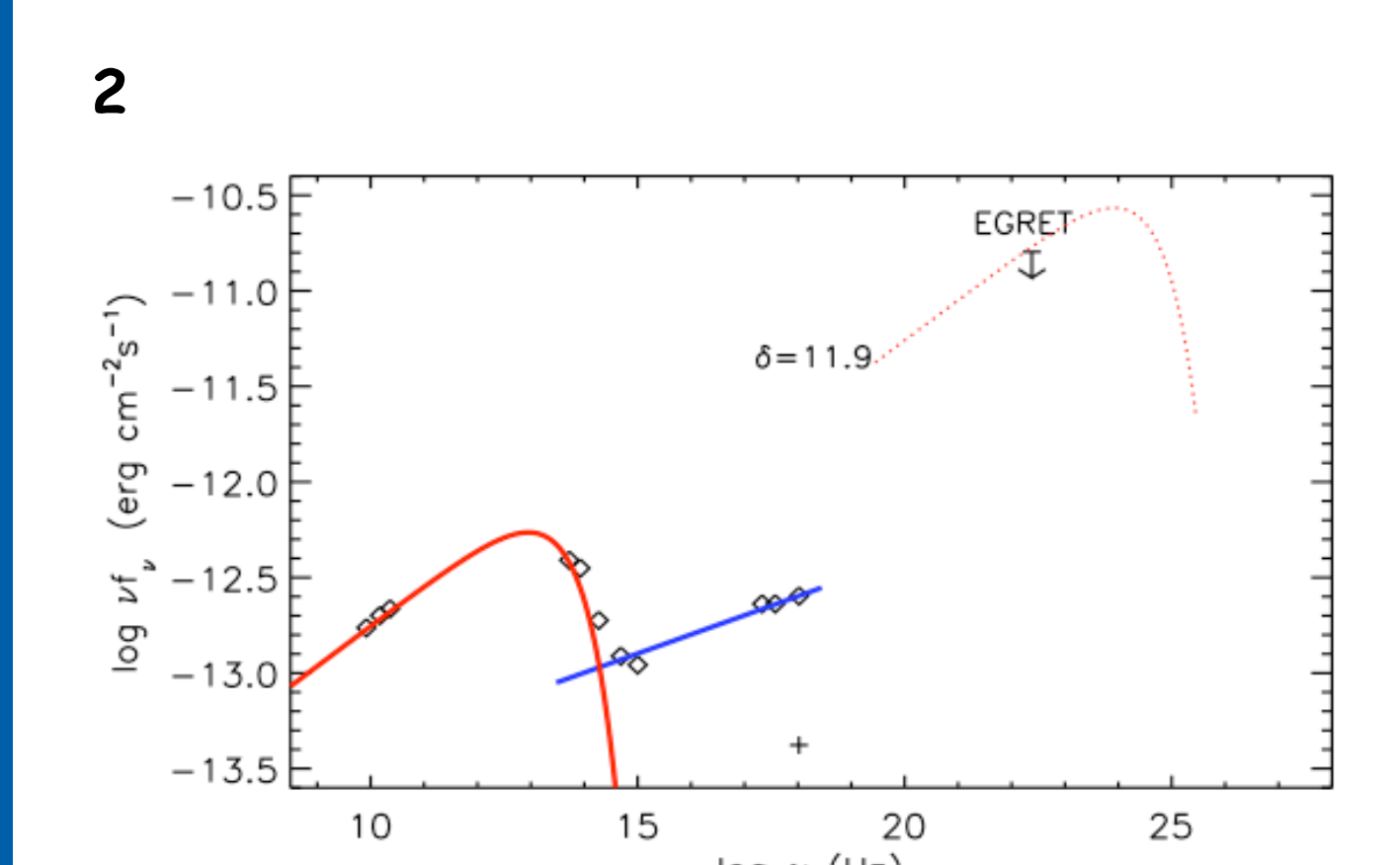
Q: Is it close enough for TeV photons to go through the extragalactic background light (EBL)?

A: Yes, it is. The Universe is more transparent than previously thought (Aharonian et al. 2006, HESS detection of $z=0.186, z=0.165$ blazars).



Add up all the flux of the jet. Red line: synchrotron emission Blue line: optical - X-ray emission of ? Nature

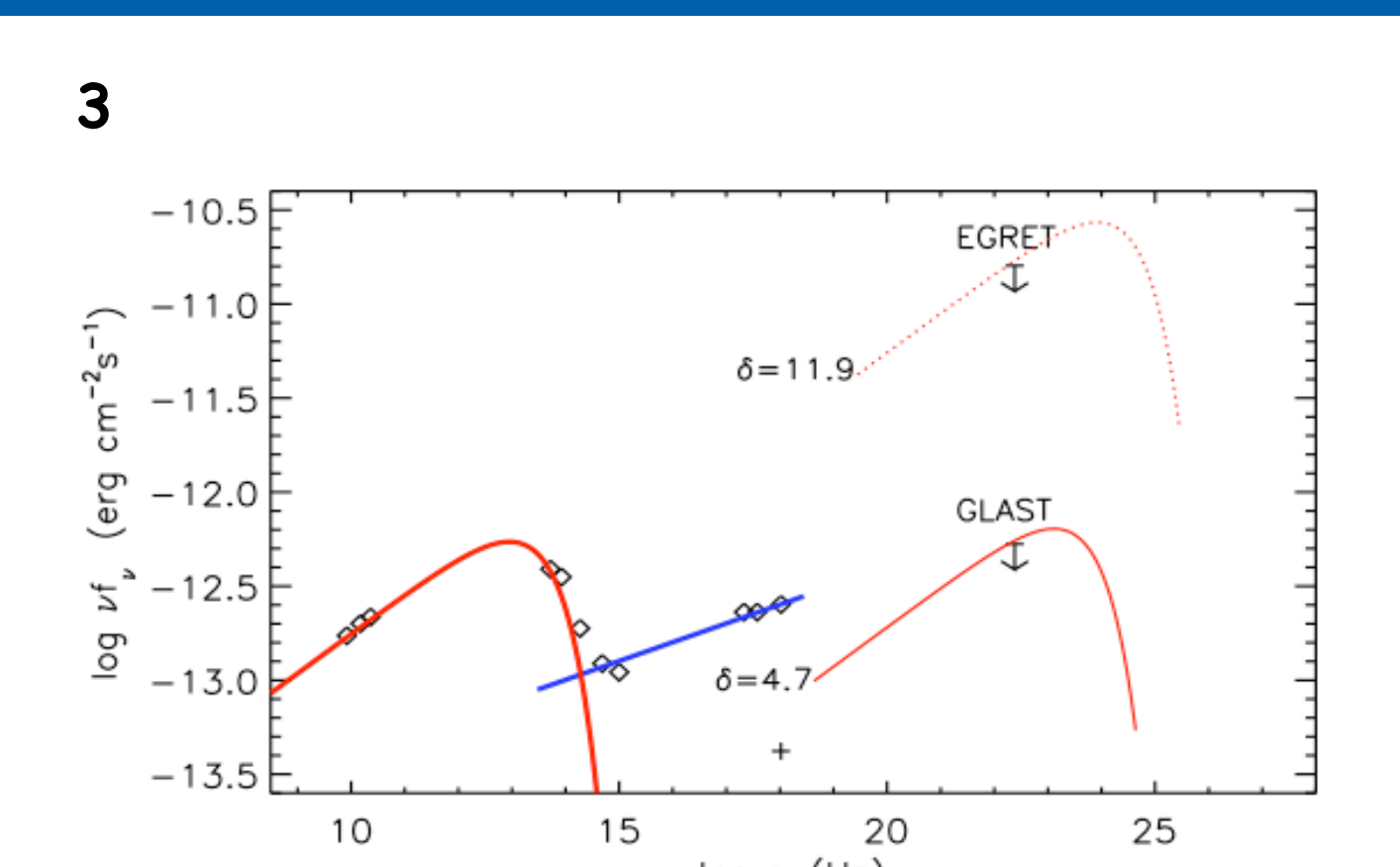
Regardless of the nature of the blue line emission, the electrons producing the red line will upscatter the CMB, producing a similar GeV component shifted in frequency and power by $\nu_{\text{CMB}}\delta^2$ and $L_{\text{syn}}\delta^4$, where δ is the jet Doppler factor.



NOTE: The blazar was below the EGRET limit about half the time the source was observed. This sets an upper limit to the jet emission.

IMPORTANT: We do not know how low the blazar drops, but it must be substantially below the EGRET limit, since it was undetectable half of the time.

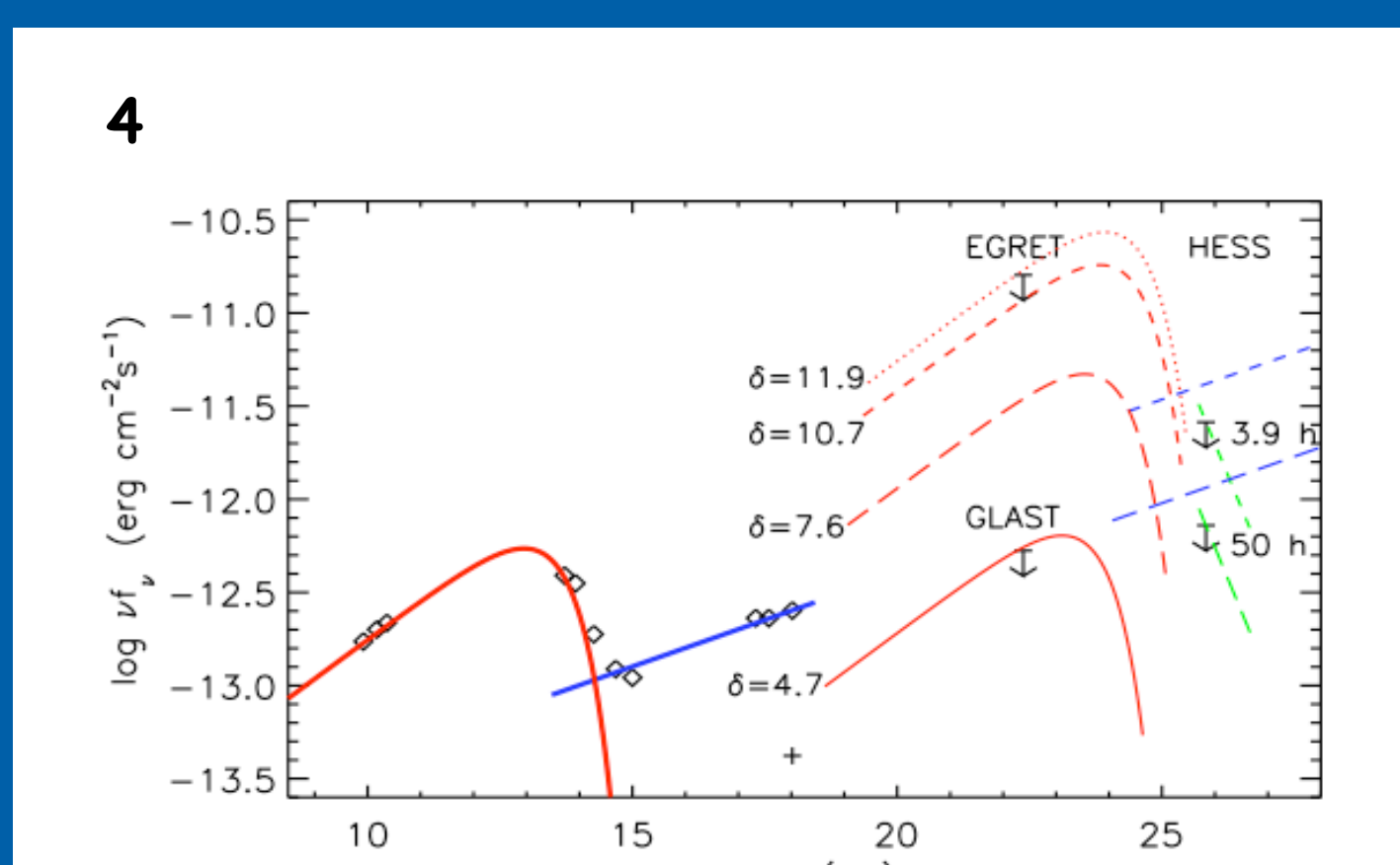
COOL: Existing good old EGRET limits require $\delta < 11.9$



In the EC/CMB hypothesis, values of $\delta < 10$ require extremely high jet power ($L_{\text{jet}} \sim 10^{47-49}$ erg/s, Uchiyama et al. 2006). This is unlikely.

GLAST will see the jet+blazar as a point source and will give us a deeper upper limit on the jet flux, therefore a smaller upper limit on δ , further disfavoring EC/CMB.

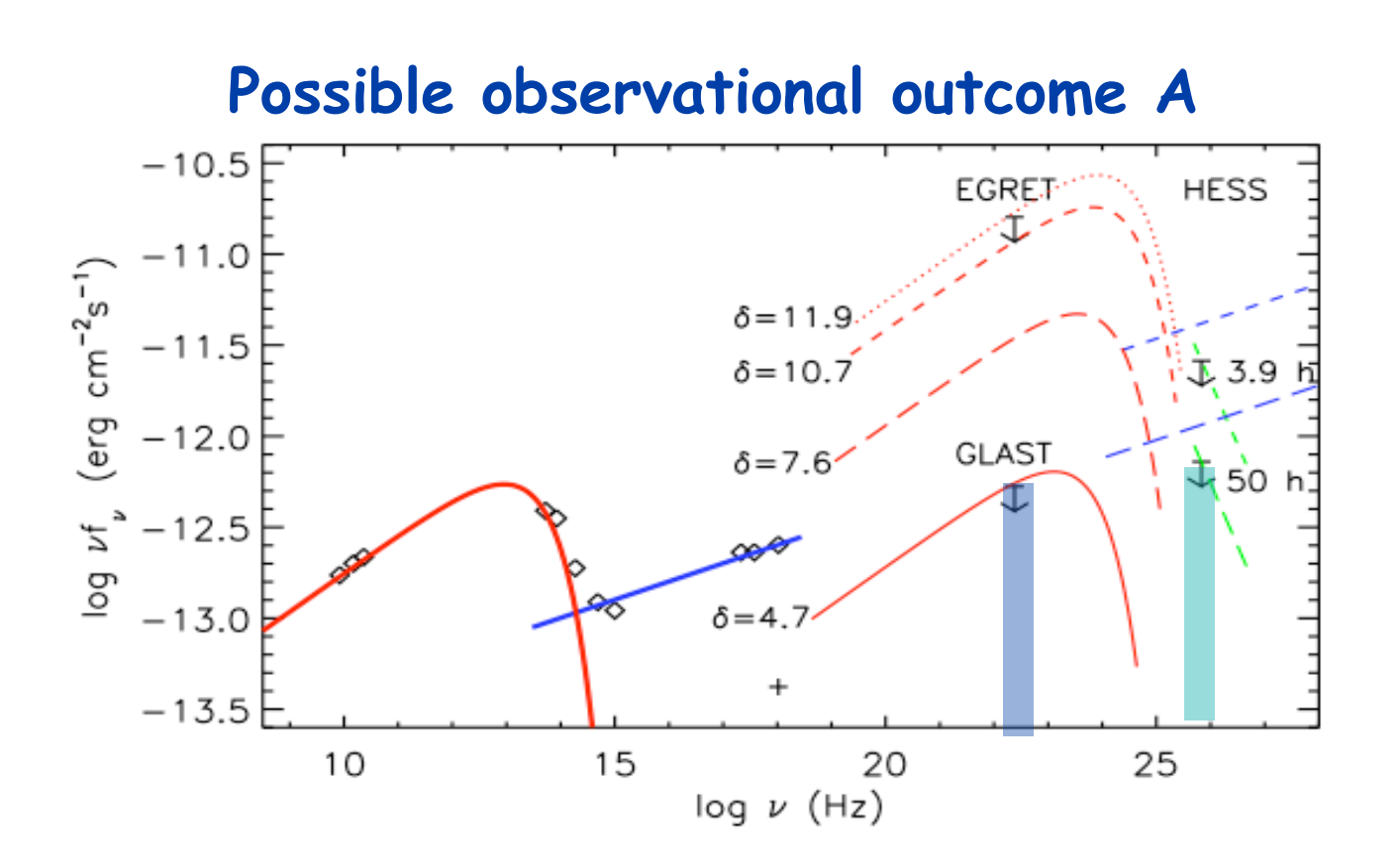
If GLAST does not detect the source, then $\delta < 4.7$. This would eliminate the EC/CMB model.



If the blue line emission is synchrotron, then the electrons producing it will also produce TeV emission.

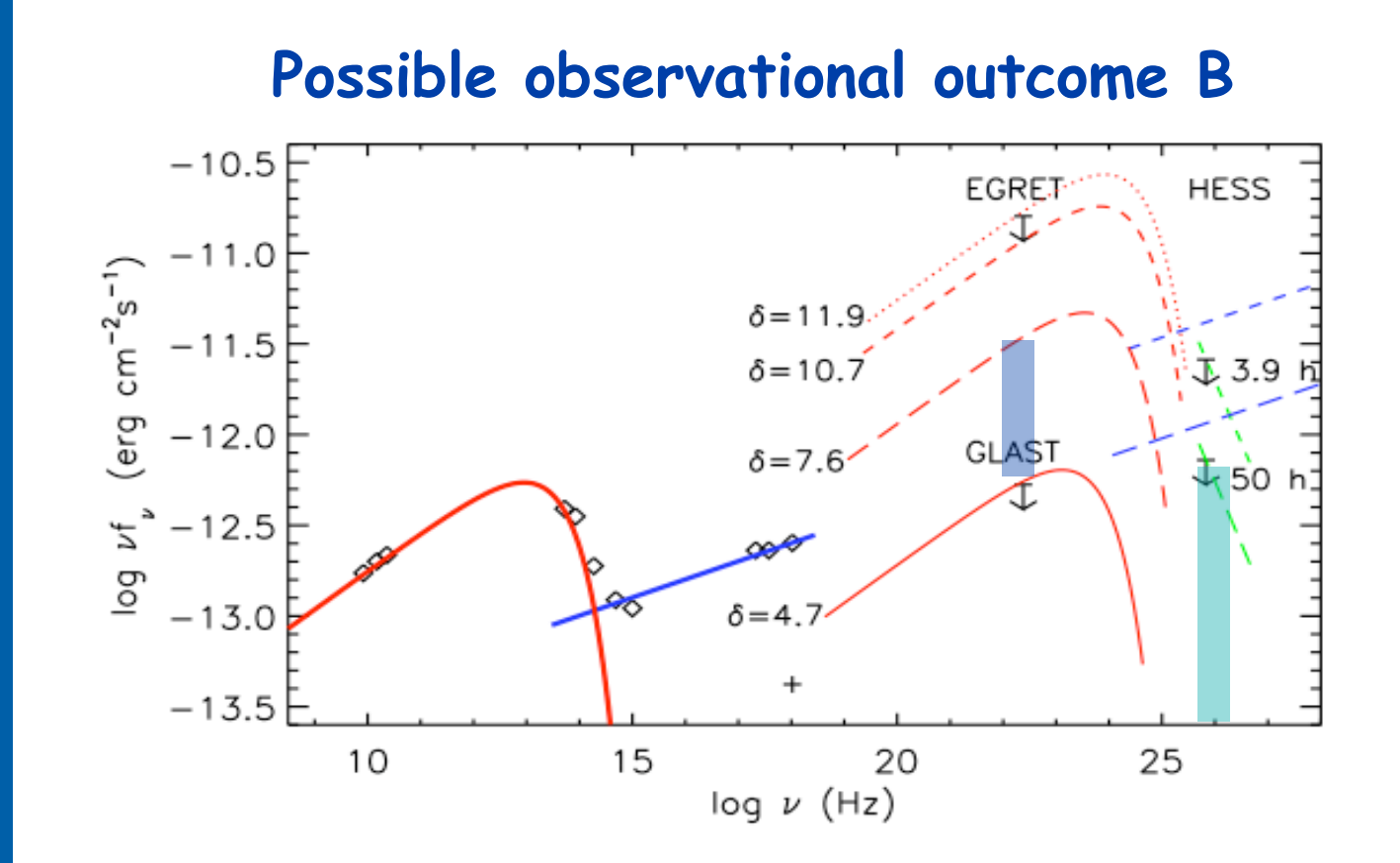
HESS shallow limit (Aharonian 2005): In the synchrotron hypothesis, $\delta < 10.7$ (blue/green broken line before/after absorption from the EBL).

50 hours HESS observations can push $\delta < 7.6$ in the synchrotron hypothesis.



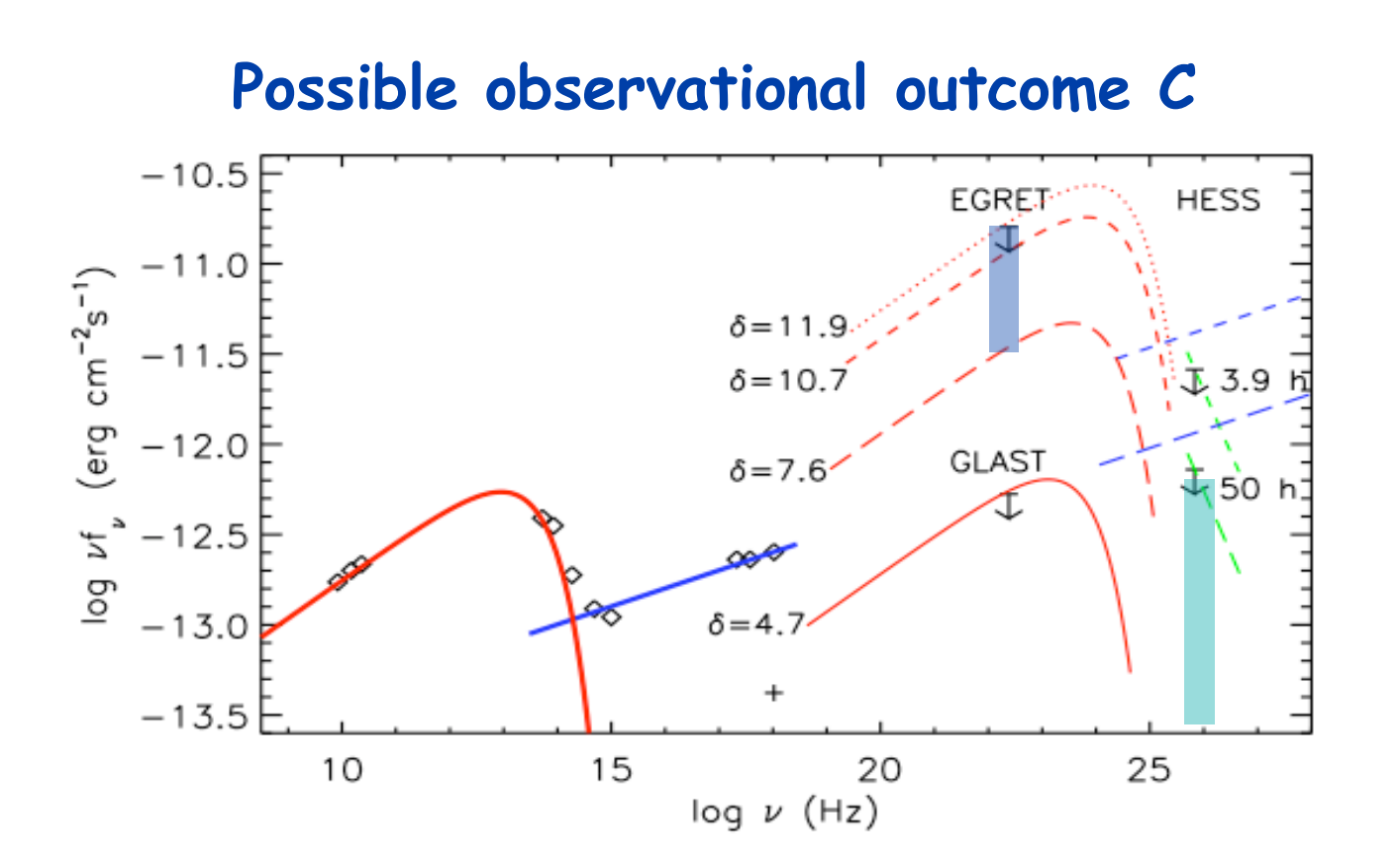
Neither GeV nor TeV emission is detected: $\delta < 4.7$
No constraints on the synchrotron model.

EC/CMB requires too much power. Strongly disfavored.



GeV emission is detected at a low level: $4.7 < \delta < 7.6$
TeV emission still is too weak to be detected.

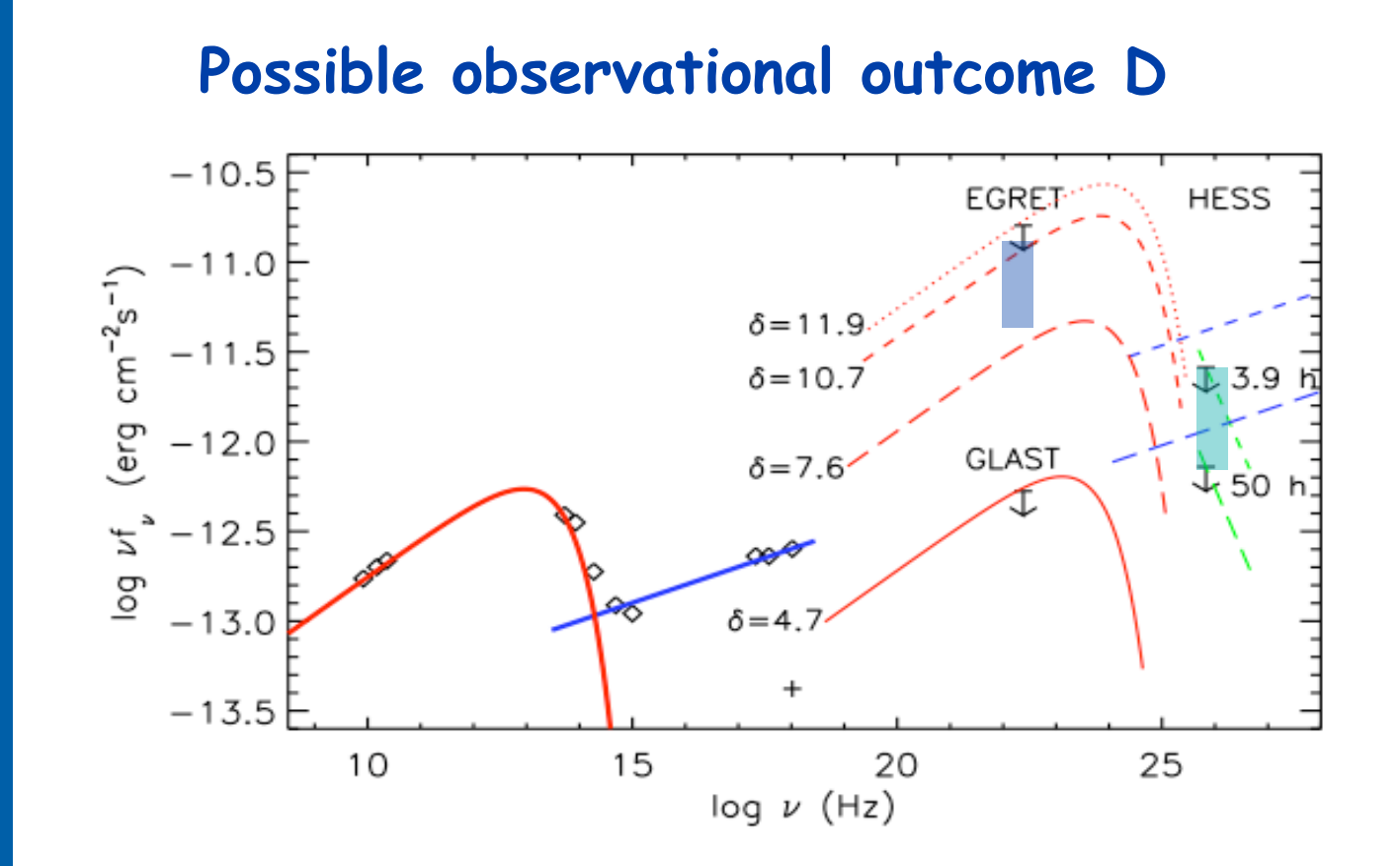
No constraints on the synchrotron model.
EC/CMB requires too much power. Disfavored.



Jet GeV emission is detected at a high level: $7.6 < \delta < 11.9$.

If we do not detect the corresponding TeV emission at a level between the 3.9 and 50 hour limits, the synchrotron model is OUT.

*hard, steady



Jet GeV emission is detected at a high level: $7.6 < \delta < 11.9$.

If we do detect the corresponding TeV emission at a level between the 3.9 and 50 hour limits, the synchrotron model is IN.

Conclusions.

- EGRET and HESS already constrain the jet Doppler factor and emission
- GLAST and further TeV observations can confirm or refute the EC/CMB and synchrotron models.

More in our paper: ApJ 653, L5