

# B2 1215+30 Long Term Study with Fermi-LAT & VERITAS

$\gamma$ -Astronomy Group

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*VERITAS analysis by:* Giuliana Noto & Reshma Mukherjee

5th June 2017

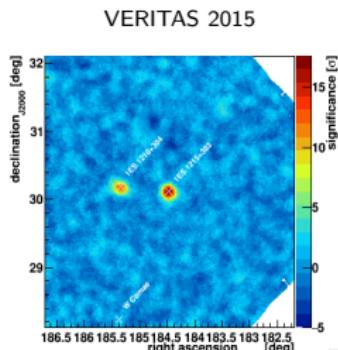
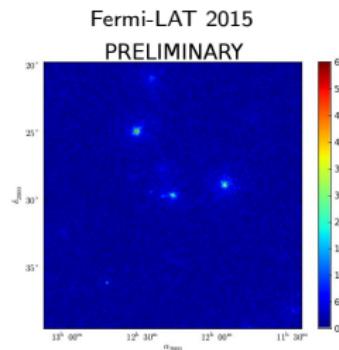
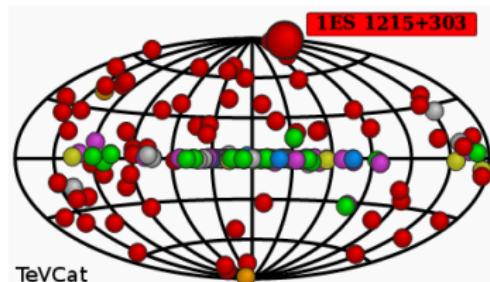
# Introduction

- 2015 analysis (internship)
  - 2016 analysis
- } VERITAS non-published data

⇒ Long Term Analysis: 2008 (Fermi START) - 2016

⇒ Include 2017 data: big flares!

Work in progress...



- **$z = 0.131$**  by Paiano et al. 2017 ([arXiv:1701.04305](#))
- High-frequency peaked BL Lac blazar.
- 1st detected in 1970 by Bologna Northern Cross telescope (B2 408 MHz)
- 1st detected at VHE by MAGIC in January 2011.

Brightest flares at HE with Fermi-LAT:

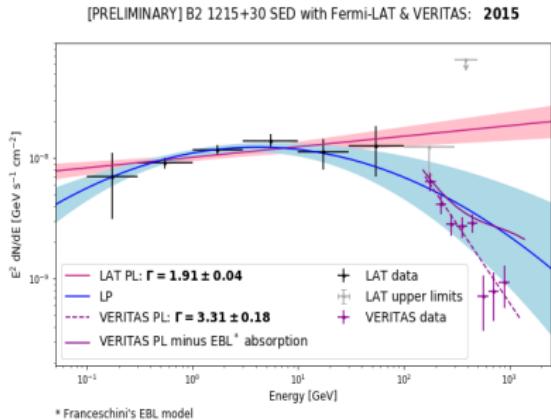
- ▶ 2008 Oct 10-15, Flux $_{(E>300\text{MeV})} \sim 15 \times 10^{-8} \text{cm}^{-2}\text{s}^{-1}$  in a weekly LC.  
“ $\gamma$ -ray Light Curves and Variability of Bright FERMI-detected Blazars”, ApJ, 722:520-542, 2010.
- ▶ 2014 Feb 08, **with a coincident VERITAS flare (below)**.  
“A Luminous and Isolated  $\gamma$ -ray Flare from the Blazar B2 1215+30”, arXiv: 1701.01067v1, 2017.

Brightest flares at VHE:

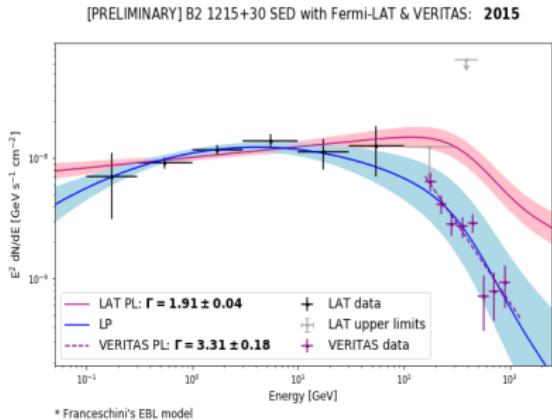
- ▶ 2013 Feb 07, ([arXiv: 1701.01067v1](#), 2017)
- ▶ **2014 Feb 08**, ([arXiv: 1701.01067v1](#), 2017)

# 2015 SED

## De-absorb EBL

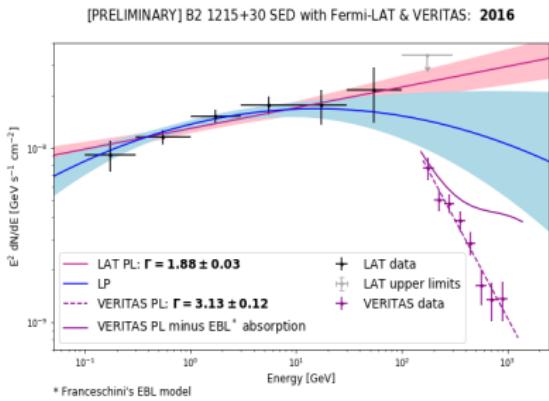


## With EBL absorption

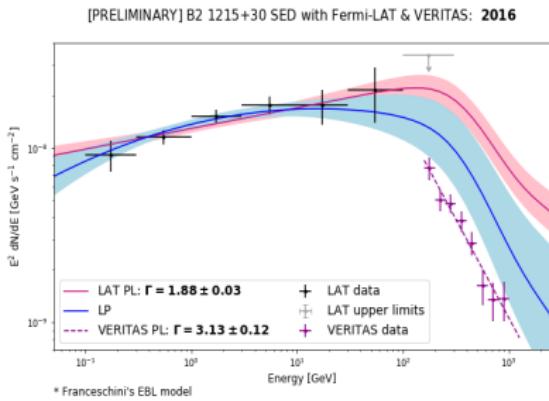


# 2016 SED

## De-absorb EBL



## With EBL absorption



# Long Term study SED with Fermi-LAT

[PRELIMINARY] B2 1215+30 SED with Fermi-LAT: 04 Aug 2008 - 30 Jun 2016

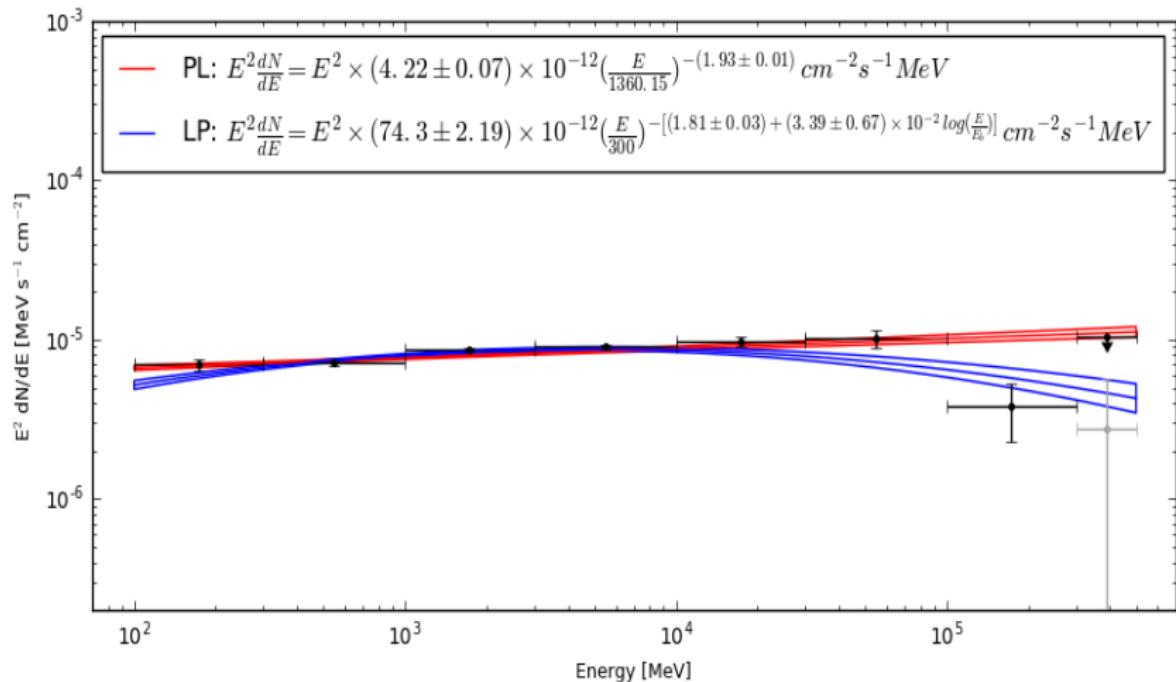


Figure 1:  $5.26\sigma$  rejection of the PL in favor of the LP model!

# 2015 and 2016 lightcurves

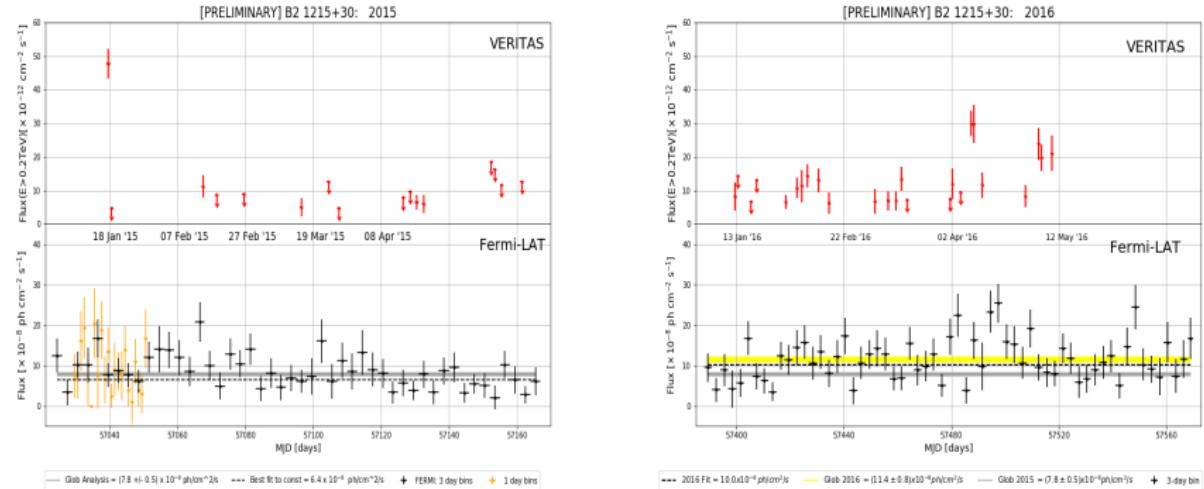
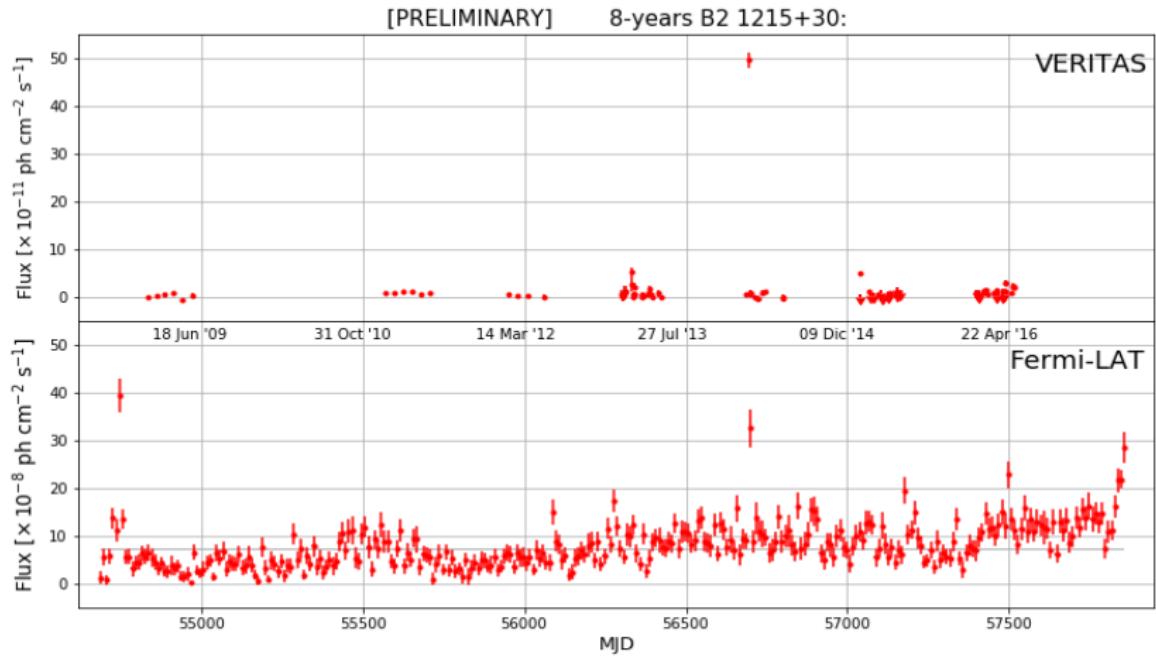
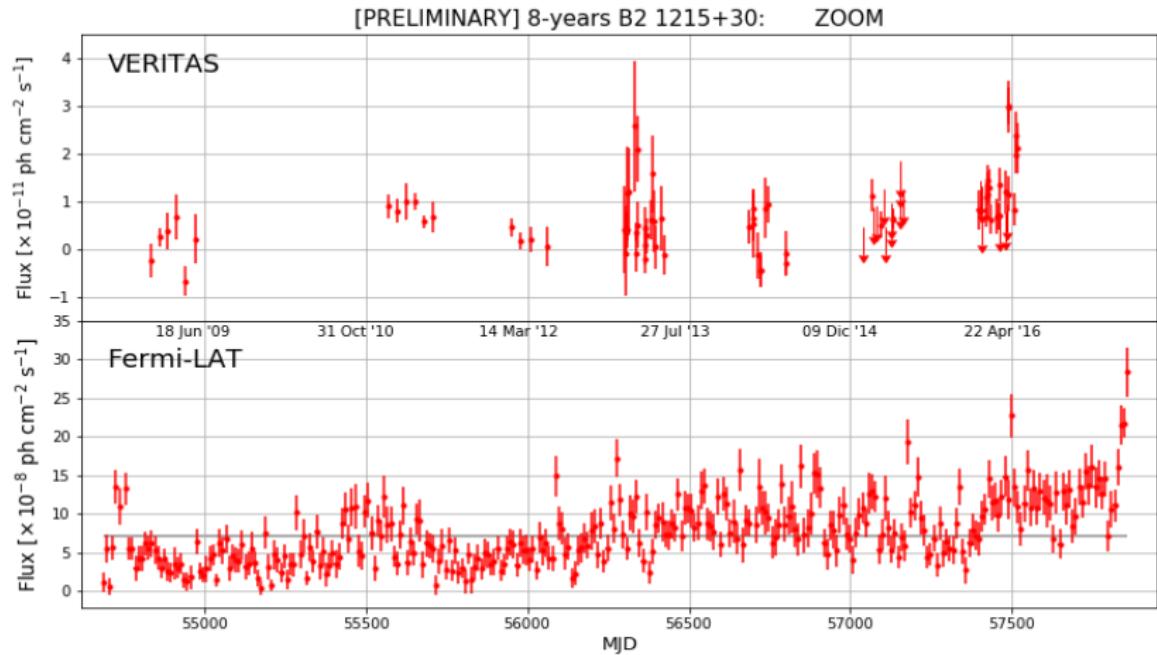


Figure 2: *VERITAS and Fermi-LAT joint lightcurves.*

# Fermi-LAT and VERITAS Long term LC



# Zoom: Increasing flux pattern



# 2017 Flares!

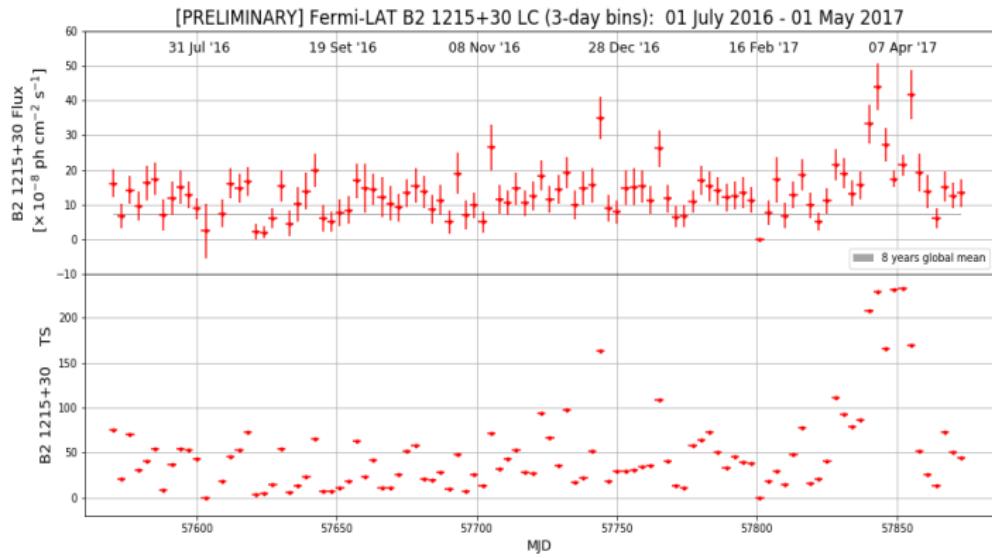
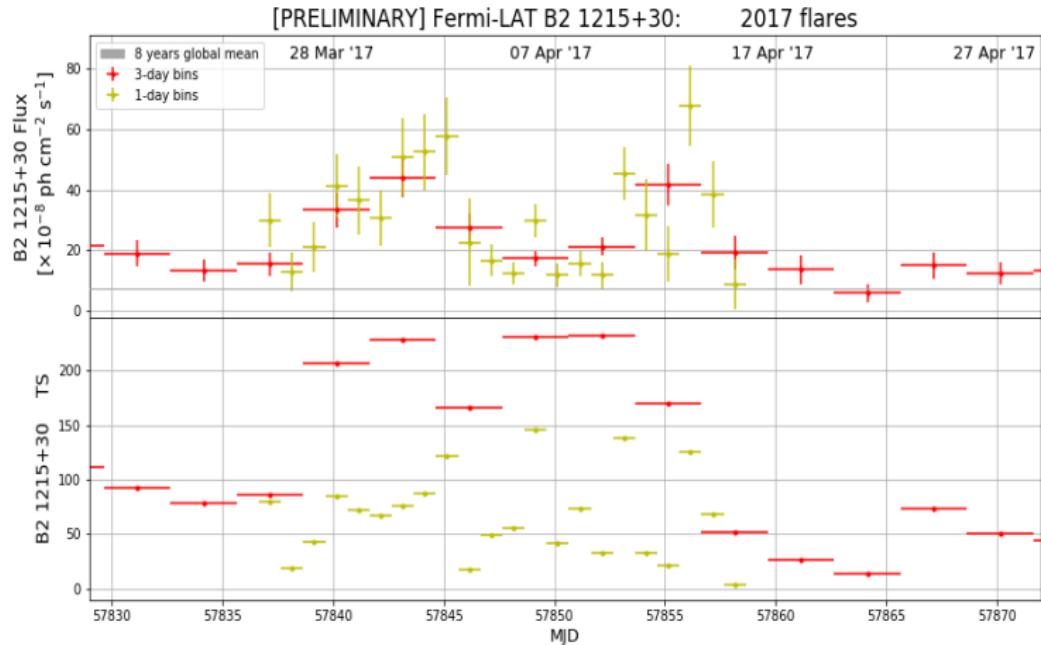


Figure 3: Coincidental VERITAS flare on Apr 1st. Includes Apr 13th ATel#10270.

# 2017 Flares: 1-day bins



# Summary of results

Table 1: Preliminary GeV TeV results for B2 1215+30

Instrument	Energy range	Dates	Live time	Sign.	$\Gamma$	Flux [ph cm $^{-2}$ s $^{-1}$ ]
VERITAS	>0.2 TeV	2015 Jan 17 – 2015 May 19	979.4 min	17.1 $\sigma$		
		2016 Jan 12 – 2016 May 09	1324 min	26.4 $\sigma$		
		2015 Jan 17	114.3 min	17.1 $\sigma$		(4.7 $\pm$ 0.5) $\times$ 10 $^{-11}$
		2016 Apr 09	120.1 min	13.4 $\sigma$		(2.3 $\pm$ 0.4) $\times$ 10 $^{-11}$
Fermi-LAT	0.1 – 500 GeV	2015 Jan 01 – 2015 Jun 30		37.4 $\sigma$	1.91 $\pm$ 0.04	(9.1 $\pm$ 0.7) $\times$ 10 $^{-8}$
		2016 Jan 01 – 2015 Jun 30		44.4 $\sigma$	1.88 $\pm$ 0.03	(11.2 $\pm$ 0.8)x 10 $^{-8}$
		2008 Aug 04 – 2016 Jun 30		119 $\sigma$	1.93 $\pm$ 0.01	(7.1 $\pm$ 0.2) $\times$ 10 $-8$
		2017 Apr 01-02		11.2 $\sigma$	1.75 $\pm$ 0.15	(4.9 $\pm$ 1.4) $\times$ 10 $^{-7}$
		2017 Apr 12-13		11.5 $\sigma$	1.78 $\pm$ 0.14	(5.9 $\pm$ 1.4) $\times$ 10 $^{-7}$

That's it, thank you!

# Long Term study

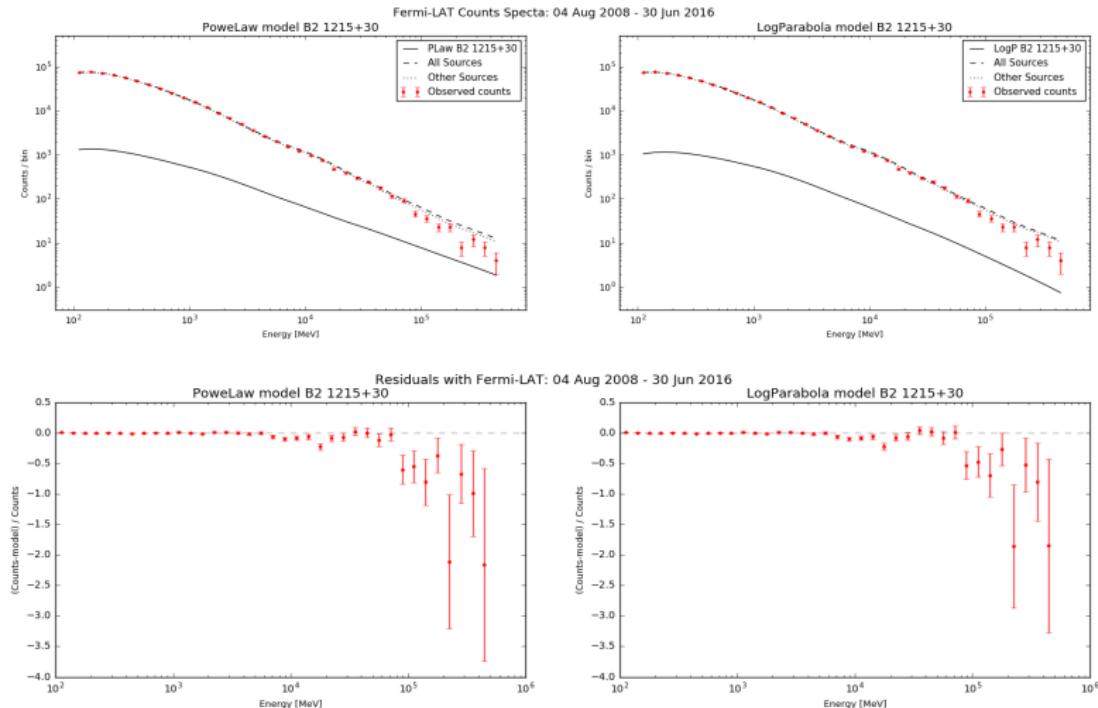
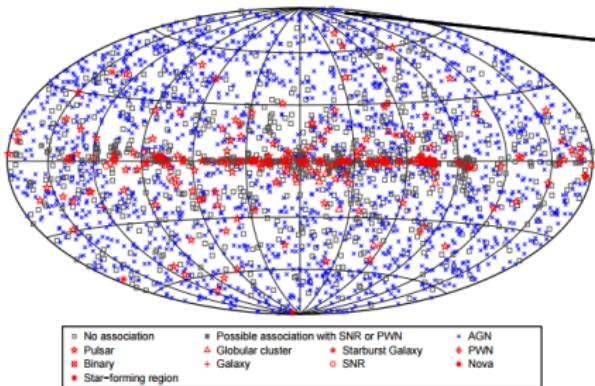


Figure 4: Counts Spectra and residuals of B2 1215+30

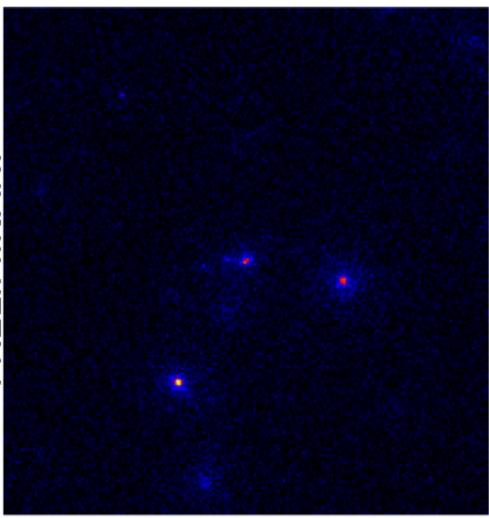
# Data and model

3FGL



- Aug. 2008-July 2012: 3033 sources.
- > 1100 blazars
- 660 BL Lacs.
- 1010 unassociated sources.
- $TS > 4\sigma$

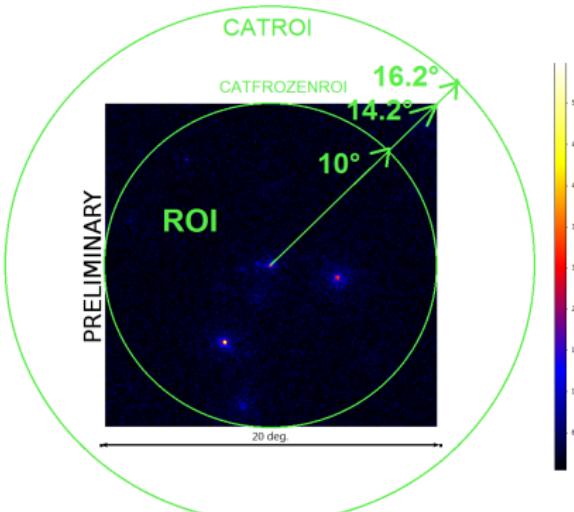
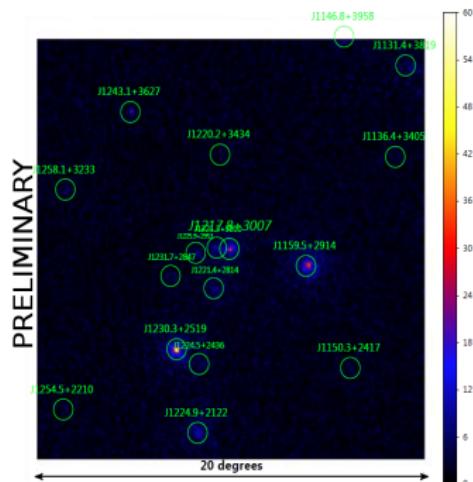
PRELIMINARY



data around  
**B2 1215+30**

2015 Counts map

# Data and Model



Maximum likelihood fit to 3FGL model:  
127 sources considered in the region.

- 117 PowerLaw modeled sources
- 8 LogParabola modeled sources
- 2 diffuse sources

\* 44 srcs have free parameters.

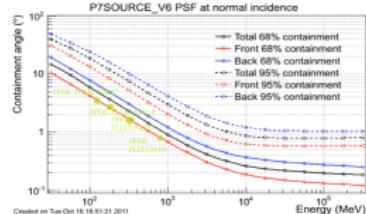


Figure 5: Point Spread Function at normal incidence